

NEXOTRANS - Next-generation EXOplanet reTRieval and ANalysiS

A detailed analysis of the 0.6 - 12 µm JWST observations of WASP-39b







Tonmoy Deka^{1,2}, Tasneem Basra Khan^{1,2}, Swastik Dewan^{1,2}, Priyankush Ghosh^{1,2}, Debayan Das^{1,2,3}, Liton Majumdar^{1,2}

¹Exoplanets and Planetary Formation Group, School of Earth and Planetary Sciences, National Institute of Science Education and Research, Jatni 752050, Odisha, India

²Homi Bhabha National Institute, Training School Complex, Anushaktinagar, Mumbai 400094, India ³Department of Physical Sciences, Indian Institute of Science Education and Research, Kolkata, Mohanpur 741246, West Bengal, India

1. Introduction

JWST is offering observations with higher-resolution and greater precision than ever before. Accurate analysis and appropriate interpretation of these datasets require advanced retrieval frameworks capable of navigating the complex parameter space. Here we introduce "NEXOTRANS", an atmospheric retrieval framework that

"**NEXOTRANS**", an atmospheric retrieval framework that utilizes both Bayesian and machine learning methods for a comparative and independent interpretation of the observations. In this poster, we provide a glimpse of this retrieval framework and few of the results from **Deka et al. 2025**.

2. The NEXOTRANS Retrieval Framework

Some of the salient features of NEXOTRANS are listed as follows:

- Capable of performing retrieval with traditional Bayesian nested sampling algorithms such as *PyMultiNest, UltraNest*.
- Does machine learning guided retrievals using four different algorithms: Random Forest, Gradient Boosting, k-Nearest Neighbor and Stacking Regressor (an ensemble of the first three, shows enhanced performance).
- Incorporates its own equilibrium chemistry grids from a newly developed equilibrium chemistry code -"NEXOCHEM".
- Models approximate disequilibrium chemistry effects on the spectrum using two new methods: Modified Hybrid Equilibrium and Modified Equilibrium offset chemistry.

3. 0.6 - 12.0 μm Retrieval of WASP-39b with NEXOTRANS

We demonstrated NEXOTRANS on the full available JWST observations of the hot Saturn-mass exoplanet WASP-39 b. The 0.6 - 2.8 μm observations (Feinstein et al. 2023) were obtained with the NIRISS instrument as a part of the JWST ERS program. The 0.5 - 5.5 μm observations (Rustamkulov et al. 2023) are from the NIRSpec PRISM, also obtained as a part of the ERS program. Finally, we also included the latest 5 - 12 μm observations from Powell et al. (2024) taken with the MIRI LRS instrument. This enabled retrievals with the combined datasets of the full JWST transmission spectrum of WASP-39b.

We performed a series of retrievals using the different chemistry models available in NEXOTRANS along with the presence of aerosols. We performed the PyMultiNest retrievals at a model resolution of R=20,000 and 2000 live points. We also performed retrievals with NEXOTRANS's machine learning mode which was trained on 60,000 model spectra, that showed consistent results with the Bayesian retrievals. It is worth mentioning that we found NIRISS and MIRI datasets have at least an offset of 57 ppm and 311.13 ppm with respect to the PRISM dataset.

References

- 1. **Deka et al**. arXiv preprint arXiv:2504.18815 (2025).
- 2. Feinstein et al. Nature 614.7949 (2023): 670-675.
- 3. **Rustamkulov et al**. Nature 614.7949 (2023): 659-663.
- 4. Powell et al. Nature 626.8001 (2024): 979-983.

Input / System Figure 1: A schematic overview of the retrieval framework parameters of NEXOTRANS. **Observed** Chemical species, spectrum Atomic/Molecular absorption crosssections, Star-Planet properties. Forward Model Retrieval Framework Machine Radiative Transfer **Extinction Calculation** Atmospheric Profiles Bayesian Learning Chemical species High resolution PT-Profiles Random Forest model spectrum Madhusudhan Patchy cloud profile/Guillot UltraNest **PyMultiNest** deck & haze profile/Isothermal kNN profile) Binning to Sigmoid clouds observation Gradient resolution Boosting Mie Aerosols Chemistry(free chemistry/NEXOCHEM/ Evidence Stacking Posterior Hybrid/Offset Ackerman & Model Spectrum Model chemistry) Marley clouds

4. Results

Below, we show one of the results from a set of retrievals performed on the combined data. Of all the retrievals, the modified hybrid equilibrium chemistry model achieved the best-fit with a reduced chi² value of 2.97.

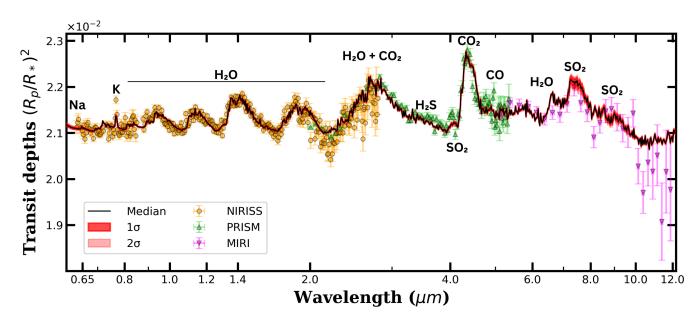


Figure 2: Best-fit retrieved spectrum using Bayesian retrieval for a model assuming hybrid equilibrium chemistry and aerosols.

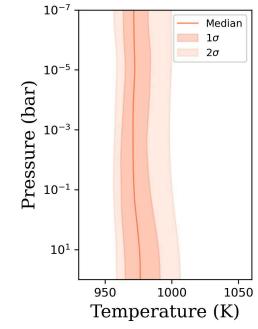


Figure 3: Retrieved PT profile.

5. Retrieved Parameters

Table 1: Retrieved values for selected parameters from the best-fit hybrid chemistry model including aerosols.

	log(H ₂ O)	log(CO ₂)	log(CO)	log(H ₂ S)	log(SO ₂)	log(K)	C/O	[M/H]
Bayesian	-2.89	-4.59	-1.99	-3.93	-5.80	-8.74	0.80	1.19
	+0.01/	+0.05/	+0.06/	+0.03/	+0.11/	+0.15/	+0.03/	+0.05/
	-0.03	-0.05	-0.04	-0.02	-0.13	-0.16	-0.01	-0.04
ML	-2.60	-4.42	-2.25	-3.81	-6.24	-8.25	0.72	1.18
	+0.08/	+0.17/	+0.27/	+0.05/	+1.14/	+0.01/	+0.14/	+0.17/
	-0.08	-0.55	-0.49	-0.20	-0.32	-0.01	-0.13	-0.30

6. Conclusions

Some of the conclusions that we make from this work are:

- Chemical models relying purely on equilibrium chemistry cannot produce the observed spectrum, especially the SO₂ feature in the PRISM and MIRI wavelengths (for combined retrievals).
- Approximate non-equilibrium chemistry models such as, Hybrid equilibrium chemistry or Equilibrium offset chemistry models, that combines equilibrium and free chemistry with multiplicative offsets to the equilibrium VMRs are at least needed for proper retrievals of chemical abundances.
- Non-grey cloud or aerosol models are preferred over grey patchy/uniform cloud and haze models. ZnS is preferred in most cases.
- All non-equilibrium chemical models hint at a super-solar C/O (0.72-0.89) ratio along with super-solar Metallicity, [M/H], (1.18-1.66) .