Exploring beyond! Advancements in exoplanet detection in the SHINE High-Contrast Imaging survey

through RSM framework

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## SHINE survey & Motivation

### SpHere INfrared Survey for Exoplanets (SHINE, Desidera et al. 2021)

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SPHERE instrument at VLT, using SAXO extreme AO & Lyot Coronagraph



F150 sample: 53 BA stars, 77 FGK, 20 M stars



IRDIS narrowband filters: H2-H3, with 9" diameter field of view Science observations: ADI image sequence

In this poster, we revisit the complete F150 SHINE sample using RSM, an advanced post-processing algorithm. We observe notable improvements in the detection limits compared to standard postprocessing algorithms, leading to the identification of new exoplanet candidates that are currently under analysis.

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### Advanced post-processing algorithm !

Angular Differential Imaging (ADI, Marois et al. 2006)

#### Regime Switching Model (RSM, Dahlqvist et al. 2021)



## Detection threshold under log-norm assumption

## New point sources in SHINE survey!



Using datasets sharing similar environmental conditions, independent realisations at each separation fitted a log norm distribution. The detection threshold is assumed at  $3 \times 10^{-7}$  false alarm probability. The contrast curve based on this detection threshold and 50% completeness is equivalent to the 5 sigma contrast under the assumption of gaussian noise.

**RSM** probability map PCA SNR map 0.6 0.0 -0.6 -1.2 -0.6 1.2 0.6 0.0 -1.2 1.2 0.6 -0.6 0.0 -1.2 ΔRA["] ΔRA["] 0.6 0.0 -0.6 -1.2

Numerous Candidates are under analysis..

-1.2

-0.6

Detection limits using different PSF subtraction methods

### **New detection limits in SHINE !**

1.2

0.6

0.0

ΔRA["]

-0.6

-1.2

1.2

0.6

0.0

ΔRA["]

