Secondary eclipses and phase curves

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The APEx Department

Atmospheric Physics of Exoplanets

- New department at the Max Planck Institute of Astronomy in Heidelberg, Germany
- Focus is observations, theory, and instrumentation for exoplanet atmospheres!
- Advertisement: we are hiring at all levels! Jobs are advertised in the fall on the AAS job register





What is a secondary eclipse?



Thermal emission and reflected light from the planet are blocked during secondary eclipse

Reveals average spectrum of the dayside hemisphere

The first detections of thermal emission Deming et al. 2005, Charbonneau et al. 2005





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Intro to thermal emission — blackbody approximation

- Luminosity L ~ T^4
- Wavelength of peak emission IS $\lambda_{ ext{peak}}$ \overline{T}

where b = 3000 micron*Kelvin

credit: <u>sun.org</u>



Key facilities

Atmosphere characterisation is easier when Earth's atmosphere is not in the way! ** note: ground-based observations are also important and complementary



Spitzer Space Telescope Launch date: 2003 85 cm mirror Infrared; 3.6 - 160 micron



Hubble Space Telescope Launch date: 1990 2.4 m mirror UV-near-IR: 0.1 - 1.7 micron



JWST Launch date: 2021 6.5 m mirror Optical - IR: 0.6 - 30 micron

Typical thermal emission signal

Hot Jupiter orbiting a Sun-like star: Tp = 1500, Ts =5000, Rp/Rs = 0.01

Warm rocky planet orbiting an M-dwarf:

Tp = 400, Ts = 3000,Rp/Rs = 0.01

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Wavelength





Wavelength



Temperature

Credit: Tom Mikal-Evans





Wavelength





Wavelength





Wavelength



Temperature

Credit: Tom Mikal-Evans





Wavelength



Intro to phase curves



Animation available at github.com/lkreidberg

- Observe a complete orbital revolution
- Changing viewing angle reveals different regions of the atmosphere over time
- Can map the global climate and chemical composition!

A new frontier: eclipse mapping!



More on this from Emily Rauscher on Thursday!

During ingress and egress, the visible disk of the planet changes incrementally

De Wit et al. 2012



Science from eclipses and phase curves

- How is heat recirculated in the atmosphere?
- What is the temperature structure? Are there thermal inversions?
- How do chemical composition and clouds vary as a function of longitude?
- What is the atmospheric metallicity?
- Is an atmosphere present?



Showman & Guillot 2002



How is heat circulated in the atmosphere? The first thermal phase curve, HD 189733b – Knutson et al. 2007



Peak brightness 16 +/- 6 degrees of the substellar point, day-night temp contrast ~ 240 K



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How is heat circulated in the atmosphere? Hotter planets have larger day-night temperature contrast



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This trend is due to the decreasing ability with increasing incident stellar flux of waves to propagate from day to night and erase temperature differences.



How is heat circulated in the atmosphere? A sharp rise in brightness at 1730 K



Onset of magnetic drag? and/or the rapid dissipation of day side clouds?

Deming et al. 2023



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How does the temperature change with altitude? Inversions increase in strength with increasing irradiation



Mansfield et al. 2021



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How does chemistry and cloud coverage change with longitude? **Evidence for nightside clouds**



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Ask me later about WASP-43b!

Uniform nightside temperature, possibly due to clouds (or the relatively long radiative timescale on the nightside)

Keating et al. 2019, Beatty et al. 2019, Parmentier et al. 2021











How does chemistry and cloud coverage change with longitude? Sometimes we see clouds on the dayside!

Total = Symmetric Reflection + Asymmetric Reflection + Thermal Emission





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Evidence for reflective clouds west of the substellar point for Kepler-7b

Hu et al. 2015

(see also Demory et al. 2013, Parmentier et al. 2016)







What is the atmospheric metallicity?

Gas giants show a diversity of compositions



Hot Saturn HD 149026b has approximately ~100x solar metallicity

- Bean et al. 2023



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Hot Jupiter WASP-77A b has a subsolar metallicity

- August et al. 2023

Sub-Neptune GJ 1214b has a ~100 -1000x solar metallicity atmosphere





Kempton et al. 2023



Do rocky planets have atmospheres? Thick atmospheres transport heat to the nightside

TRAPPIST-1c

Thick atmosphere -> full heat redistribution -> 340 K

Bare rock —> no heat redistribution —> 430 K



First results indicate that rocky planets do not have thick atmospheres



LHS 3844b-Kreidberg et al. 2019

More on this from Natasha Batalha on Thursday!



1500

TRAPPIST-1c — Greene et al. 2023 TRAPPIST-1b — Zieba et al. 2023



											PhaseC	NIRISS.SOSS
	ots more to come!									LTT-9779	PhaseC	NIRISS.SOSS
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1 WASP-69 Eclipse	NIRCAM.GRISMR+F322W2	L HD-189733	Eclipse	MIRI.LRS	L LHS-1140	Eclipse	1 LHS-1140	Eclipse	MIRI.F1500W			
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1 GJ-436 Eclipse	NIRCAM.GRISMR+F444W	L 55CNC	Eclipse	NIRCAM.GRISMR+F444W	L LTT-3780	Eclipse	1 LTT-3780	Eclipse	MIRI.F1500W	LTT9779	PhaseC	NIRSPEC.BOTS+G39
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1 WASP-77A Eclipse	NIRSPEC.BOTS+G395H	2								1453811	PhaseC	NIRSDEC BOTS+C20
1 GJ1132 Eclipse		<u></u>								LI155044	rhasec	MINSFEC. DUISTUS

