

Effect of stellar inclination on planet detectability

GROUP 3

Jennifer Carter ¹ Jacob Shapiro ² Umberto Simola ³

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Hypothesis

What happens if
you change the
stellar inclination?

How does
inclination
influence the
stellar activity?

How does stellar
inclination
influence the
ability to detect a
planetary signal in
RV data?

Conclusions

Future Perspective

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²Cornell University

³University of Padua, Yale University

Questions

- ▶ What happens if you change the **stellar inclination**? Is the activity RV signal more or less important?

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What happens if you change the stellar inclination?

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- ▶ What happens if you change the **stellar inclination**? Is the activity RV signal more or less important?
- ▶ Is it therefore more difficult or easier to detect planets? Do you prefer to observe equator-on stars, or pole on stars for planet detection?

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We expect:

- ▶ as the stellar inclination decreases, the magnitude of the RV signal decreases as well.

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Hypothesis

We expect:

- ▶ as the stellar inclination decreases, the magnitude of the RV signal decreases as well.
- ▶ as the host star is viewed more pole on the influence of stellar activity decreases, leading to an increased planetary detectability.

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Question 1: Methodology

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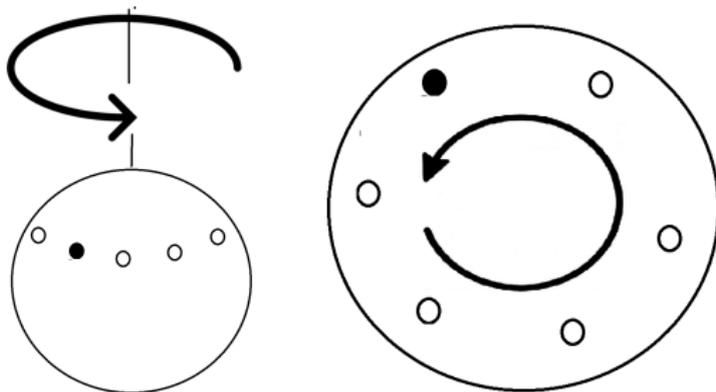


Figure: (a) Equator-on scenario, (b) Pole-on scenario.

Question 1: Analysis

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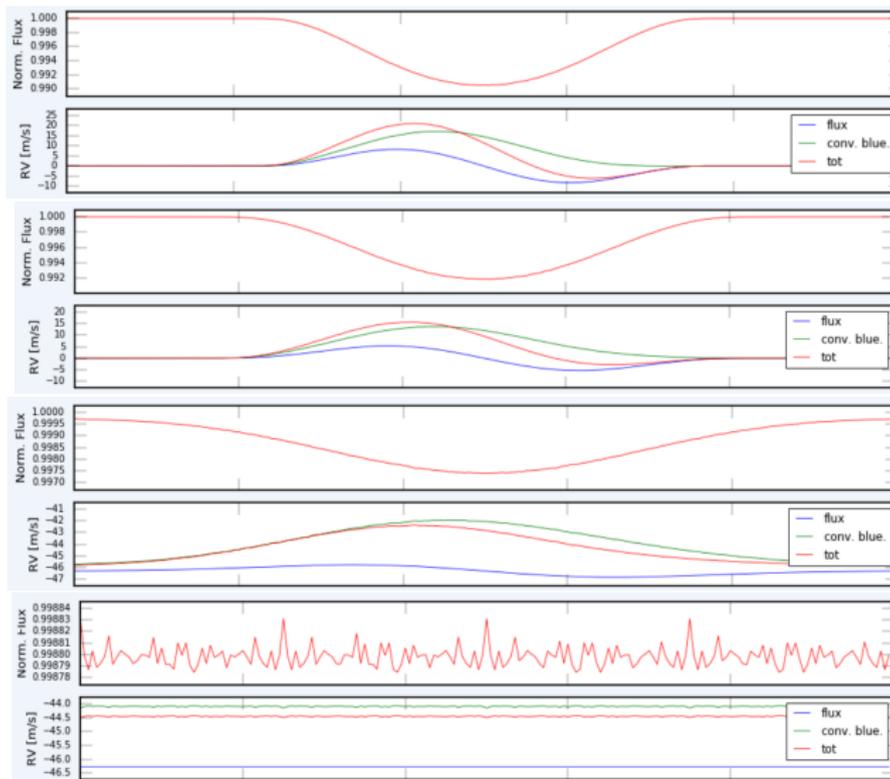


Figure: from the top to the bottom: inclination = (90, 45, 10, 0).

Question 2: Methodology

- ▶ Stellar parameters: same as those of the sun.

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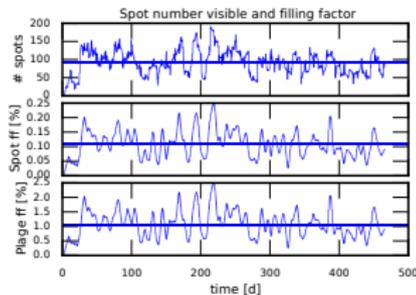
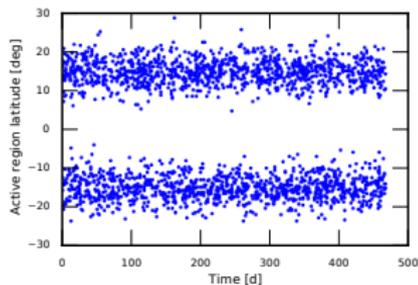
How does stellar inclination influence the ability to detect a planetary signal in RV data?

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Question 2: Methodology

- ▶ Stellar parameters: same as those of the sun.
- ▶ Stellar inclination: is allowed to vary (90, 60, 45 and 30 degrees).



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Question 2: Analysis

Compare the extreme cases of inclination.

- ▶ At 90 degrees SPOT AND PLAGE correlations are wide.

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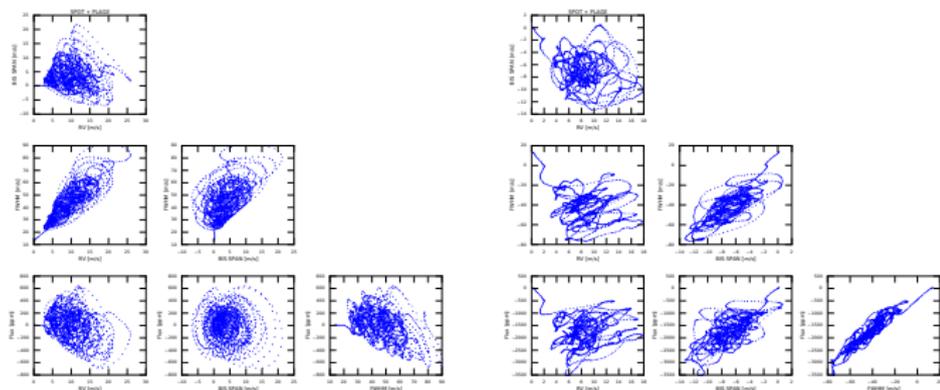
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Compare the extreme cases of inclination.

- ▶ At 90 degrees SPOT AND PLAGE correlations are wide.
- ▶ At 30 degrees SPOT AND PLAGE correlations show ridge structure for FLUX vs FWHM.



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Compare the flux and RV effects.

- ▶ at 90 degrees the flux vs time total floats around zero and the RV total ranges from near zero to 25 m/s.

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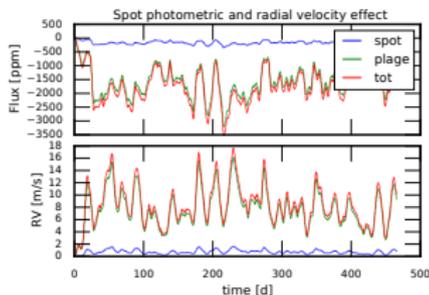
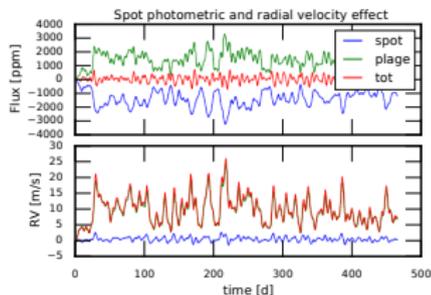
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Question 2: Analysis

Compare the flux and RV effects.

- ▶ at 90 degrees the flux vs time total floats around zero and the RV total ranges from near zero to 25 m/s.
- ▶ at 30 degrees the flux total a plage flux are both very negative. The spot flux appears unchanged (perhaps because it was not updated in SOAP2). The RV range is between 0 and 17 m/s.



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Question 2: Results

- ▶ As the stellar inclination approaches zero (pole on), the structure of the correlations change, becoming more ridge like in some dimensions.
- ▶ As the stellar inclination decreases the magnitude of the RV signal decreases

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Question 3 Methodology

- ▶ Stellar parameters are kept constant except the inclination.

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- ▶ Stellar parameters are kept constant except the inclination.
- ▶ Planet has a RV semi-amplitude of 3 m/s and period of 10 days .

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Question 3: Analysis

Compare the RV vs time plots showing 25 samples characterizing the planetary signal and the stellar signal.

- ▶ At 90 degrees SPOT AND PLAGE correlations are wide.

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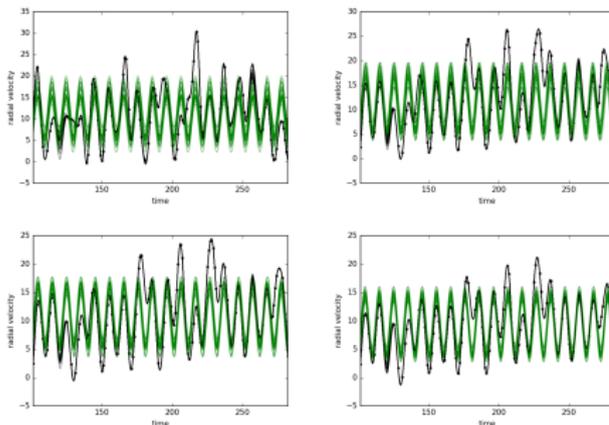
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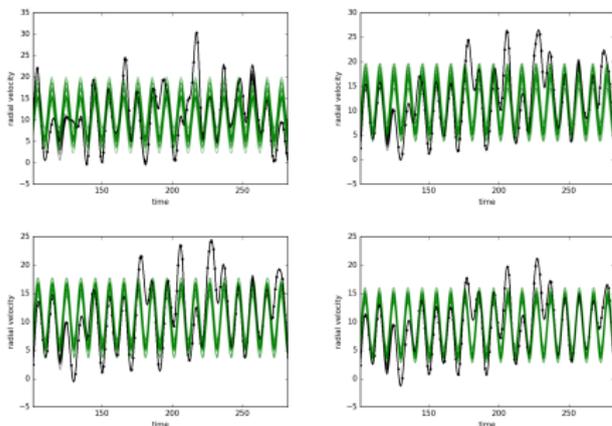
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Compare the RV vs time plots showing 25 samples characterizing the planetary signal and the stellar signal.

- ▶ At 90 degrees SPOT AND PLAGE correlations are wide.
- ▶ At 30 degrees SPOT AND PLAGE correlations show ridge structure for FLUX vs FWHM.



- ▶ Amplitude of stellar activity induced RV increases as inclination increases.

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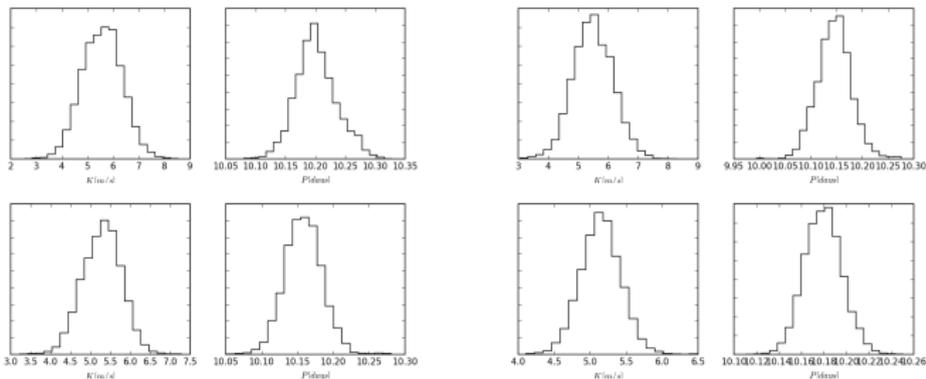
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Compare planet parameter histograms.

- ▶ width decreases as inclination decreases.



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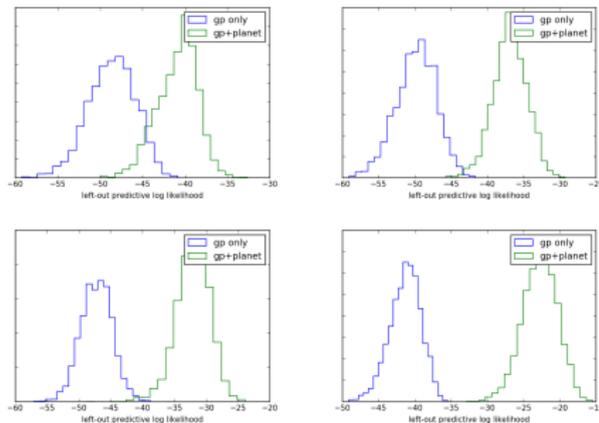
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Compare $\ln Z$ of GP only and GP + planet histograms.

- ▶ Distance between the two increases as inclination decreases.



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Question 3: Results

- ▶ As the host star is viewed more pole on the influence of stellar activity decreases. This in turn increases planetary detectability.

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- ▶ Our initial hypothesis have been confirmed by the analysis.

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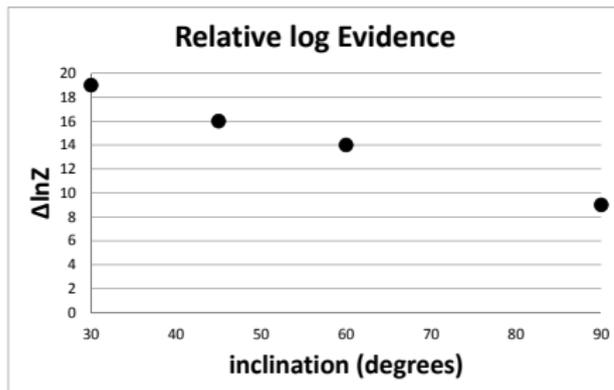
How does stellar inclination influence the ability to detect a planetary signal in RV data?

Conclusions

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Conclusions

- ▶ Our initial hypothesis have been confirmed by the analysis.
- ▶ The variation in the activity radial velocity decreases the further away we get from the equator-on case.



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Future Perspective

- ▶ Examination of the pole-on case.

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- ▶ Examination of the pole-on case.
- ▶ To change the planet equatorial plane.

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Future Perspective

- ▶ Examination of the pole-on case.
- ▶ To change the planet equatorial plane.
- ▶ Different types of stars.

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