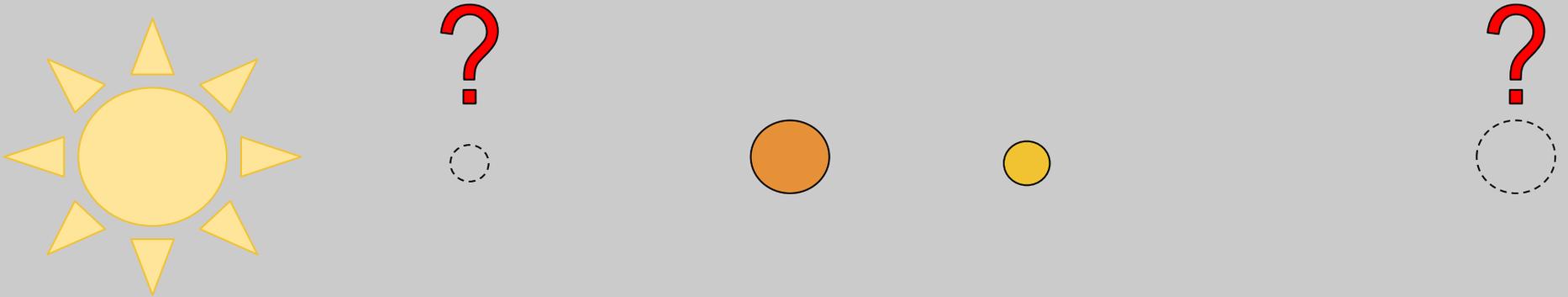


Bayesian model comparison for radial velocity: 1, 2, 3, or many planets?



Ben Nelson

Data Science Scholar at Northwestern University

@exobenelson

Outline

1. “Evidence”, Bayes factors, and decision making
2. How to efficiently compute “evidence”
3. Cross-validation as an alternative to BIC/BFs

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posterior



prior



likelihood



$$p(\theta|d) = \frac{p(\theta)p(d|\theta)}{p(d)}$$

d : data

θ : parameters

posterior



prior



likelihood



$$p(\theta | \mathbf{d}, \mathcal{M}) = \frac{p(\theta | \mathcal{M}) p(\mathbf{d} | \theta, \mathcal{M})}{p(\mathbf{d} | \mathcal{M})}$$

\mathbf{d} : data

θ : parameters

\mathcal{M} : model

$$p(\theta | \mathbf{d}, \mathcal{M}) = \frac{p(\theta | \mathcal{M}) p(\mathbf{d} | \theta, \mathcal{M})}{\int p(\theta | \mathcal{M}) p(\mathbf{d} | \theta, \mathcal{M}) d\theta}$$

\mathbf{d} : data

θ : parameters

\mathcal{M} : model

fully marginalized likelihood
or “evidence”



$p(\theta)$

Without something to compare to,
FML is not very useful...

$p(d|\mathcal{M})$

$$\int p(\theta|\mathcal{M})p(d|\theta, \mathcal{M})d\theta$$

d : data

θ : parameters

\mathcal{M} : model

fully marginalized likelihood
or “evidence”



Let's have two models compete!

\mathcal{M}_A : model A

\mathcal{M}_B : model B

Let's have two models compete!

\mathcal{M}_A : model A

\mathcal{M}_B : model B

$$\frac{p(d|\mathcal{M}_A)}{p(d|\mathcal{M}_B)} = \frac{\int p(\theta_A|\mathcal{M}_A)p(d|\theta_A, \mathcal{M}_A)d\theta_A}{\int p(\theta_B|\mathcal{M}_B)p(d|\theta_B, \mathcal{M}_B)d\theta_B}$$



Bayes Factor

Let's have two models compete!

\mathcal{M}_A : model A

\mathcal{M}_B : model B

$$\frac{p(d|\mathcal{M}_A)}{p(d|\mathcal{M}_B)} = \frac{\int p(\theta_A|\mathcal{M}_A)p(d|\theta_A, \mathcal{M}_A)d\theta_A}{\int p(\theta_B|\mathcal{M}_B)p(d|\theta_B, \mathcal{M}_B)d\theta_B}$$



Bayes Factor

Prior on
models

Posterior
Odds Ratio



$$\frac{p(d|\mathcal{M}_A)}{p(d|\mathcal{M}_B)} \times \frac{p(\mathcal{M}_A)}{p(\mathcal{M}_B)} = \frac{p(\mathcal{M}_A|d)}{p(\mathcal{M}_B|d)}$$

“So, what do I do with this?”

$$\frac{p(\mathbf{d}|\mathcal{M}_A)}{p(\mathbf{d}|\mathcal{M}_B)} \times \frac{p(\mathcal{M}_A)}{p(\mathcal{M}_B)} = \frac{p(\mathcal{M}_A|\mathbf{d})}{p(\mathcal{M}_B|\mathbf{d})}$$

“So, what do I do with this?”

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If BF/POR is really huge, favor \mathcal{M}_A

If BF/POR is really small, favor \mathcal{M}_B

Otherwise, it's not very decisive.

“So, what do I do with this?”

$$\frac{p(d|\mathcal{M}_A)}{p(d|\mathcal{M}_B)} \times \frac{p(\mathcal{M}_A)}{p(\mathcal{M}_B)} = \frac{p(\mathcal{M}_A|d)}{p(\mathcal{M}_B|d)}$$

favor \mathcal{M}_A

Takeaway #1:
model comparison \neq model selection

If BF/POR is 10

Otherwise, it's not very decisive.

Why model comparison \neq model selection

Model comparison just gives probabilities.

Model selection is a decision based on other (outside) factors, i.e. a cost function/utility.

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Model comparison just gives probabilities.

Model selection is a decision based on other (outside) factors, i.e. a cost function/utility.

Most rigorous thing to do is **average** all models, not select the most probable.

$$p(\theta|\mathbf{d}) = \sum_{k=1}^K p(\theta|\mathbf{d}, \mathcal{M}_k) p(\mathcal{M}_k|\mathbf{d})$$

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Computing FML in practice

$$p(\mathbf{d}|\mathcal{M}) = \int p(\theta|\mathcal{M})p(\mathbf{d}|\theta, \mathcal{M})d\theta$$

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$$p(\mathbf{d}|\mathcal{M}) = \int p(\theta|\mathcal{M})p(\mathbf{d}|\theta, \mathcal{M})d\theta$$



Takeaway #2:
This integral is **HARD**

Computing FML in practice

$$p(\mathbf{d}|\mathcal{M}) = \int p(\theta|\mathcal{M})p(\mathbf{d}|\theta, \mathcal{M})d\theta$$



Takeaway #2:
This integral is **HARD***

* but there's an entire literature on how to compute this efficiently

Thermodynamic Integration (Theory)

1. Start with **parallel-tempering MCMC**.

i.e. multiple MCMCs with likelihoods taken to different powers: $p_{\beta}(\theta|\mathbf{d}) \propto p(\theta)p^{\beta}(\mathbf{d}|\theta)$

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2. FML at β is $p_{\beta}(\mathbf{d}) = \int p(\theta)p^{\beta}(\mathbf{d}|\theta)d\theta$

3. Ultimately, derive...

$$p(\mathbf{d}) \approx \exp \left[\int_0^1 d\beta \langle \log p(\mathbf{d}|\theta) \rangle_{\beta} \right]$$

“average” log-likelihood at β



Thermodynamic Integration (Practice)

$$p(\mathbf{d}) \approx \exp \left[\int_0^1 \mathbf{d}\beta \langle \log p(\mathbf{d}|\theta) \rangle_{\beta} \right]$$

Advantages:

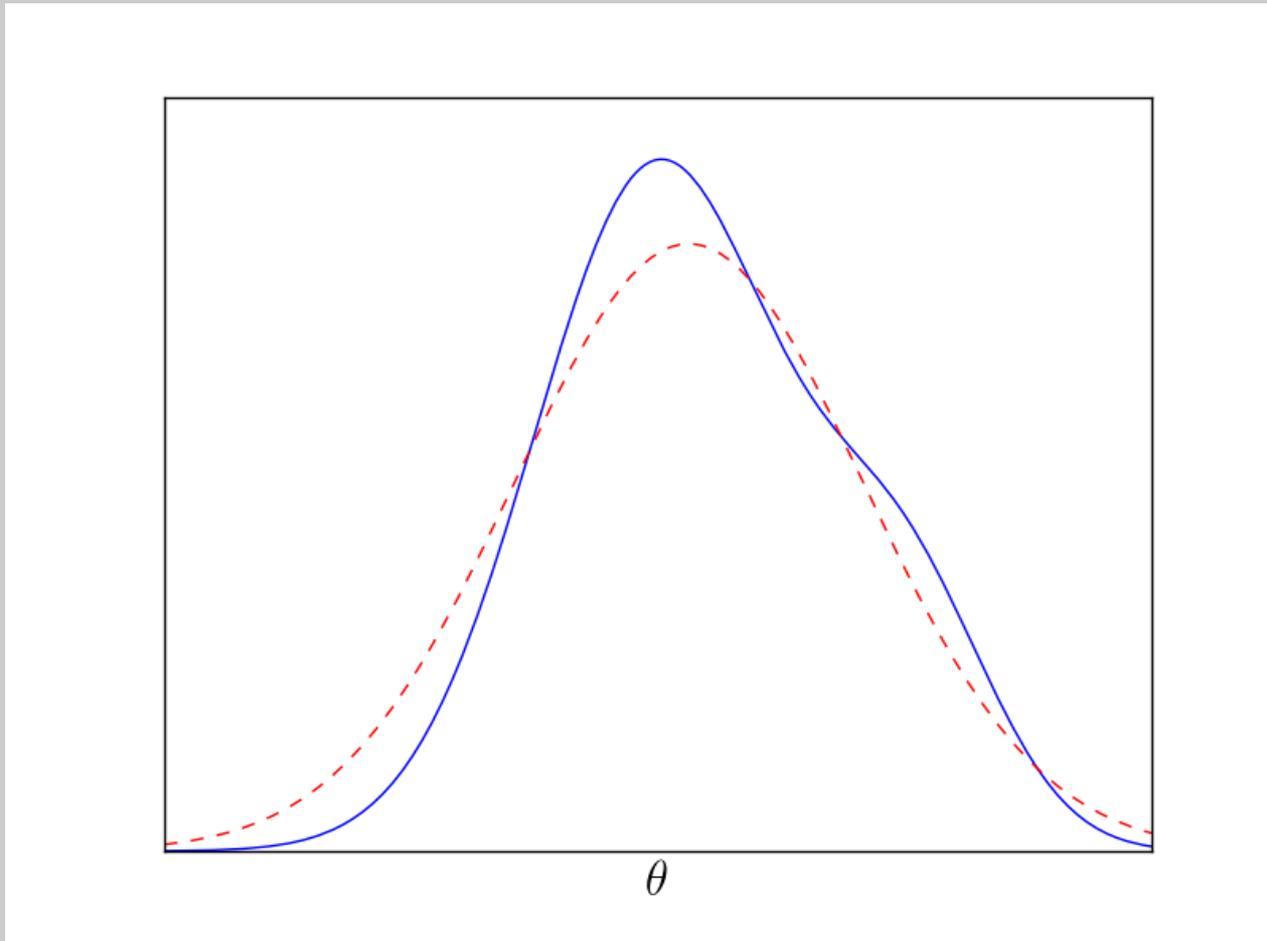
1. A nice side effect of performing PTMCMC
2. Already implemented in emcee*

Caveats:

Need a robust estimate of $\langle \log p(\mathbf{d}|\theta) \rangle_{\beta}$
at every β

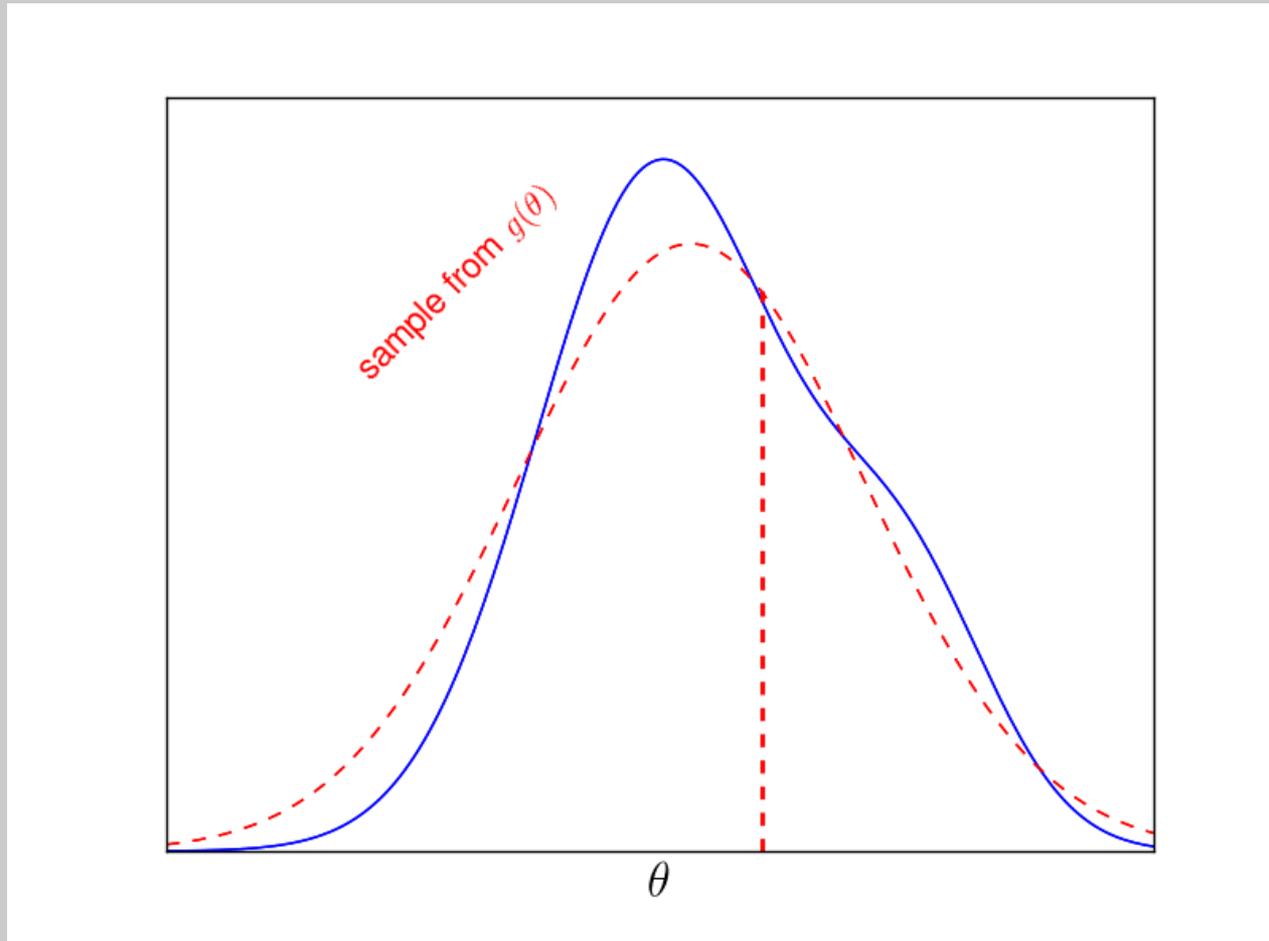
* dan.iel.fm/emcee/current/user/pt

Importance Sampling (Theory)



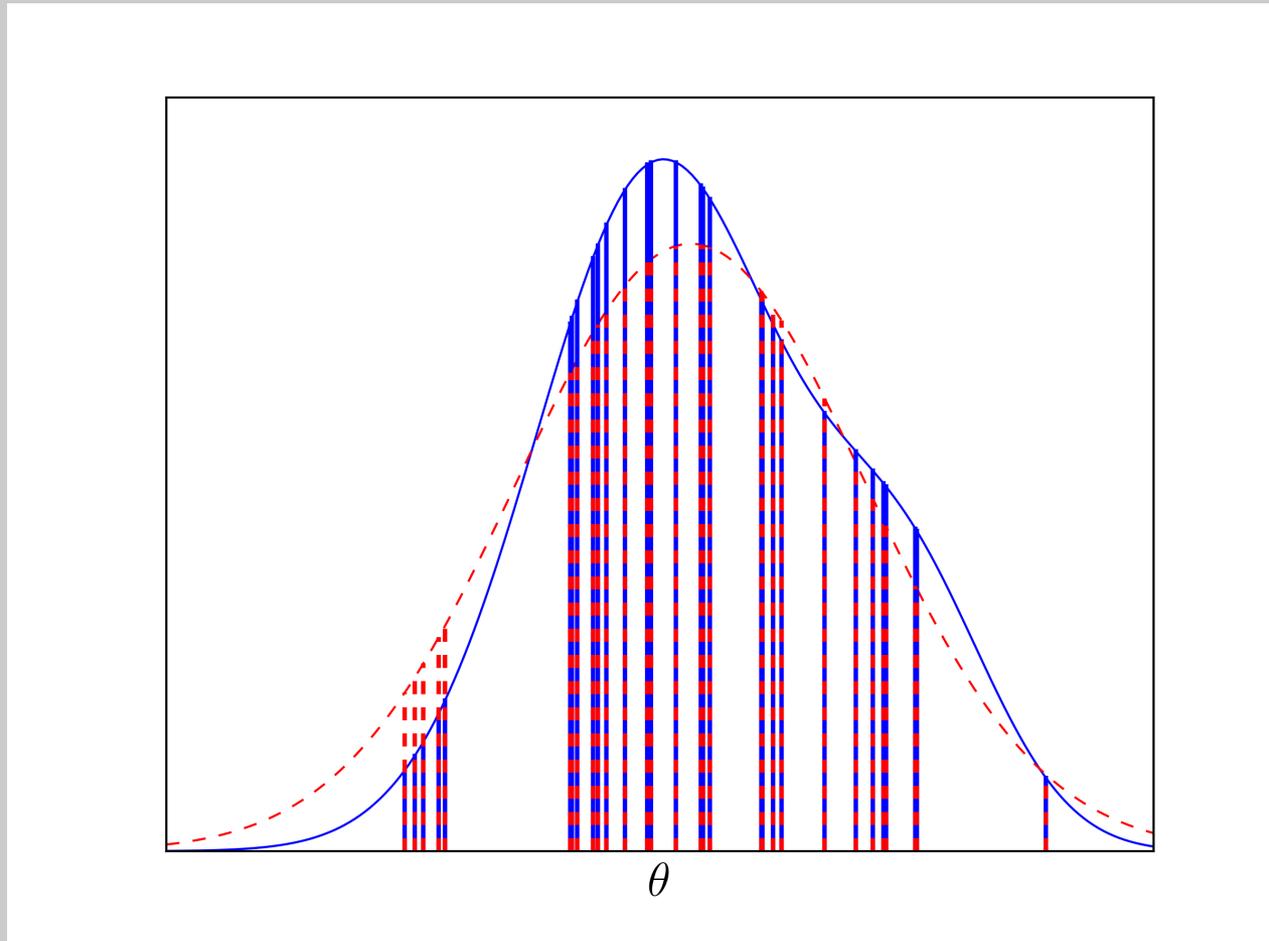
$$p(d) = \int \frac{p(\theta)p(d|\theta)}{g(\theta)} g(\theta) d\theta$$

Importance Sampling (Theory)



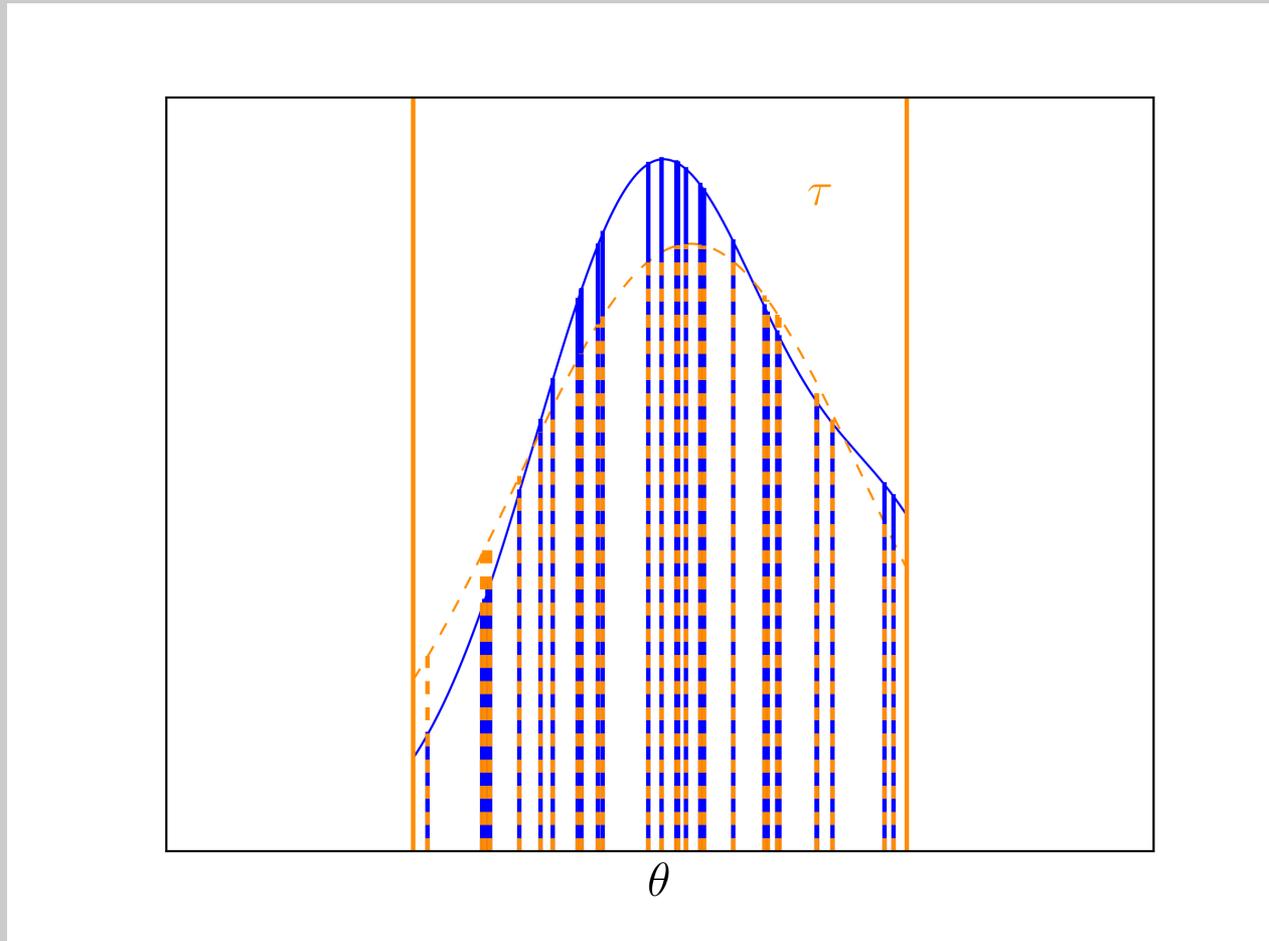
$$p(d) = \int \frac{p(\theta)p(d|\theta)}{g(\theta)} g(\theta) d\theta$$

Importance Sampling (Theory)



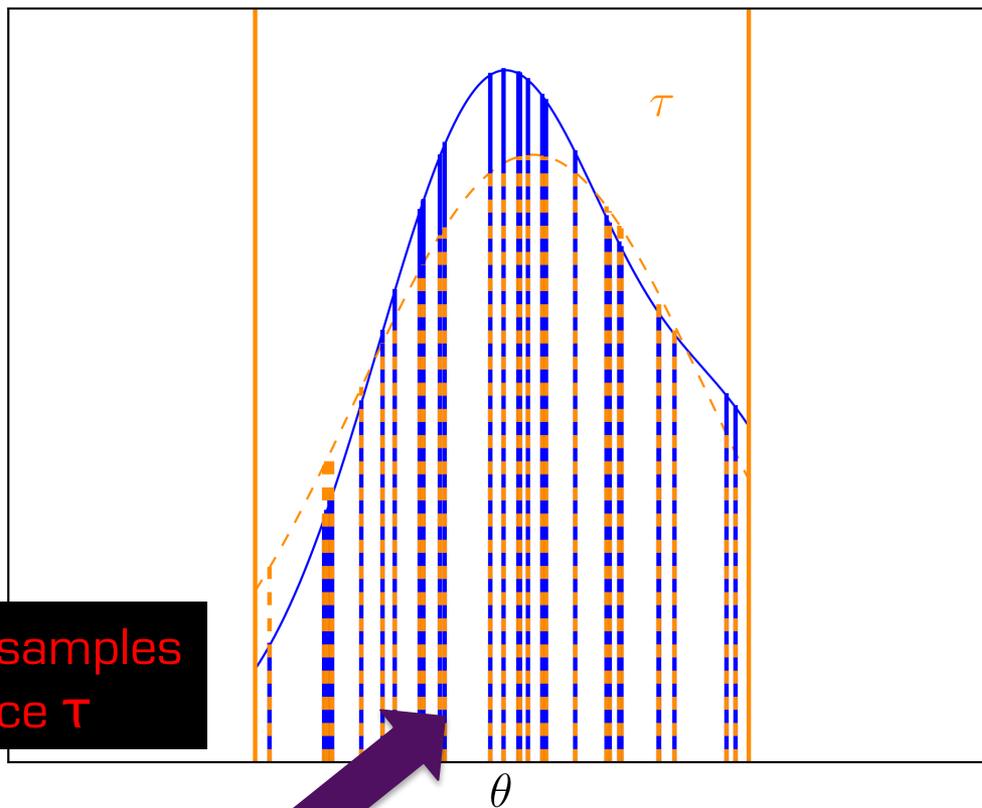
$$\widehat{p(d)} \approx \frac{1}{N} \sum_{\theta_i \sim g(\theta)} \frac{p(\theta_i)p(d|\theta_i)}{g(\theta_i)}$$

Importance Sampling (Theory)



$$\dots = \frac{1}{N} \sum_{\theta_i \sim g_\tau(\theta)} \frac{p(\theta_i)p(d|\theta_i)}{g_\tau(\theta_i)}$$

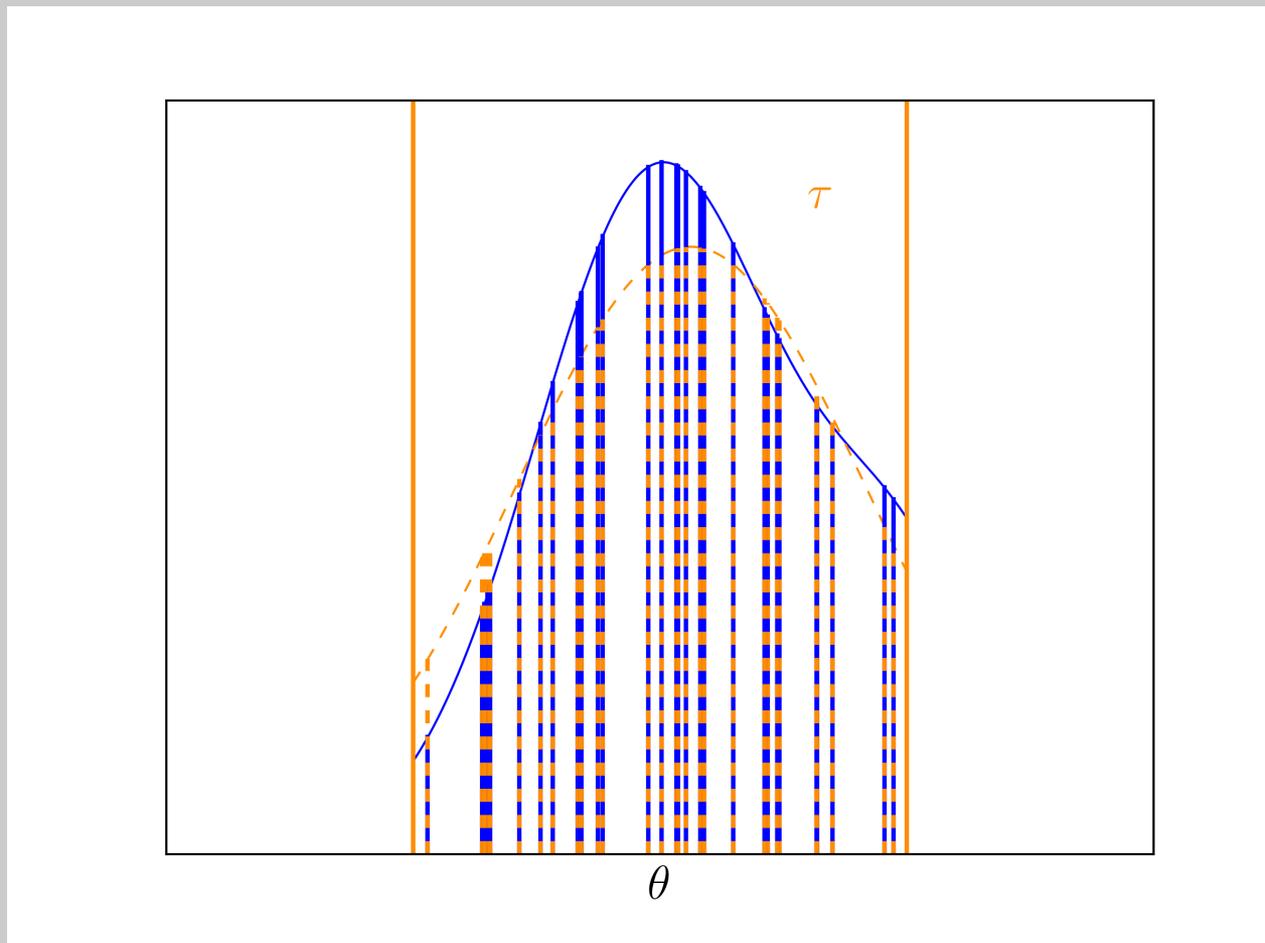
Importance Sampling (Theory)



Fraction of MCMC samples to reside in subspace \mathcal{T}

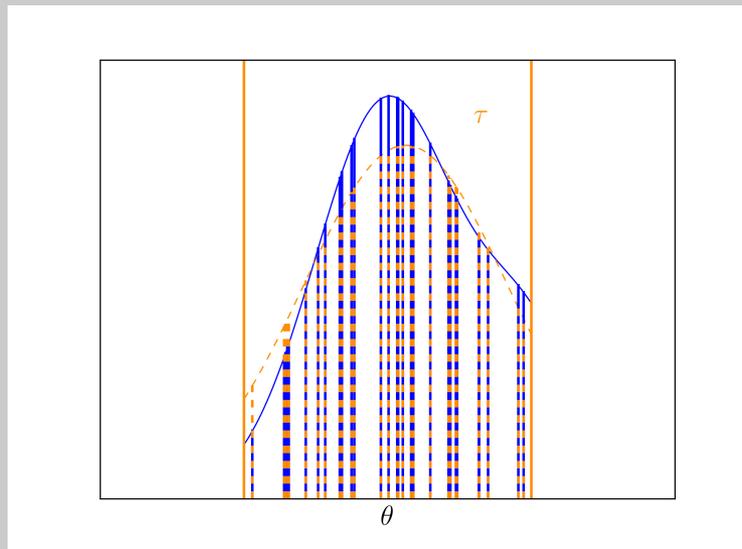
$$f_{\text{MCMC}} \times \widehat{p(d)} \approx \frac{1}{N} \sum_{\theta_i \sim g_{\mathcal{T}}(\theta)} \frac{p(\theta_i)p(d|\theta_i)}{g_{\mathcal{T}}(\theta_i)}$$

Importance Sampling (Theory)



$$\widehat{p(d)} \approx \frac{1}{N \times \mathbf{f}_{\text{MCMC}}} \sum_{\theta_i \sim g_\tau(\theta)} \frac{p(\theta_i)p(d|\theta_i)}{g_\tau(\theta_i)}$$

Importance Sampling (Practice)



$$\widehat{p(d)} \approx \frac{1}{N} \sum_{\theta_i \sim g(\theta)} \frac{p(\theta_i)p(d|\theta_i)}{g(\theta_i)}$$

$$\widehat{p(d)} \approx \frac{1}{N \times f_{\text{MCMC}}} \sum_{\theta_i \sim g_{\tau}(\theta)} \frac{p(\theta_i)p(d|\theta_i)}{g_{\tau}(\theta_i)}$$

Advantages:

1. Embarrassingly parallel
2. Have a posterior sample? Already partway there!

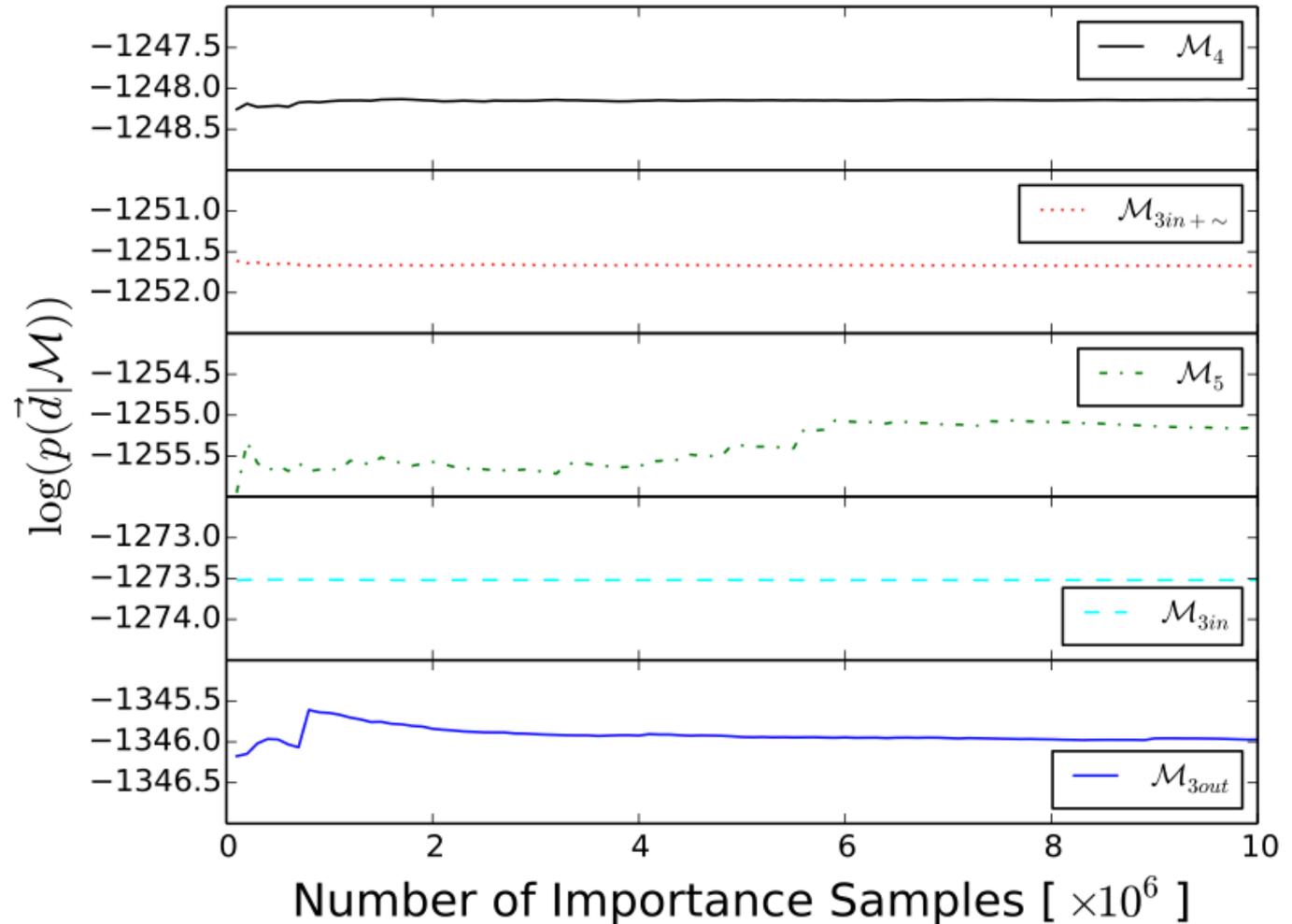
Caveats:

1. Performance depends on chosen $g(\theta)$ or $g_{\tau}(\theta)$
2. Needs a robust value of f_{MCMC}

Importance Sampling (Gliese 876)



Seth Pritchard
(undergrad at
UT San Antonio)



Importance Sampling (Tutorial)

github.com/benelson/FML

Features:

- generate synthetic RVs of input planetary system
- MCMC with n-body model
- step-by-step importance sampling tutorial

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Features:

- generate synthetic RVs of input planetary system
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Also check out John Boisvert's (UNLV) poster

Uncovering System Architectures Near 2:1 Resonance

More methods

Nested Sampling

Science: determining evidence for exomoons (Kipping+ 2013), functional form of eccentricity distribution (Kipping 2013), testing n-planets in RV observations (Brewer & Donovan 2015)

Publicly available code: Multinest (Feroz & Hobson 2008, Feroz 2009), DNest3/4 (Brewer+ 2010), Transdimensional MCMC (Brewer & Donovan 2015)

Geometric Path Monte Carlo

Science: testing n-planets in RV observations (Hou, Goodman, Hogg 2014)

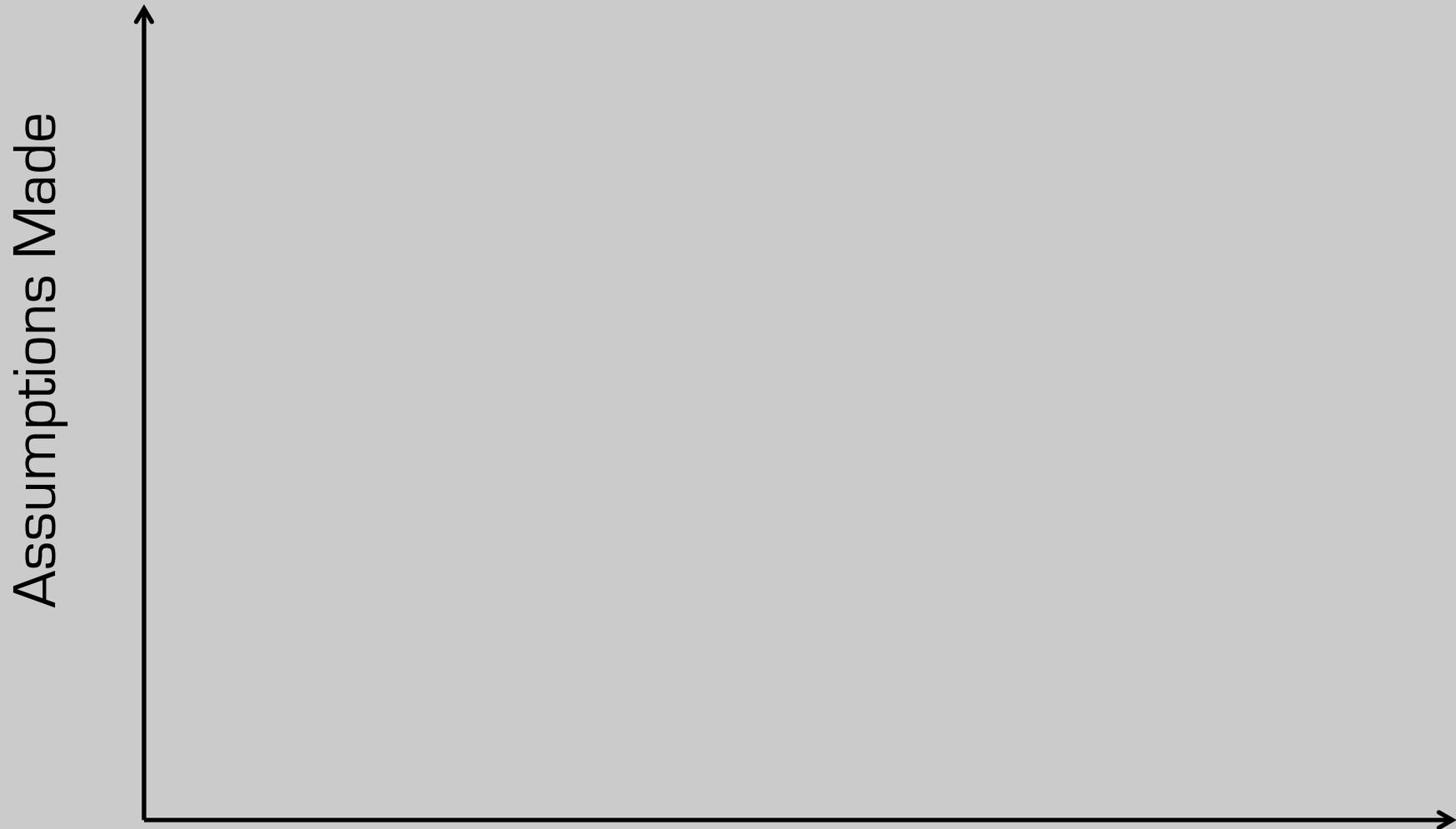
Savage-Dickey Density Ratio

Specializes in comparing nested models with 1-2 parameter difference

Science: Mass of Mars-sized Kepler-138b (Jontof-Hutter+ 2016)

Outline

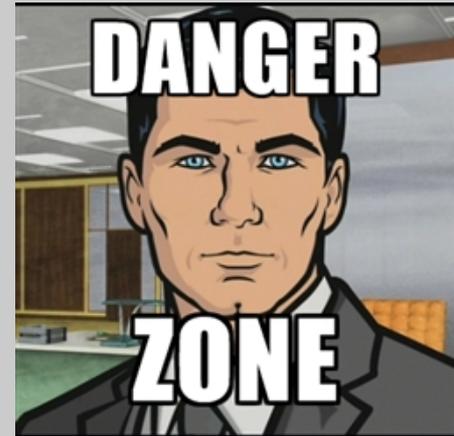
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Computational Difficulty

Assumptions Made

AIC
BIC

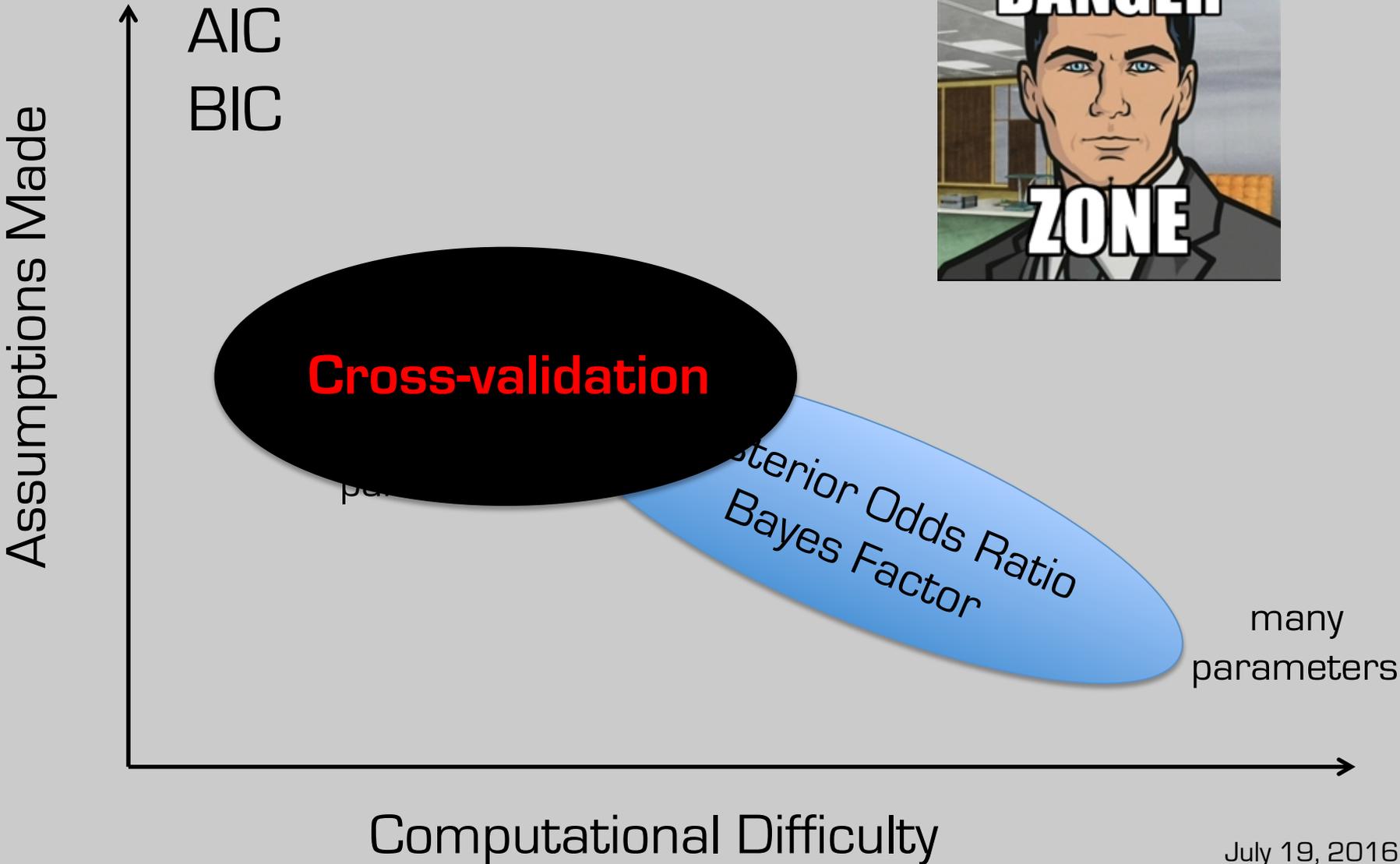
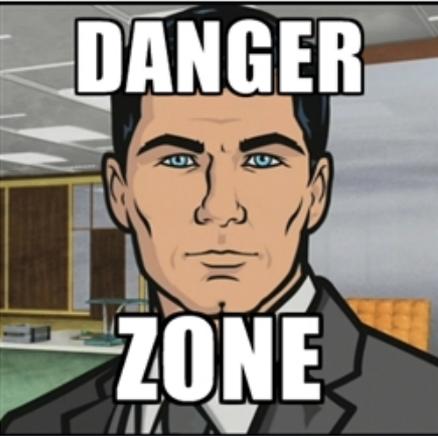


few
parameters



many
parameters

Computational Difficulty



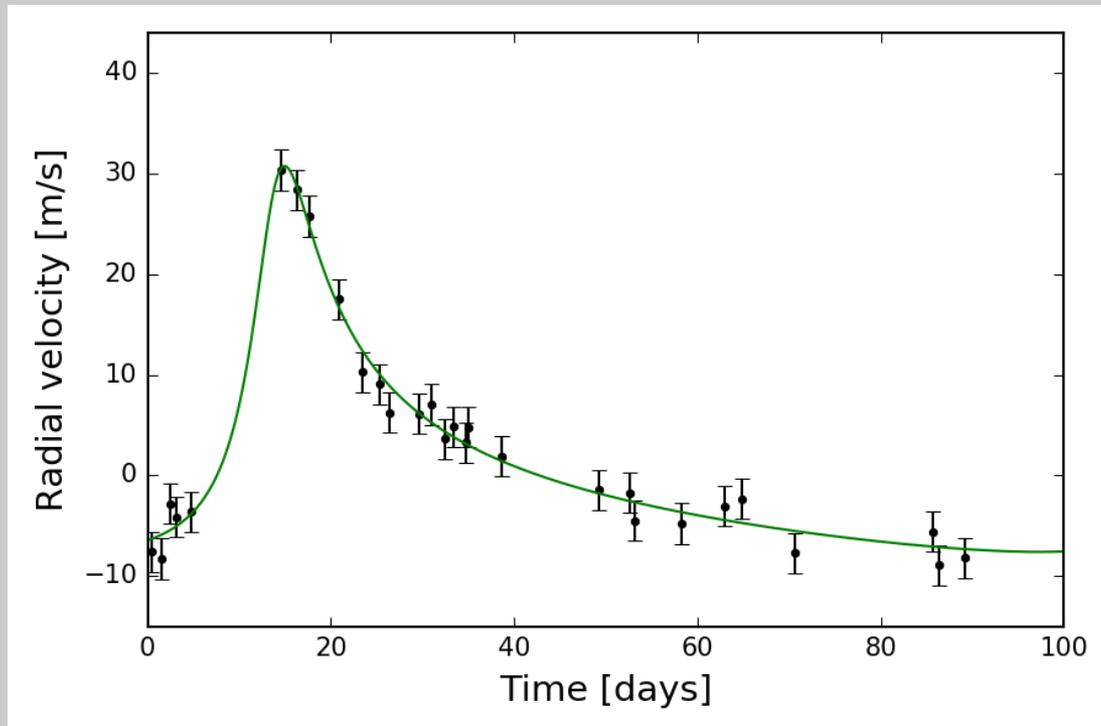
Cross-validation

```
cvl = 1.;
```

```
for (d in data){
```

```
  get parameters  $\theta_{(d)}$  that optimize on data WITHOUT d;
```

```
  cvl *= p(d |  $\theta_{(d)}$ ,  $\mathcal{M}$ ); }
```



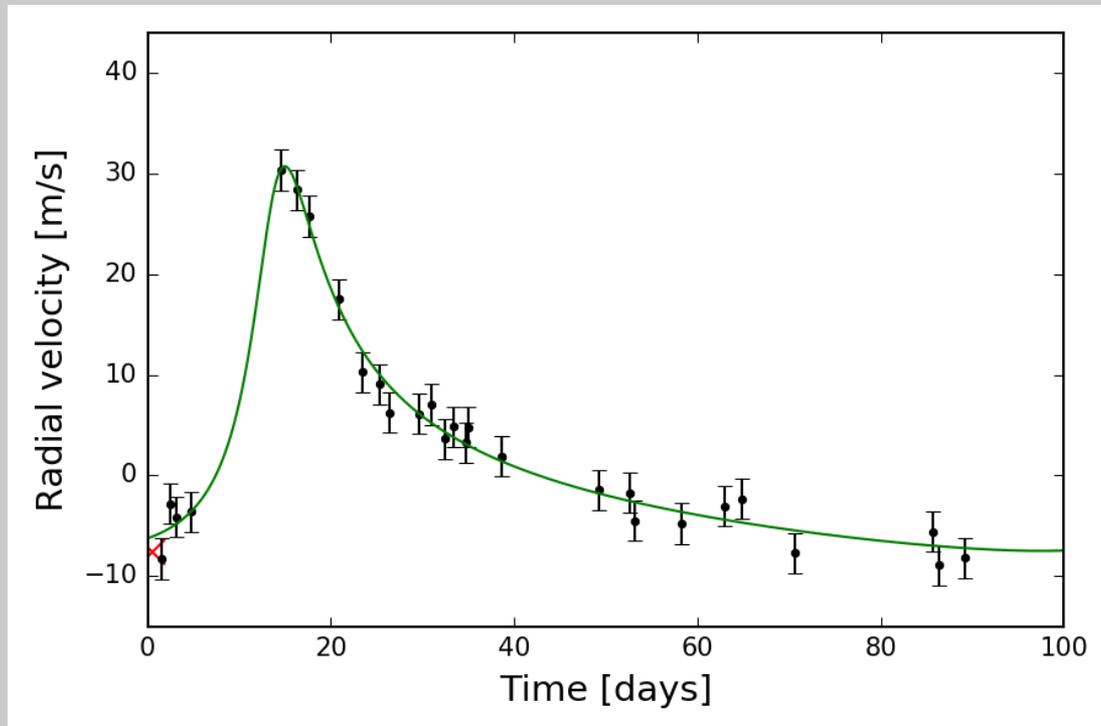
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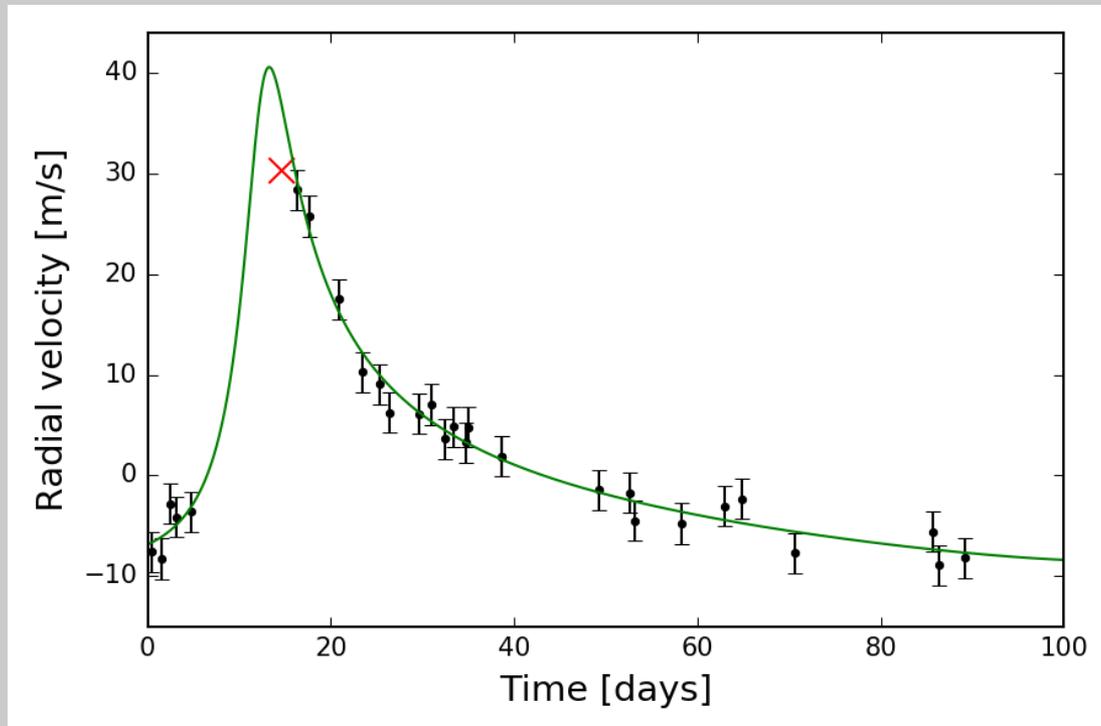
Cross-validation

```
cvl = 1.;
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```

```
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```

```
  cvl * = p(d |  $\theta_{(d)}$ ,  $\mathcal{M}$ ); }
```



Cross-validation

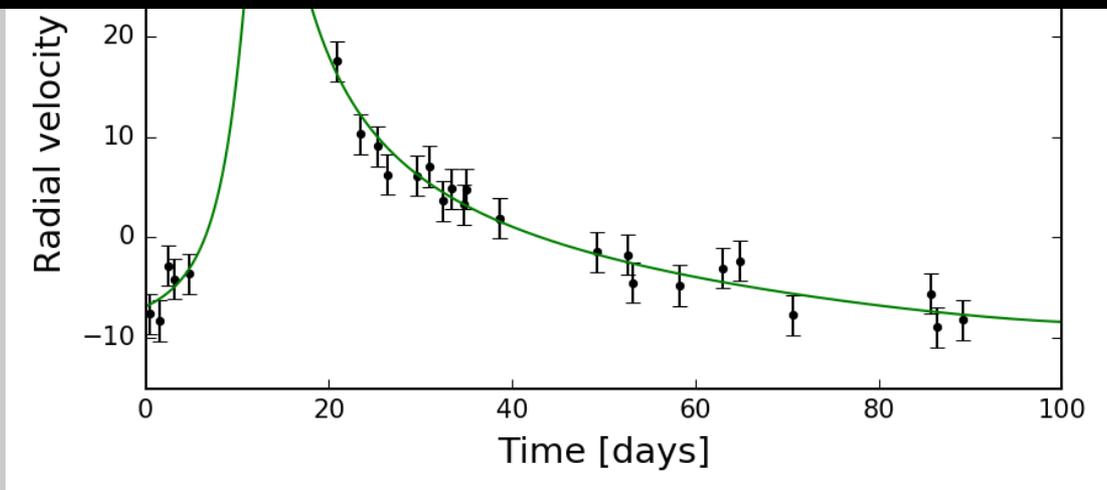
```
cvl = 1.;
```

```
for (d in data){
```

```
  get parameters  $\theta_{(-d)}$  that optimize on data WITHOUT d;
```

```
cvl * =
```

The model with the largest cross-validation likelihood (cvl) is preferred.



Takeaway #3: General Recommendations

1. Do you need to make many rough decisions quickly (i.e. milliseconds)?

AIC/BIC

2. Do you have decent computational resources and really understand your priors/utility?

Bayes factor/posterior odds ratio

3. Is your problem somewhat in between?

Cross-validation

Conclusions

model comparison \neq model selection: how to decide depends on your utility

For 3+ parameter models, computing FML is hard. But it's an active problem in exoplanet research.

For tutorial on using importance sampling to compute FMLs: github.com/benelson/FML