

The Virtual Observatory and The Virtual Astronomical Observatory

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The Virtual Observatory

Astronomy being transformed by Information Technology

- Generate large data sets
 - High speed digital signal processing
 - Large format CCDs
 - HPC-based simulations



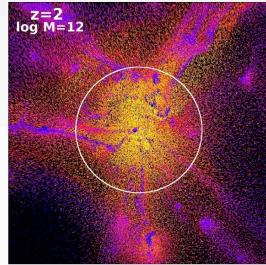
- Access, analyze, archive
- New telescopes only going to exacerbate challenge















The Virtual Observatory

The VO is a data discovery, access, and integration facility

- Image, spectral, time series data
- Catalogs, databases
- Transient event notices
- Software and services
- Distributed computing authentication, authorization, process management



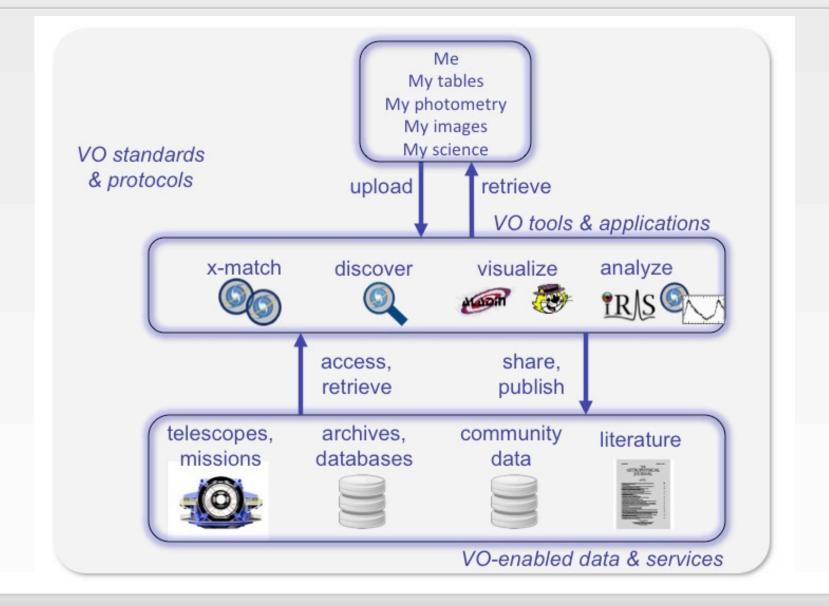
Application inter-communication

International coordination and collaboration





The Virtual Observatory







VO impact and penetration

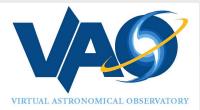
Astronomers use VO every day without realizing it

- A few examples
 - Major archives and data centers are VO compliant
 - Hubble Legacy Archive is built with "VO Inside"
 - Dark Energy Survey pipeline and archive
 - WIYN One-Degree Imager pipeline and archive
 - Pan-STARRS
 - LSST development
- ~1M VO-compliant service requests every 2 weeks from several sites









- National Virtual Observatory (NVO) development effort, 2001-08
 - \$14M, 17 organizations
 - NSF Information Technology Research program
- Virtual Astronomical Observatory (VAO) operational facility, 2010-2015
 Managed by VAO,LLC, co-owned by AUI and AURA
 - VAO Board of Directors (J. Gallagher, chair)
 - Science Council, G. Fabbiano (CfA, acting)
 - Users Committee, C. Miller (U. Mich.)



http://www.usvao.org/





Research Collaborations

 Call for Proposals issued in January

VAO to provide in-kind resources to further a well-defined research project and/or expose valuable data to the community

Think about next year!



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Call for Proposals

Science Collaborations with the Virtual Astronomical Observatory

8 January 2012 Proposals due: 1 March 2012

The US Virtual Astronomical Observatory (VAO) seeks proposals from the astronomical research community for science collaborations that utilize virtual observatory capabilities or develop novel virtual observatory tools to produce new research results. Proposals may be in any area of astronomy, but projects that benefit from the integration of data in multiple wavelength regions, that involve analysis of time-domain data sets, and/or utilize large databases or data sets are the most germane to the VAO. The VAO Science Council has suggested, for example, that multi-wavelength studies of M31 would be a good candidate for such a collaboration, but proposals in any area of astronomy are welcome and will be considered.

The VAO cannot directly fund these collaborations, but will provide in-kind support in using and adapting VO technologies in the selected projects. Such support may include VAO technical staff visiting a research team or covering travel costs for members of a research team to visit a VAO organization. Ongoing support will be in the form of telecons, e-meetings (e.g., using WebEx), a Wiki site, or other on-line collaborative tools.

Accepted

- Serving data and models of the Magellanic Clouds from Spitzer and Herschel legacy programs (PI: Meixner)
- Serving light curves from the American Association of Variable Star Observers (PI: Templeton)
- VO image interfaces to the NASA Extragalactic Database (PI: Schombert)
- VO-compliant databases for the Evolutionary Map of the Universe and Variable and Slow Transient Survey Science Projects (PI: Murphy)





VAO Science Tools

http://www.usvao.org/







Time Series Tool



- Discover time series data
 - Harvard Time Series Center (TSC)
 - NASA Exoplanet Archive at IPAC/Caltech
 - Catalina Real-Time Transient Survey at CACR/Caltech
- Analyze with the NASA **Exoplanet Archive's** periodogram application
- Pathfinder for developing utility that interconnects time series data repositories





Time Series Tool II

Кер	ler planetary candidates	
star_id	9941662	
koi	13.01	
field	The Kepler Field	
ra	286.971159	
dec	46.868401	
pm_ra	0.000000	
pm_dec	0.00000.0	
pm_total	0.00000.0	
gal_lat	16.806619	
gal_lon	77.505951	
umag	null	
gmag	9.697000	
rmag	10.045000	
imag	10.548000	
zmag	10.020000	
gredmag	null	
d51mag	9.669000	
jmag	9.465000	
hmag	9.455000	
kmag	9.425000	
kepmag	9.958000	
filter	Kepler Magnitude	
transit duration	3.202900	

TrES (Trans-atlantic Exc	oplanet Su	ırvey) observ	ations of the Kepler field
	star_id	TrES_Lyr1_00173	
	region	TrES_Lyr1	
	ra	286.971250	
	dec	46.868330	
	bmag	10.704000	
	vmag	10.476000	
	rmag	11.504000	
	rerr	0.039100	
	starthjd	2453541.787548	
	endhjd	2453616.734163	
	lcfil	R	
	npts	11398	
	lcdisp	0.140000	
	lechisq	13356.299805	
	n5sig	0	
	f5sig	0.000000	
	public	Y	
	reference_ids	8	
	etss_id	173	
	download	download	
	plot	plot	
	periodogram	periodogram	

Kepler

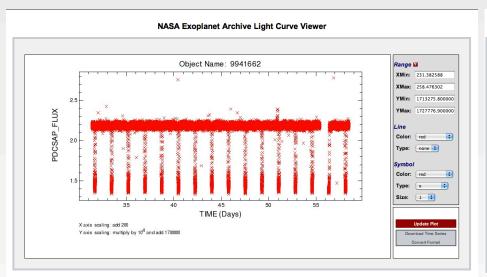
TrES

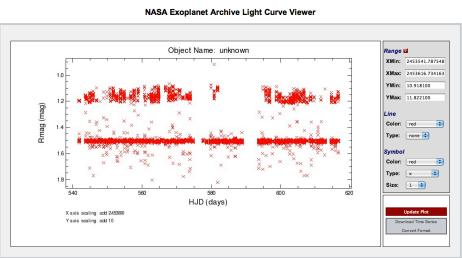
After searching on an object, examine its entries in a time series data set ...





Time Series Tool III





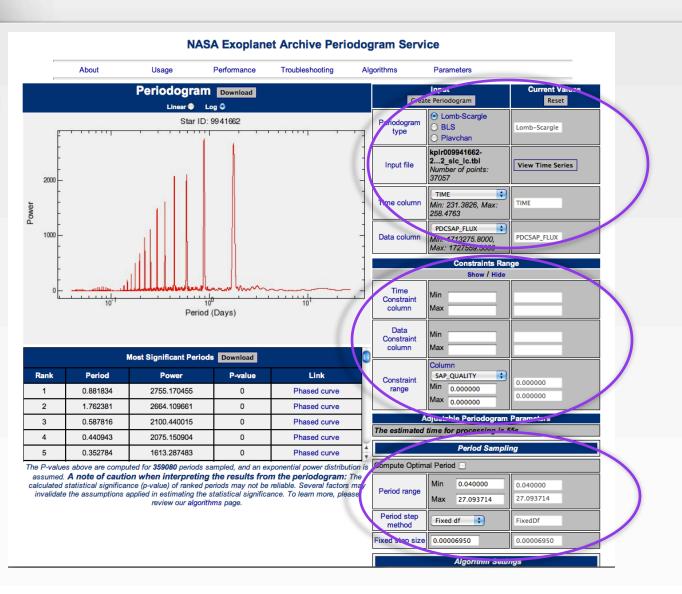
Kepler

TrES

Plot the time series as a "quick look" ...



Tool Series IV



Analyze!





Data Discovery Tool



arXiv.org > astro-ph > arXiv:1207.1715

Search or Article-id

Astrophysics > Earth and Planetary Astrophysics

SOPHIE velocimetry of Kepler transit candidates VII. An additional companion in the KOI-13 system

A. Santerne, C. Moutou, S. C. C. Barros, C. Damiani, R. F. Díaz, J.-M. Almenara, A. S. Bonomo, F. Bouchy, M. Deleuil, G. Hébrard

(Submitted on 6 Jul 2012)

We report the discovery of a new stellar companion in the KOI-13 system. KOI-13 is composed by two fast-rotating A-type stars of similar magnitude. One of these two stars hosts a transiting planet discovered by Kepler. We obtained new radial velocity measurements using the SOPHIE spectrograph at the Observatoire de Haute-Provence that revealed an additional companion in this system. This companion has a mass between 0.4 and 1 Msun and orbits one of the two main stars with a period of 65.831 \pm 0.029 days and an eccentricity of 0.52 \pm 0.02. The radial velocities of the two stars were derived using a model of two fast-rotating line profiles. From the residuals, we found a hint of the stellar variations seen in the Kepler light curve with an amplitude of about 1.41 km/s and a period close to the rotational period. This signal appears to be about three order of magnitude larger than expected for stellar activity. From the analysis of the residuals, we also put a 3-sigma upper-limit on the mass of the transiting planet KOI-13.01 of 14.8 Mjup and 9.4 Mjup, depending on which star hosts the transit. We found that this new companion has no significant impact on the photometric determination of the mass of KOI-13.01 but is expected to affect precise infrared photometry. Finally, using dynamical simulations, we infer that the new companion is orbiting around KOI-13B while the transiting planet candidate is expected to orbit KOI-13A. Thus, the transiting planet candidate KOI-13.01 is orbiting the main component of a hierarchical triple system.

Comments: Accepted in A&A. 4 pages including 4 figures and the RV table

Subjects: Earth and Planetary Astrophysics (astro-ph.EP); Solar and Stellar Astrophysics (astro-ph.SR)

Cite as: arXiv:1207.1715v1 [astro-ph.EP]

Question:

1.Can I combine my data with existing data to understand an object better?

E.g., KOI-13 system

2.Do I need to acquire additional data or do they exist already?

a.k.a. observational or proposal planning





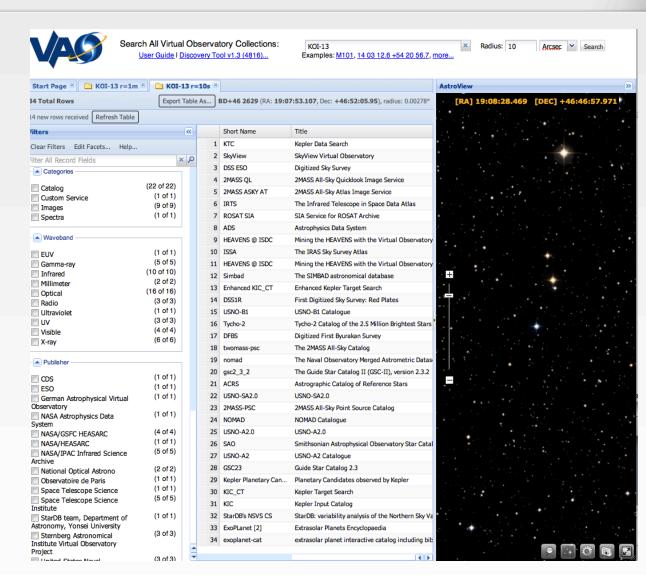
Data Discovery Tool

Data Discovery Tool searches VO-aware collections for data

- Catalog, images, spectra
- Waveband

• . . .

➤Do your data need to be in a VO-aware service?

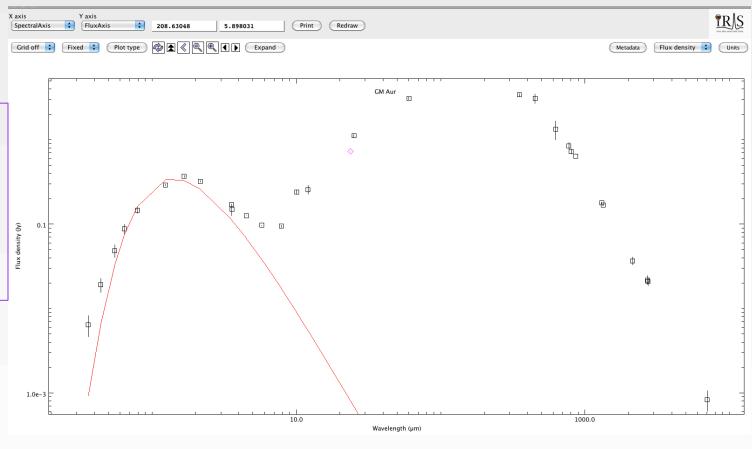






SED tool: Iris

SEDs needed to understand emission physics e.g., star vs. disk

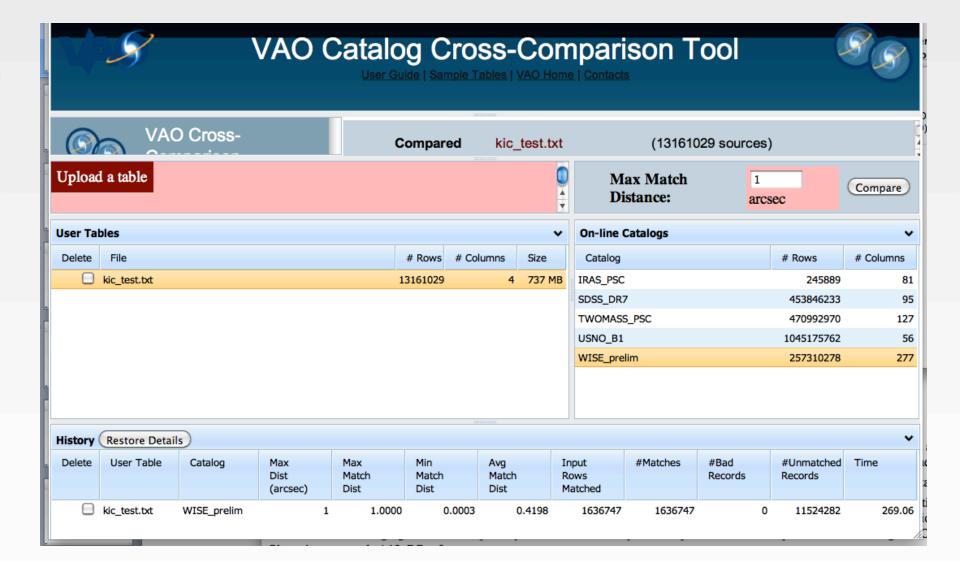


GM Aur





Cross-Matching







Cross-Matching and Cross-Identification



arXiv.org > astro-ph > arXiv:1207.0521

earch or Article-id

Astrophysics > Earth and Planetary Astrophysics

Confusion limited surveys: using WISE to quantify the rarity of warm dust around Kepler stars

G. M. Kennedy, M. C. Wyatt

(Submitted on 2 Jul 2012)

We describe a search for infra-red excess emission from dusty circumstellar material around 180,000 stars observed by the Kepler and WISE missions. This study is motivated by i) the potential to find bright warm disks around planet host stars, ii) a need to characterise the distribution of rare warm disks, and iii) the possible identification of candidates for discovering transiting dust concentrations. We find about 8,000 stars that have excess emission, mostly at 12um. The positions of these stars correlate with the 100um background level so most of the flux measurements associated with these excesses are spurious. We identify 271 stars with plausible excesses by making a 5MJy/sr cut in the IRAS 100um emission. The number counts of these excesses, at both 12 and 22um, have the same distribution as extra-Galactic number counts. Thus, although some excesses may be circumstellar, most can be explained as chance alignments with background galaxies. The one exception is a 22um excess associated with a relatively nearby A-type star that we were able to confirm because the disk occurrence rate is independent of stellar distance. Despite our low detection rate, these results place valuable upper limits on the distribution of large mid-infrared excesses; e.g. fewer than 1:1000 stars have 12um excesses (F_ obs/F_star) larger than a factor of five. In contrast to previous studies, we find no evidence for disks around 1790 stars with candidate planets (we attribute one significant 12um excess to a background galaxy), and no evidence that the disk distribution around planet hosts is different to the bulk population. Higher resolution imaging of stars with excesses is the best way to rule out galaxy confusion and identify more reliable disk candidates among Kepler stars. A similar survey to ours that focusses on nearby stars would be well suited to finding the distribution of rare warm disks.

VAO Cross-				Compared kic_test.txt			(13161029 sources)						
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Delete	File				# Rows	# Column	ns Size		Catalog			# Rows	# Column
	kic_test.txt				13161029		4 737	мв	IRAS_PSC			245889	
									SDSS_DR7			453846233	
									TWOMASS	_PSC		470992970	
									USNO_B1			1045175762	
									WISE_prel	im		257310278	
listory	Restore Detai	IC)											
	(Hestore Detail												
Delete	User Table	Catalog	Max	Max	Min	A			put	#Matches	#Bad	#Unmatched	Time

Distinguish cross-matching and cross-identification

•Cross-matching → two objects are at the same position on the sky (within uncertainties)

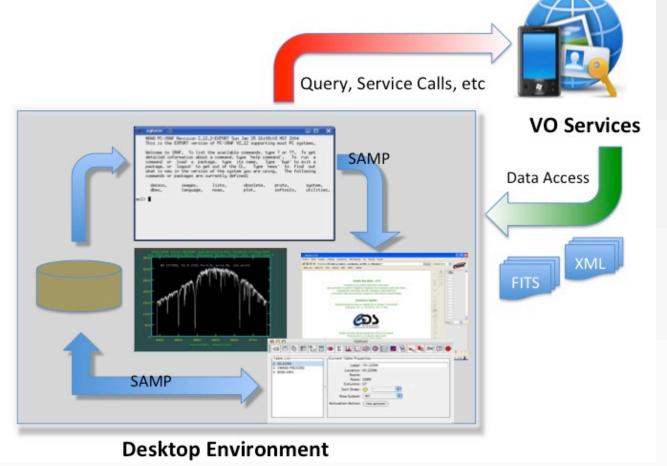
Cf. brown dwarf searches

 Cross-identification → two sources at different wavelengths represent the same object

On-going research into probabilistic approaches

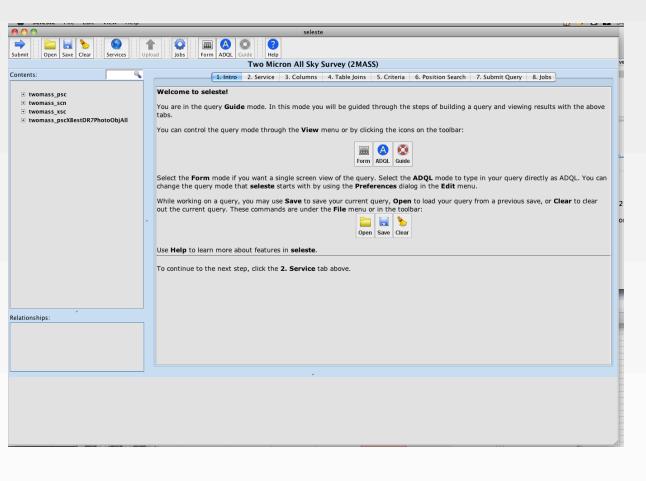


VAO-IRAF Integration



- 2000 registered **IRAF** users
- ~ 5000 total users
- > 700 IRAF tasks now VO-aware
- > 1000 downloads of **IRAF** v. 2.16 as of 2012 April 1

seleste



- Intuitive yet flexible and powerful access to catalogs, or any tabular data
- Discover and browse catalogs and databases
- Full access to data and metadata
- Quick look results
- Customizable for finetuned results
- SQL queries, with clickand-drag GUI
- Scriptable via command line interface
- Query results saved or passed to other VO tools



VOStat

- VOStat, rev 2 is a Web-service providing a suite of statistical analyses for VO datasets.
 - Developed at Penn State
 - http://vostat.org/
- Java-based GUI to the R public domain statistical software environment
 - > 60,000 functions in > 3500 packages
 - http://r-project.org
- Input files obtained via SAMP, URL, or user (ASCII, FITS, VOTable)
- R code provided to assist user in further analysis on local computer.
- VOStat provides ~ 60 statistical functions
 2D & 3D plots, smoothers (kernel, ASH), hypothesis tests (K-S, A-D), regressions (linear, robust, local), PCA, normal mixtures, hierarchical clustering, spatial (k-NN, Moran's I, Ripley's K, 2-pt correlation), directional data, non-detections (Kaplan-Meier, Lynden-Bell-Woodroofe, 2-sample tests), time series (ACF, ARMA, Fourier)



Coming Attractions

- Renewed emphasis on Time Series Tool
- Python interfaces to VO libraries and services
- Inter-operability
 - E.g., DDT → Iris
 - In alpha/pre-beta stage currently
- VO-aware versions of other software packages
 - e.g., CIAO (Chandra Interactive Analysis of Observations), CASA (Common Astronomical Software Applications
- Data publishing initiative
 - How do I share my data with the community?
 - viz. NSF Data Management and NASA ROSES policies
 - Currently in pilot stage

