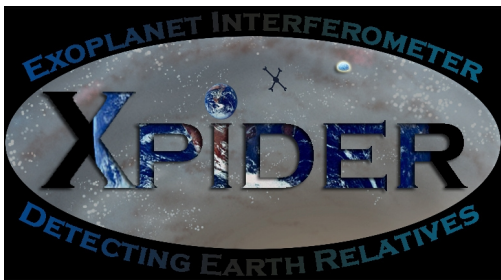


EXOPLANET INTERFEROMETER

XPIDER

DETECTING EARTH RELATIVES





Big-picture

Past and Present: HST, Spitzer, Kepler, Ground-based

The Next Decade: Ground-based, Kepler, JWST, TPF-Darwin

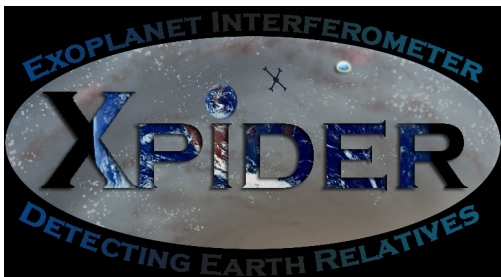
What has been done: physical properties, atmospheres of hot Jupiters

What can be done: Atmospheres of transiting super-Earths (JWST)

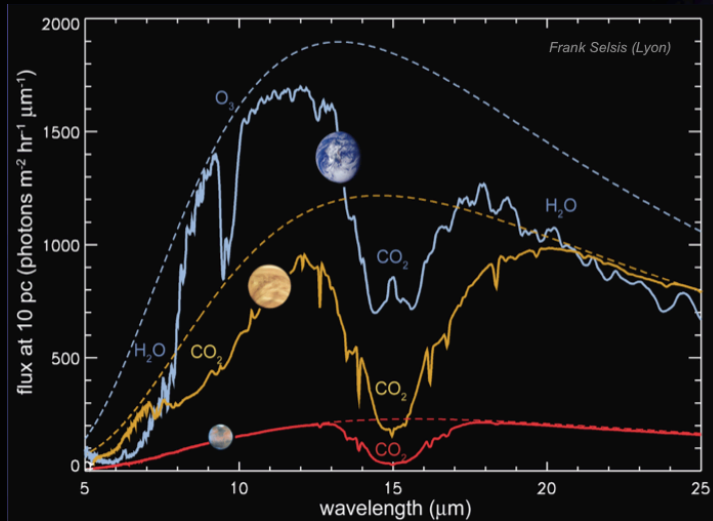
What cannot be done ?

The Question to ask:

What will humanity look forward to in 10 years ?



The exo-Earth argument

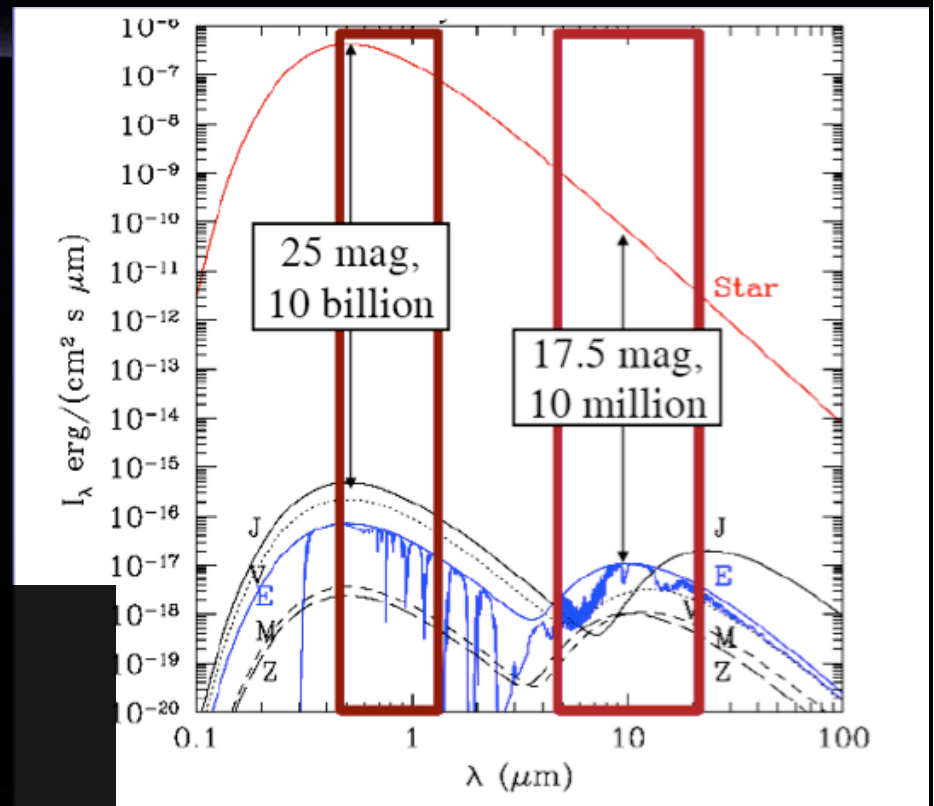


For detection HZ (SNR~5):

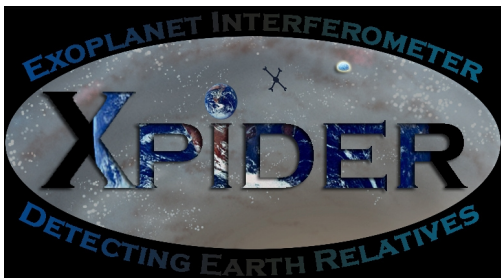
- M stars, 32 hours, 26 targets, = 35 days total integration
- K stars, 64 hours, 25 targets, = 67 days total integration
- G stars, 128 hours, 38 targets, = 203 days total integration
- ~ 300 days of integration needed for detection phase, assume 50% integration time = 600 days total for detection phase

For characterization H2O (SNR~10)

- M stars, 2.5 days integration
- K stars, 25 days integration
- G stars, 250 days integration

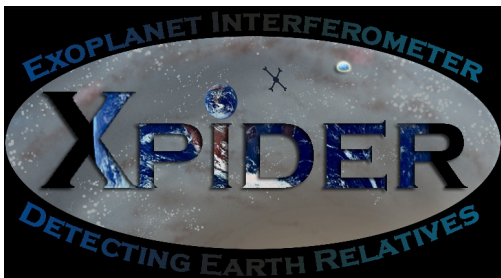


- Follow-up characterization of candidates from RV and transits (90 candidates)
- M,G,K stars within 17 pc
- Bio signatures: H2O, CO2, CH4, O3



Primary Science Matrix

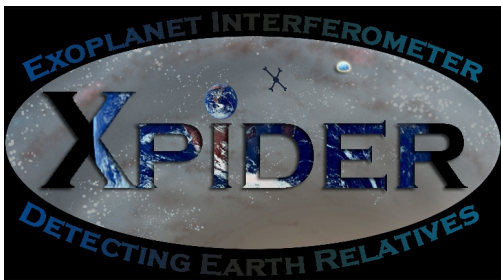
Science Objectives	Measurement Objectives	Measurement Requirements	Instruments	Mission Requirements
Detection and characterization of Habitable planets	<ul style="list-style-type: none"> • Direct imaging • Mid-IR Spectroscopy • Detection 	<ul style="list-style-type: none"> • Nulling interferometry • 10 mas ang res • 6-20 micron spectroscopy 	<ul style="list-style-type: none"> • 4 x 1m telescopes • 500 m base line • mid-IR spectrograph 	<ul style="list-style-type: none"> • Pointing stability • Thermal stability • Multi-year time span
Comparative exoplanetary science	<ul style="list-style-type: none"> • Direct imaging of multiple-planet systems • Direct imaging of range of planets 	<ul style="list-style-type: none"> • Nulling interferometry • 10 mas ang res • 6-20 micron spectroscopy 	<ul style="list-style-type: none"> • 4 x 1m telescopes • Variable base line • mid-IR spectrograph 	<ul style="list-style-type: none"> • Pointing stability • Thermal stability • Multi-year time span
Study of formation and evolution of planetary systems	<ul style="list-style-type: none"> • Direct imaging of debris disks • Direct imaging of young stars 	<ul style="list-style-type: none"> • Imaging interferometry • 10 mas ang res • 6-20 micron spectroscopy 	<ul style="list-style-type: none"> • 4 x 1m telescopes • Variable base line • mid-IR spectrograph 	<ul style="list-style-type: none"> • Pointing stability • Thermal stability
Microlensing	<ul style="list-style-type: none"> • Gravitational lensing images • post-event follow-up 	<ul style="list-style-type: none"> • Imaging interferometry • 10 mas ang res 	<ul style="list-style-type: none"> • 4 x 1m telescopes • Variable base line 	<ul style="list-style-type: none"> • Pointing stability • Thermal stability



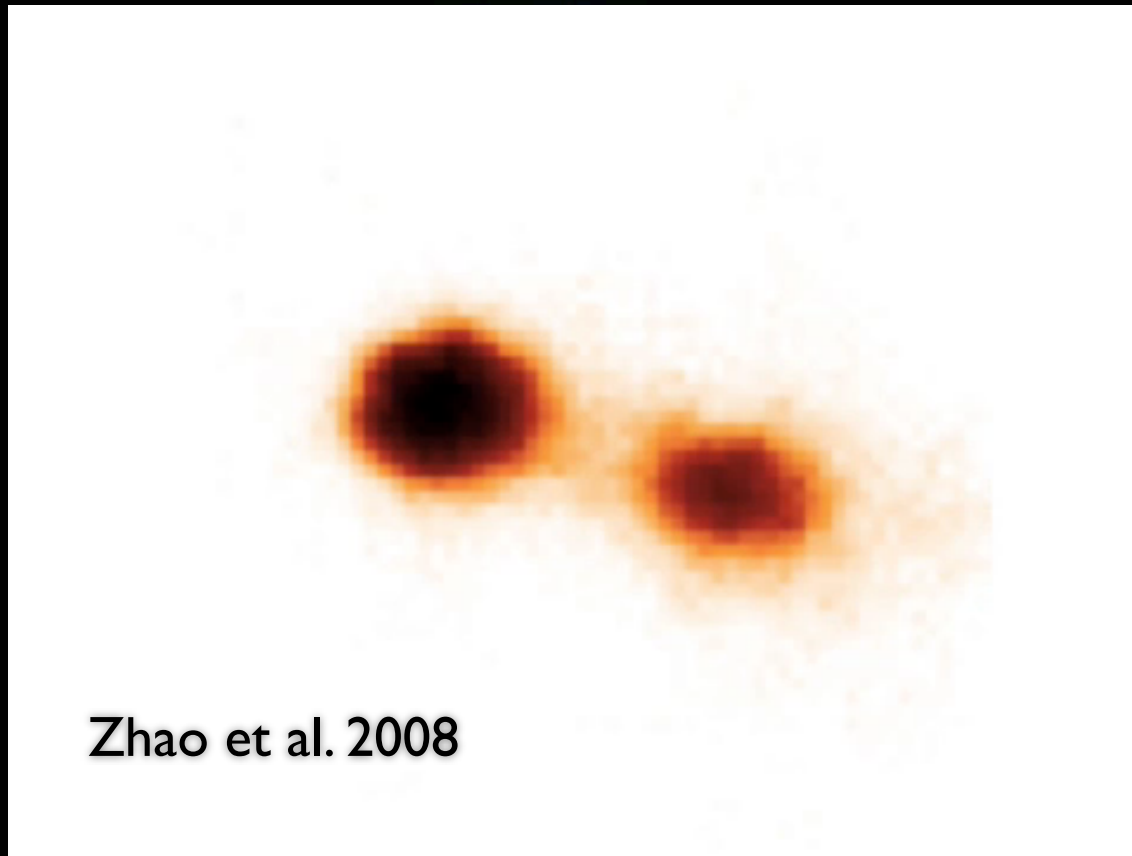
Additional Science Matrix

Science Objectives	Measurement Objectives	Measurement Requirements	Instruments	Mission Requirements
Direct Imaging of circumstellar disks	<ul style="list-style-type: none"> • Direct imaging • Mid-IR Spectroscopy • Density and T profiles 	<ul style="list-style-type: none"> • Interferometry • 10 mas ang res • 6-20 micron spectroscopy 	<ul style="list-style-type: none"> • 4 x 1m telescopes • 500 m base line • mid-IR spectrograph 	<ul style="list-style-type: none"> • Pointing stability • Thermal stability
Interacting Binaries	<ul style="list-style-type: none"> • Roche Lobe imaging • Orbits, mass-transfer • Interacting magnetic fields 	<ul style="list-style-type: none"> • Interferometry • 10 mas ang res • 6-20 micron spectroscopy 	<ul style="list-style-type: none"> • 4 x 1m telescopes • 500 m base line • mid-IR spectrograph 	<ul style="list-style-type: none"> • Pointing stability • Thermal stability
Star spots	<ul style="list-style-type: none"> • Interferometric imaging of stellar surfaces • Differential T measurements 	<ul style="list-style-type: none"> • Interferometry • 10 mas ang res • 6-20 micron spectroscopy 	<ul style="list-style-type: none"> • 4 x 1m telescopes • 500 m base line • mid-IR spectrograph 	<ul style="list-style-type: none"> • Pointing stability • Thermal stability
AGNs	<ul style="list-style-type: none"> • Direct imaging of dusty torus 	<ul style="list-style-type: none"> • Interferometry • 10 mas ang res • 6-20 micron spectroscopy 	<ul style="list-style-type: none"> • 4 x 1m telescopes • 500 m base line • mid-IR spectrograph 	<ul style="list-style-type: none"> • Pointing stability • Thermal stability

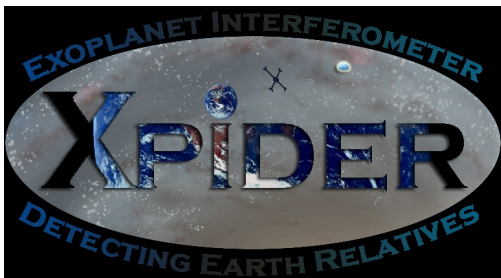
Guest Observer Programs



Imaging Roche Lobes



Zhao et al. 2008



Telescope and Instruments

Telescope : 4 x 1m aperture in formation flight

Instruments : Long baseline interferometer + mid-IR spectrometer

Baseline : 500 m

Wavelength :

- Imaging - 10 micron
- Spectroscopy: 6 - 20 micron

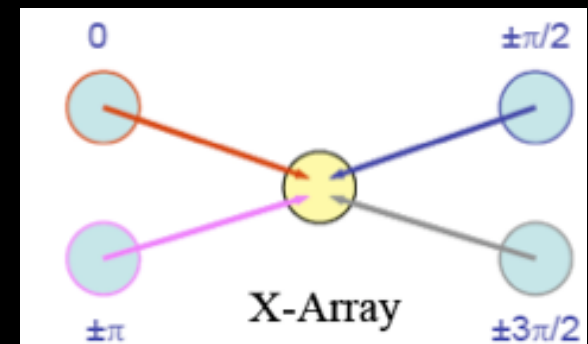
Resolution :

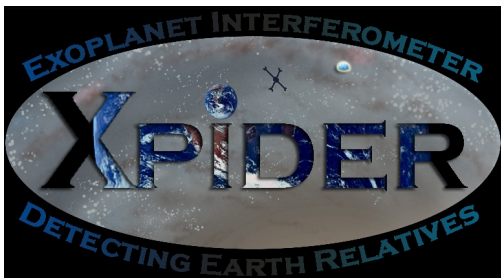
- Angular - 2 mas
- Spectral - 25

Sensitivity : 4.8 microJy, 10 microns, 15 hours

Duration :

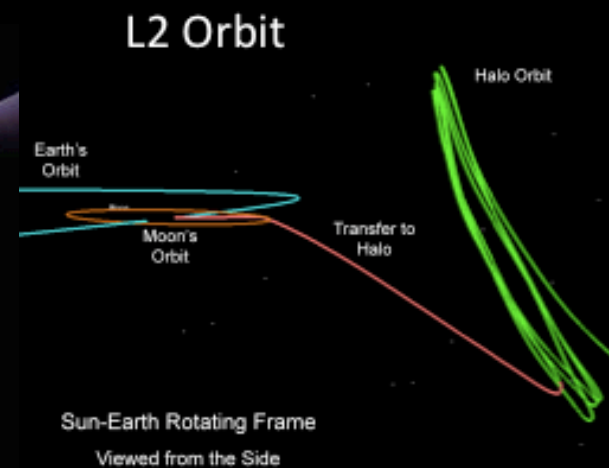
- Mission life-time - 5 years
- Extendable life-time - 3 years

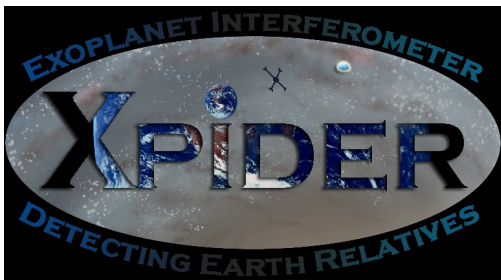




L2 Orbit

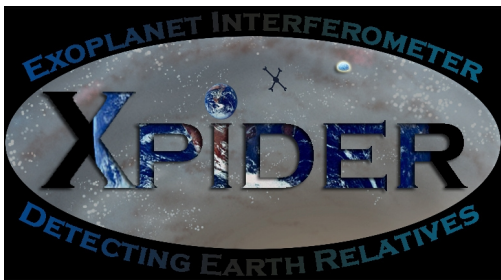
- Requirement of low thermal background : L2
- Requirement of better communication bandwidth for control of formation flight





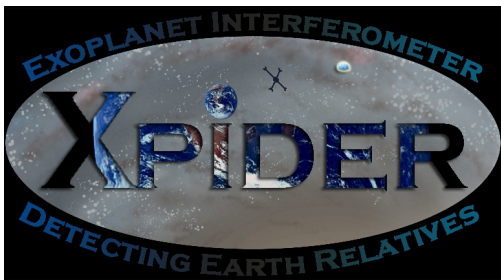
Data volume and rate

- Data volume requirements : 5 kbps x 25 (easy to meet requirements)
- Standard telecom hardware (X-band, for L2)



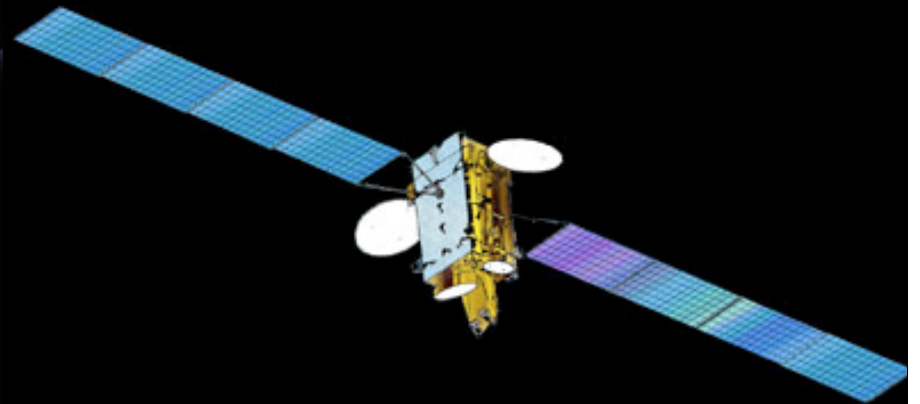
Pointing requirements

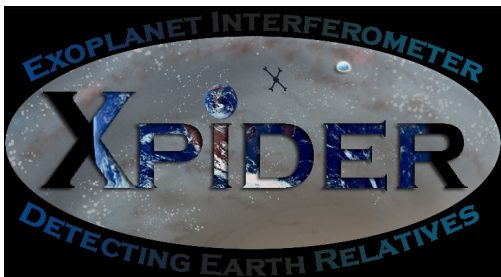
- Requirement: 1 arc sec
- 1 arc sec accuracy in attitude control of the spacecraft is available from standard star-trackers
- Delay lines allow finer control
- Minimizes costs on star-trackers



Spacecraft Buses

- Need 5 spacecraft on a single Launch vehicle: 4 x 1m + beam combiner.
- Each spacecraft = \$ 250M (Custom made bus) x 5 = \$ 1250M
- Mass: 5 x 300kg = 1500 kg

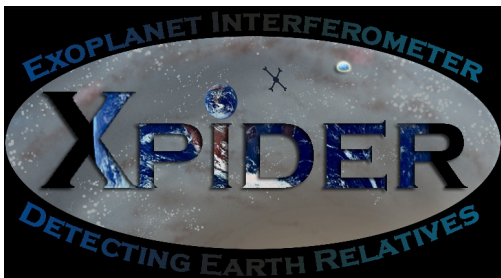




Launch Mass

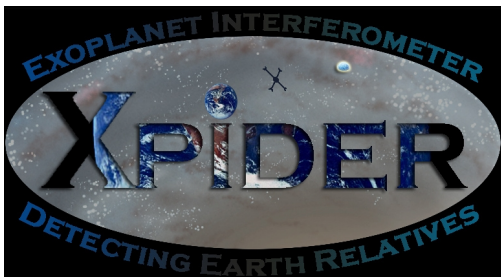
26 % margin on launch mass

SYSTEMS WORKSHEET:		FY08Q2 Templates Spacecraft			
	Mass Fraction	Mass (kg)	Subsys Cont. %	CBE+ Cont. (kg)	Mode 1 Power (W) Launch
Power Mode Duration (hours)					
Payload on this Element					
Instrument 1	26%	600.0	35%	810.0	0
Instrument 2	5%	120.0	43%	171.6	
Instrument 3	1%	20.0	43%	28.6	
Instrument 4	2%	40.0	43%	57.2	
Instrument 5	0%		43%	0.0	
Instrument 6	0%		43%	0.0	
Instrument 7	0%		43%	0.0	
Payload Total	34%	780.0	37%	1067.4	0
Spacecraft Bus					
do not edit formulas below this line, use the c					
Spacecraft	65%	1500.0	30%	1950.0	
S/C-Side Adapter	1%	25.0	5%	26.3	
Bus Total		1525.0	30%	1976.3	
Thermally Controlled Mass				1976.3	
Spacecraft Total (Dry)		2305.0	32%	3043.7	0
Subsystem Heritage Contingency		738.7	32%	32%	
System Contingency		252.5	11%	11%	0
Spacecraft with Contingency		3296	of total	w/o addl pld	0
Propellant & Pressurant1	7%	250.0			
Spacecraft Total (Wet)		3546			
L/V-Side Adapter		100.0			
Launch Mass		3646			
Launch Vehicle Capability		3495			
Launch Vehicle Margin		-151.2			
JPL Design Principles Margin		26%		30% required	



Major Risks

- Novel mission type: Lack of precursor missions
- Formation flight not demonstrated at this level
 - ▶ Navigation
 - ▶ Propulsion
- Challenges in interferometry
 - ▶ laser metrology
 - ▶ long term performance of nulling
- Cooling : Need to design details, but Spitzer is proof in action



Total Mission Cost

COST SUMMARY (FY2009 \$M)

WBS Elements	Total
Project Cost (\$ FY09)	\$2948.3 M
Development Cost (Phases A - D)	\$2639.8 M
01.0 Project Management	\$101.5 M
02.0 Project Systems Engineering	\$101.5 M
03.0 Mission Assurance	\$81.2 M
04.0 Science	\$30.0 M
05.0 Payload System	\$310.0 M
Instrument 1 (4 x 1m telescopes)	\$200.0 M
Instrument 2 (Interferometer)	\$70.0 M
Instrument 3 (Laser telemetry)	\$20.0 M
Instrument 4 (Spectrometer)	\$20.0 M
Instrument 5	
06.0 Flight System	\$1250.0 M
07.0 Mission Operations Preparation	\$15.0 M
09.0 Ground Data Systems	\$15.0 M
10.0 ATLO	\$109.2 M
11.0 Education and Public Outreach	\$10.2 M
12.0 Mission and Navigation Design	\$7.0 M
Development Reserves	\$609.2 M
Operations Cost (Phases E - F)	\$172.5 M
Operations	\$150.0 M
Operations Reserves	\$22.5 M
8.0 Launch Vehicle (L/V B)	\$136.0 M

5% of development
 5% of development
 4% of development

\$15M
 \$15M
 7% of Payload and Flight
 0.5% of development
 \$7M

30%

\$30M/yr **5**
 15%

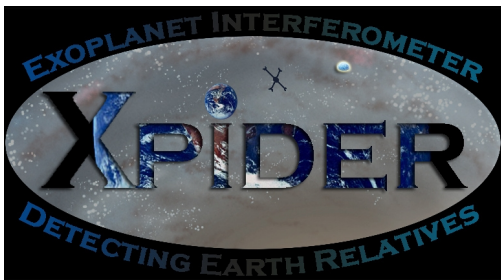
The logo for the Exoplanet Interferometer (XPiDER) is centered on a black background. It features a large, stylized oval containing a space scene with stars, a small planet, and a satellite. The word "XPiDER" is written across the oval in large, bold, blue letters with a white outline. The letters 'X', 'P', 'D', 'E', and 'R' are solid black, while the 'i' is filled with a blue and white Earth-like planet texture. Above the oval, the text "EXOPLANET INTERFEROMETER" is written in a blue, sans-serif font, following the curve of the top edge. Below the oval, the text "DETECTING EARTH RELATIVES" is written in the same blue, sans-serif font, following the curve of the bottom edge.

EXOPLANET INTERFEROMETER

XPiDER

DETECTING EARTH RELATIVES

This is what humanity will look forward to in 10 years.



Team Xpider

- ✦ Nikku Madhusudhan
- ✦ Yamina Touhami
- ✦ Jessie Christiansen
- ✦ David Bernat
- ✦ Andrew Fittingoff
- ✦ Krista Soderlund
- ✦ Padma Yanamandra-Fisher

Scientific Advisors

- ✦ Rachel Akeson
- ✦ Keith Wafield