

Operational Experiences with Vulcan



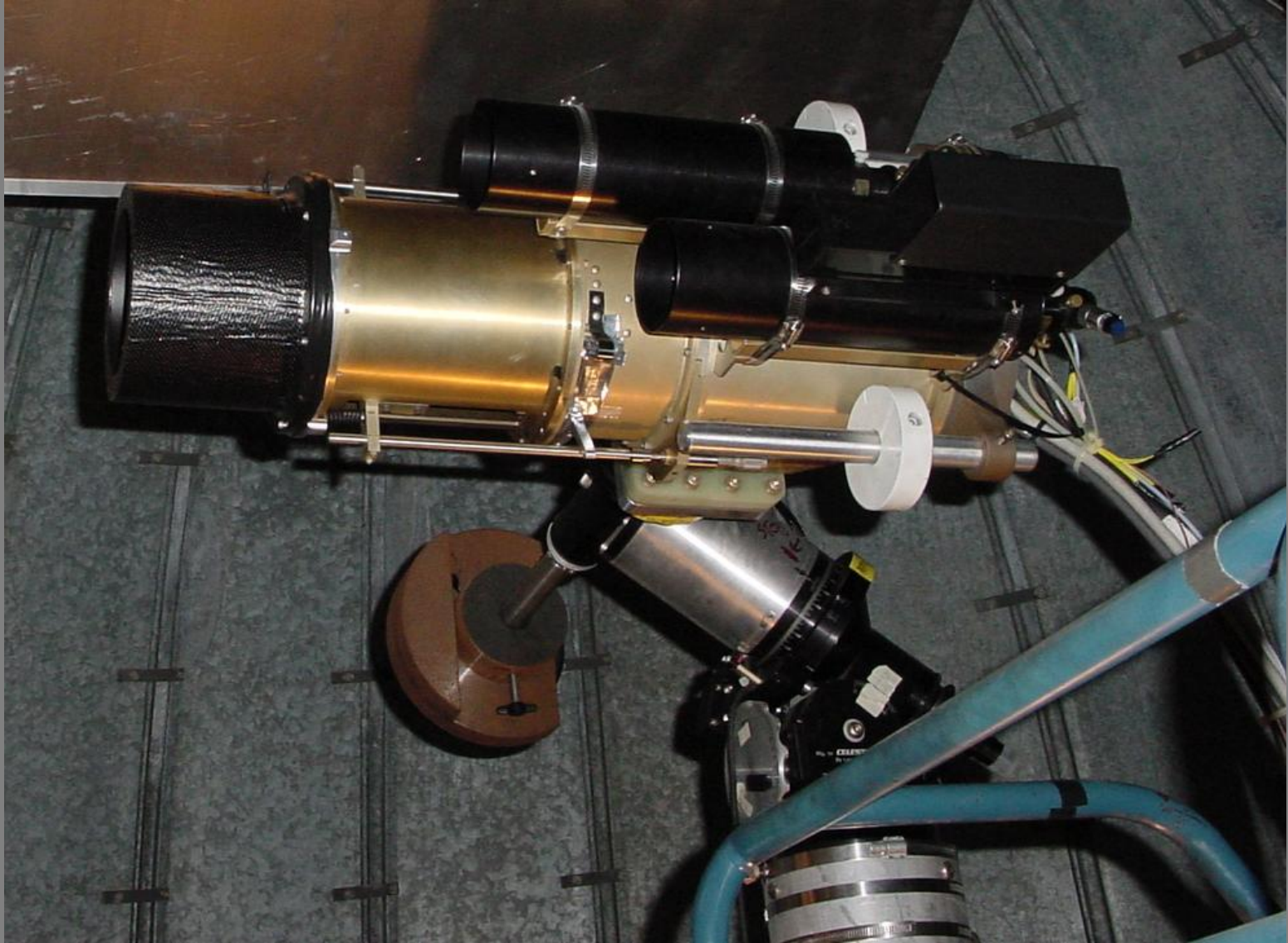
N. Batalha

San Jose State University

NASA Ames Research Center

Vulcan, Roman god of fire

The Camera



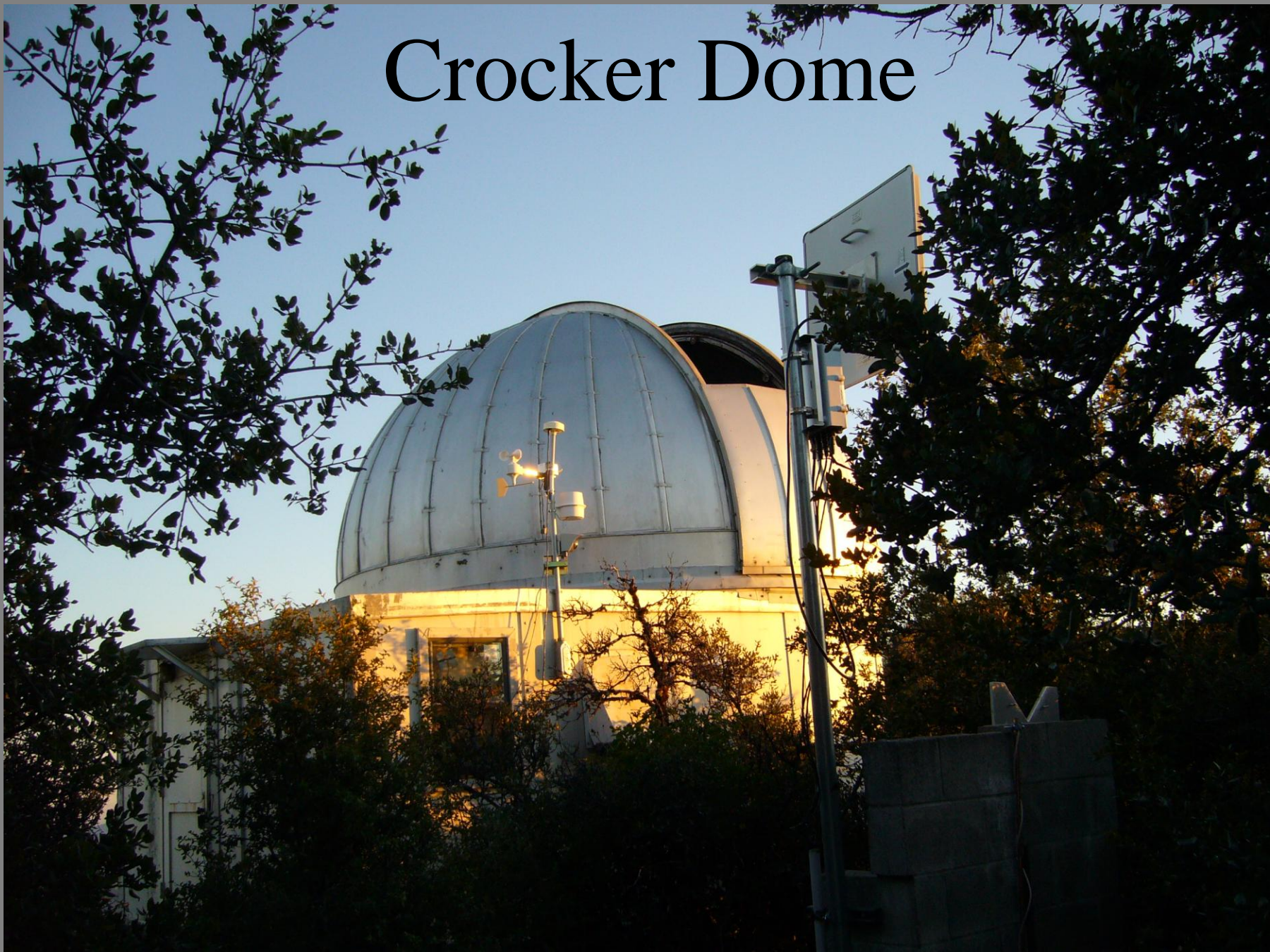
An f/2.8 Canon lens feeds a 4096x4096 Kodak CCD

Weather Station



Dome & outside air temperature, pressure, wind speed, humidity, precipitation, cloud cover (via IR sensor)

Crocker Dome



Data Transfer

2.4 GHz Solectek MP1100E
wireless bridge routers yield
500 kbytes per second.

Ames to Crocker dome:
~25 miles





Automated Vulcan Image Acquisition

Redbox Bluebox



Calibration List BFD4kavia.lst

1	Bias	x7	Bias4k	
2	Flat	x7	Bias4k	15s
3	Dark	x3	Bias4k	180s

Observation List IyrIVcygVI.lst

1	IyrIV	x4	Bias4k	180s	f:VR
2	CygVI	x4	Bias4k	180s	f:VR

Sequence Status Wait:21600

Sequence In Progress

Alarm Status No Alarms

Image Acquisition Progress Image# 17 of 17

Sky and Guider Status

- Dark?
- Elev OK?
- LinkEnabled?
- GuiderExpose?
- GuiderTrack?
- GuiderStop?

Redbox time

LST 03:56:41
UT 16:46:22
UT Date 11/07/2007

CCD Status

CCD Temp (°C) -22.80
Backplate Temp (°C) 11.30
Vacuum Pressure (torr) 2.46

SEI Encoder Status

RA (hours) Dec (degrees)
18.30 47.10
HA Encoder Range: 0 to 3599
HA Encoder Reading 3245
DEC Encoder Range: 0 to 3599
DEC Encoder Reading 2179

Dome Status

In Position?
Sunny?
Slit Open?
Rainy?
Dome Azimuth 294

Latest Marginal Values

MargSigma (X) NaN StarSkew 184.87
MargSigma (Y) NaN Delta X (pixels) -1
XCorrPeak (X) NaN Delta Y (pixels) 0
XCorrPeak (Y) NaN Rubin Marginals?

Manual Control

Guider Control

WARP Control

Dome Control

Guider Control Command

MaximDL parameters

param 1 4.00

param 2 1.00

Use MaximDL?

Guider GO



storage folder Observer

C:\VulcanData

Notes

THStarSkew (110)

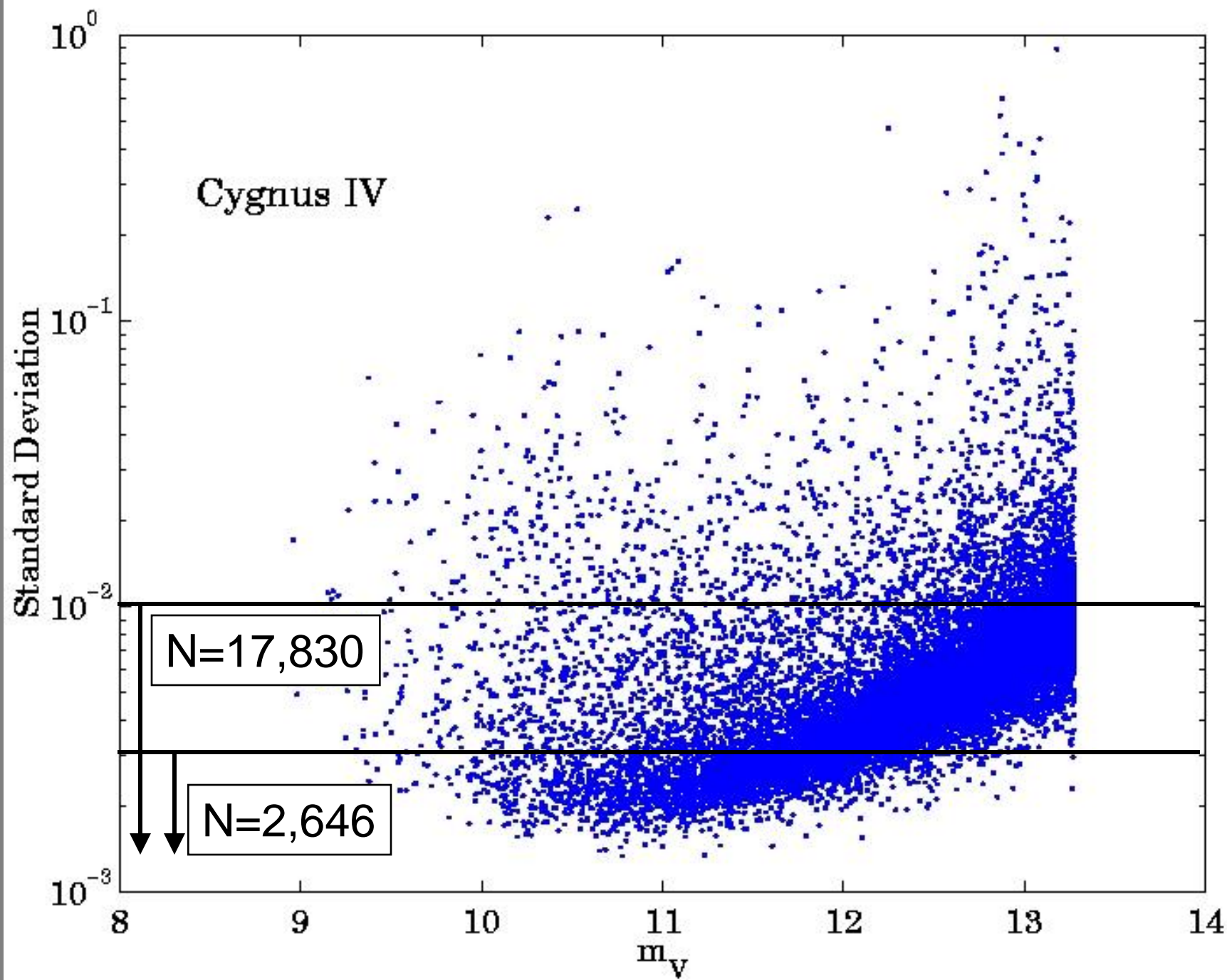
58.00

THXCorrPeak (.6)

0.60

THMarginalSigma (5)

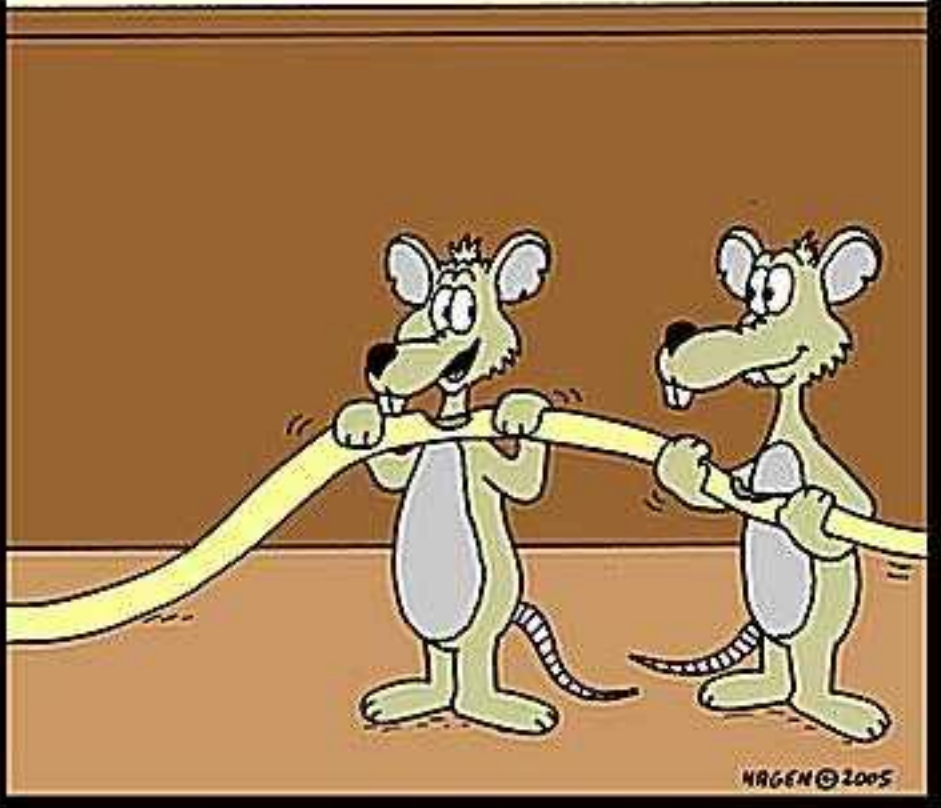
5.00







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The power is off, I've checked:
We can chew as much as we want...



The Dome Leak Diverter



Good thing it never snows in California...

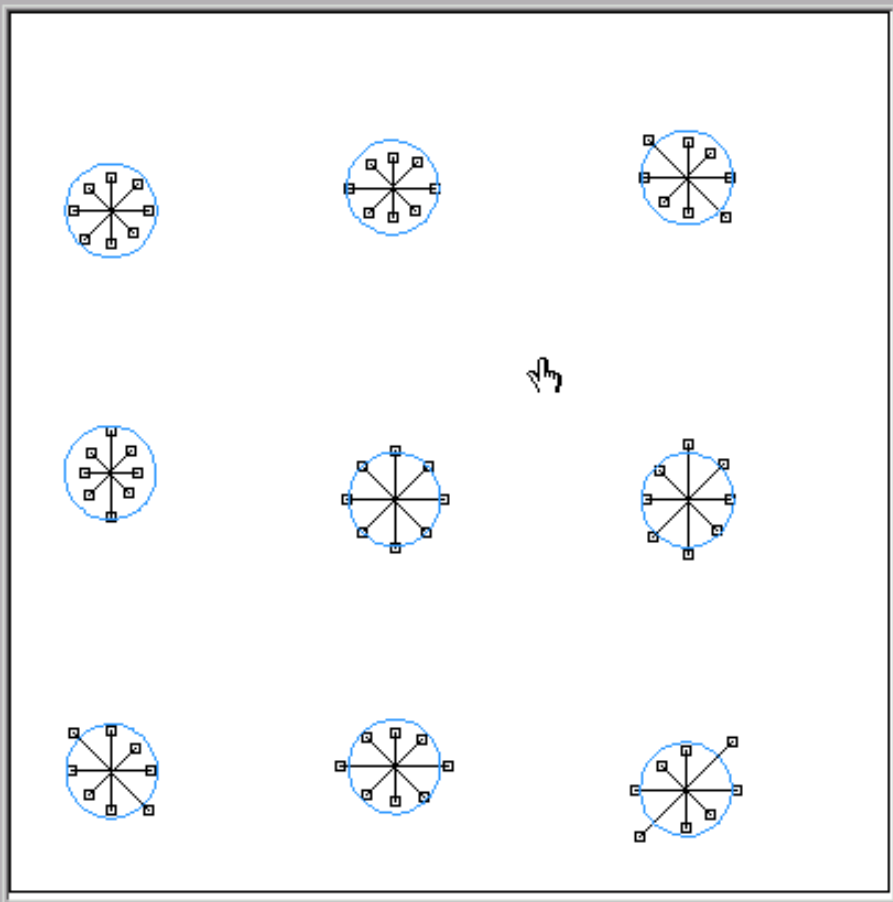


Current Image

C:\VulcanData\Data-2007-07-10\Cygvi168497.fits

Finding star 9 of 9...

FWHM Plot



Ideal FWHM: 3.00
Mean FWHM: 2.65
SD of FWHM: 0.59

Stars from Whole Image

FWHMs

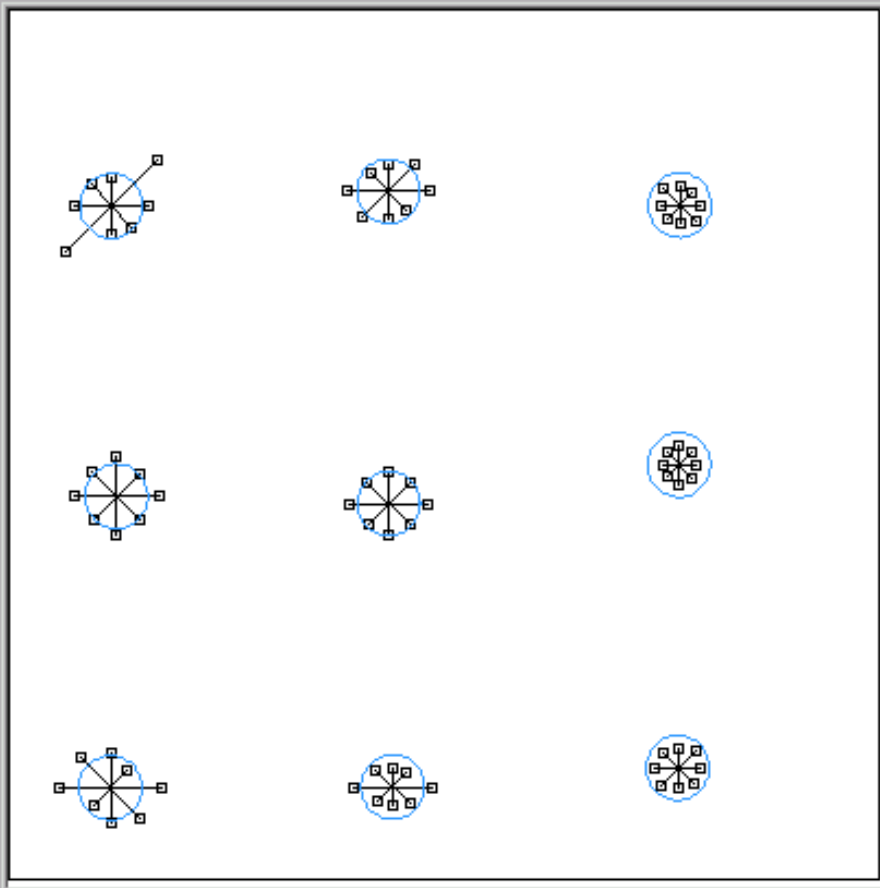
2.58	2.14	2.54	3.48
3.56	2.55	2.17	2.63
3.27	4.24	2.49	2.18
1.75	1.96	2.75	1.80
3.20	3.07	3.09	3.02
2.71	3.24	3.52	2.72
2.44	2.49	2.11	1.99
2.82	2.29	1.86	2.12
2.81	2.15	2.30	3.51

Current Image

C:\WulcanData\CygVI167305.fits

Finding star 9 of 9...

FWHM Plot



Ideal FWHM
3.00

Mean FWHM
2.78

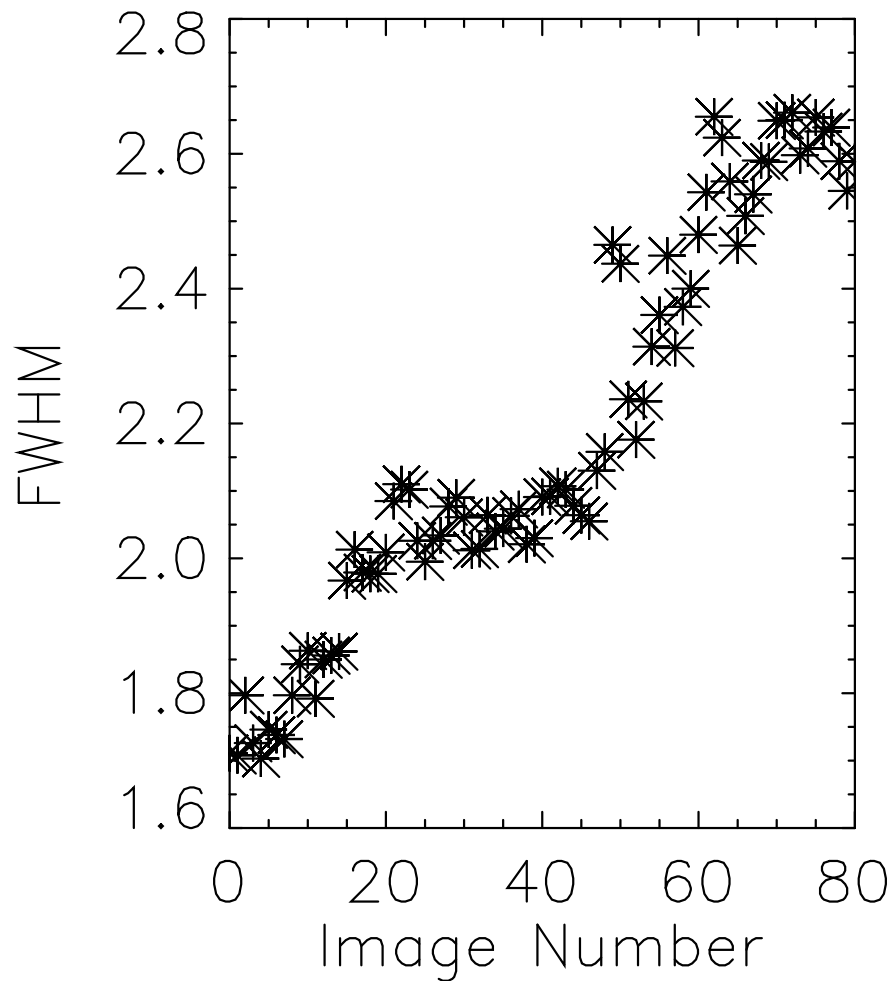
SD of FWHM
1.04

Stars from Whole Image

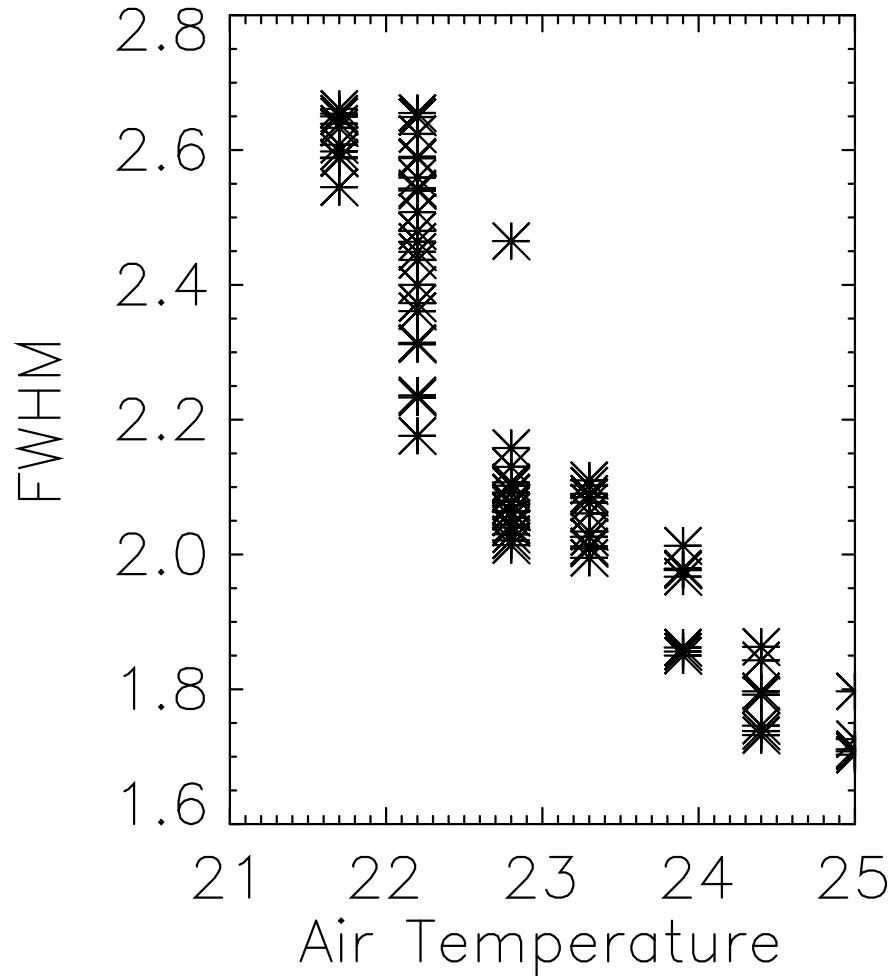
FWHMs

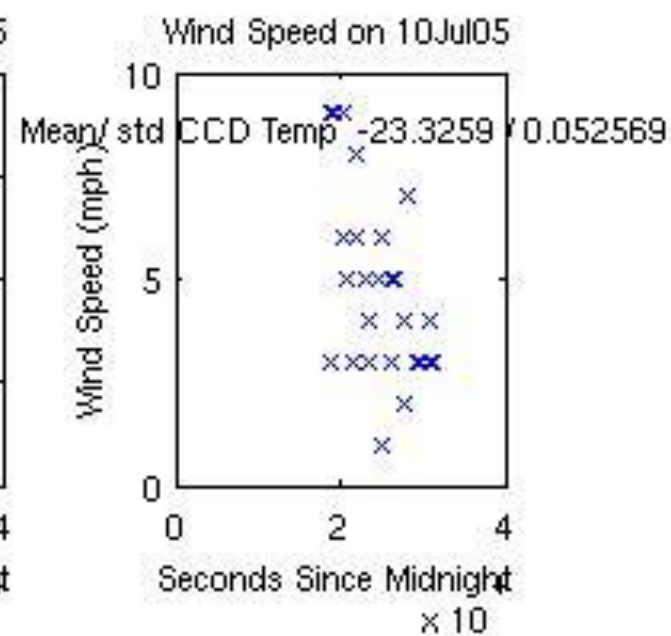
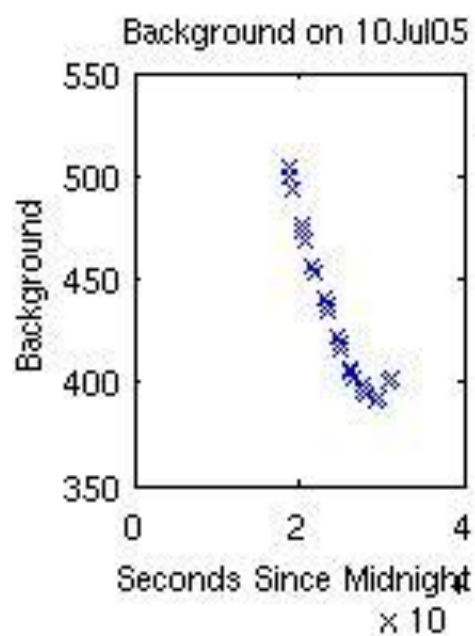
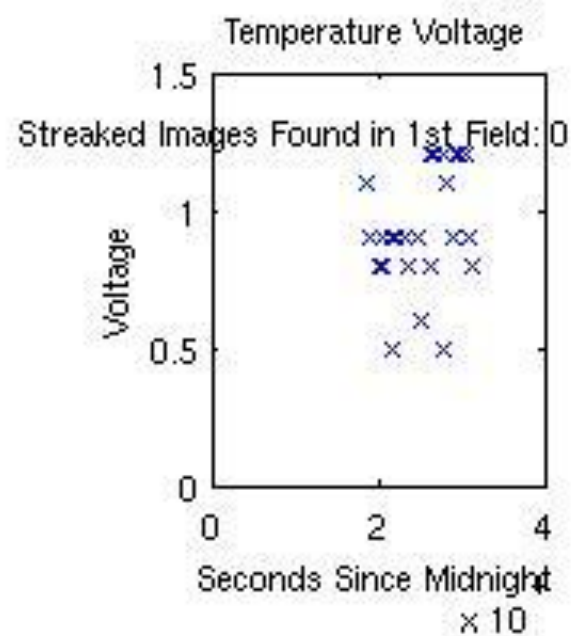
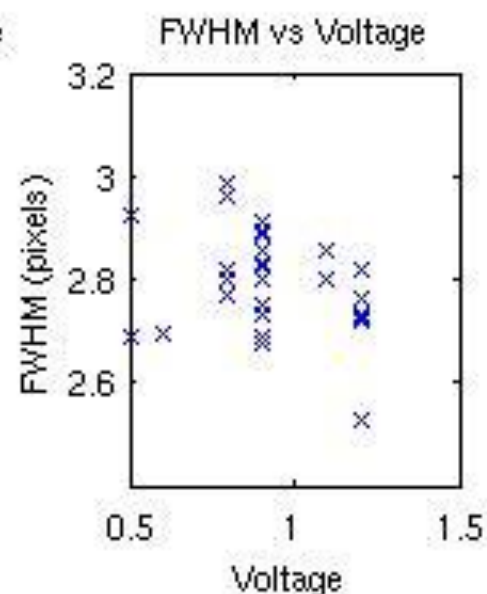
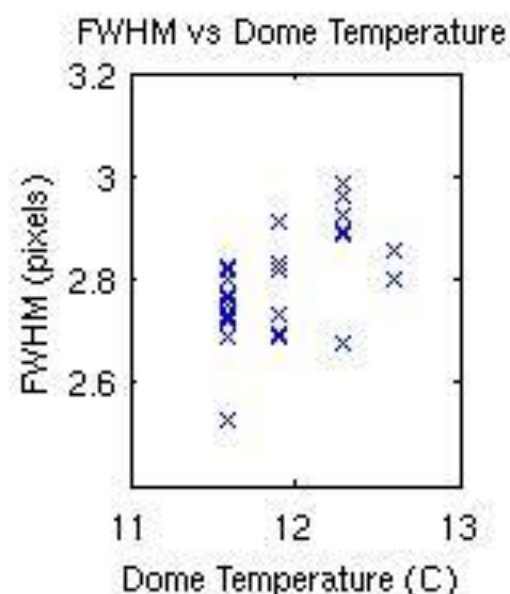
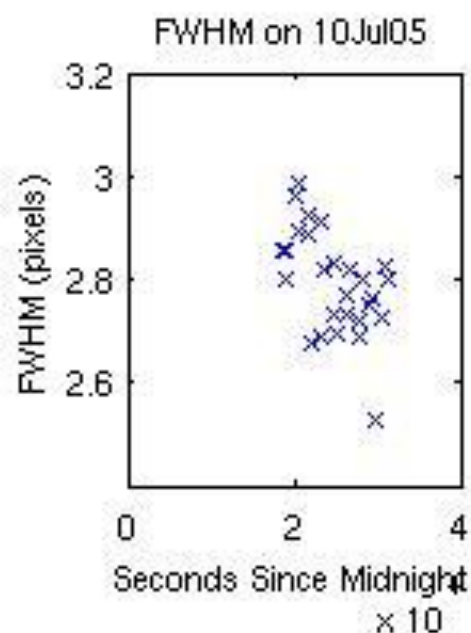
4.99	2.21	3.15	3.99
3.67	1.90	1.65	2.26
2.22	2.39	1.90	2.05
4.07	3.08	3.67	3.19
3.76	2.80	2.88	2.83
1.52	1.61	1.84	1.68
3.55	6.09	2.56	2.77
3.93	3.45	2.51	2.43
1.87	1.66	1.74	2.19

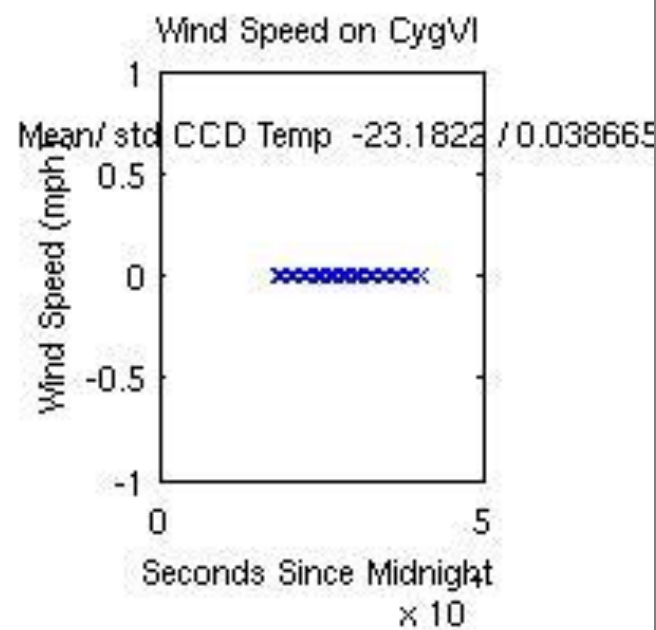
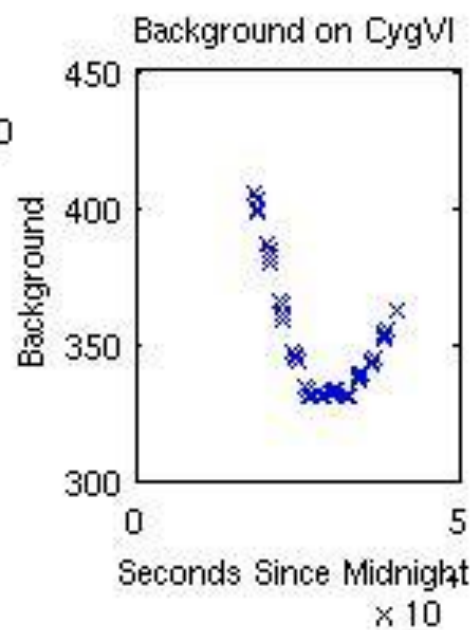
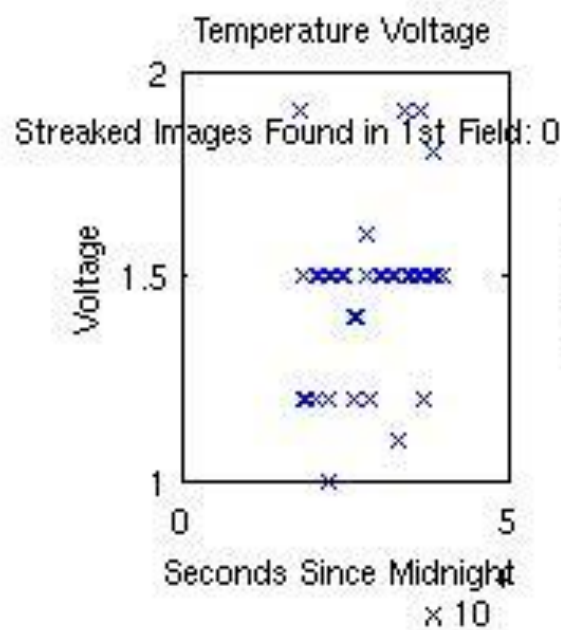
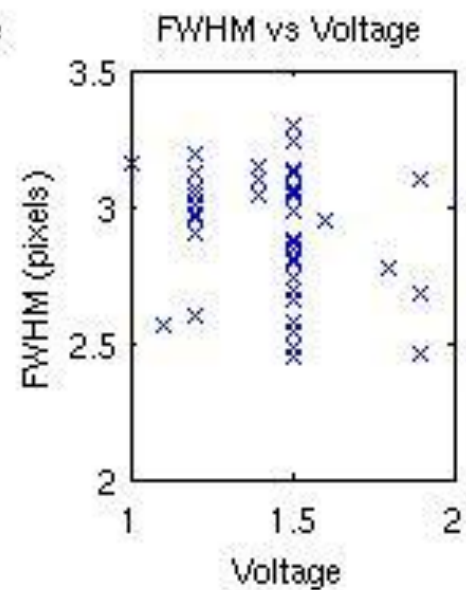
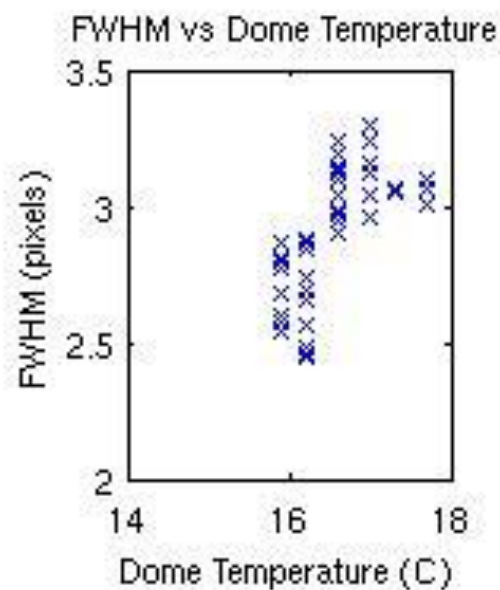
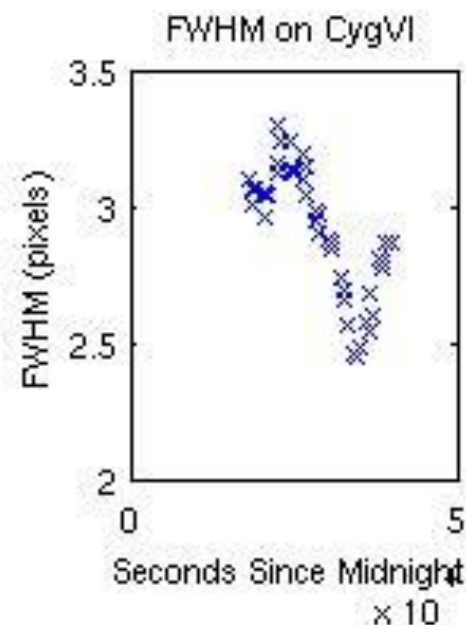
Focus Stability

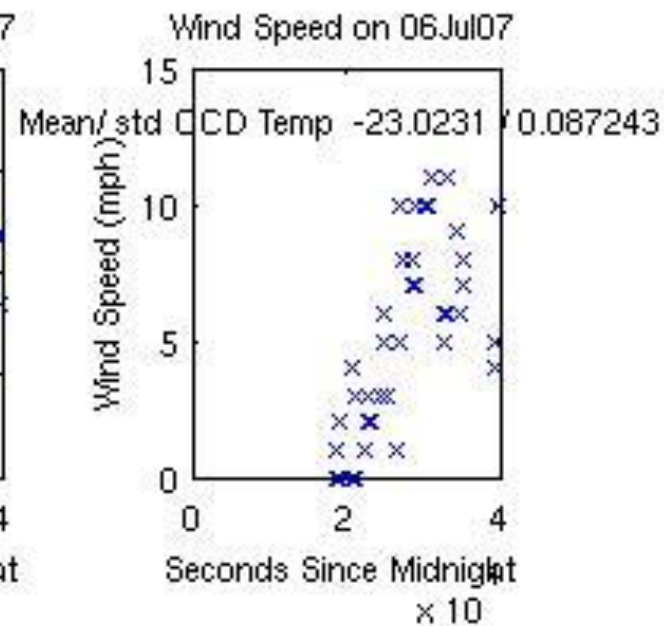
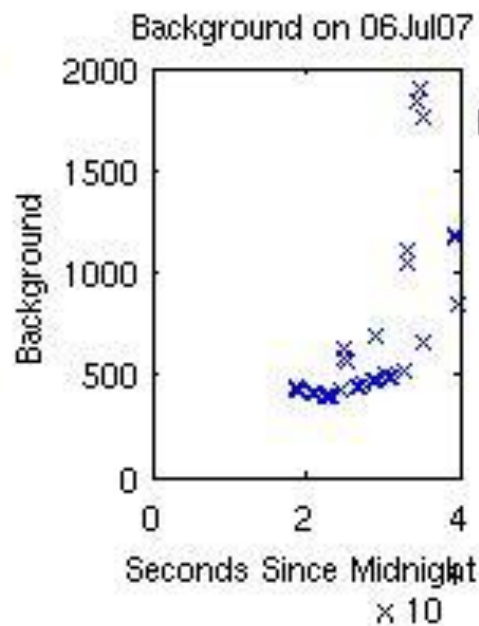
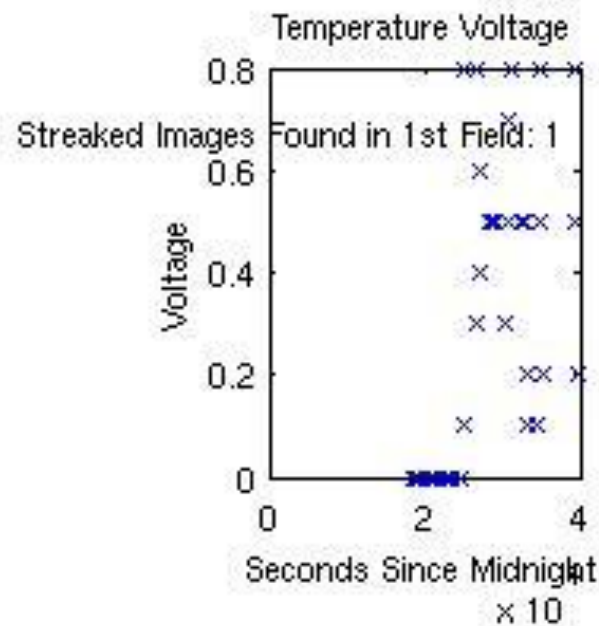
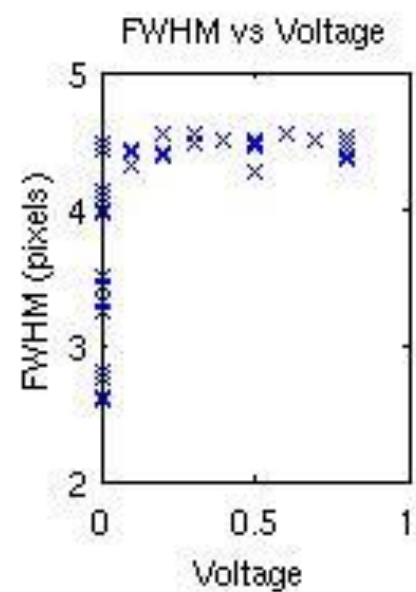
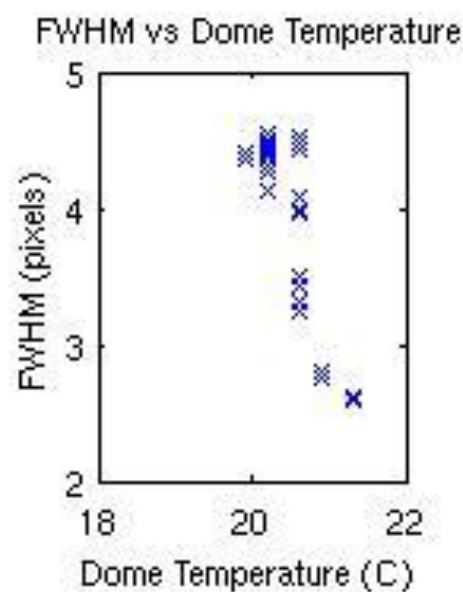
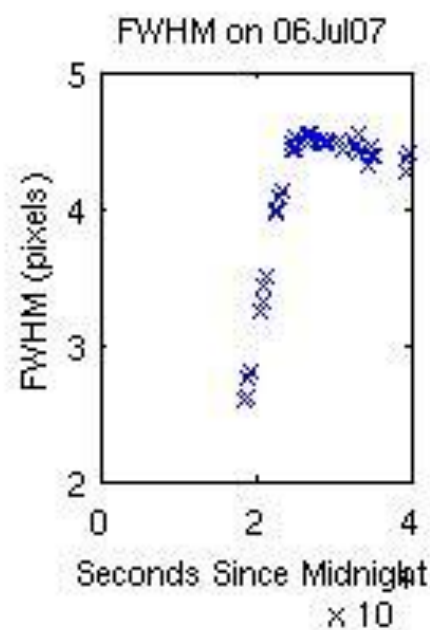


FWHM correlated with both air and CCD Temperature









Things that can go wrong

- Mice chew cables
- Mice nest in warm equipment
- Motors fail
- Circuit breakers blow
- Power fails
- Domes leak
- Hard drives fill up
- Memory leaks cause glitches
- Air temperature is not stable

What can help

- Nightly diagnostics
- Thorough observing logs
- Web camera
- Internet-controlled dome light
- Internet-controlled switches
- CRON scripts to archive data
- Linux/Unix wherever possible
- Absolute encoders on drive axes
- Thermal control heaters

The Elephant in the Closet



Number of Expected Detections

$$N = P_d \times P_p \times P_a \times P_3 \times N_{\text{stars}}$$

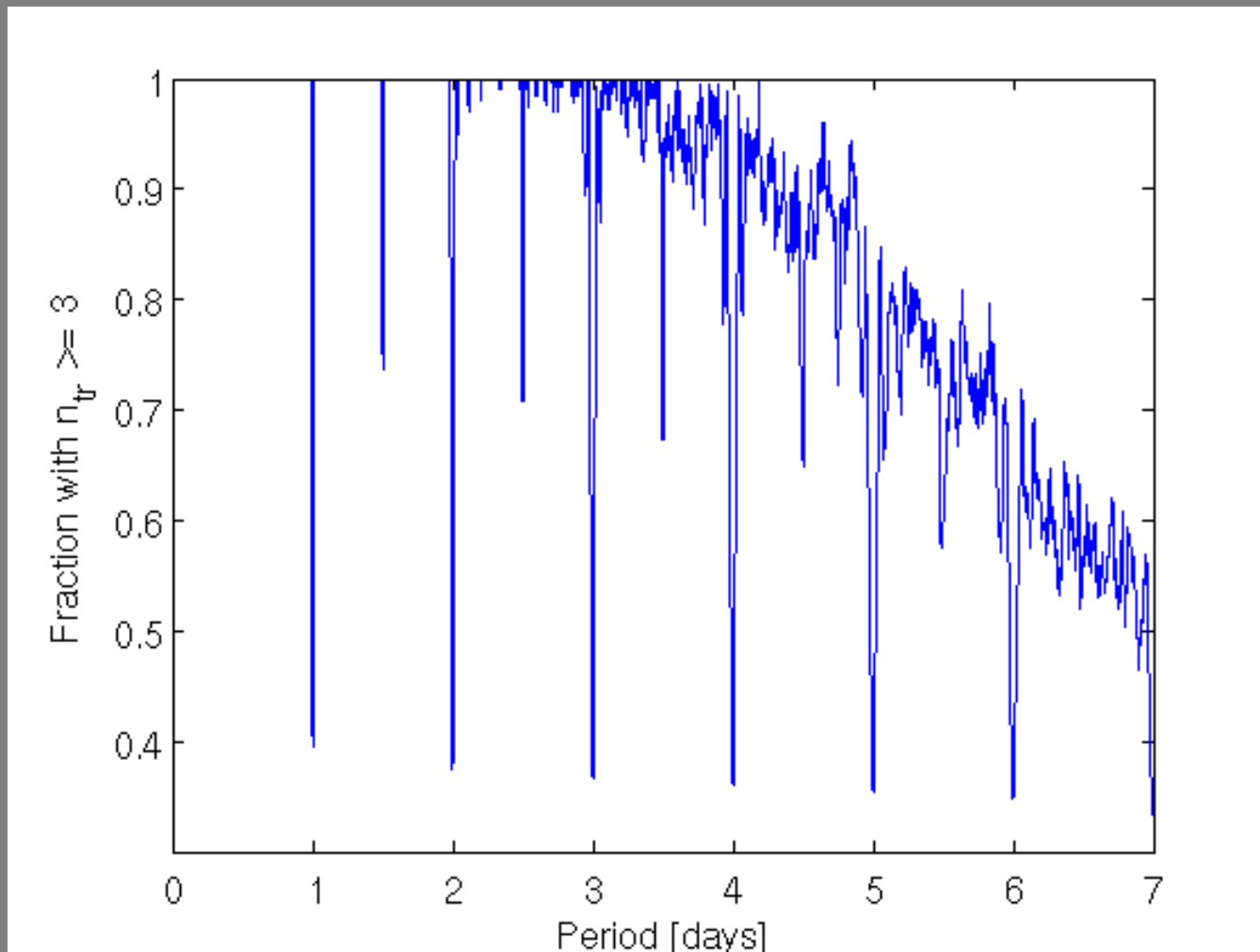
P_d = probability star is MS dwarf (50%)

P_p = probability has short – period planet (1%)

P_a = probability of alignment (10%)

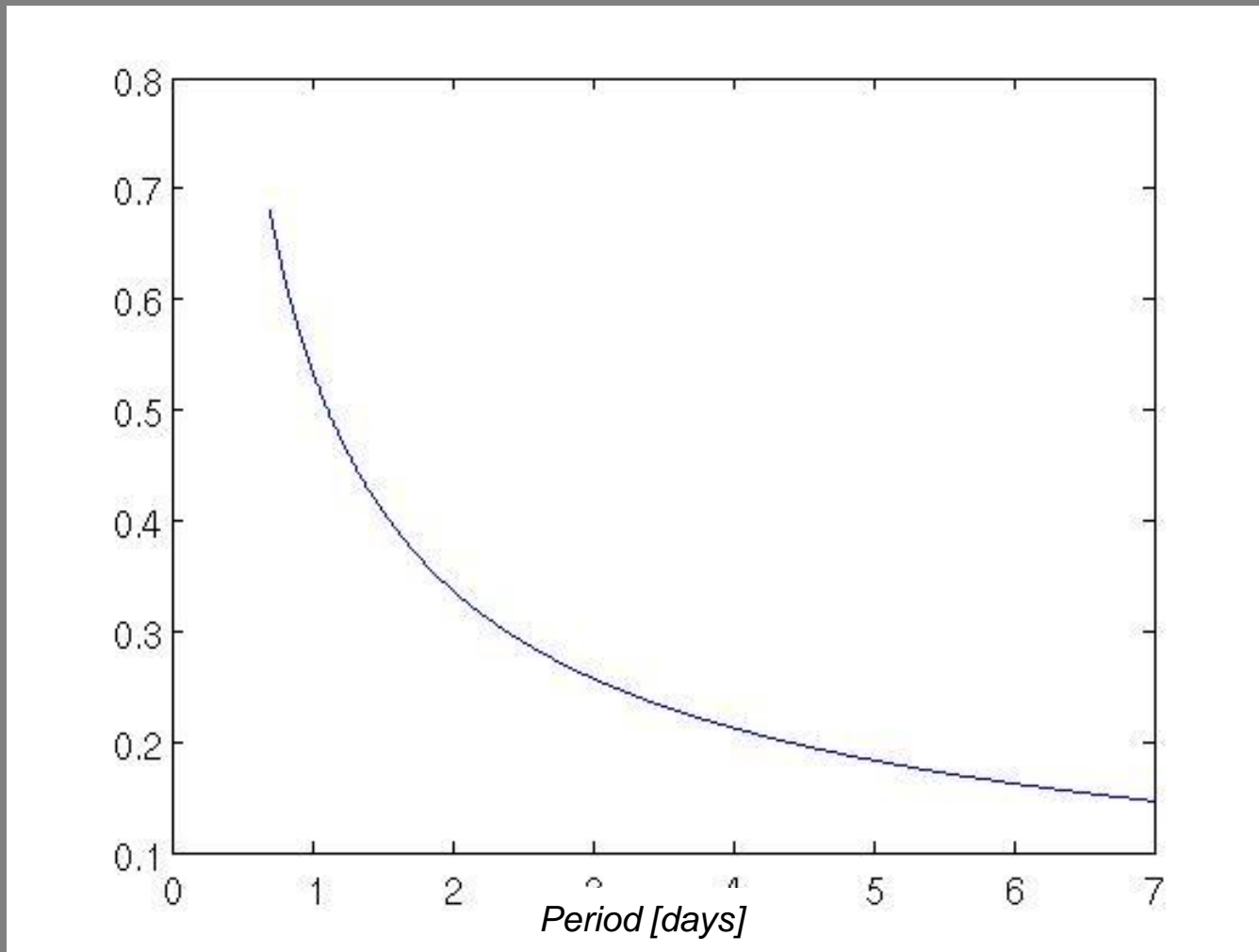
P_3 = probability of observing 3 transits (60-80%)

Assume milli-mag precision is achieved for about 3,000 stars per field. This translates to one detection per field.



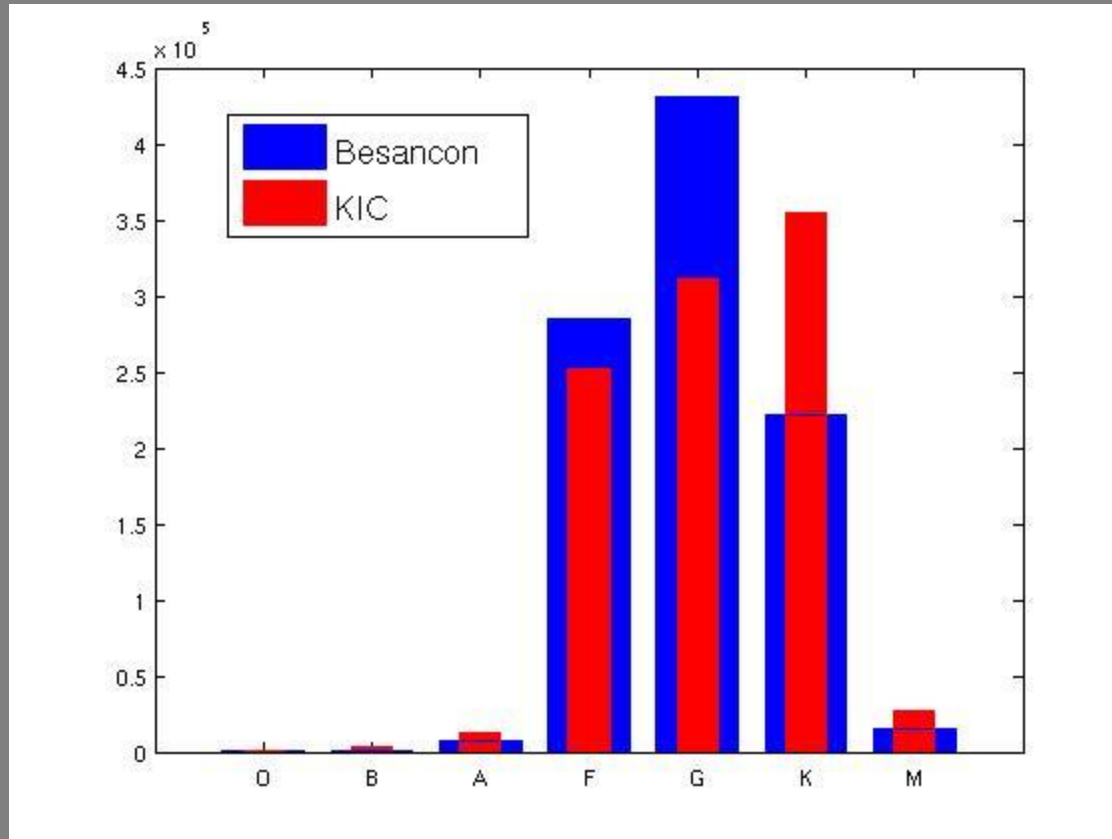
Does not imply 100% detectability probability for $P < 2$ days

Alignment probability also decreases with increasing period.



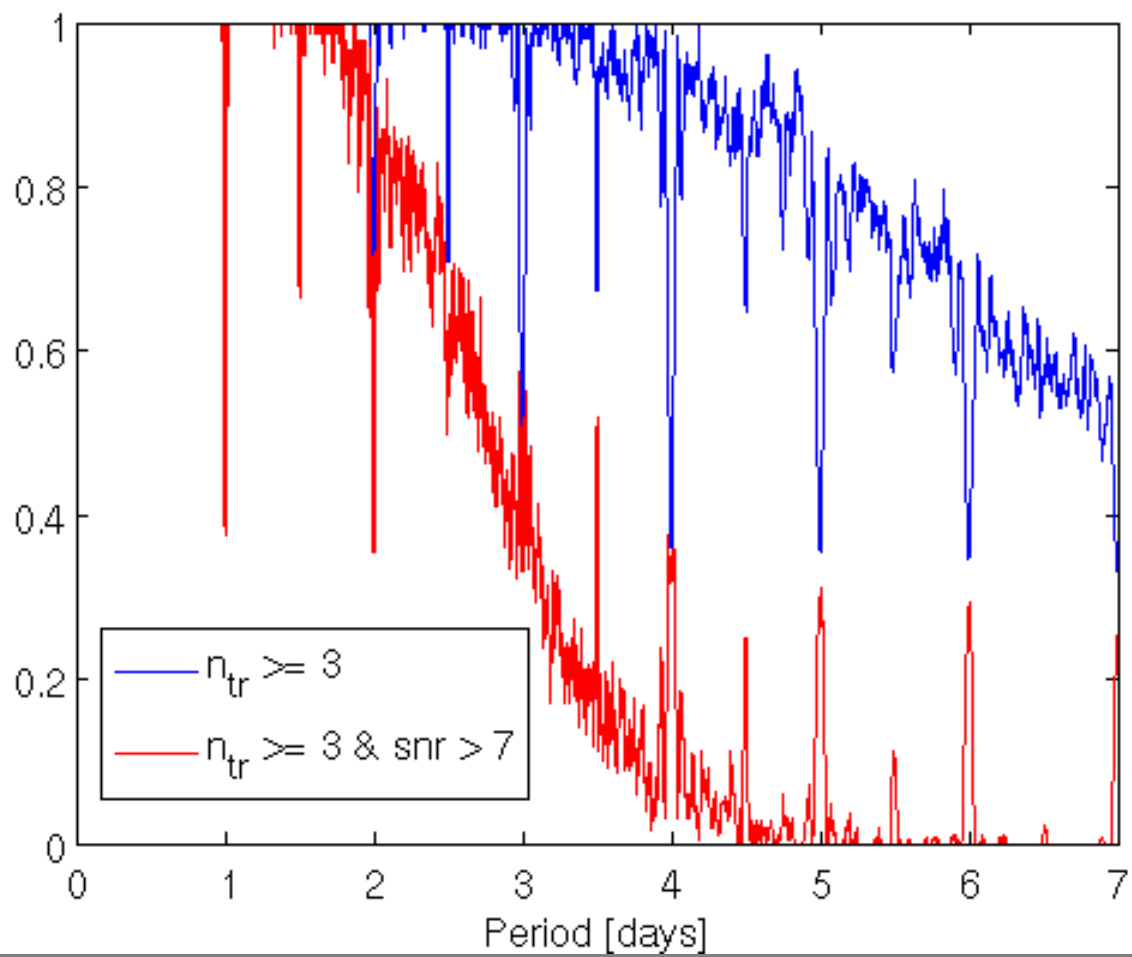
Probability depends on stellar characteristics.

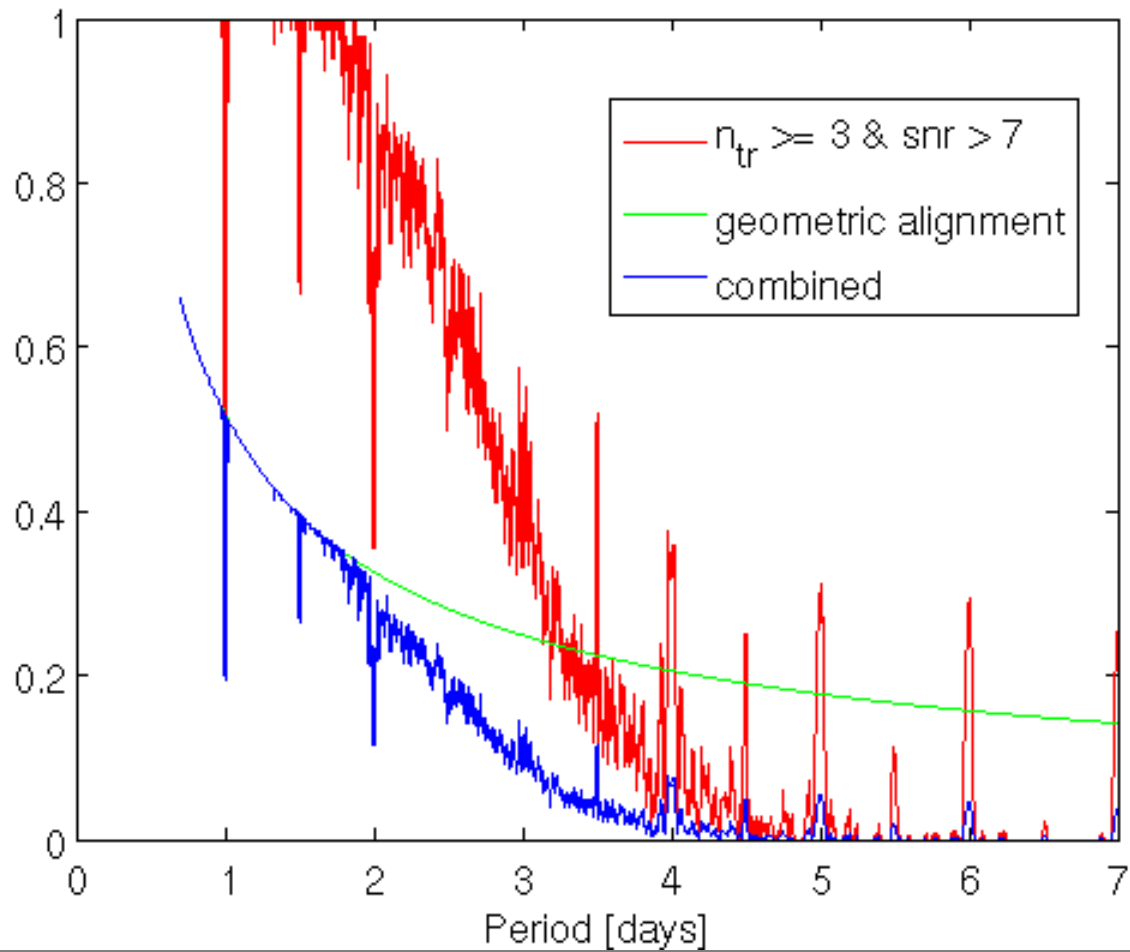
Modeled versus Observed Stellar Populations



<http://bison.obs-besancon.fr/modele>

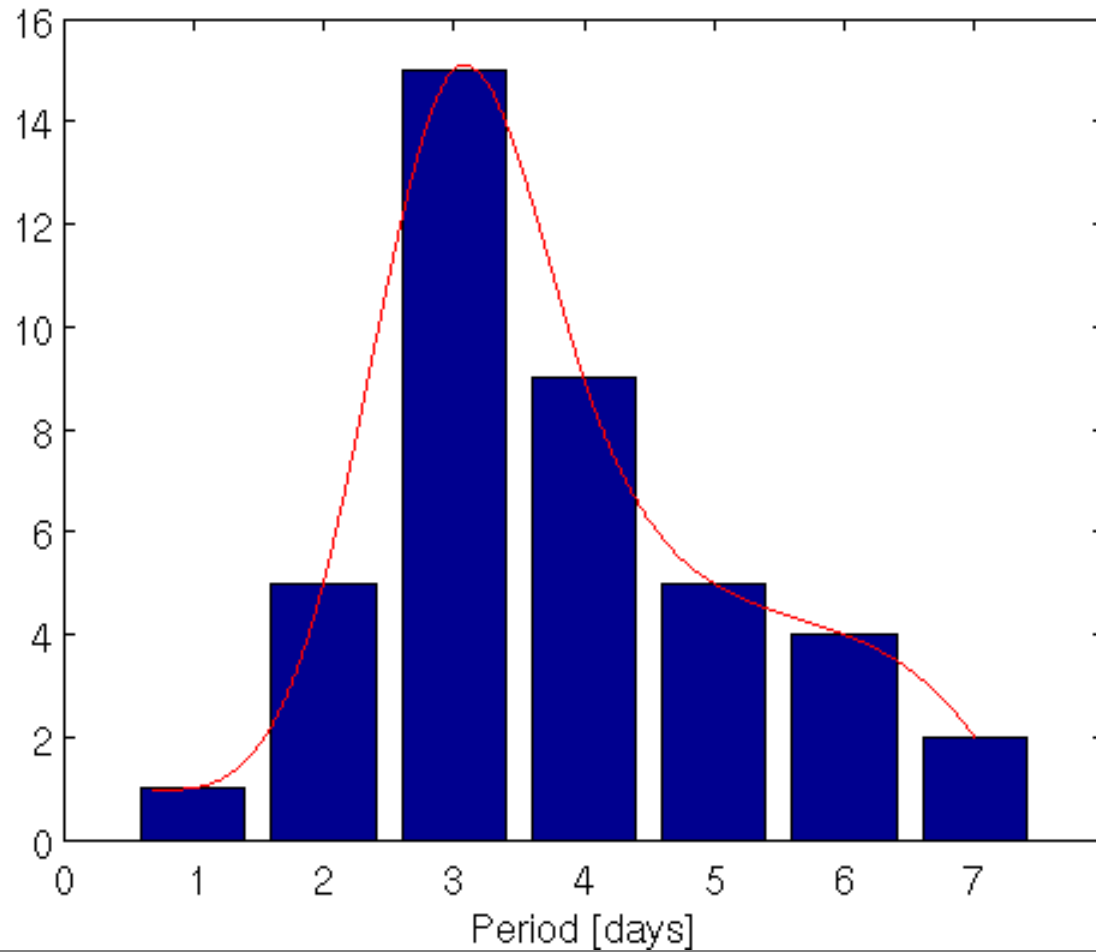
Robin, Reyle, Derriere, Picaud, 2003, AA 409 523

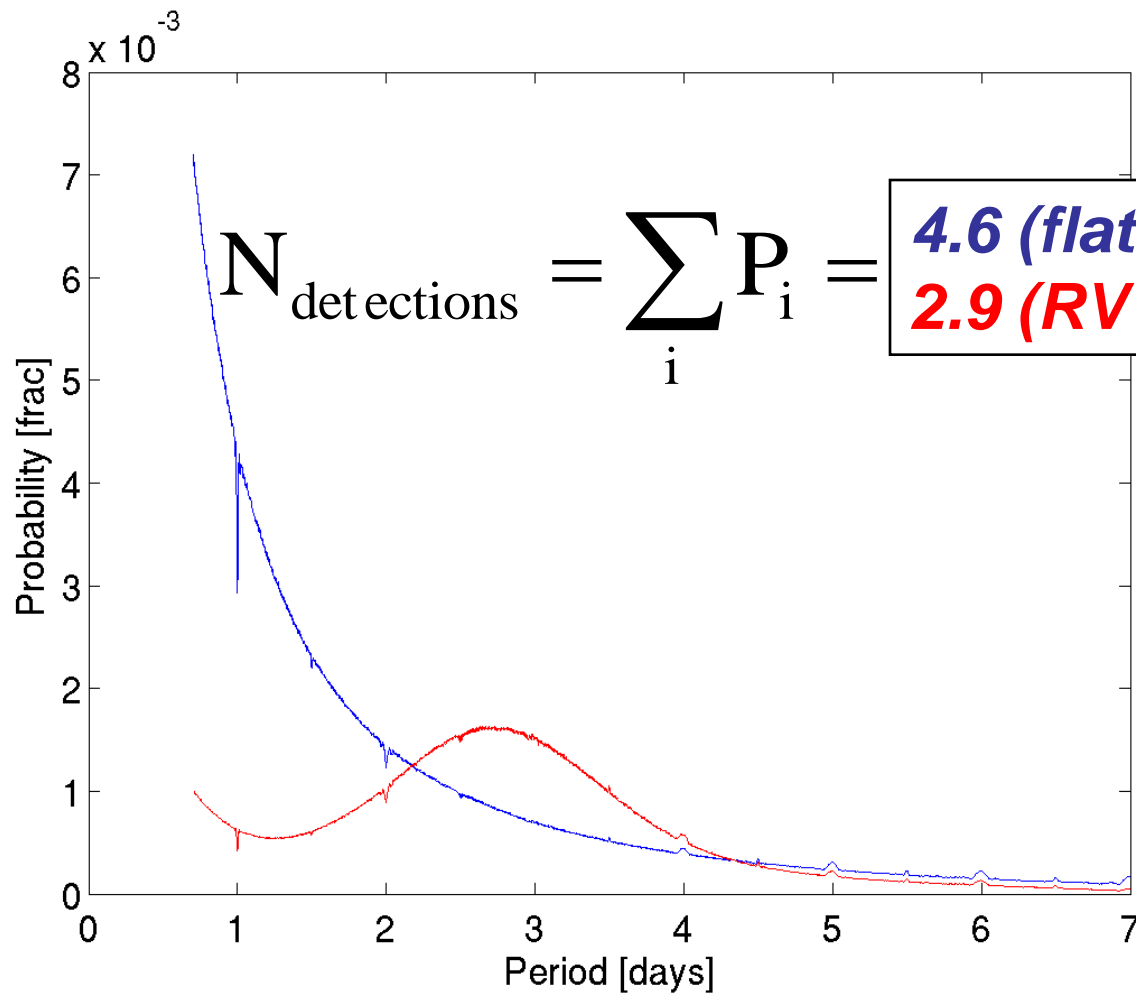




Note: blue and green traces applicable to one specific star

RV Period Distribution





Does not consider binarity or dilution effects due to crowding.

More cameras, more fields, more telescopes...

$$N_{\text{detections}} = \sum_i P_i =$$

One camera, one field,
one telescope:

4.6 (flat period distr)
2.9 (RV period distr)

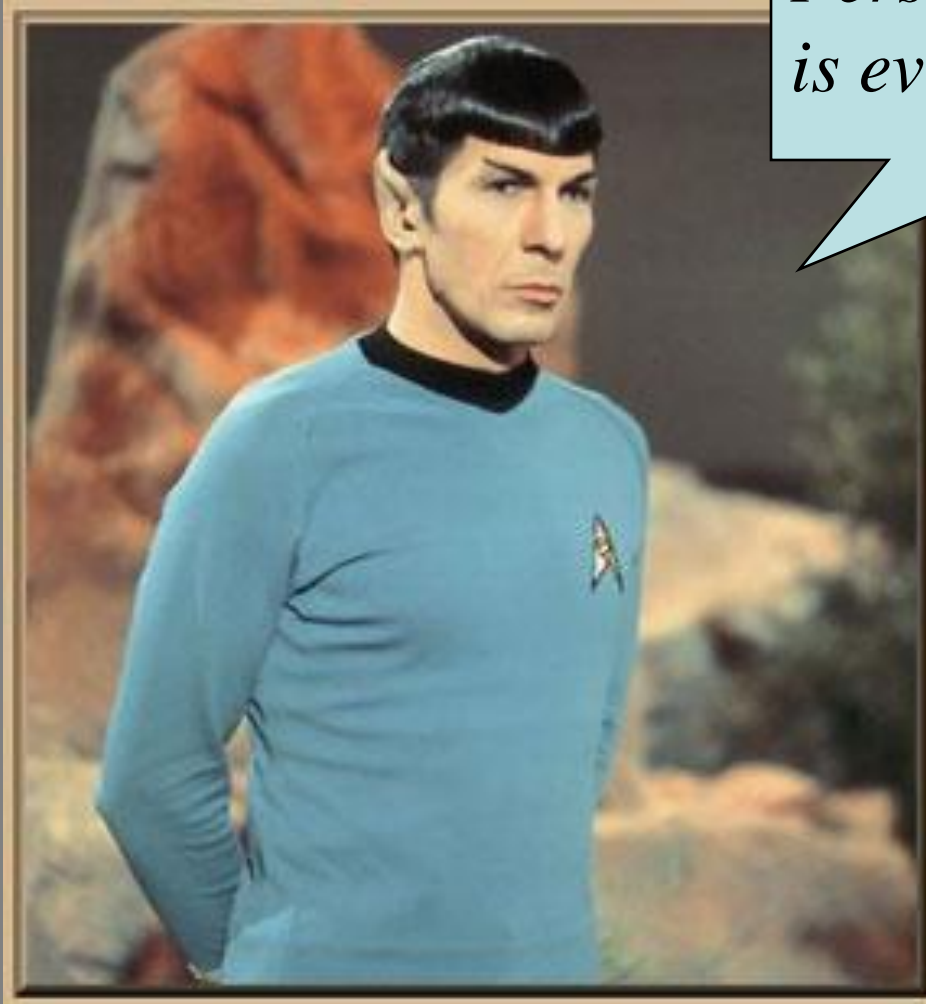
One camera, two fields,
one telescope:

Per field:

4.3 (flat period distr)
2.6 (RV period distr)

Total:

8.6 (flat period distr)
5.2 (RV period distr)



*Persistence, Captain,
is everything.*