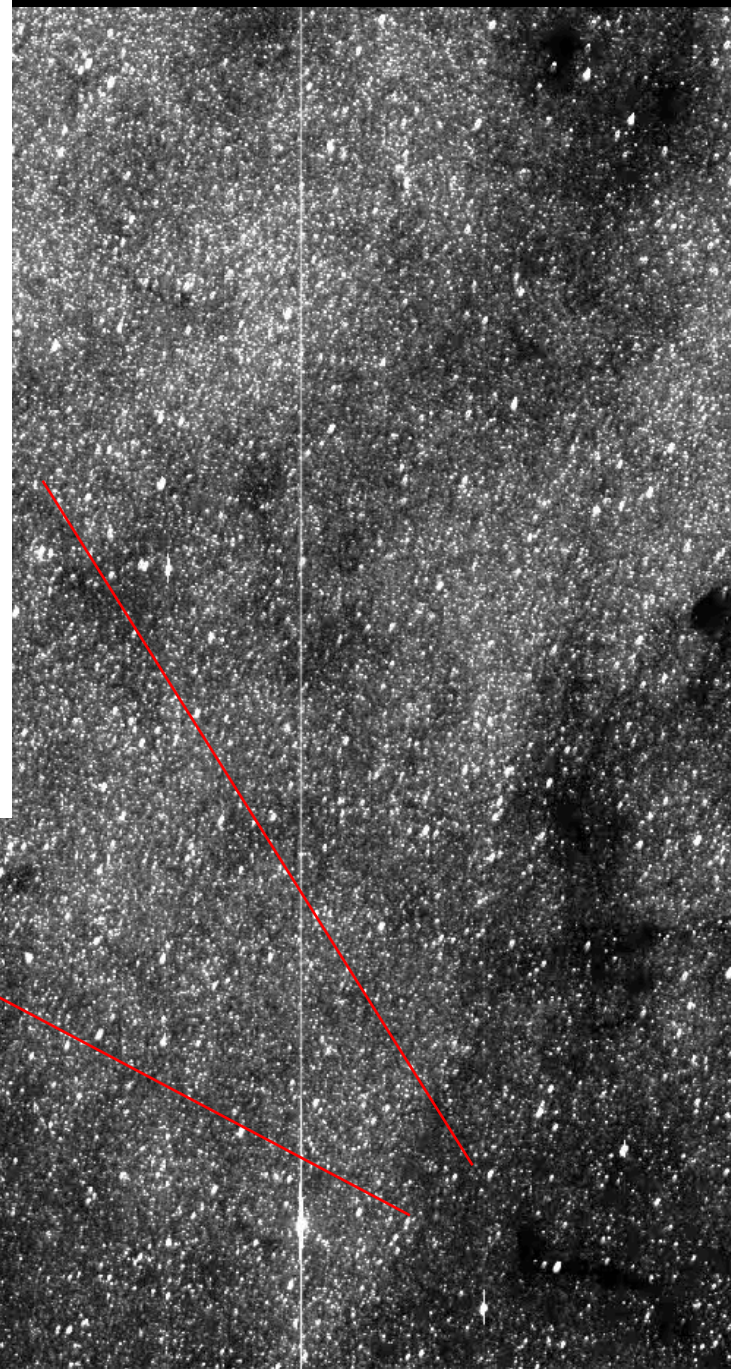
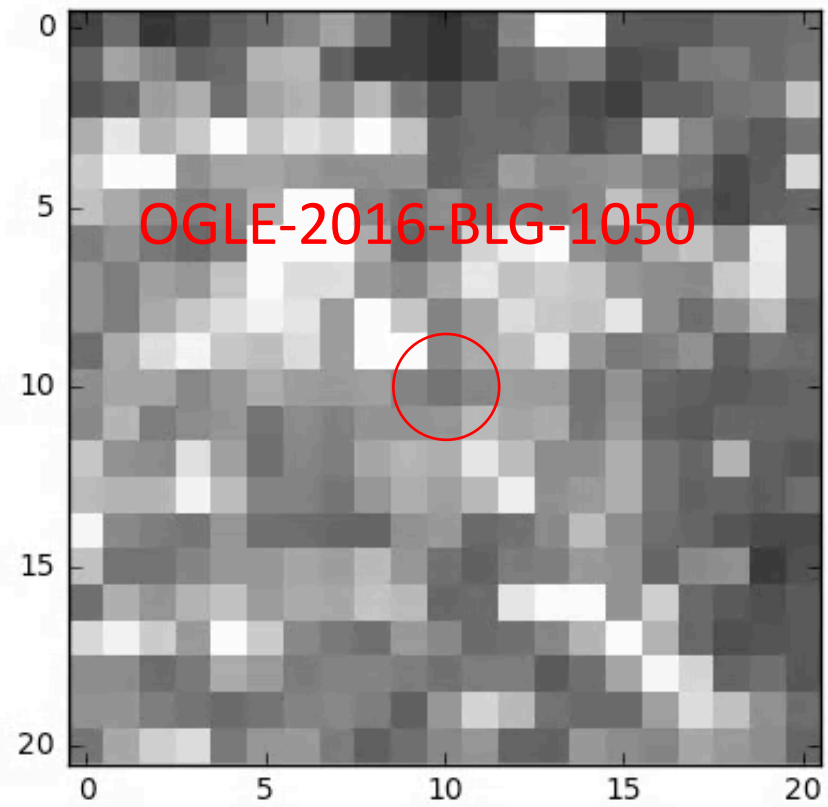
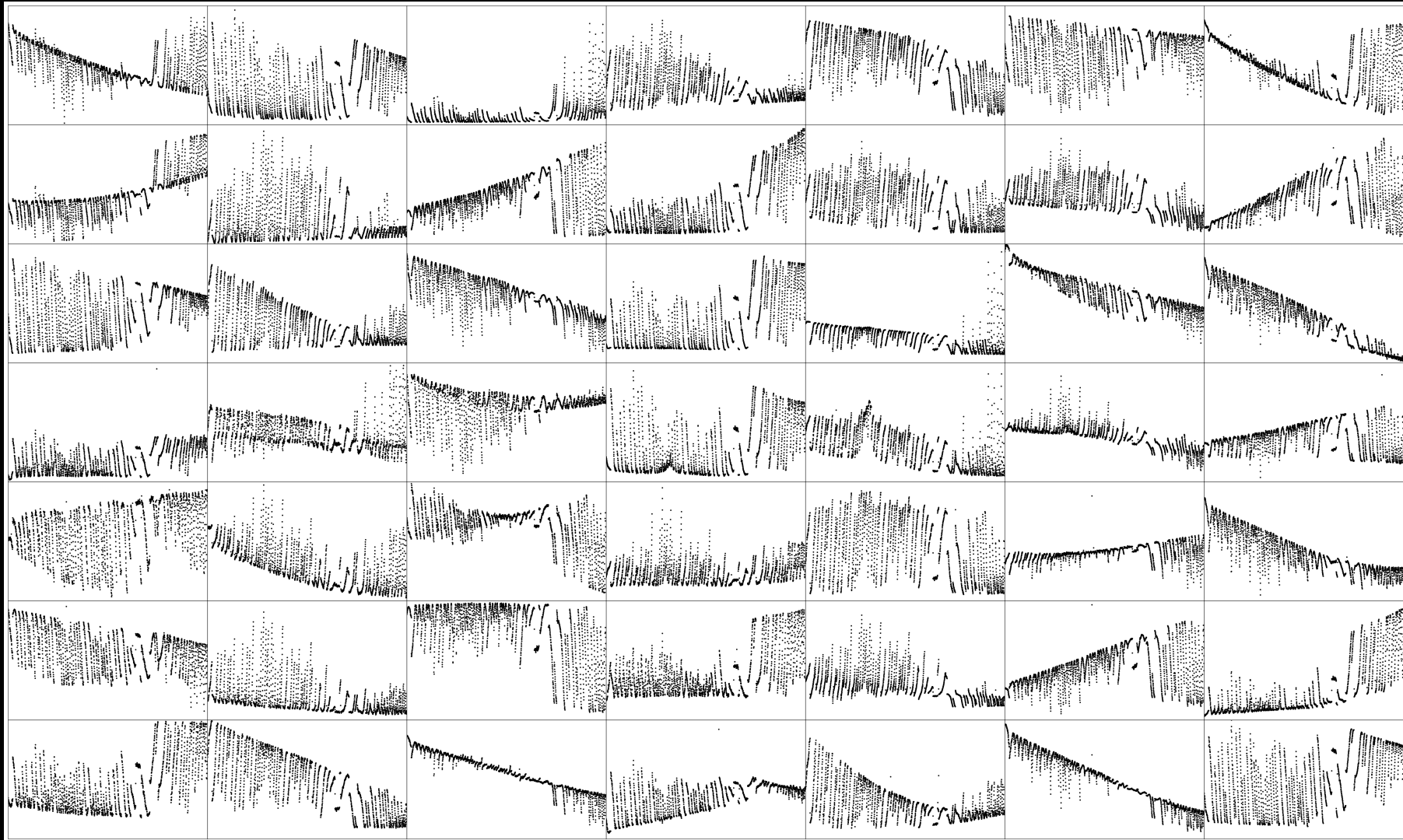


K2 photometry: Causal Pixel Model

Dun Wang
New York University



OGLE-2016-BLG-1050



- Crowded field Differencing
- Systematics (pointing, inter- and intra-pixel variation) CPM

Causal Pixel Model



David Hogg
NYU



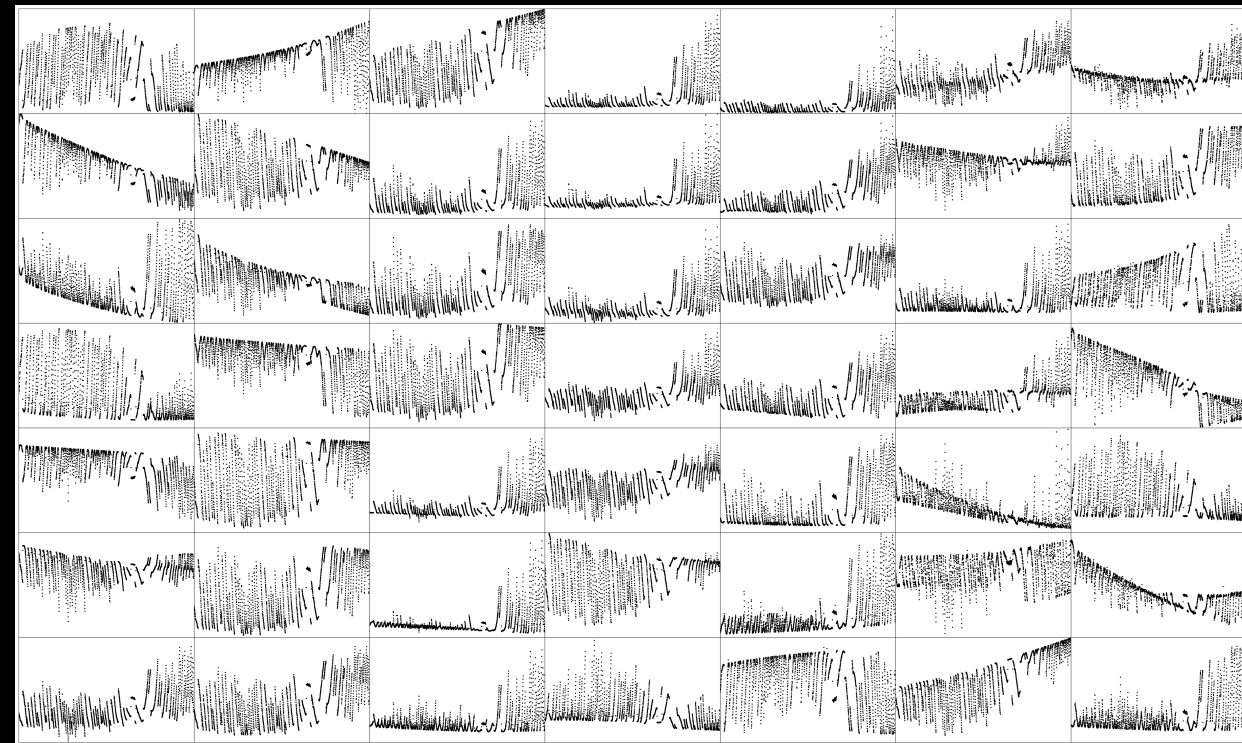
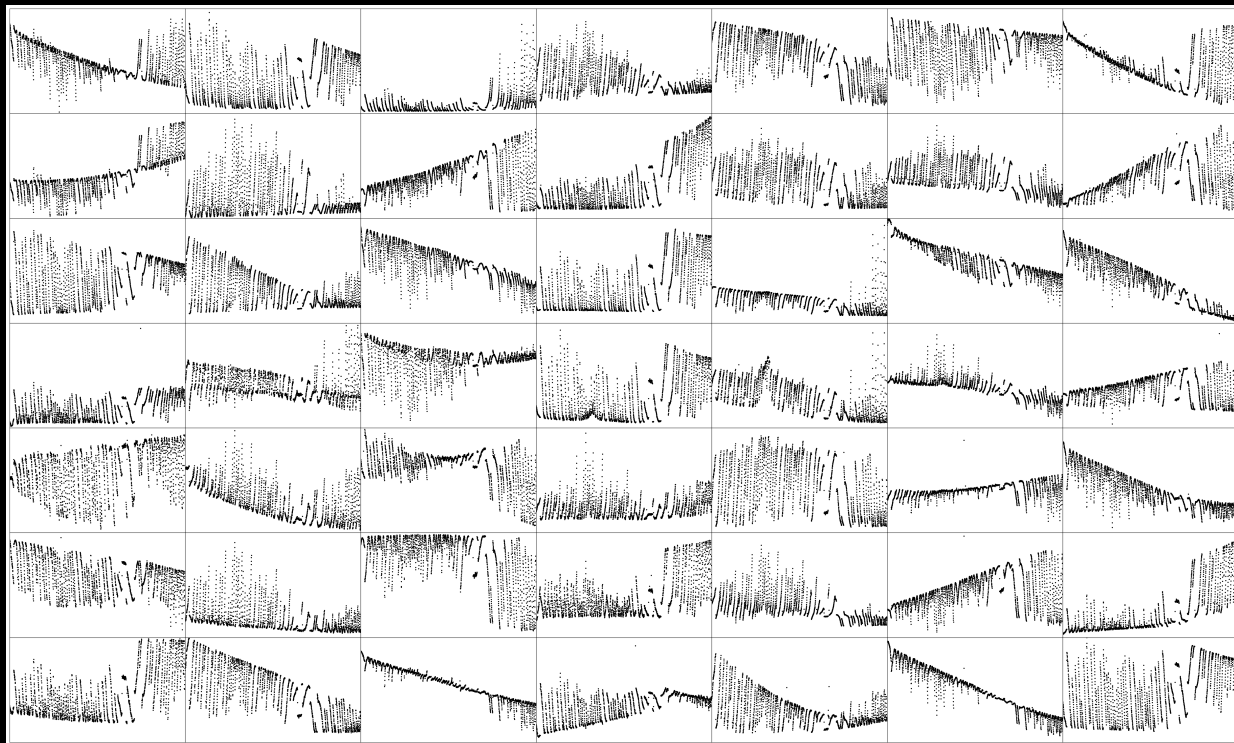
Daniel Foreman-Mackey
UW

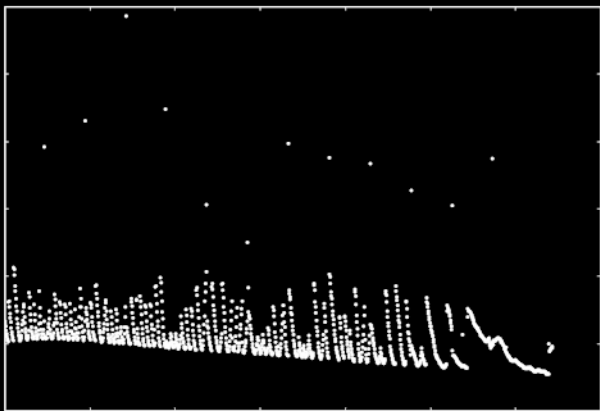


Bernhard Schölkopf
MPI-IS

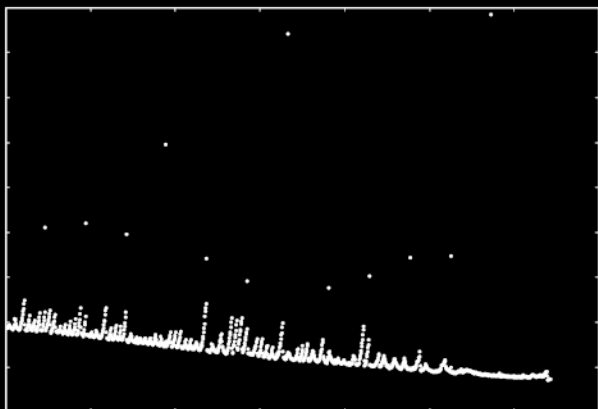
Causal Pixel Model

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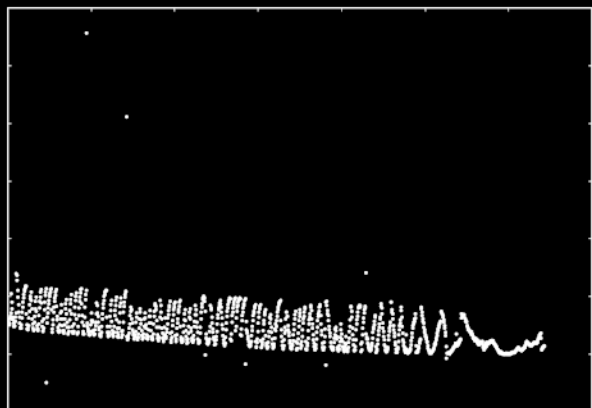




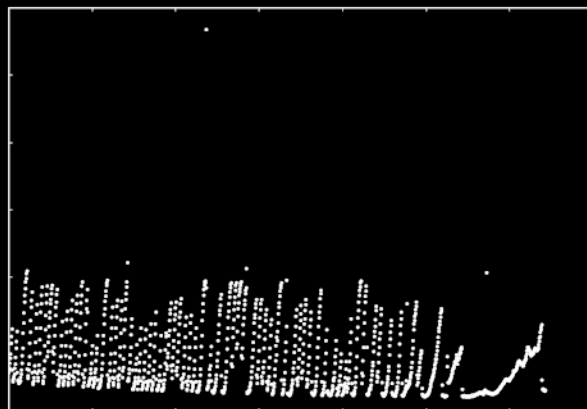
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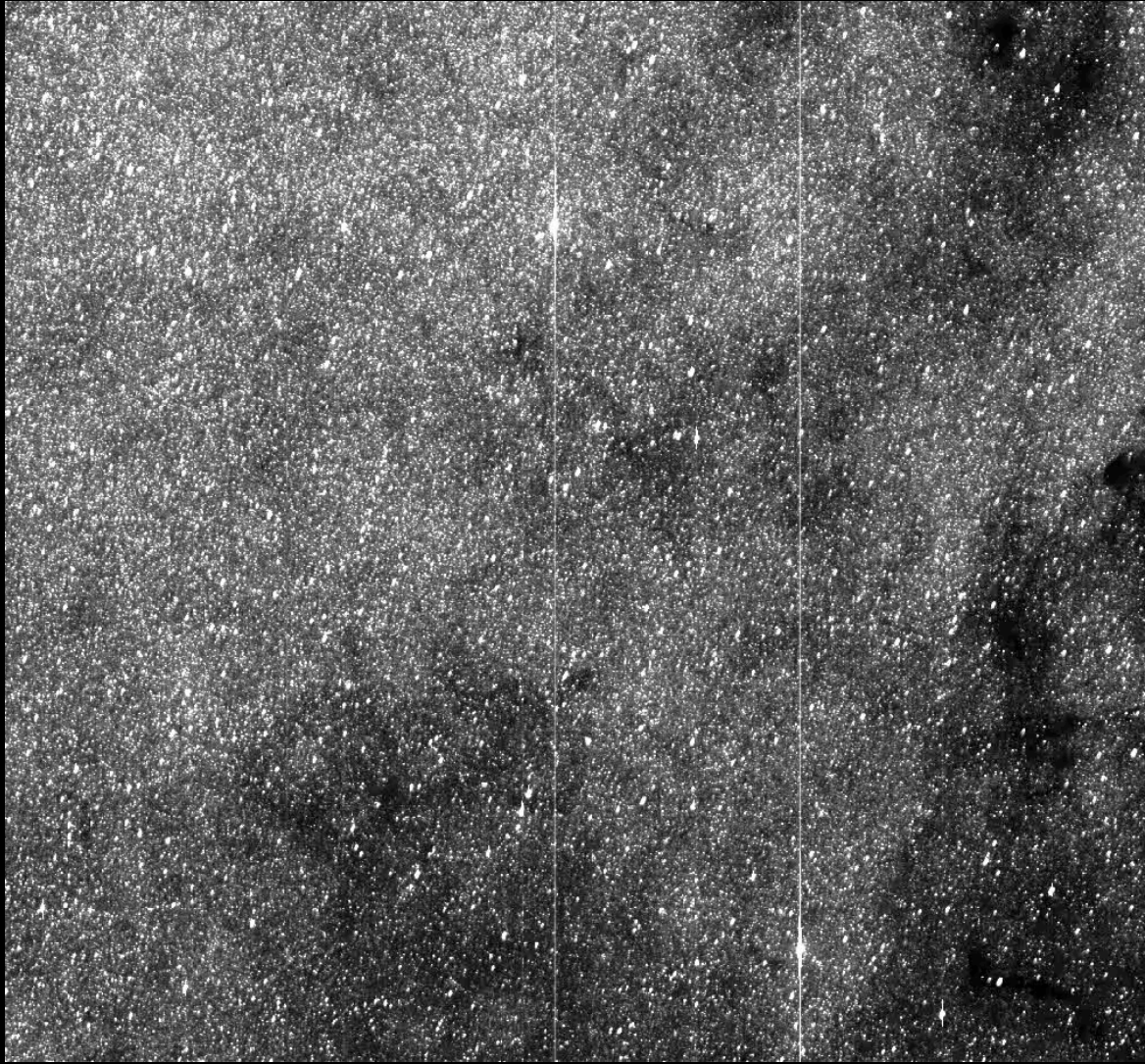
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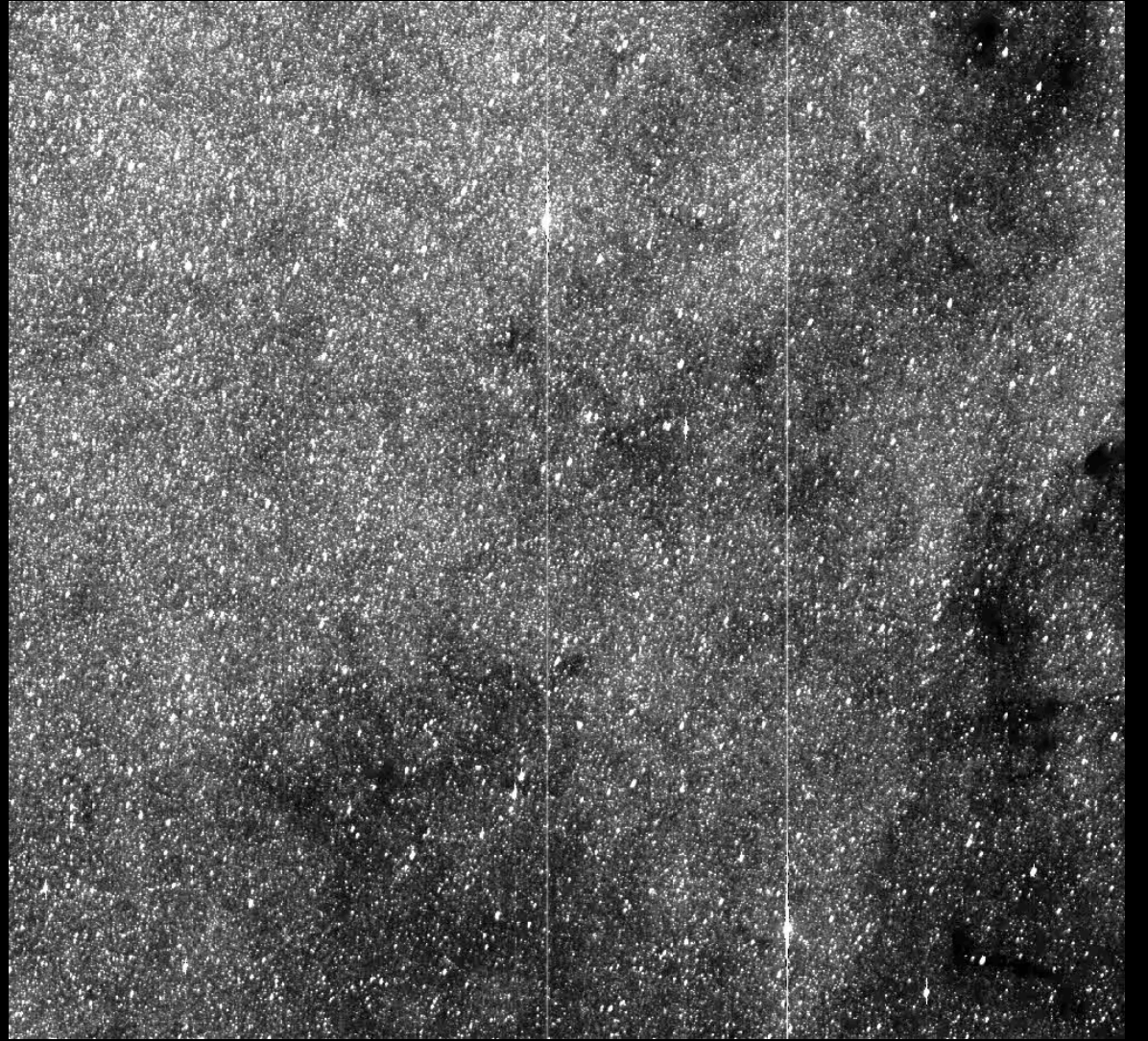
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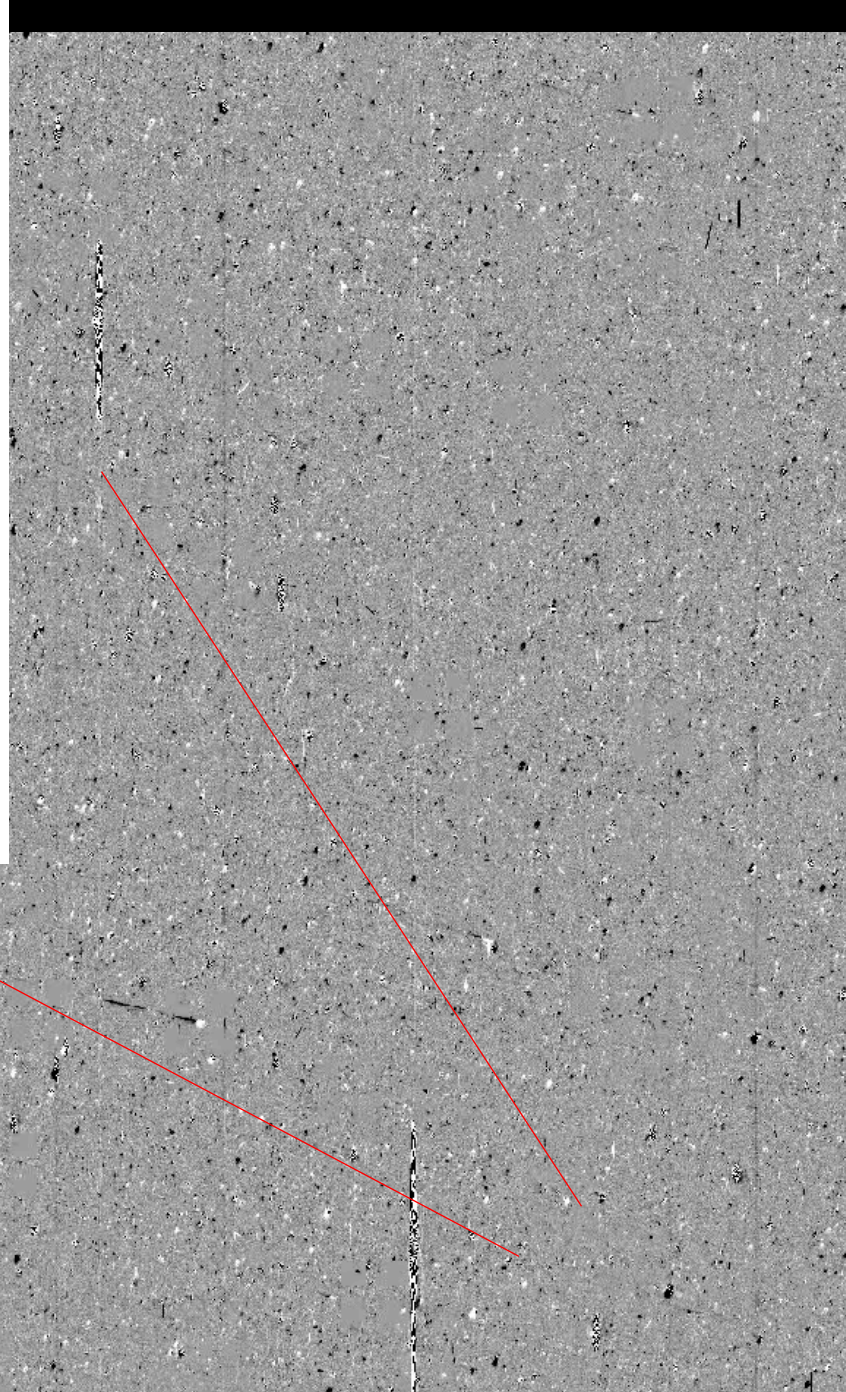
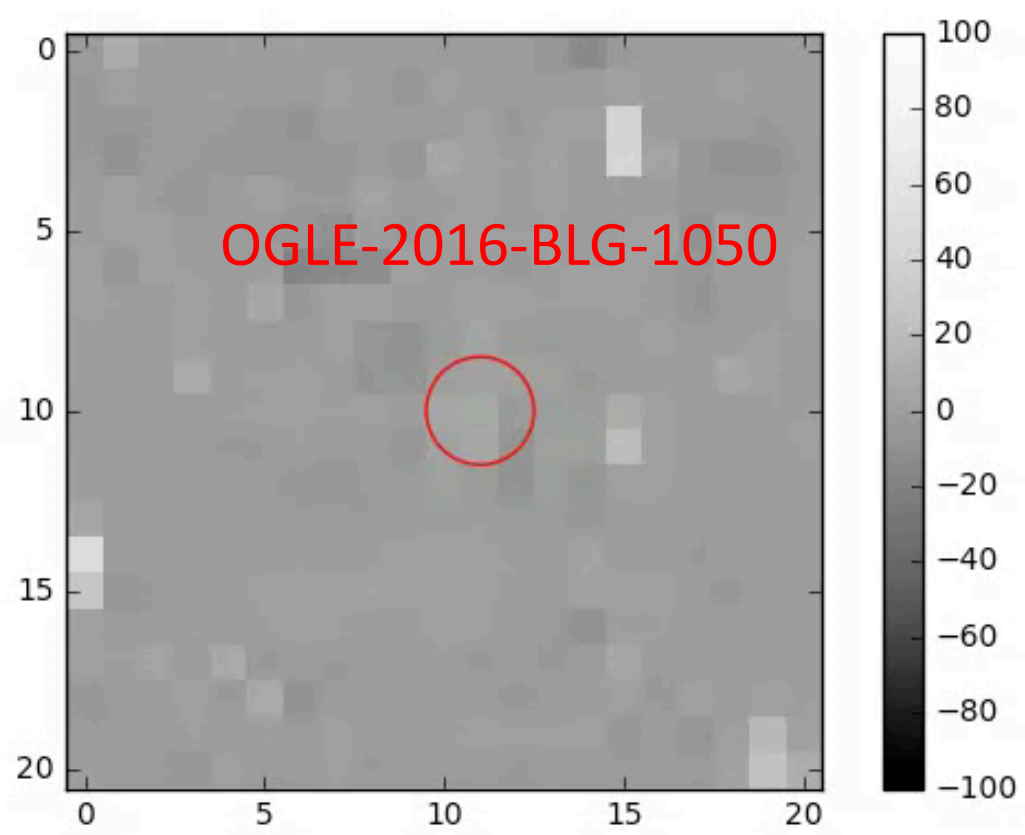
$$I_{m,n} = I_{m,n}^* + e_{m,n}$$

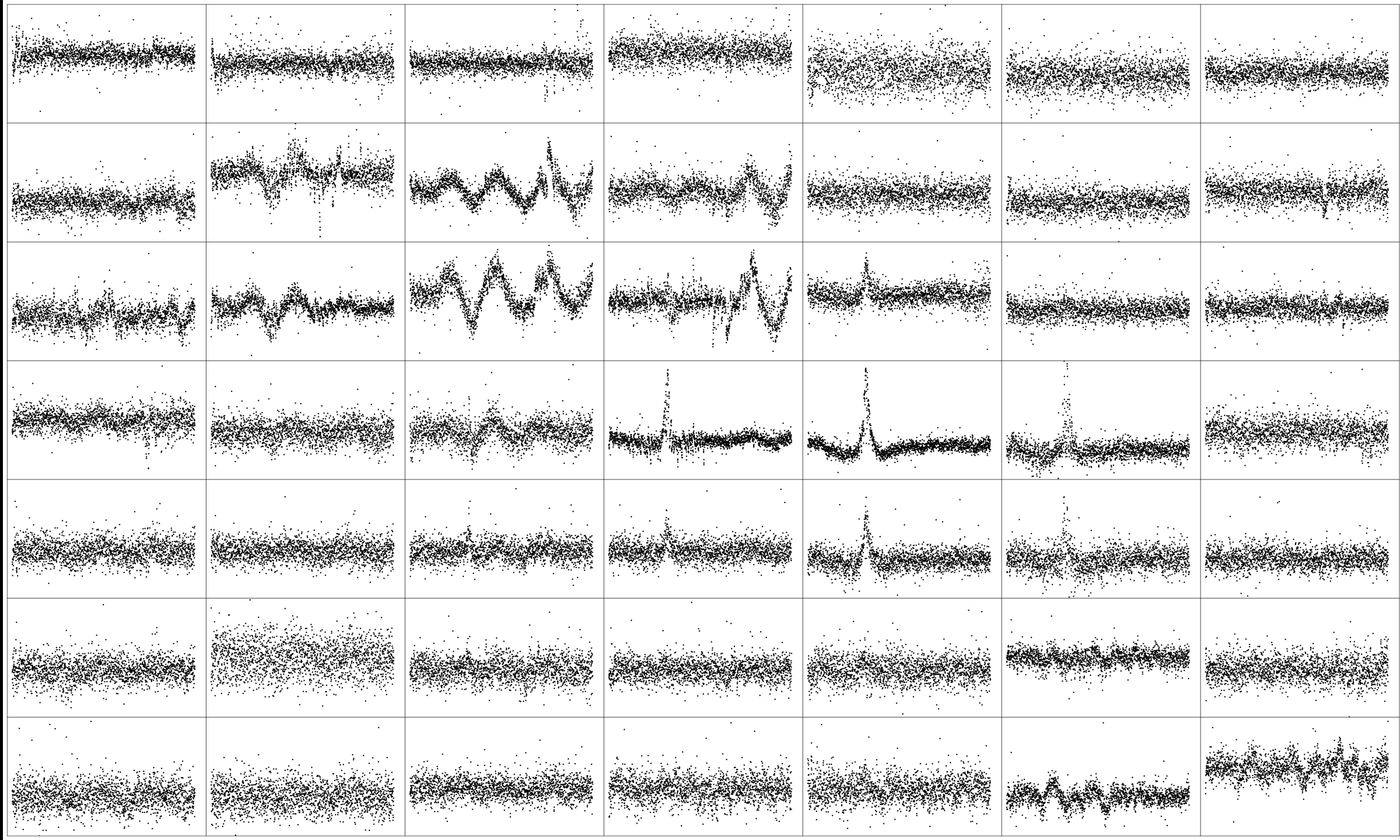
$$I_{m,n}^* = \sum_{m' \in \mathcal{M}_m} a_{m,n,m'} I_{m',n}$$

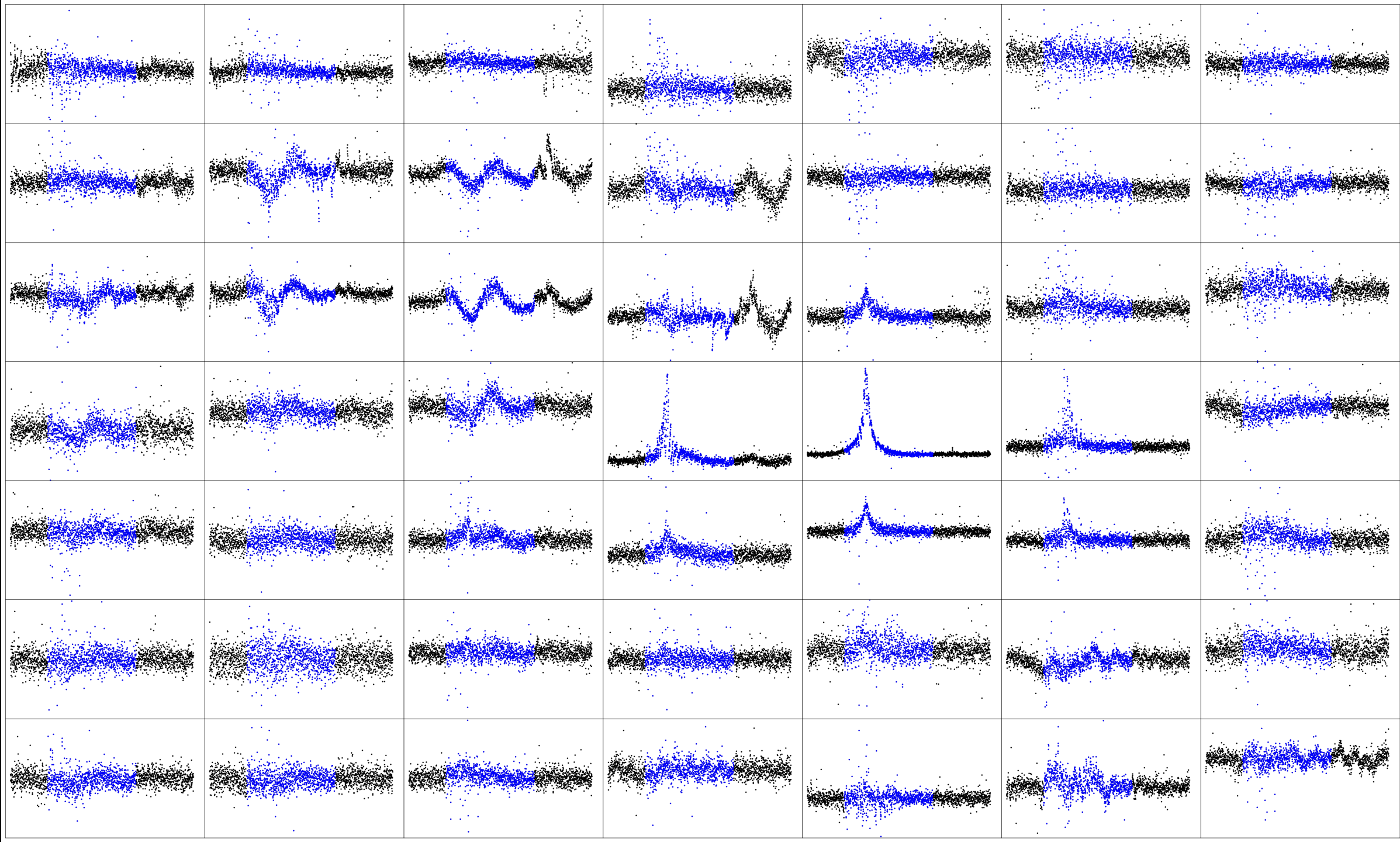


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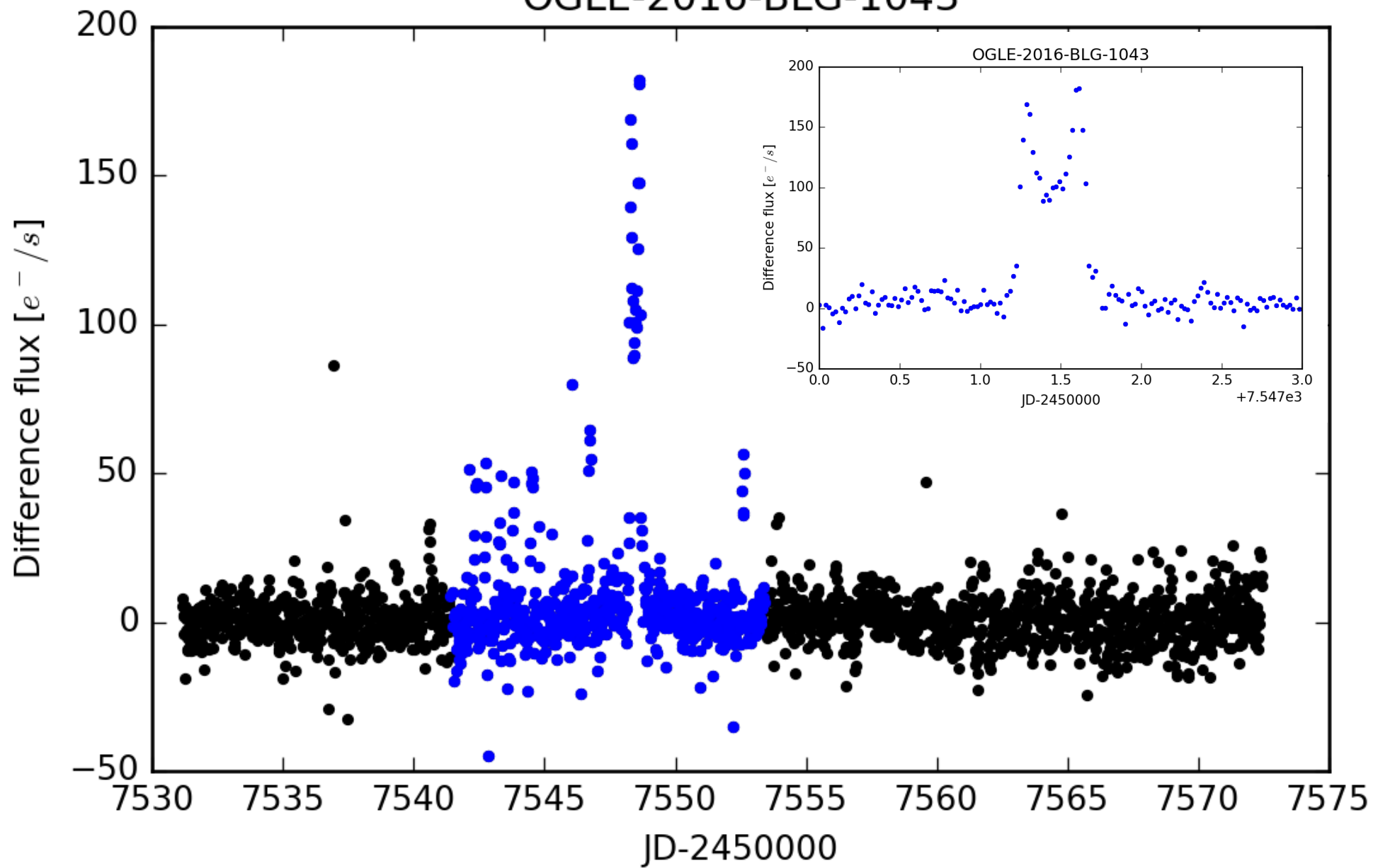


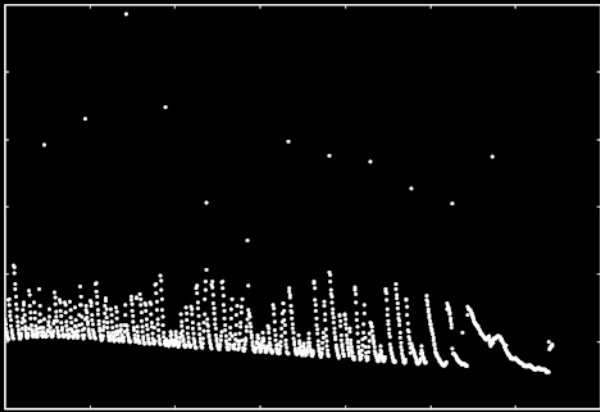






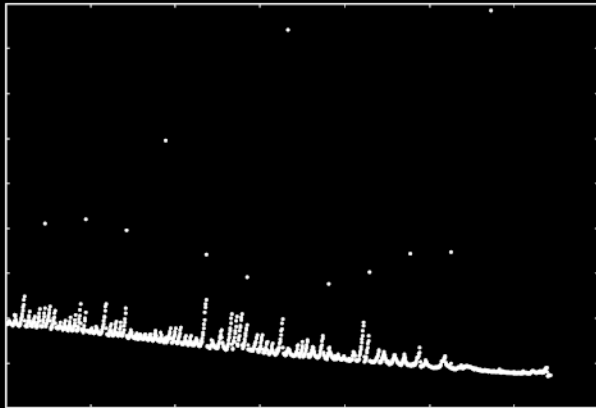
OGLE-2016-BLG-1043



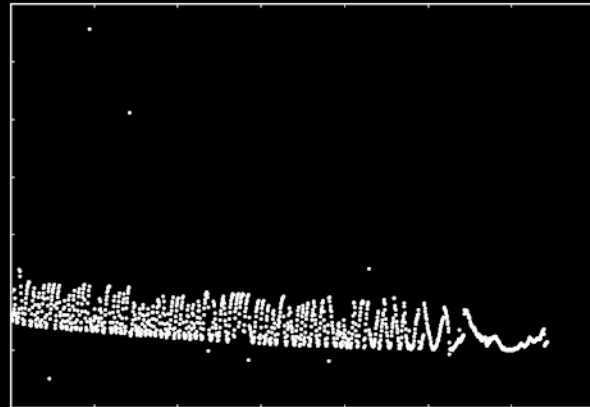


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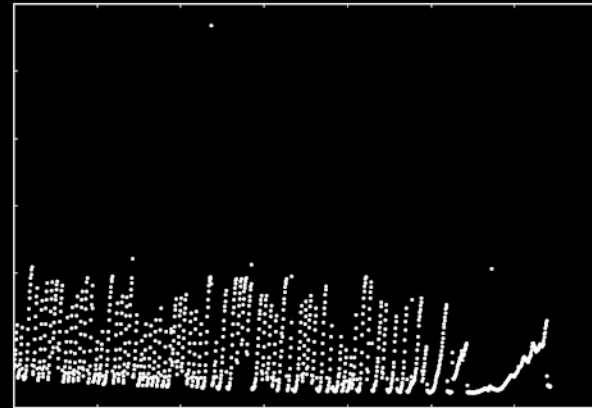
Systematics Model
(CPM)



+

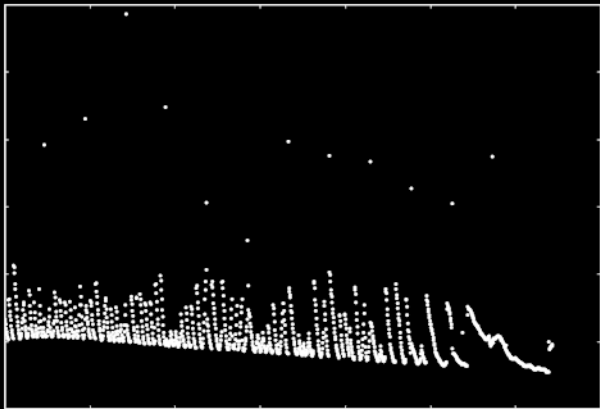


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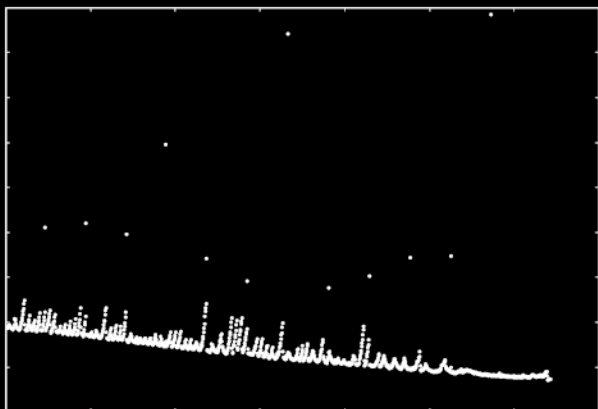
+ ...

Uncompleted
Model!

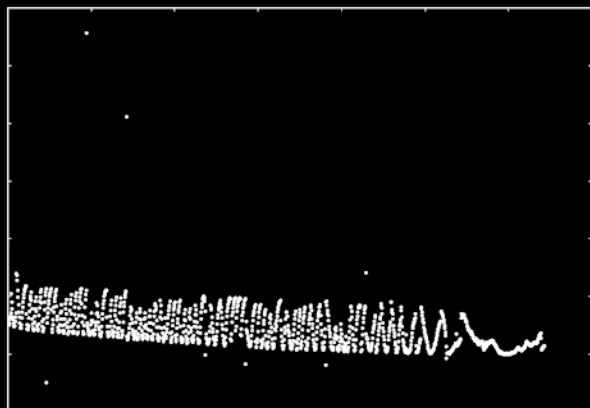


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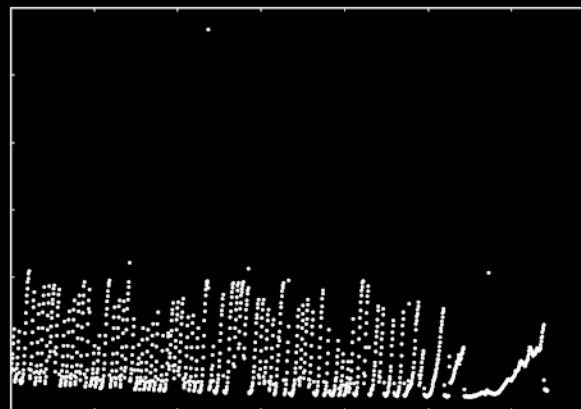
Systematics Model
(CPM)



+

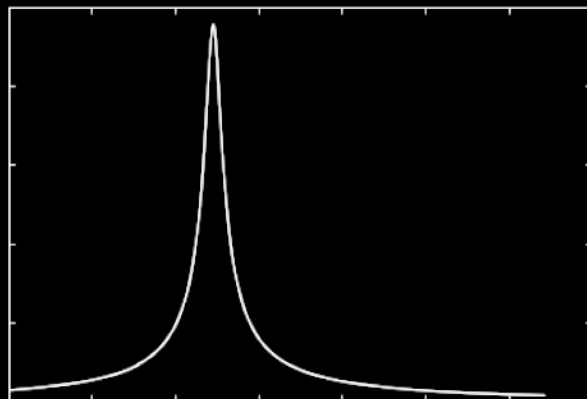


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Astronomical Model,
eg Microlensing

Time table

version usable for most people

- - better control of parameters
- - conventions better documented
- Timescale: 2 weeks from now

good understanding of CPM parameters

- - tests on selected short events including ground-based data and satellite parallax
- - tests on detached eclipsing binaries
- Timescale: 6 weeks from now

profile photometry

- - allow tabulated or analytical PRF
- - incorporate Kepler PRF
- Timescale: 10 weeks from now

photometry of the long events

- - tests on well-characterize events but with part of the data removed
- - tests on selected in- and out-of-superstamp events, also C11 events
- - tests on Mira variables (or other variable stars with light curves similar to microlensing events)
- Timescale: 15 weeks from now

fine-tuning details:

- - can we present binary events photometry without model assumed?
- - events with nearby variable stars
- Timescale: a few months from now