

# TESS

**Discovering Nearby Exoplanets  
with the  
Transiting Exoplanet Survey Satellite**

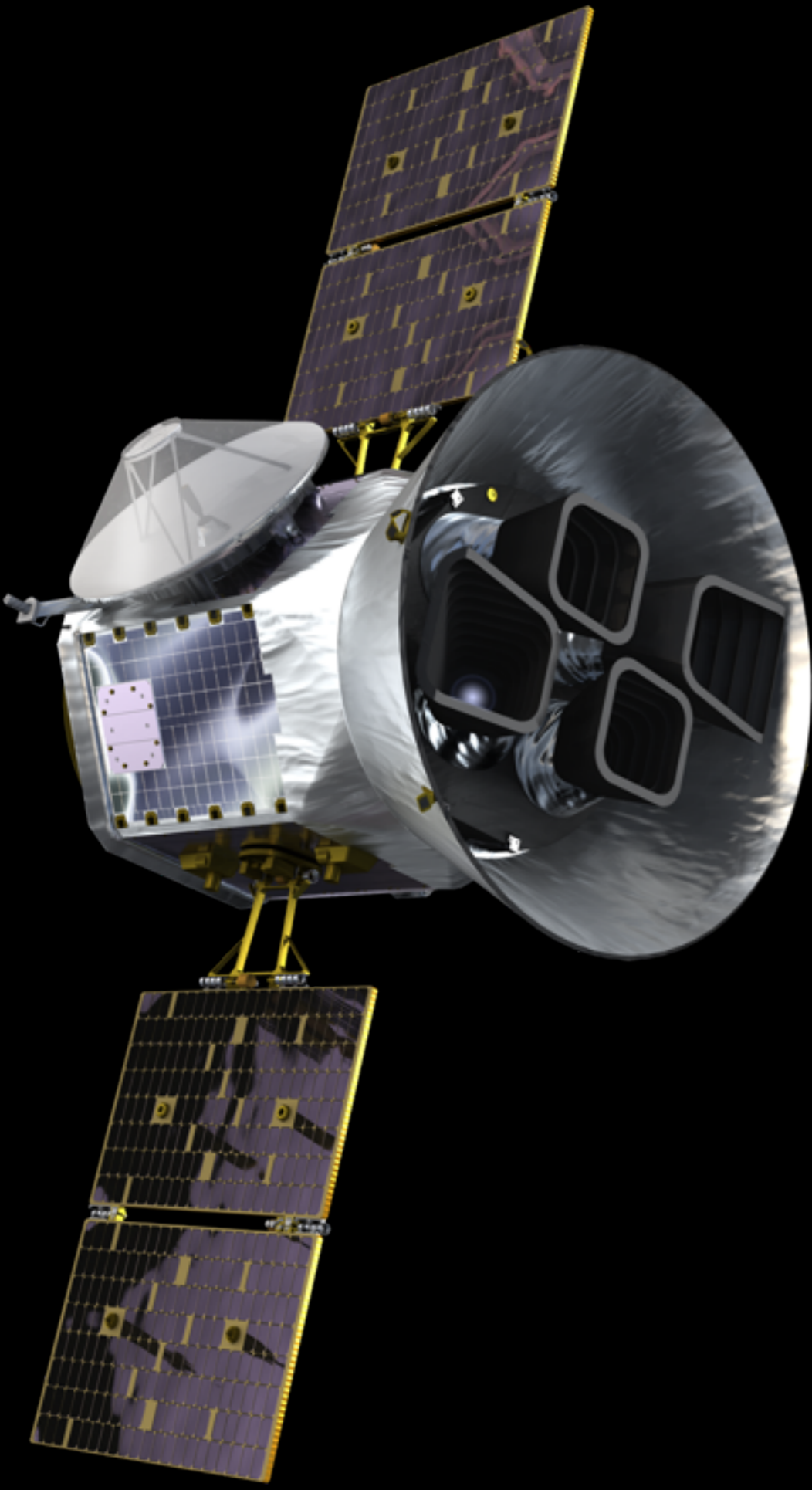
**George Ricker  
MIT**

***Know Thy Planet – Know Thy Star***  
**Pasadena CA**

**12 October 2017**

**+ Scientific Collaboration including:**

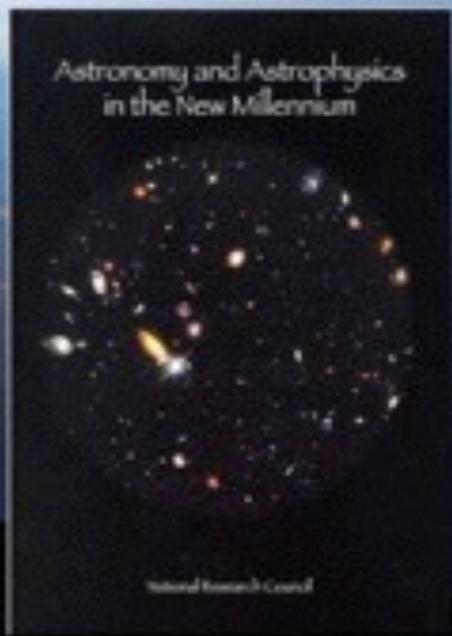
MIT/MKI, NASA Goddard, NASA Ames, STScI, SAO, Harvard University, Cornell University, Las Cumbres Observatory, Lowell Observatory, University of California, University of Maryland, Princeton University, Vanderbilt University, Yale University, Aarhus University, Geneva Observatory, OHP-France,



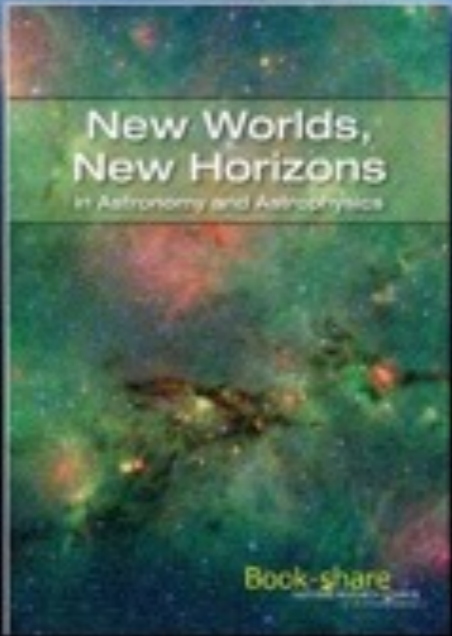
## NASA's Exoplanet Missions



Ground-based Observatories

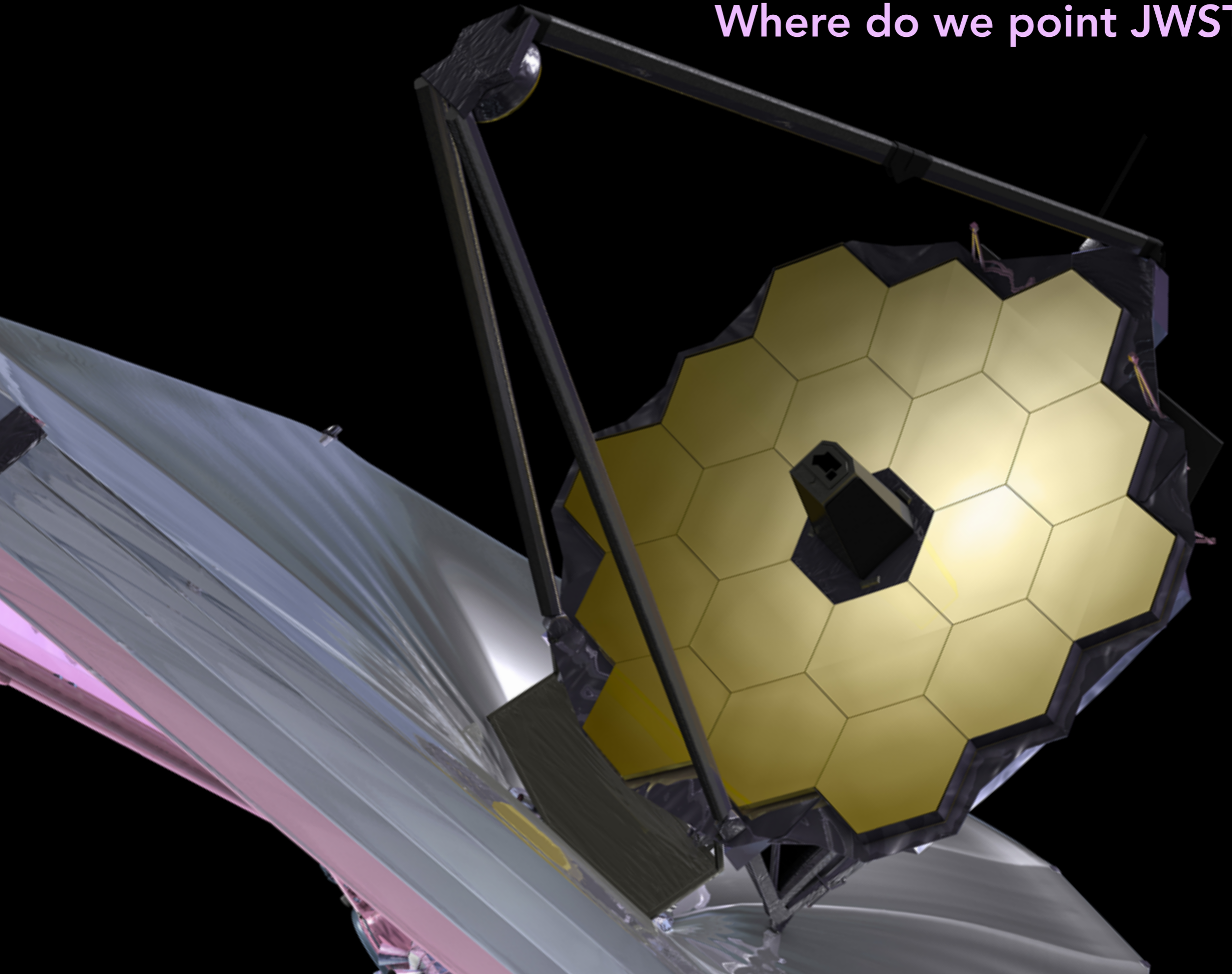


2001 Decadal Survey

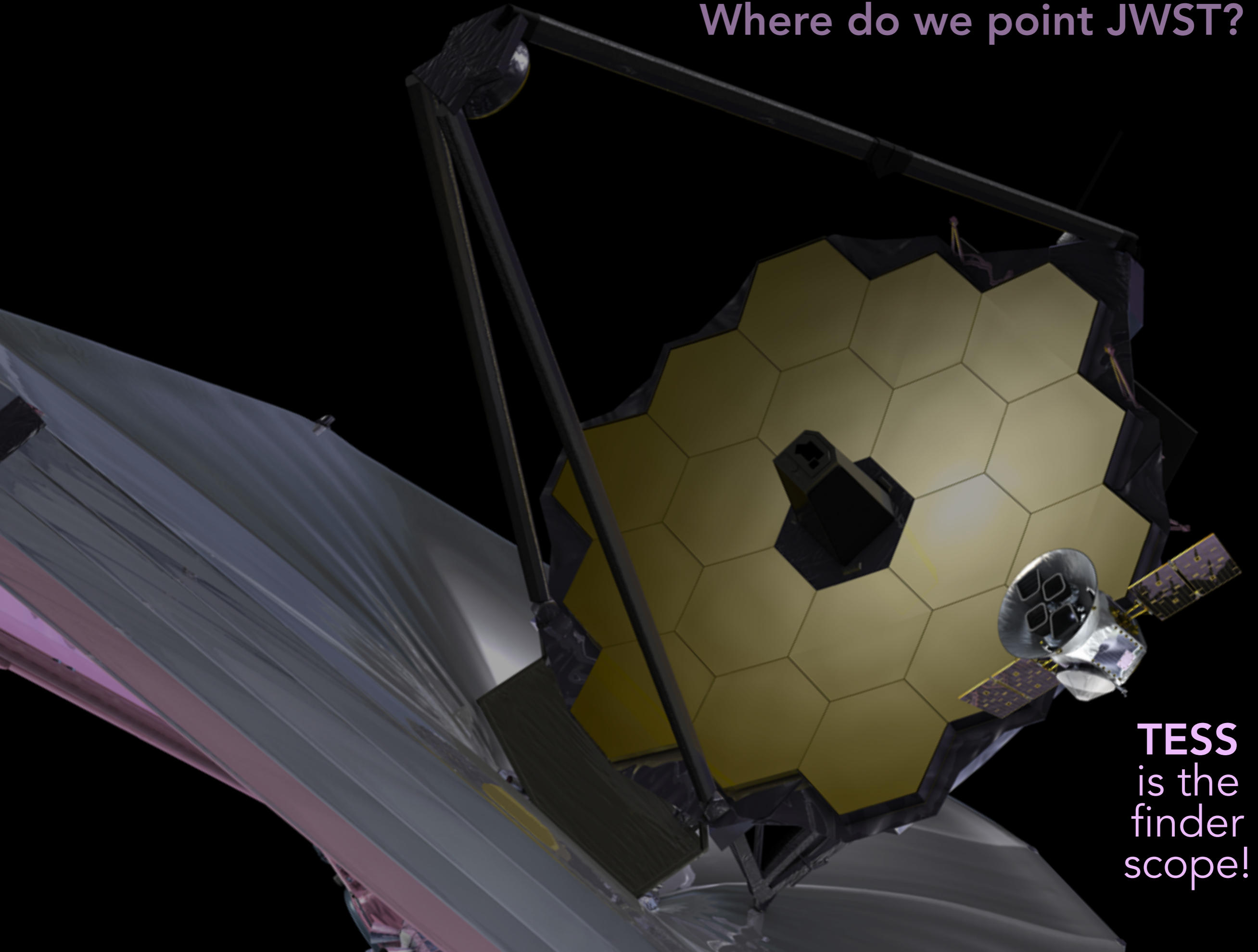


2010 Decadal Survey

Where do we point JWST?



Where do we point JWST?



**TESS**  
is the  
finder  
scope!

- ◆ Definition of Etendue [ $m^2 \cdot \text{deg}^2$ ]:

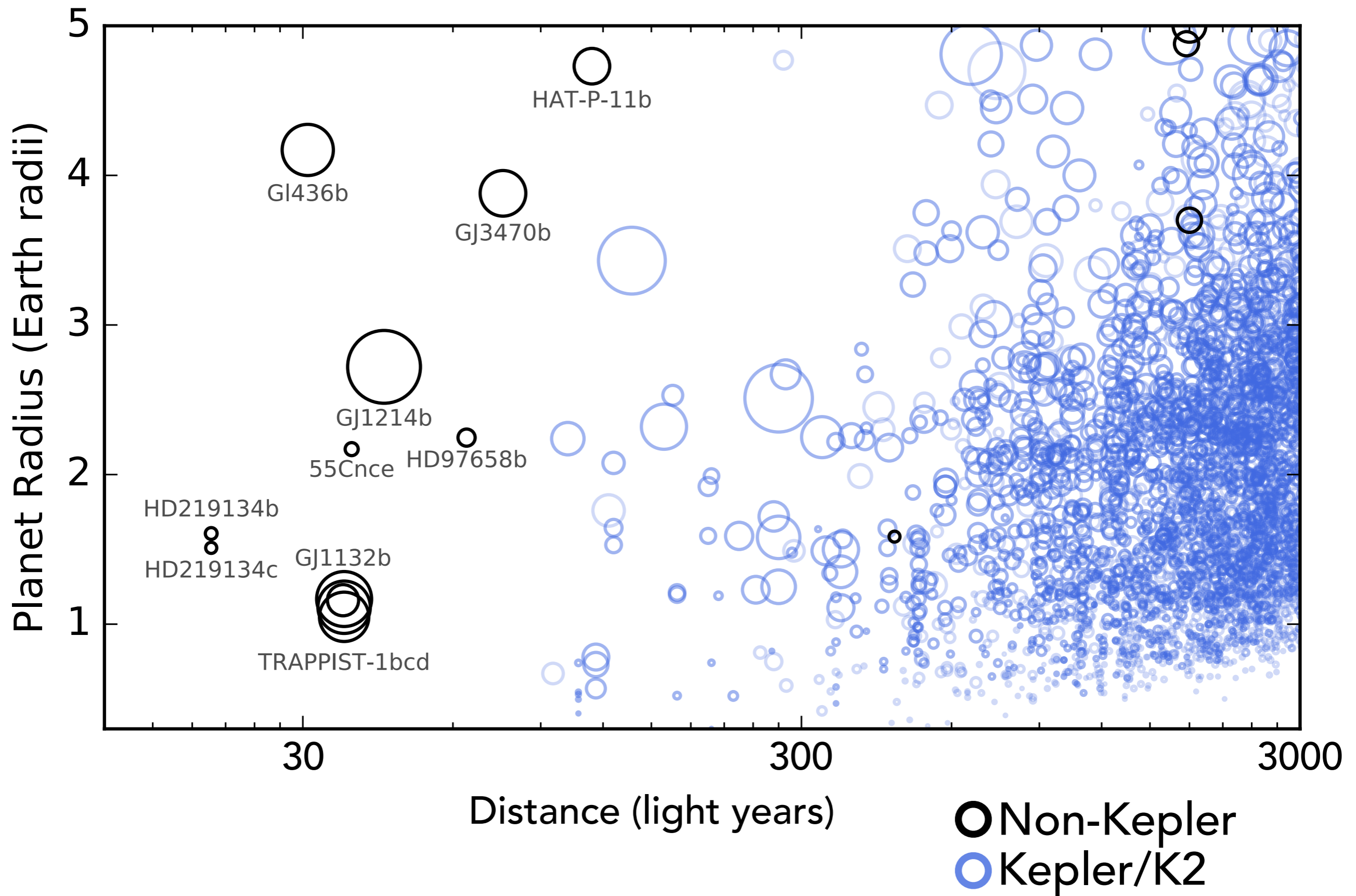
$$E = A_{\text{optics}} * \Omega_{\text{net}}$$

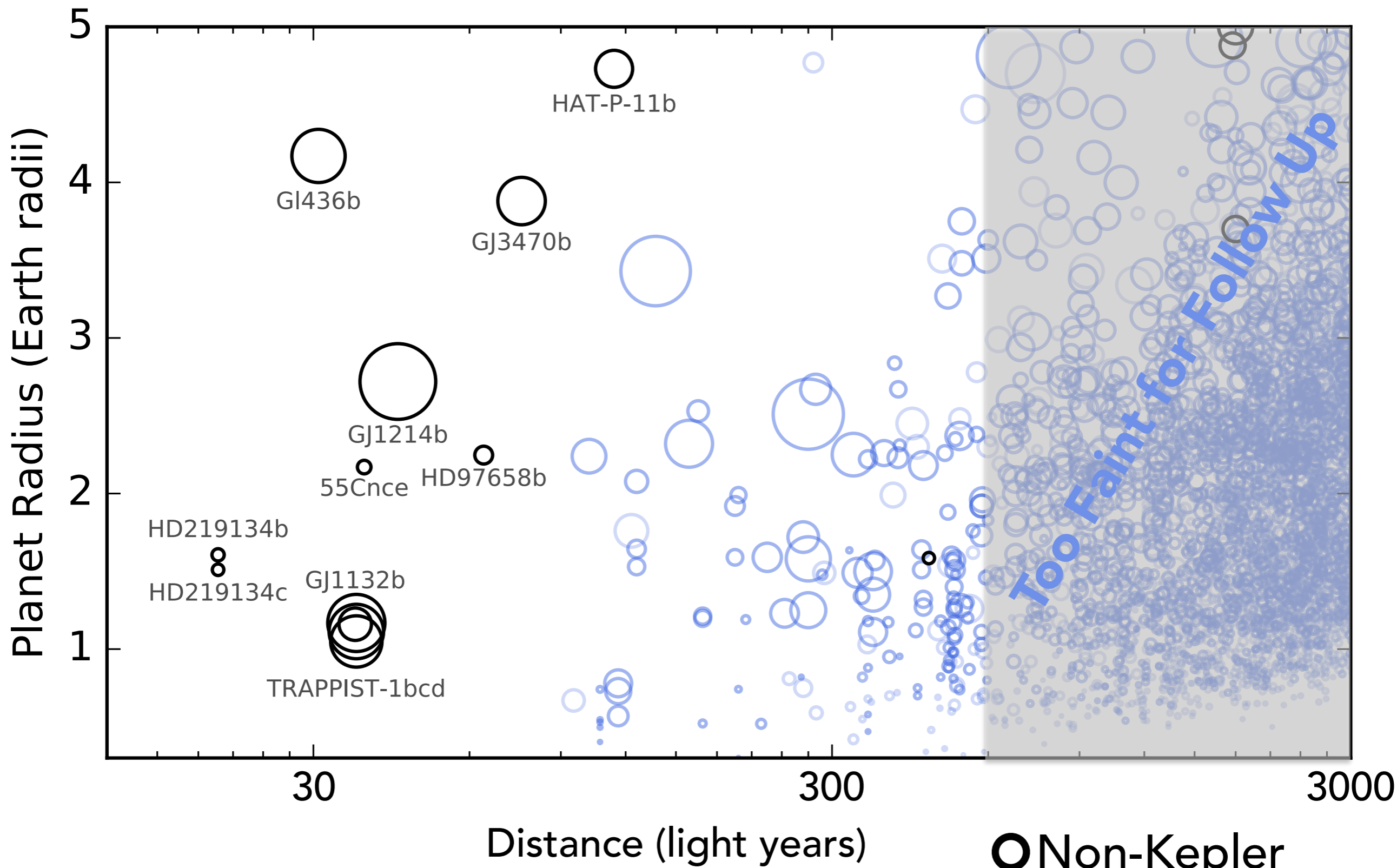
$$\text{where } \Omega_{\text{net}} = \Omega_{\text{gross}} * \left( \frac{\# \text{ pixels telemetered}}{\# \text{ pixels in focal plane}} \right)$$

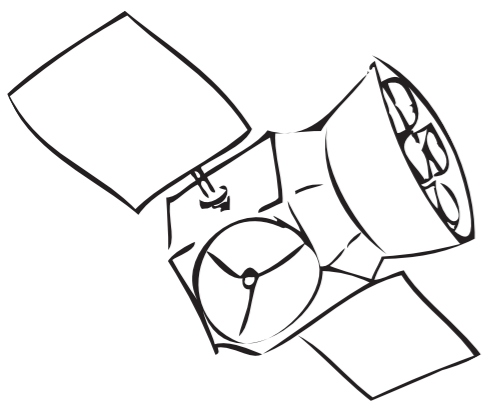
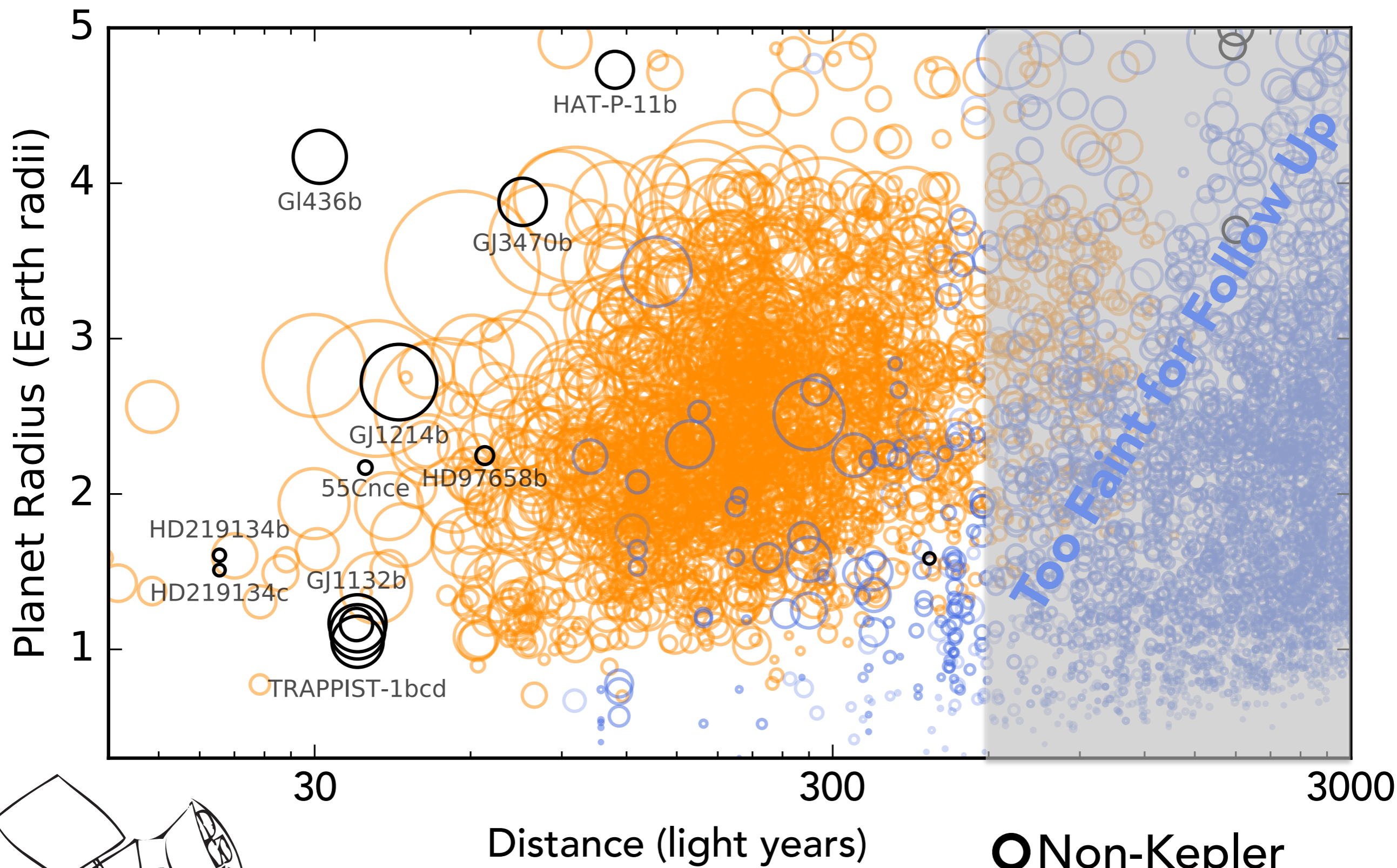
	$A_{\text{optics}}$ [ $m^2$ ]	$\Omega_{\text{gross}}$ [ $\text{deg}^2$ ]	$\frac{\# \text{ pixels telemetered}}{\# \text{ pixels in focal plane}}$	E [ $m^2 \text{ deg}^2$ ]
<b>TESS</b>	0.0095	2304	1	21.9
<b>Kepler</b>	0.71	105	0.06	4.2

**TESS is the highest etendue optical space mission ever flown:  
~5 times greater than Kepler**

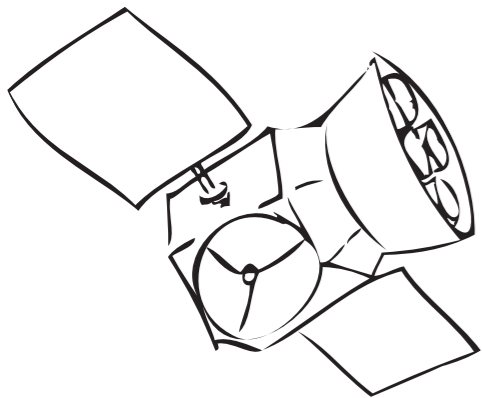
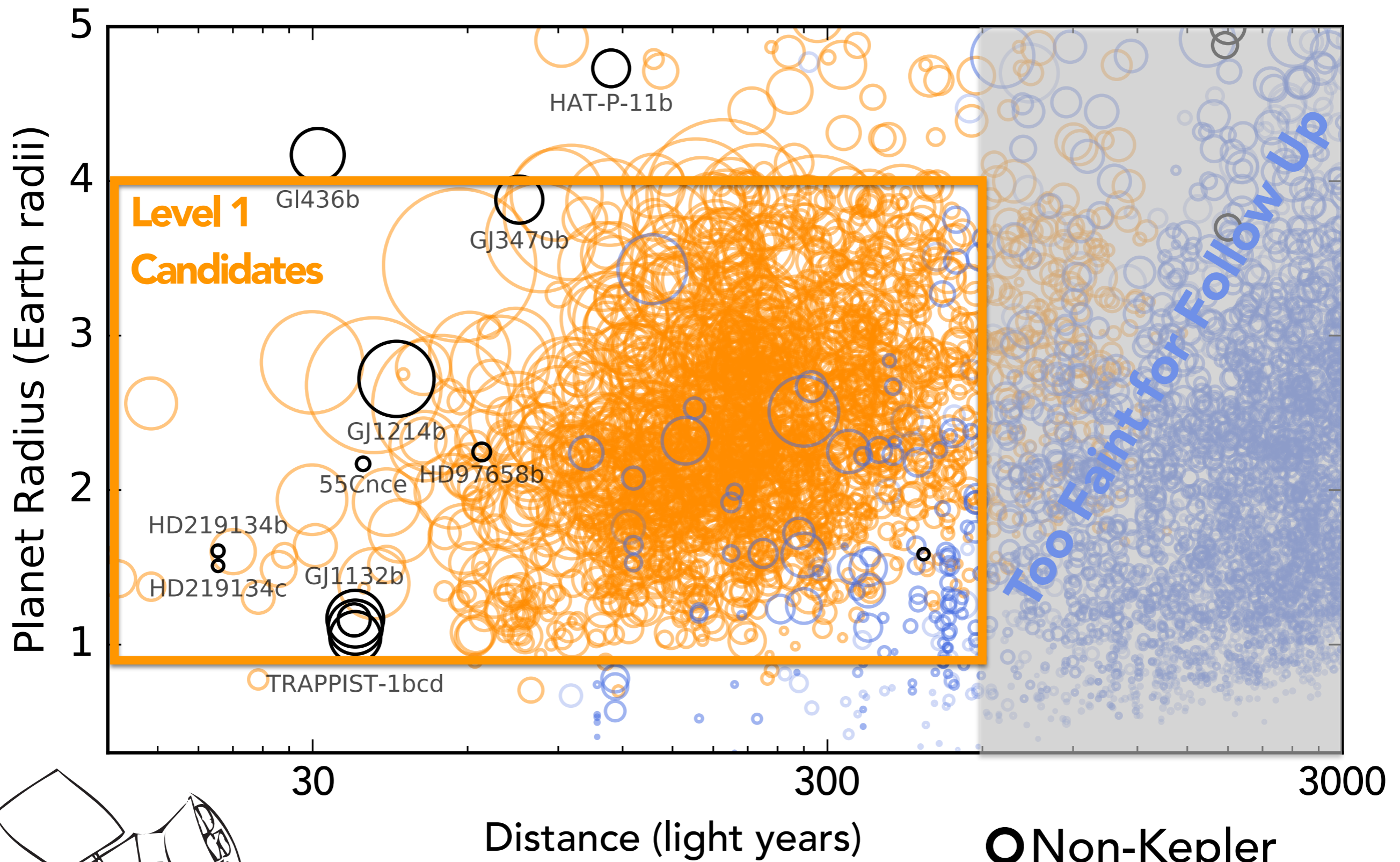
Bryson et al. 2010  
Ricker et al. 2016

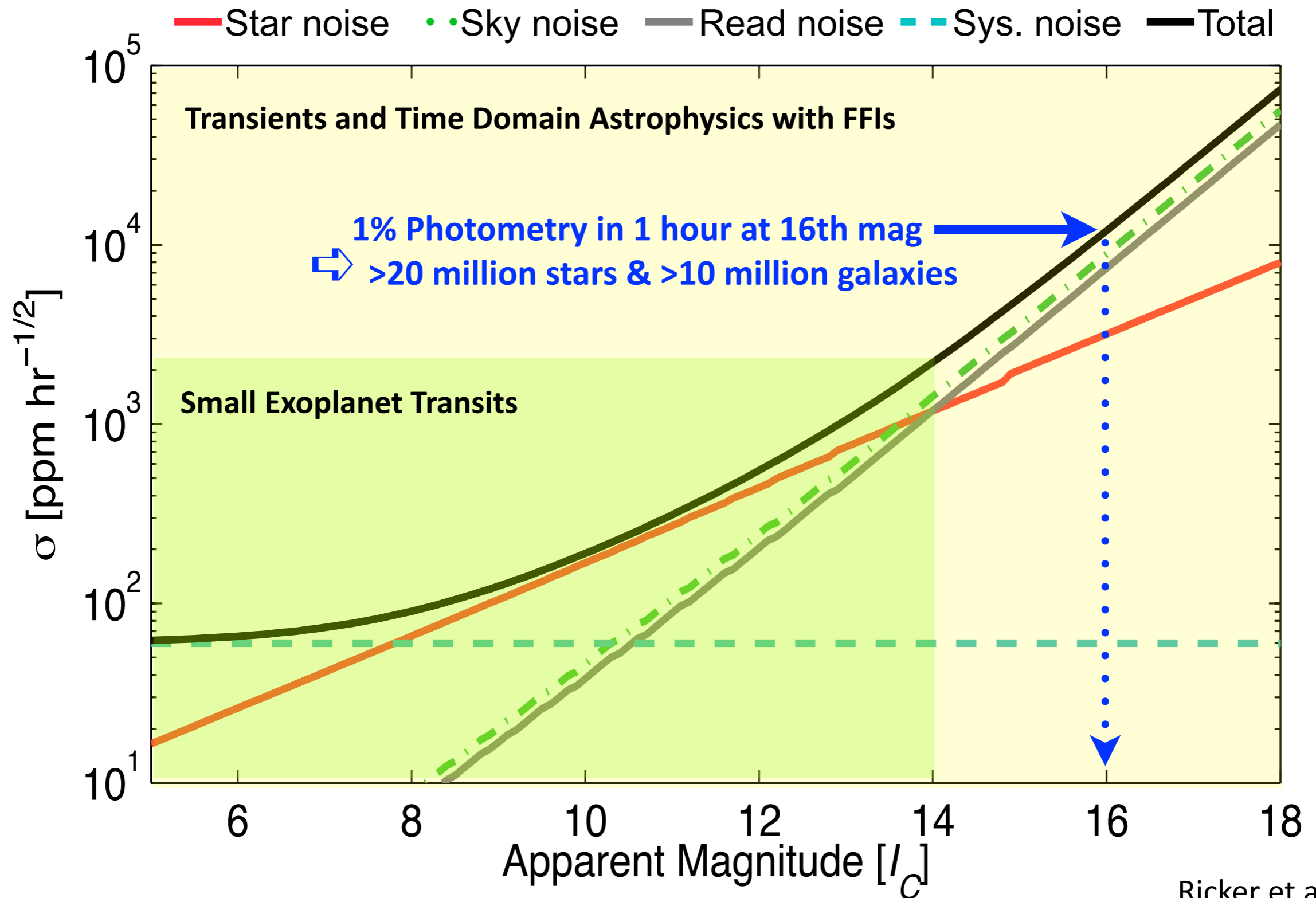




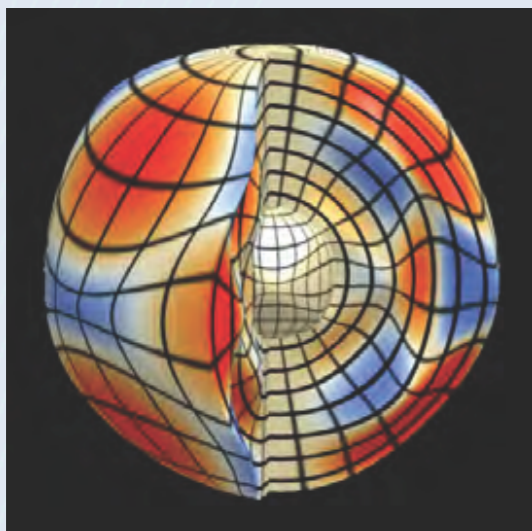








Ricker et al. 2016



## Non-Transiting Exoplanets

- ◆ Microlensing events

## Solar System

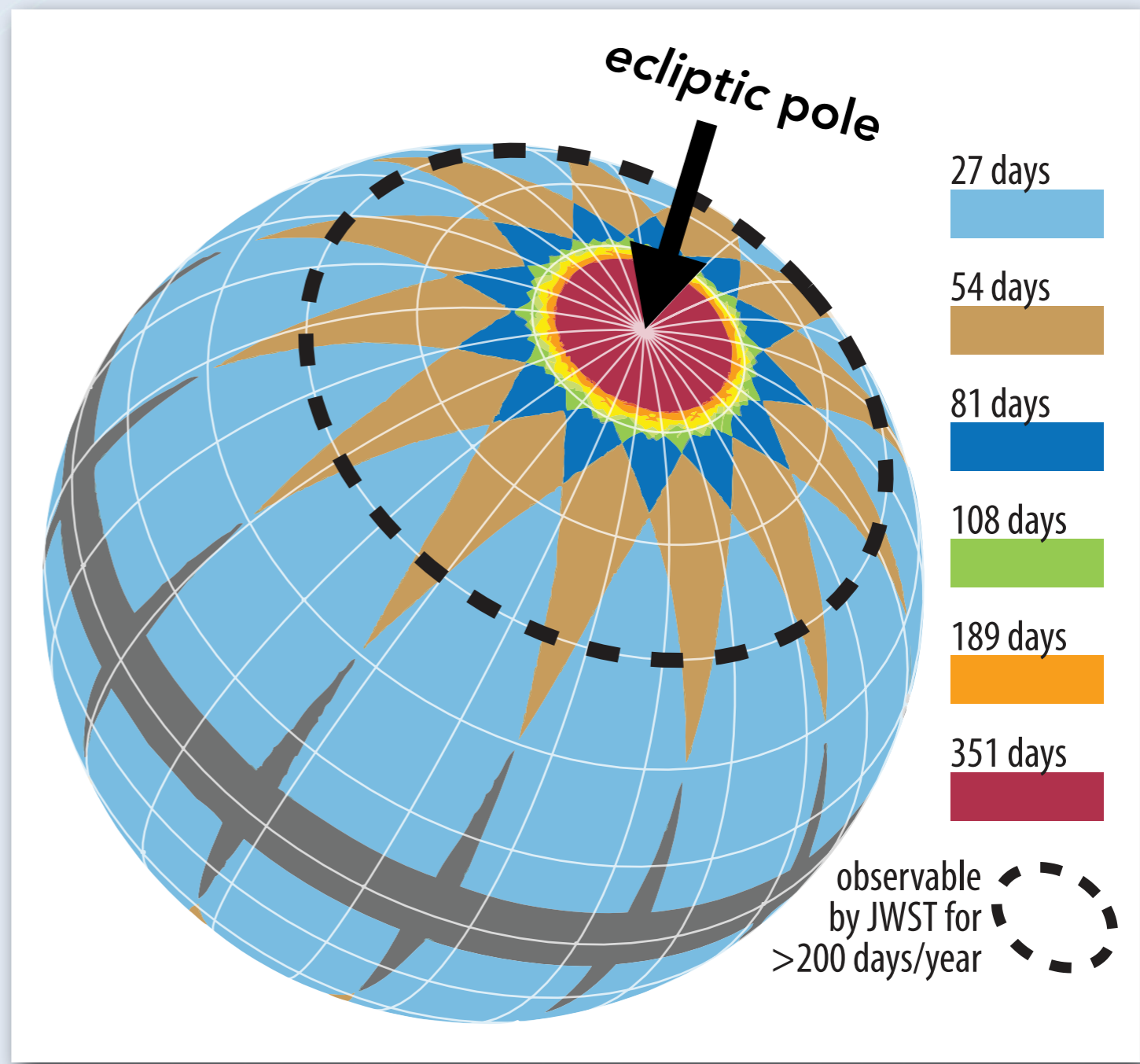
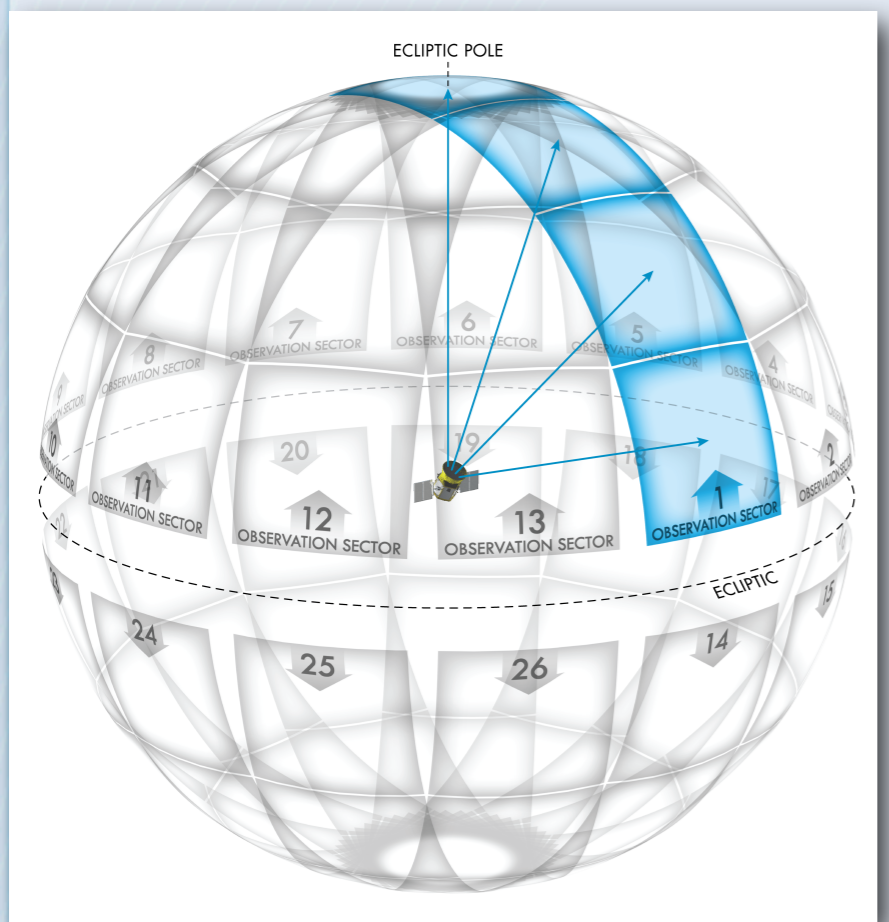
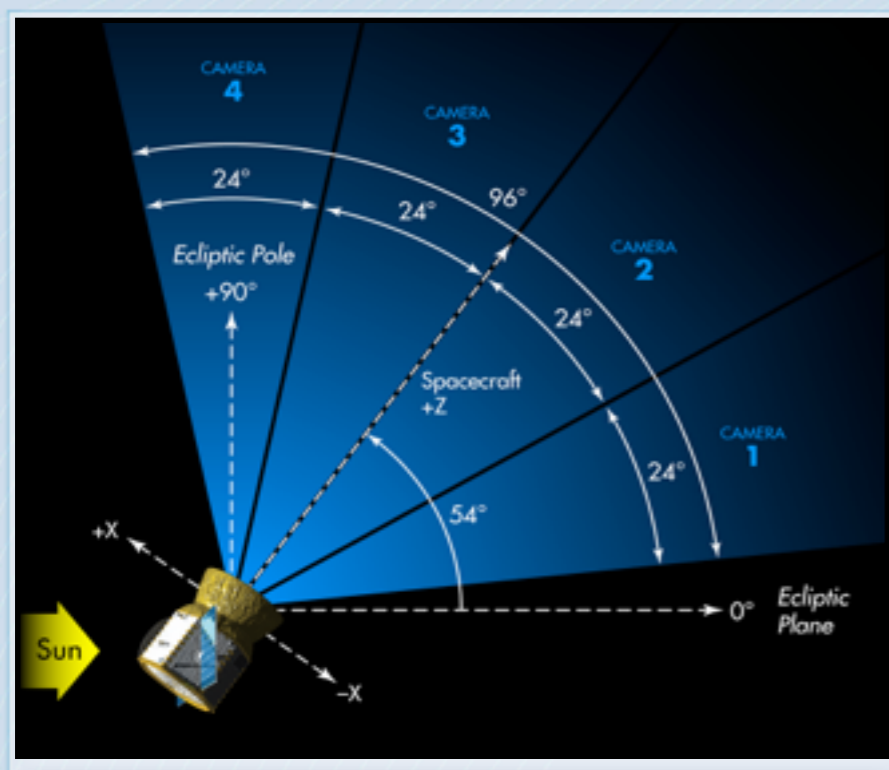
- ◆ Occultation Events
- ◆ Comets
- ◆ Asteroids

## Extragalactic Sources

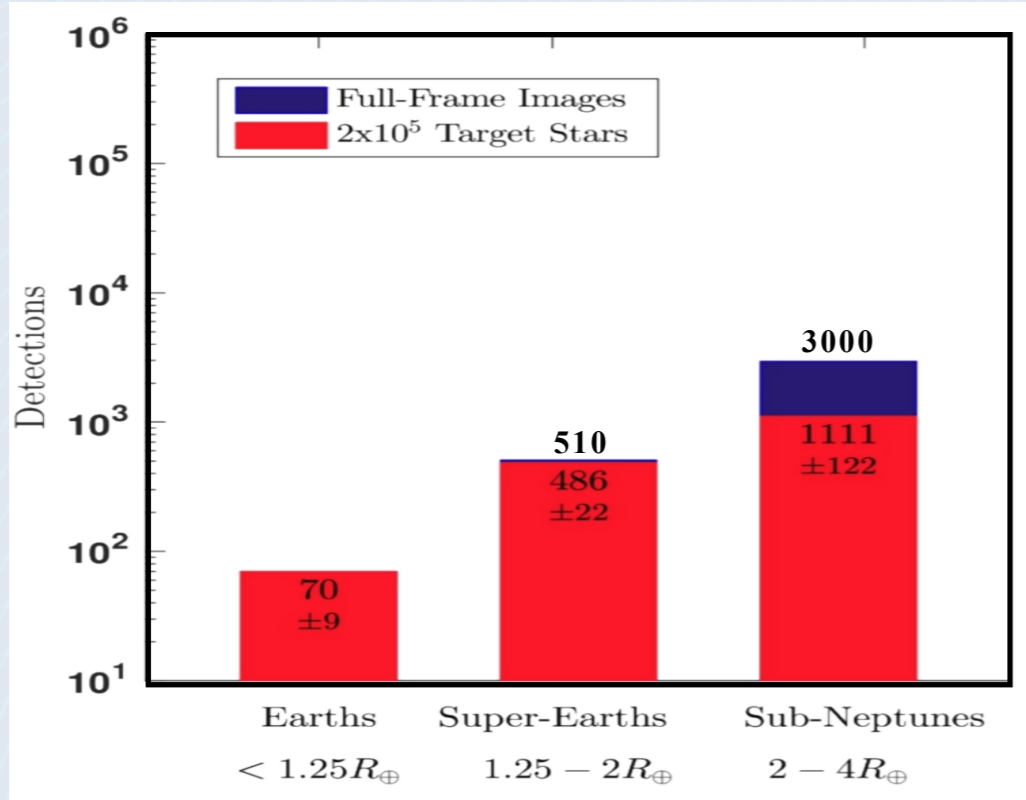
- ◆ Supernovae
- ◆ AGNs
- ◆ Blazars
- ◆ Quasars
- ◆ Tidal Disruption Events
- ◆ Gamma-ray Bursts
- ◆ Kilonovae
- ◆ Hypernovae

## Stars

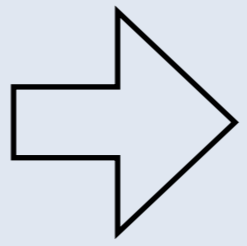
- ◆ Asteroseismology
- ◆ Brown Dwarfs
- ◆ Eclipsing Binaries
- ◆ Flare Stars
- ◆ Cepheids
- ◆ T Tauri Stars
- ◆ Cluster Gyrochronology
- ◆ White Dwarfs
- ◆ Neutron Stars
- ◆ Emission line stars (Be stars)
- ◆ RR Lyrae Stars
- ◆ WD Oscillations
- ◆ Novae
- ◆ Young Stellar Objects



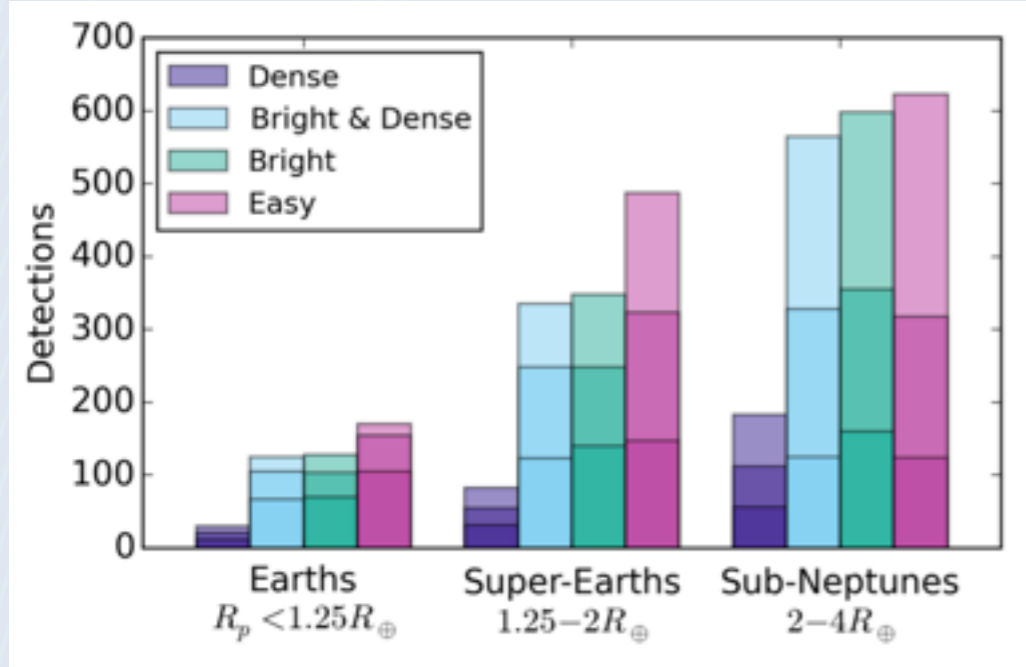
Ricker et al., JATIS, (2014)



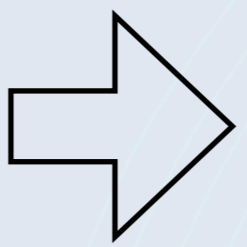
Sullivan et al. 2015



Sullivan+15	For 2 x 10 <sup>5</sup> 2min Targets	Scaled: For 5 x 10 <sup>4</sup> 2min Targets
Earths	70	18
Super-Earths	486	122
Sub-Neptunes	1111	278
<b>TOTAL</b>	<b>1667</b>	<b>417</b>



Muirhead et al. 2017



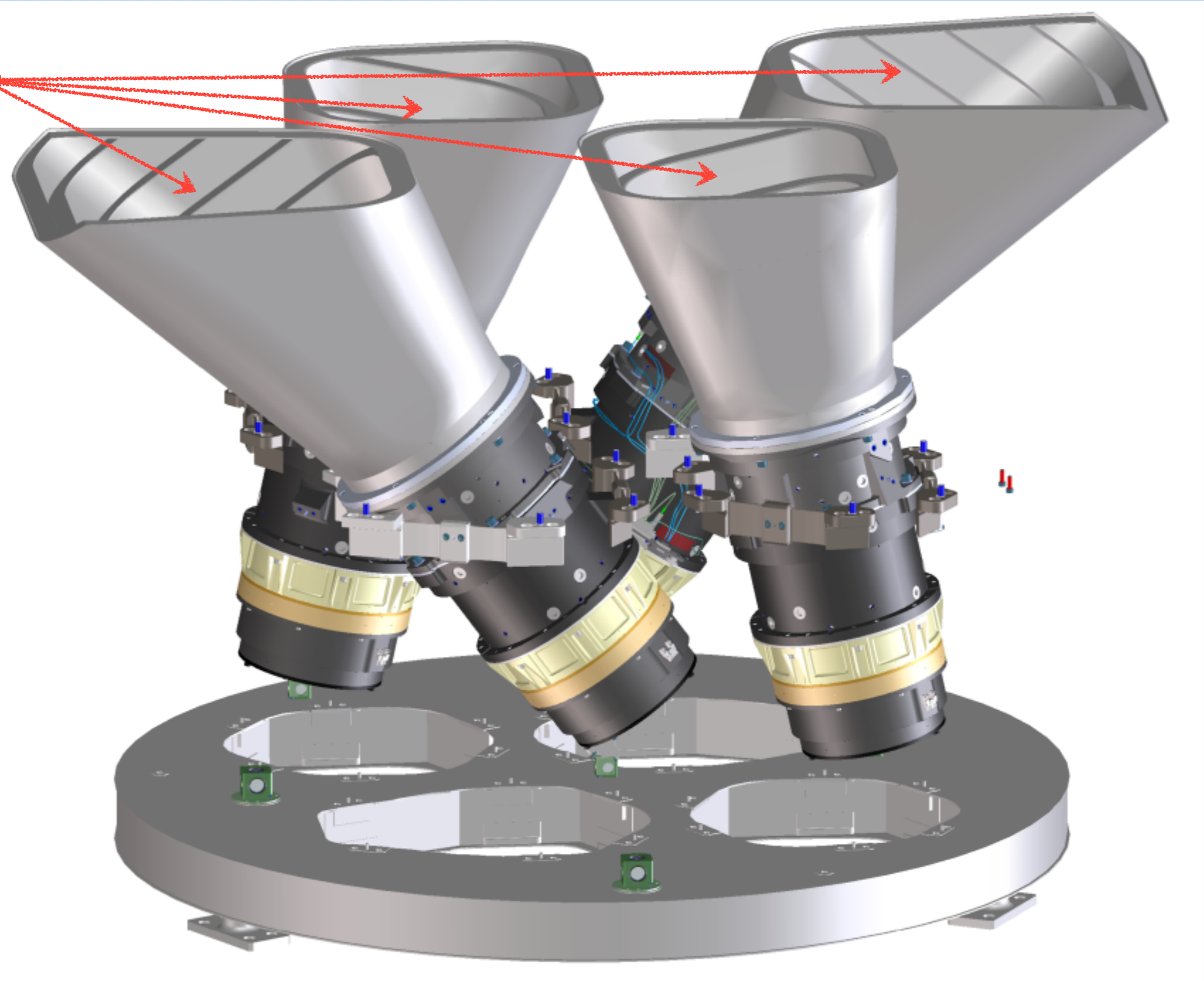
Muirhead+17	For 5 x 10 <sup>4</sup> "Bright" 2min Targets	For 5 x 10 <sup>4</sup> "Easy" 2min Targets
Earths	127	171
Super-Earths	348	489
Sub-Neptunes	599	625
<b>TOTAL</b>	<b>1,074</b>	<b>1,284</b>

**All TESS Flight Instruments  
are delivered and integrated.**

**Environmental Testing in  
Progress!**

**4 Cameras**

**FOV of  
4 camera  
ensemble  
24° x 96°  
or  
2300 deg<sup>2</sup>**



## **In Case You Have Ever Wondered:**

How Many Engineers Are Needed to Mount the Tess Cameras?

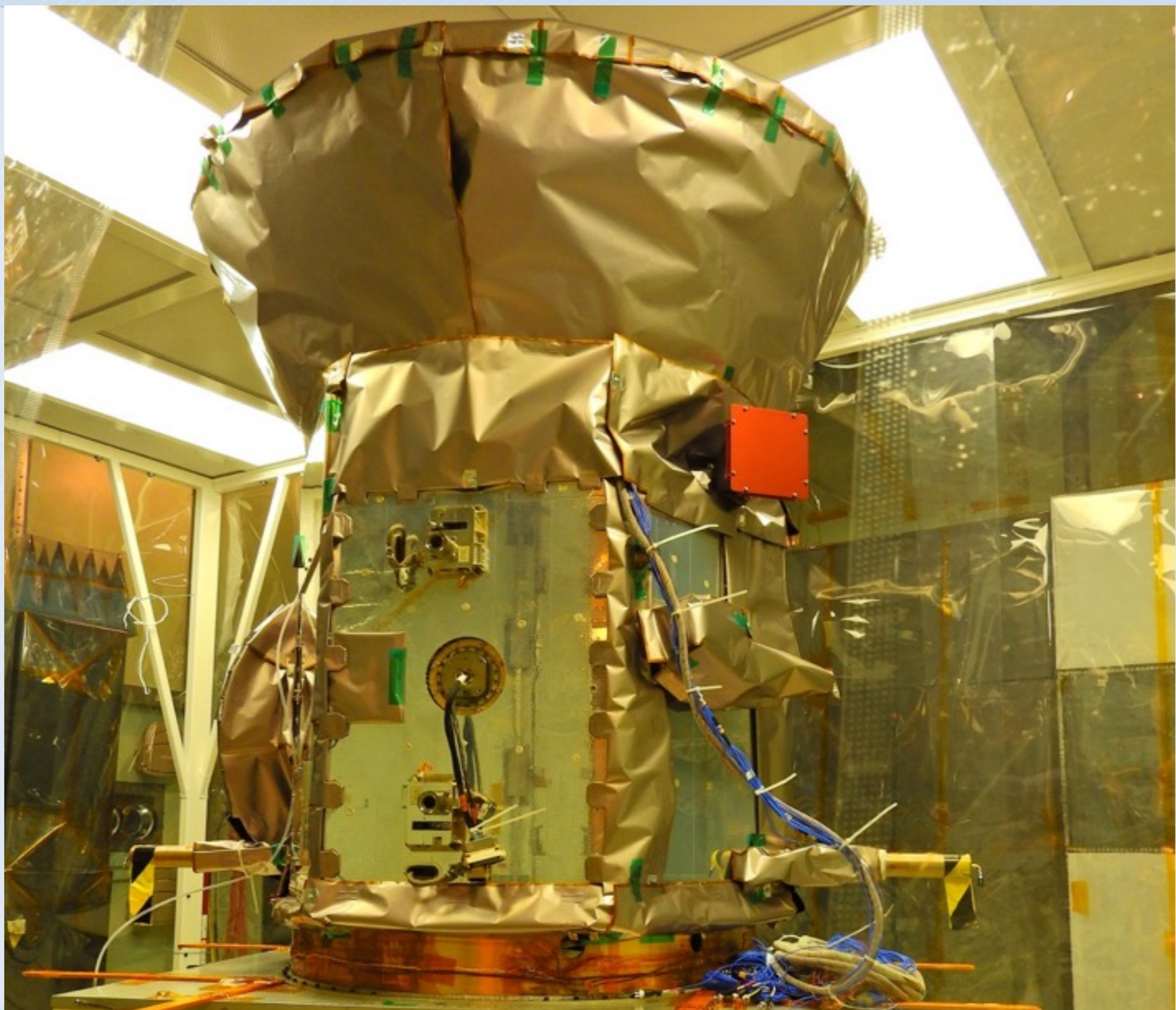
**Answer:**

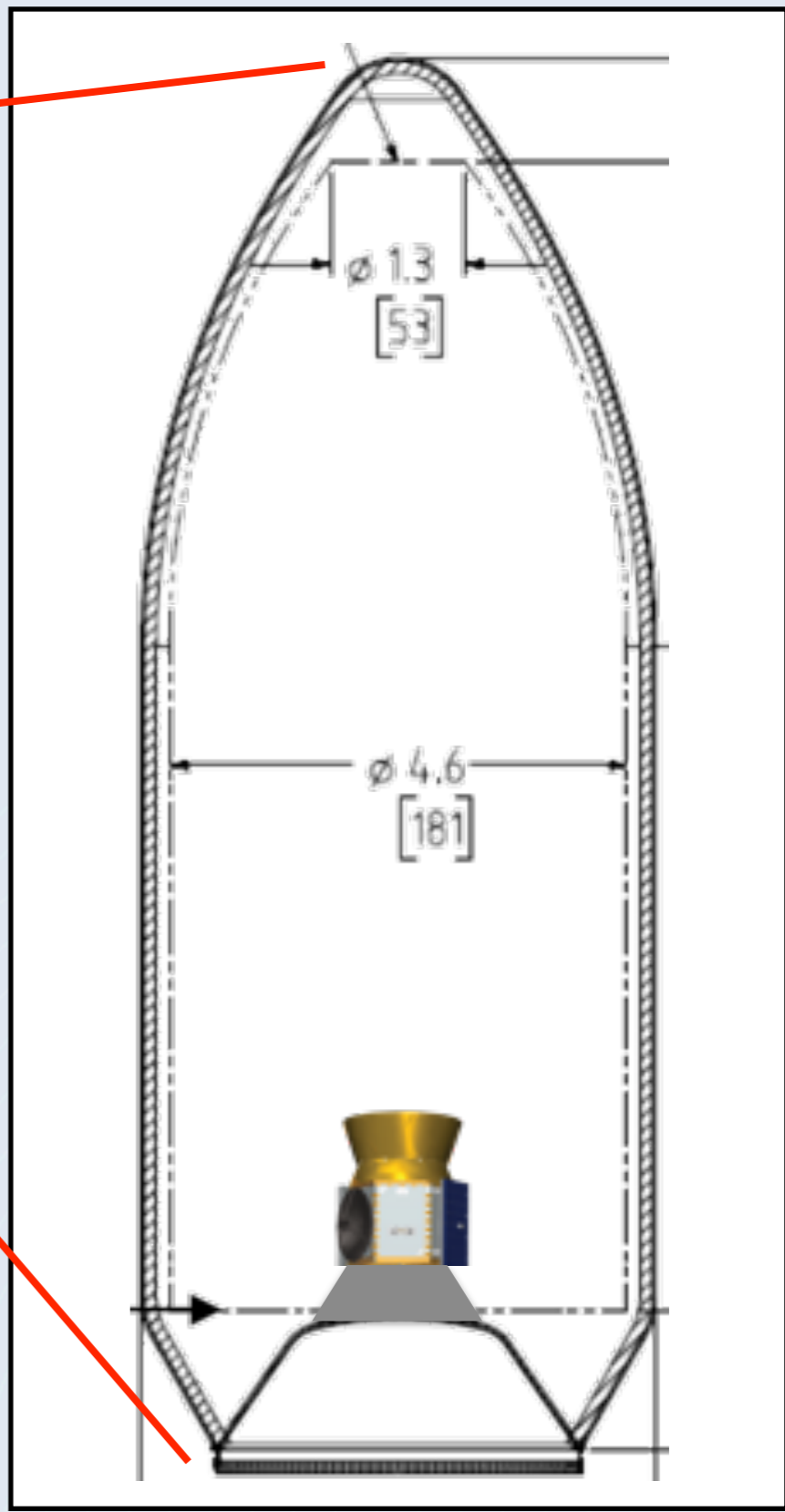
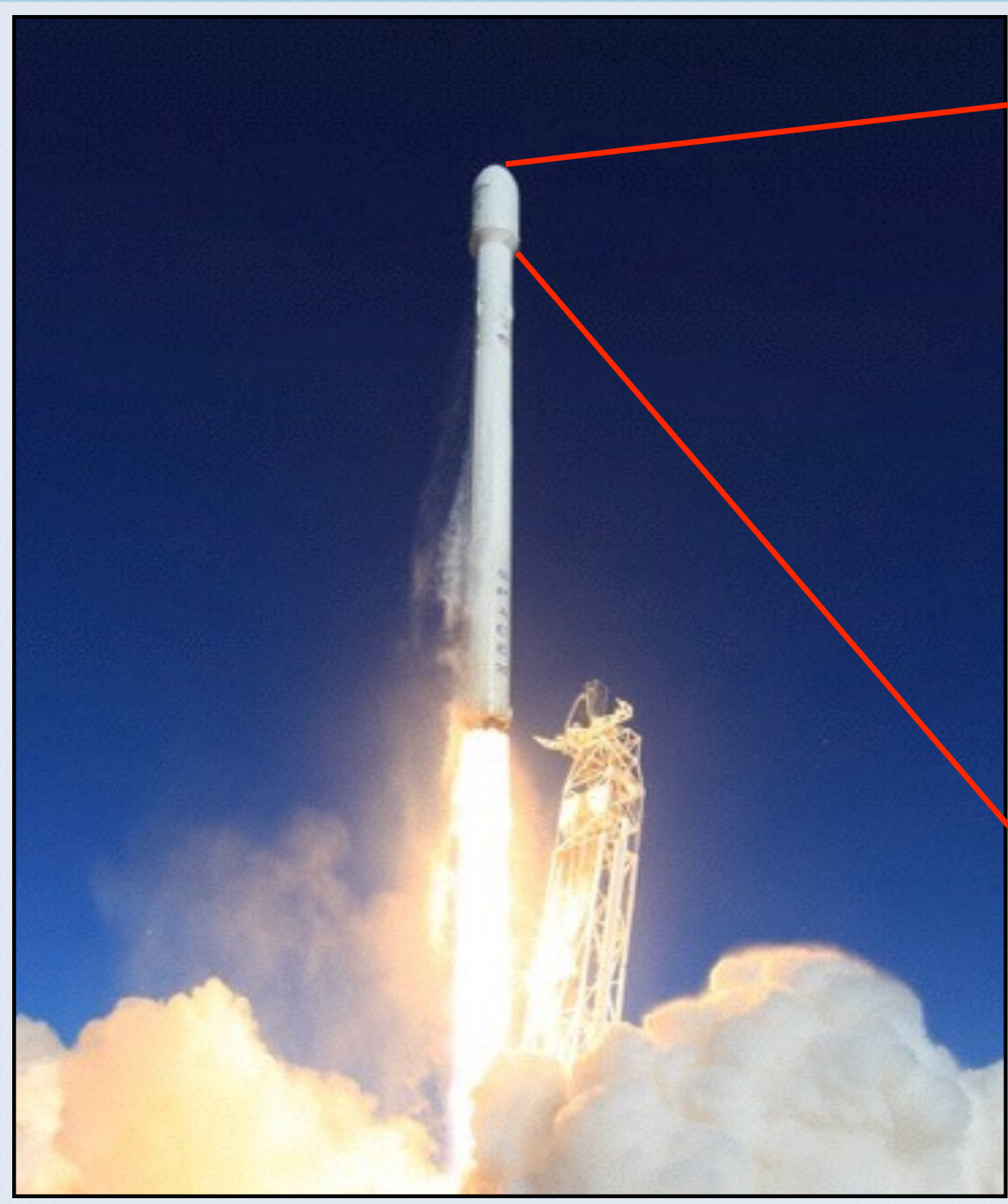
7 (plus 1 photographer)



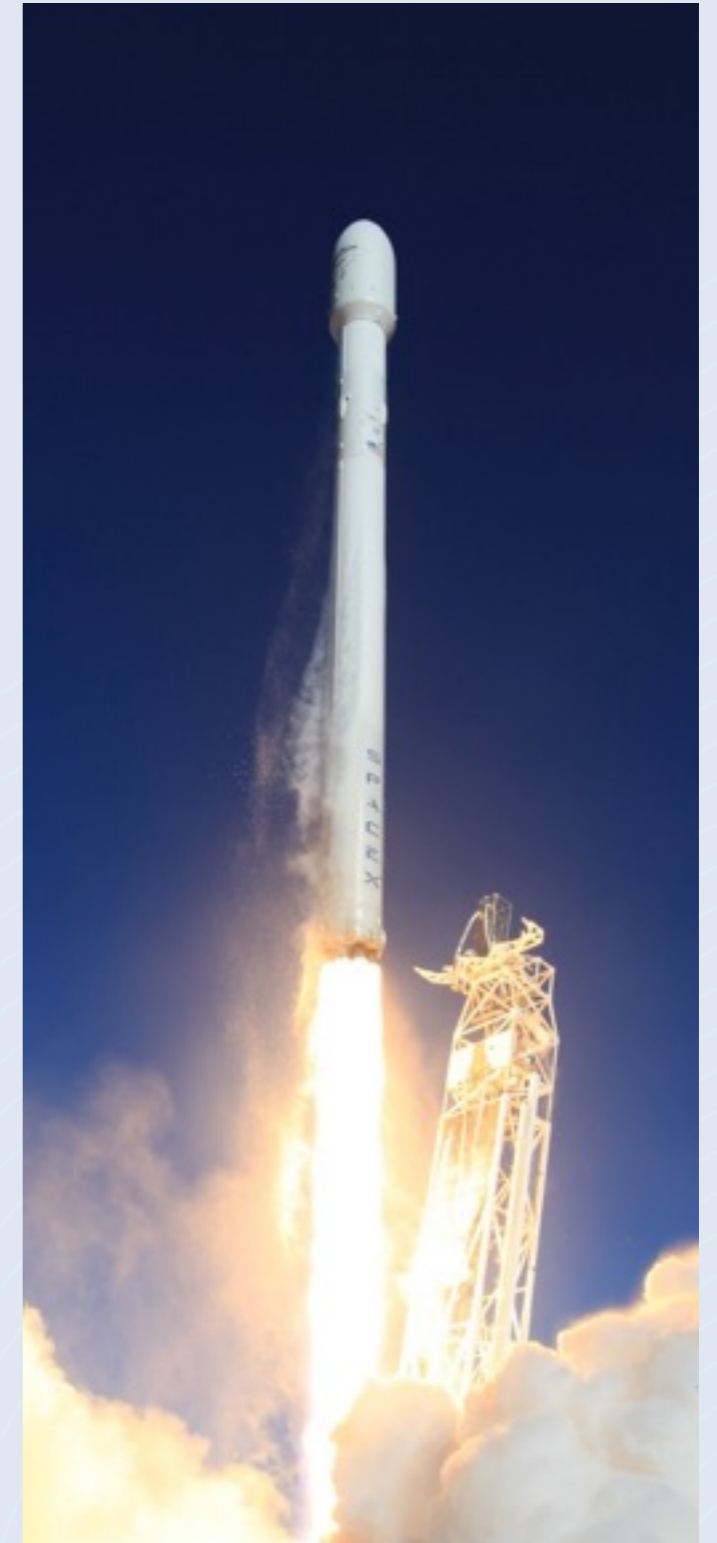




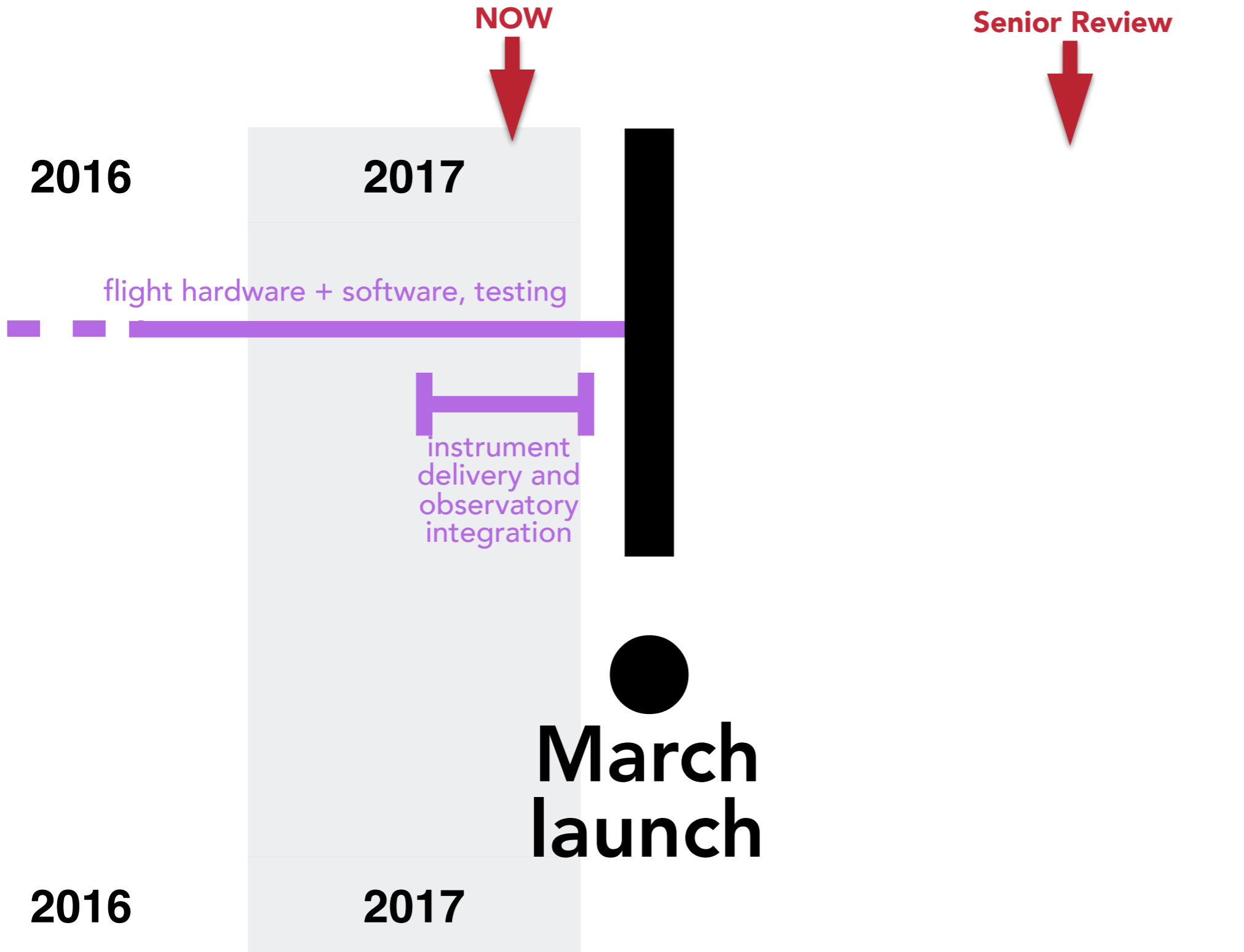




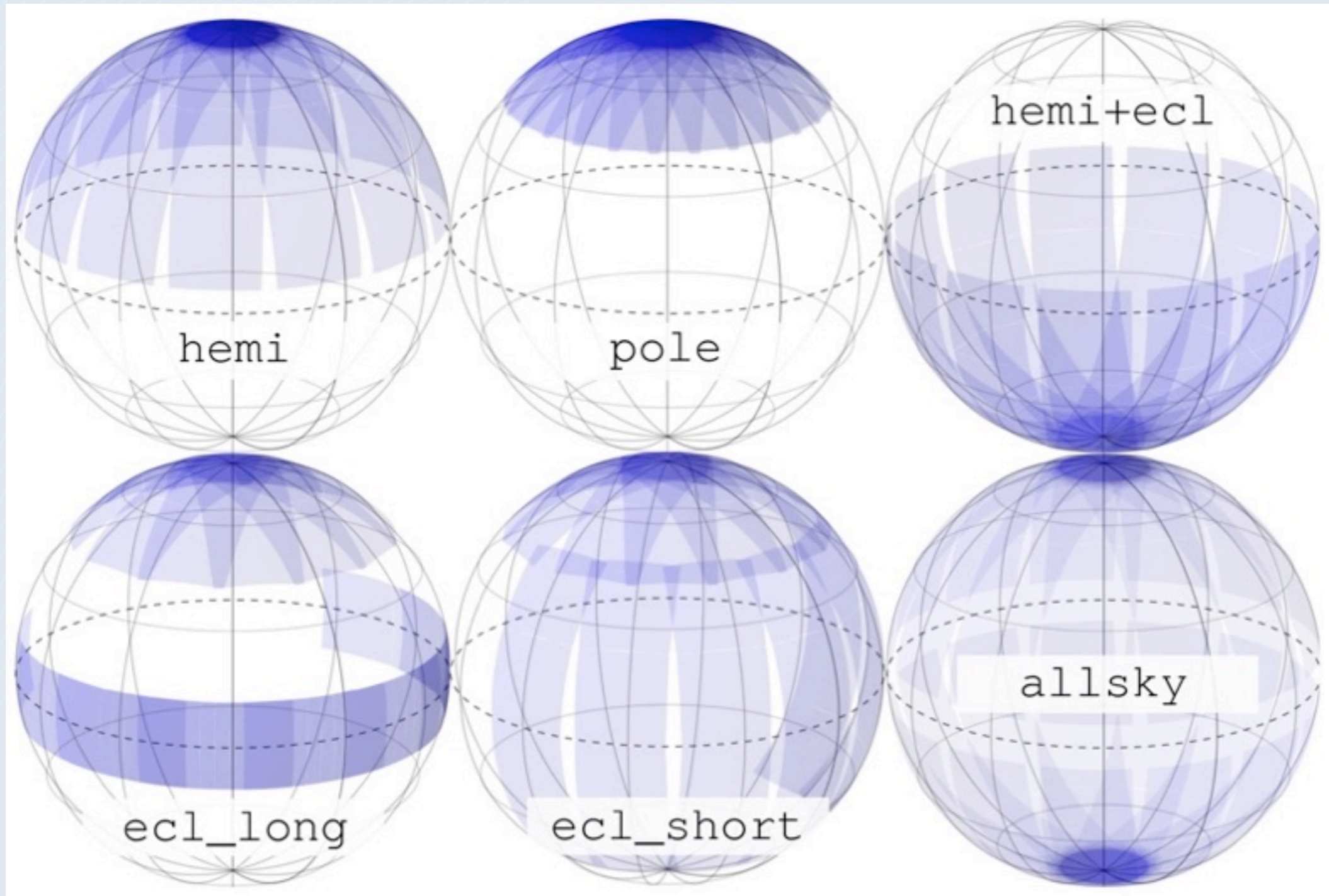
- ◆ SpaceX Falcon 9 successfully returned to flight in January 2017:
  - ~~Seven~~ ~~13~~ **15** successful launches already in 2017
- ◆ ~~~20~~ ~~15~~ **11** additional Falcon 9 flights anticipated before TESS launches
- ◆ TESS's Falcon 9 will possibly be a Block 5 vehicle
  - *Recoverable fairing?*
  - *NASA re-certification needed (launch loads)*
  - *TESS launch scheduled in March 2018*



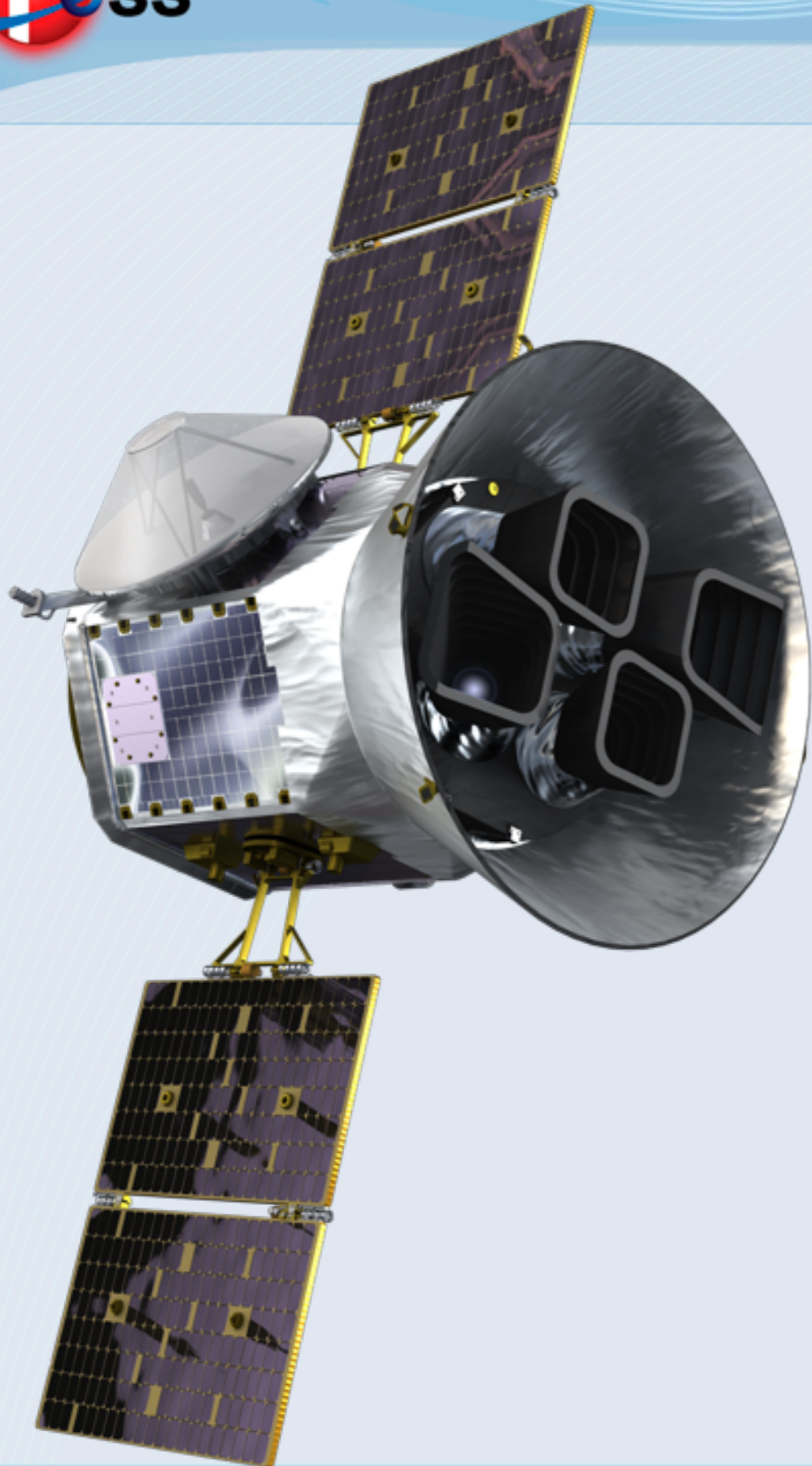
# TESS timeline:



- ◆ TESS Orbit is Stable for >20 Years....



Bouma et al. 2017 (in prep)



- TESS will find nearby bright small transiting planets for followup by JWST and upcoming giant 20-, 30-, and 40-meter ground-based telescopes.
- TESS could operate for more than two decades.
- TESS planets will endure as the best small planet targets for radial velocity mass measurements and atmospheric characterization.
- TESS will fly a highly adaptable data handling unit, enabling support of a broad range of time domain astrophysics, well beyond exoplanet studies.





Backup Slides

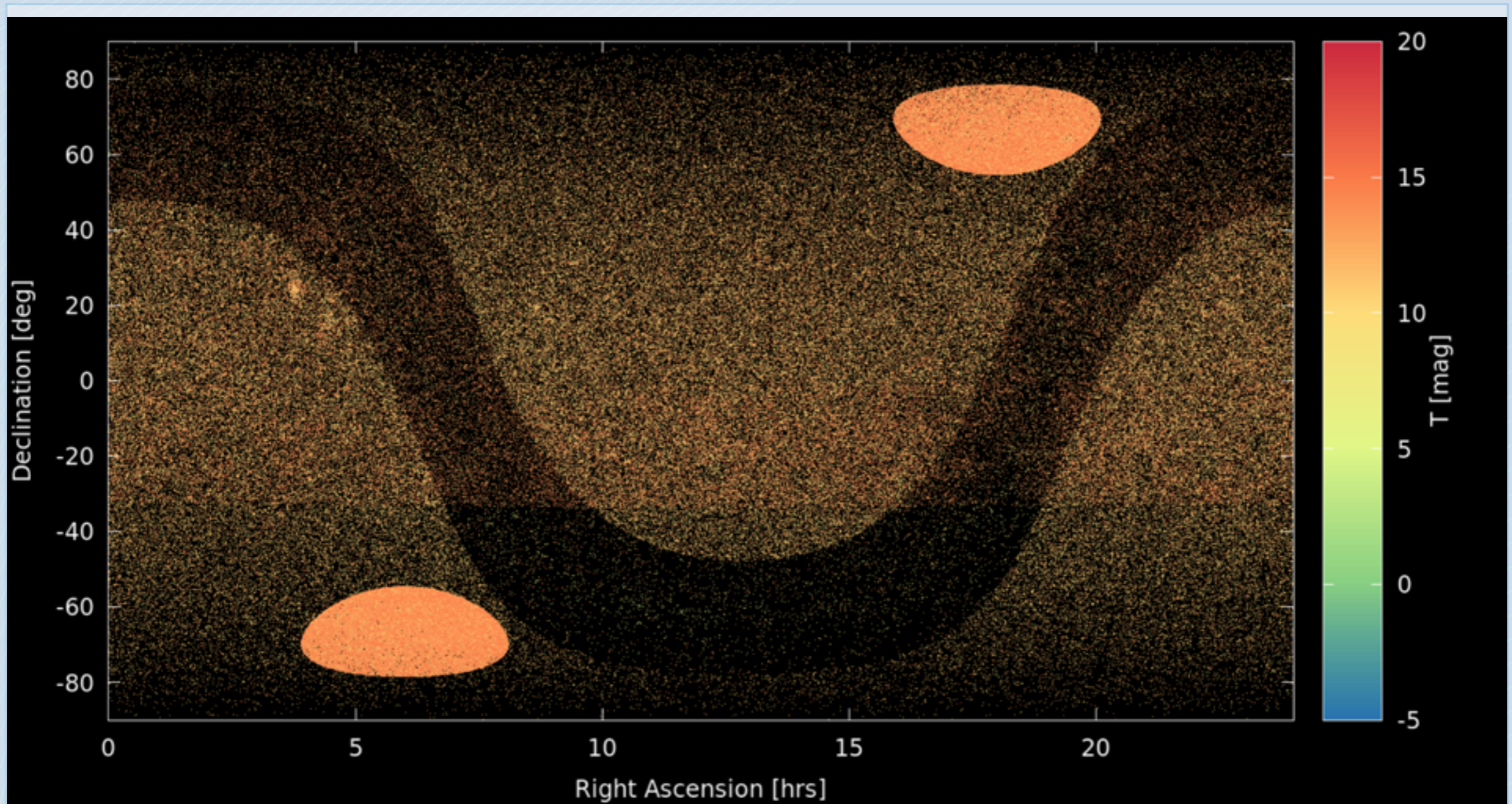


Fig. 19.— The distribution of top CTL targets in right ascension and declination, colored by T magnitude. Clear patterns arise due to de-boosting in the Galactic Plane ( $|b| < 15^\circ$ ), the special boosting within the ecliptic poles ( $|\beta| > 78^\circ$ ), and the coverage of the proper motion catalogs (the lines near 35)

Kepler

TESS

