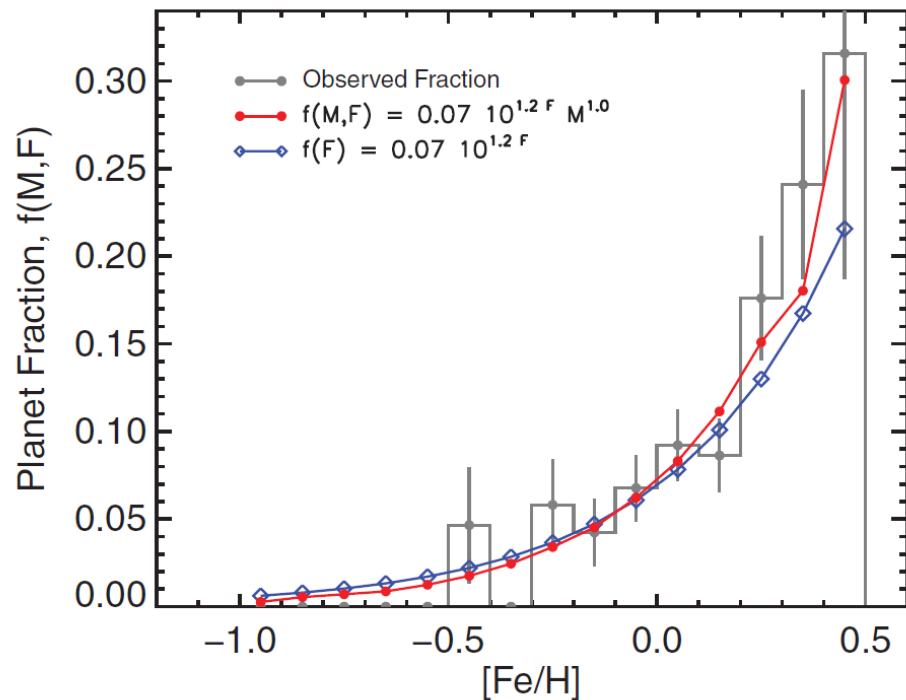


Alpha Elements' Effects on Planet Formation & the Hunt for Extragalactic Planets

Matthew Penny
Ohio State University

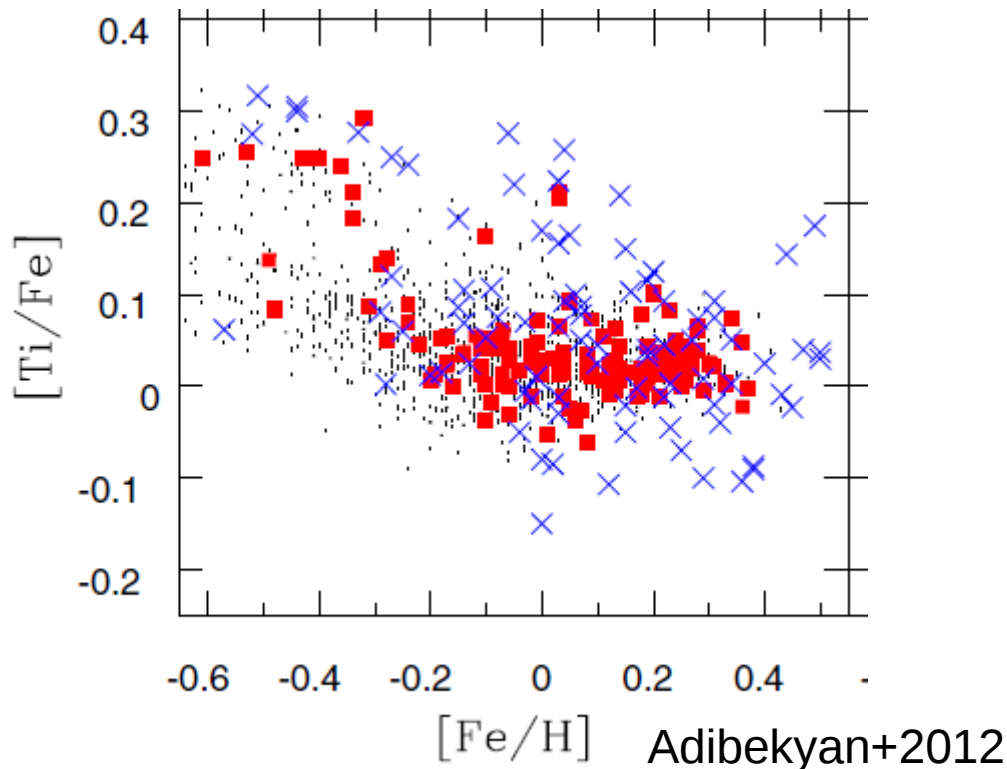
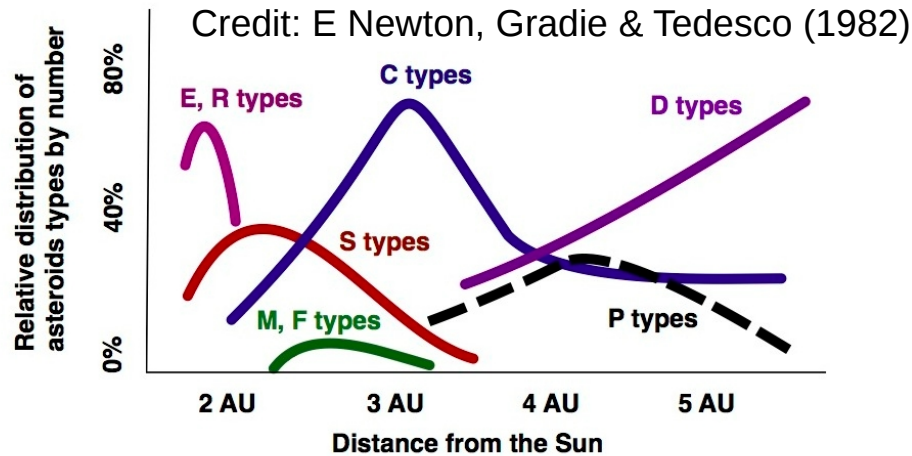
Planet Frequency & Metallicity



Johnson+2010

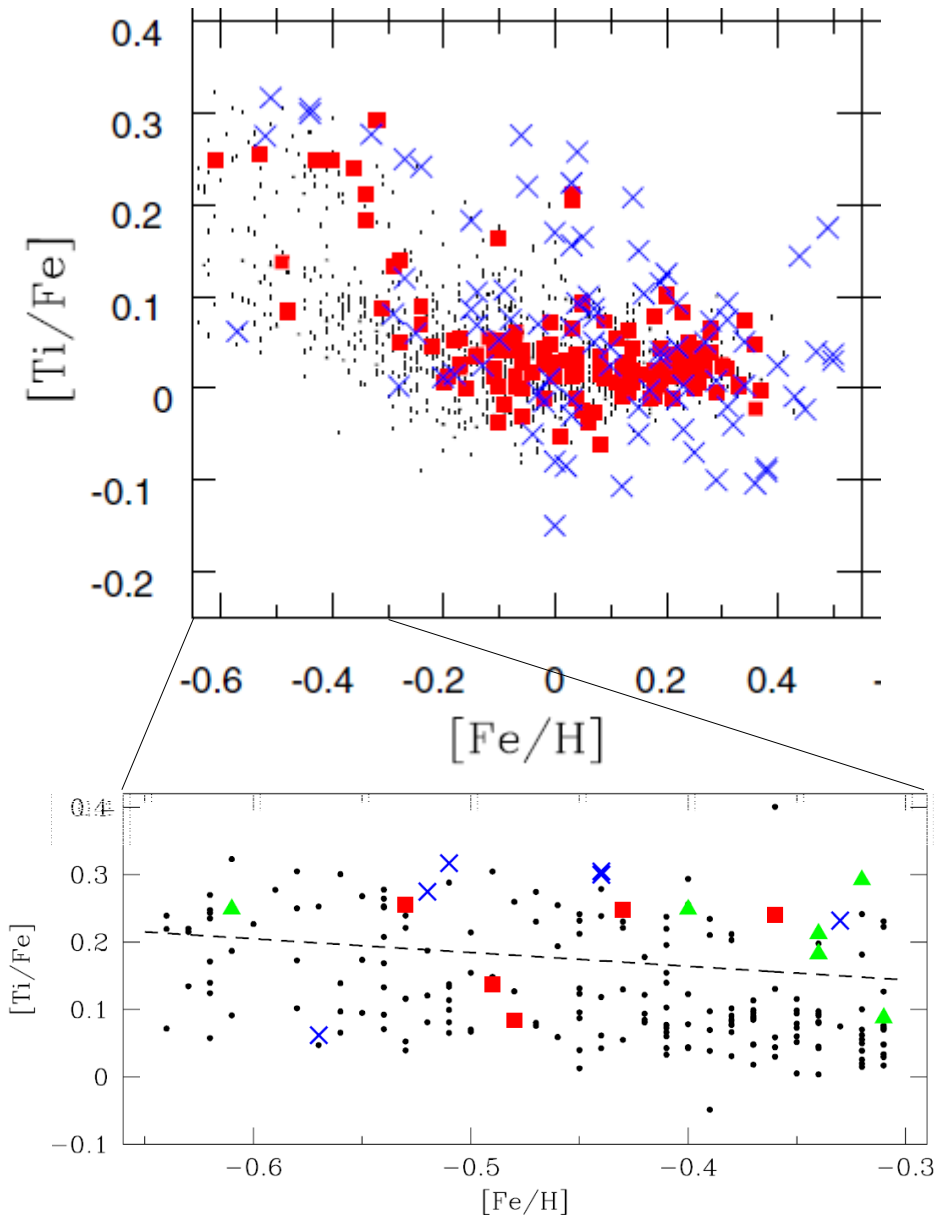
- Well known, strong correlation for RV giant planets (Gonzalez 1998, Santos 2004, Fischer & Valenti 2005, Johnson+2010, ...)
- Similar results for hot Jupiters (e.g., Guo+2016)
- Mixed results for smaller planets

Iron or Rock?



- Differentiation of asteroid composition as function of semimajor axis
- Assumed that HJs form past snowline (ices, e.g. H_2O , CH)
- Might alpha be a better correlate?

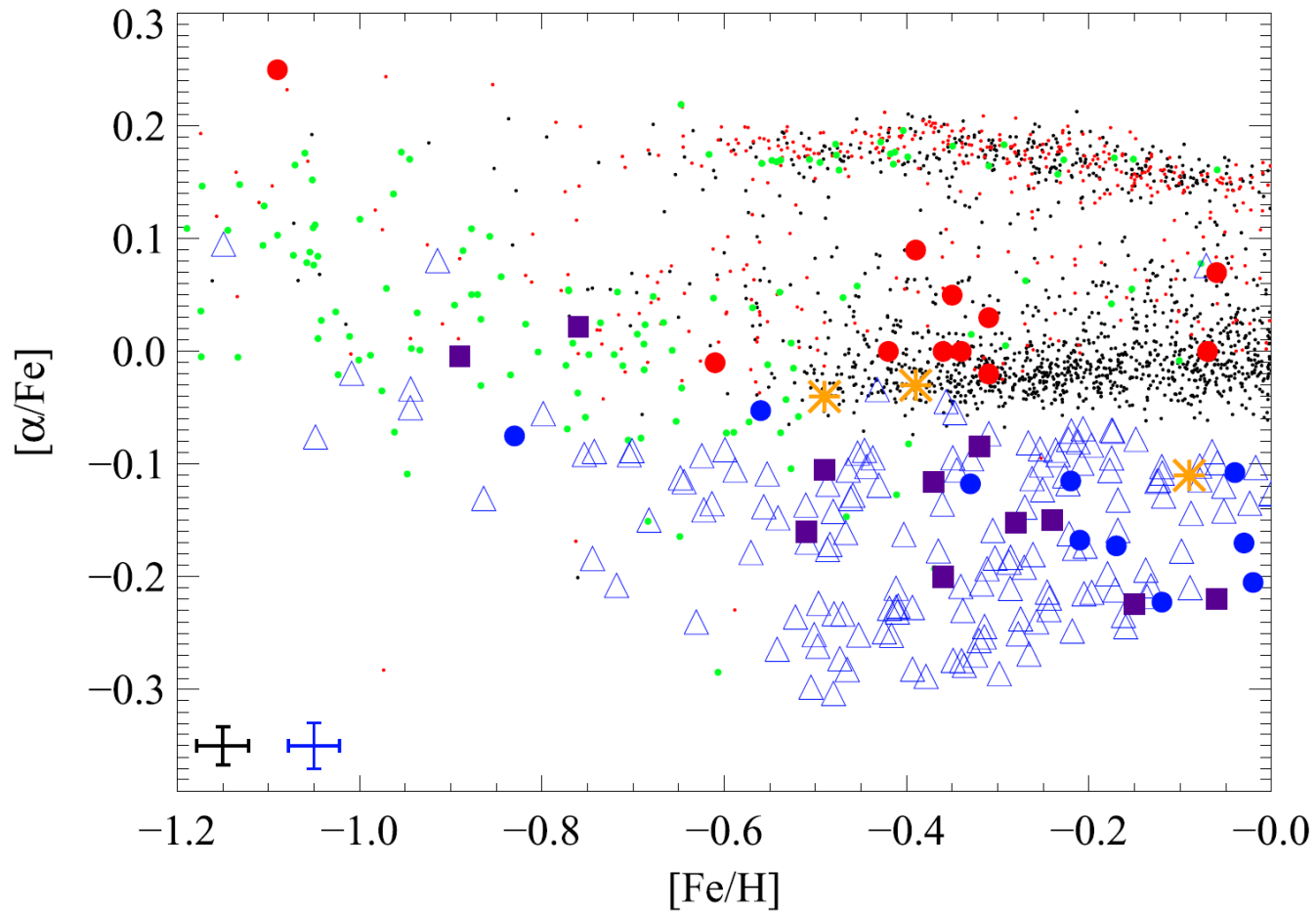
Alpha Elements are Important



- Low $[Fe/H]$ planet hosts tend to be α -enhanced, thick disk stars (e.g Adibekyan+2012)
- Does α -enhancement help higher metallicity stars?
- Too few SN & Kepler stars with high-Fe & high- α
- What about low- α ?
- Are Hot Jupiters different?

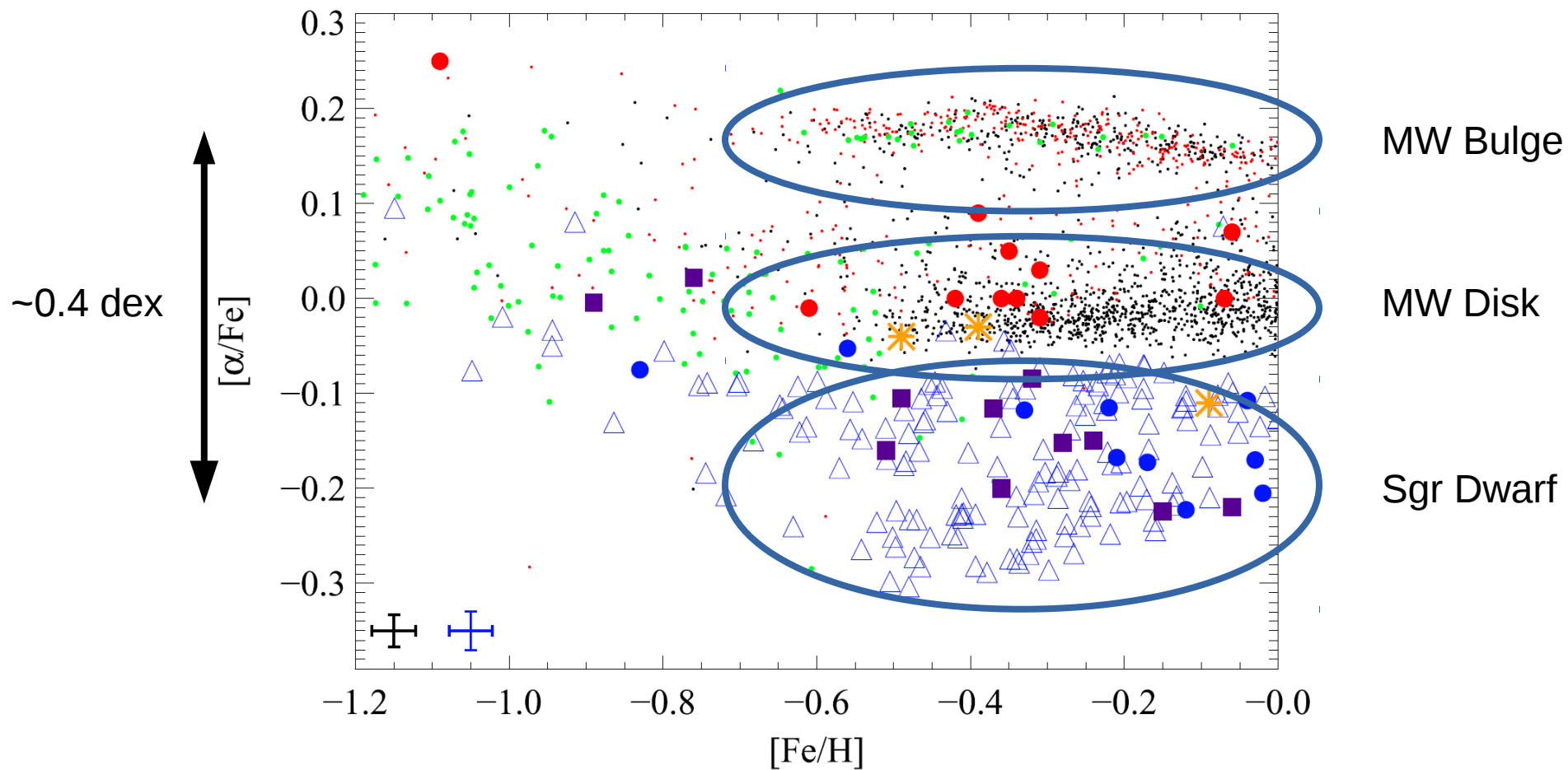
Adibekyan+2012

Stuck in the Neighborhood



Hasselquist+2017 (APOGEE)

Stuck in the Neighborhood



Hasselquist+2017 (APOGEE)

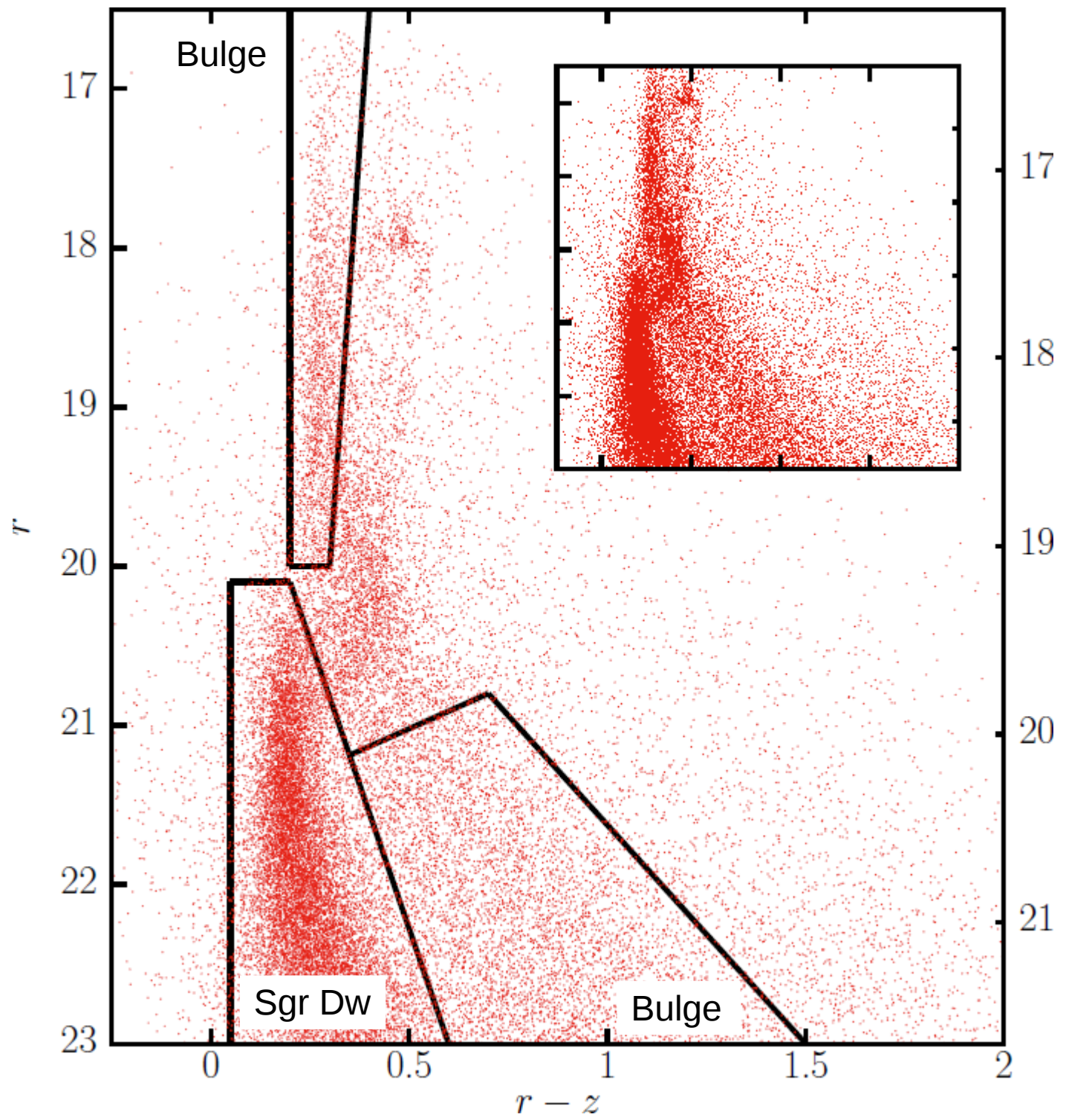
Sgr Dwarf & Bulge are Coincident

CENTER OF MILKY WAY
DISKS OF MILKY WAY

- Simultaneously survey $\Delta[\alpha/\text{Fe}]$ of ~ 0.4 dex

AREAS OF INITIAL OBSERVATIONS

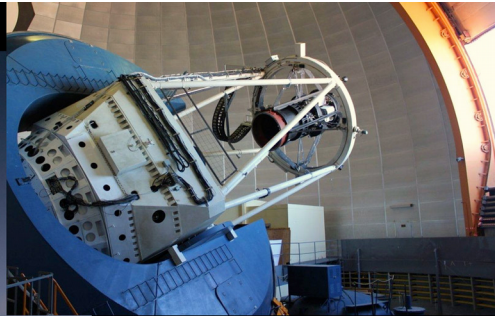
SAGITTARIUS DWARF GALAXY



A DECam Transit Survey

- 3 deg² field, 4m scope
- Efficient (20s readout, filter changes)
- Wide VR filter, red sensitive CCDs
- ~1% phot. in 3 min for Sgr main sequence
- Faint, so survey must rule out false positives alone
- 5 nights x 3 years
- $P < 3$ days, simplifies interpretation (e.g., can assume $e=0$)

The NOAO Blanco
4-meter telescope



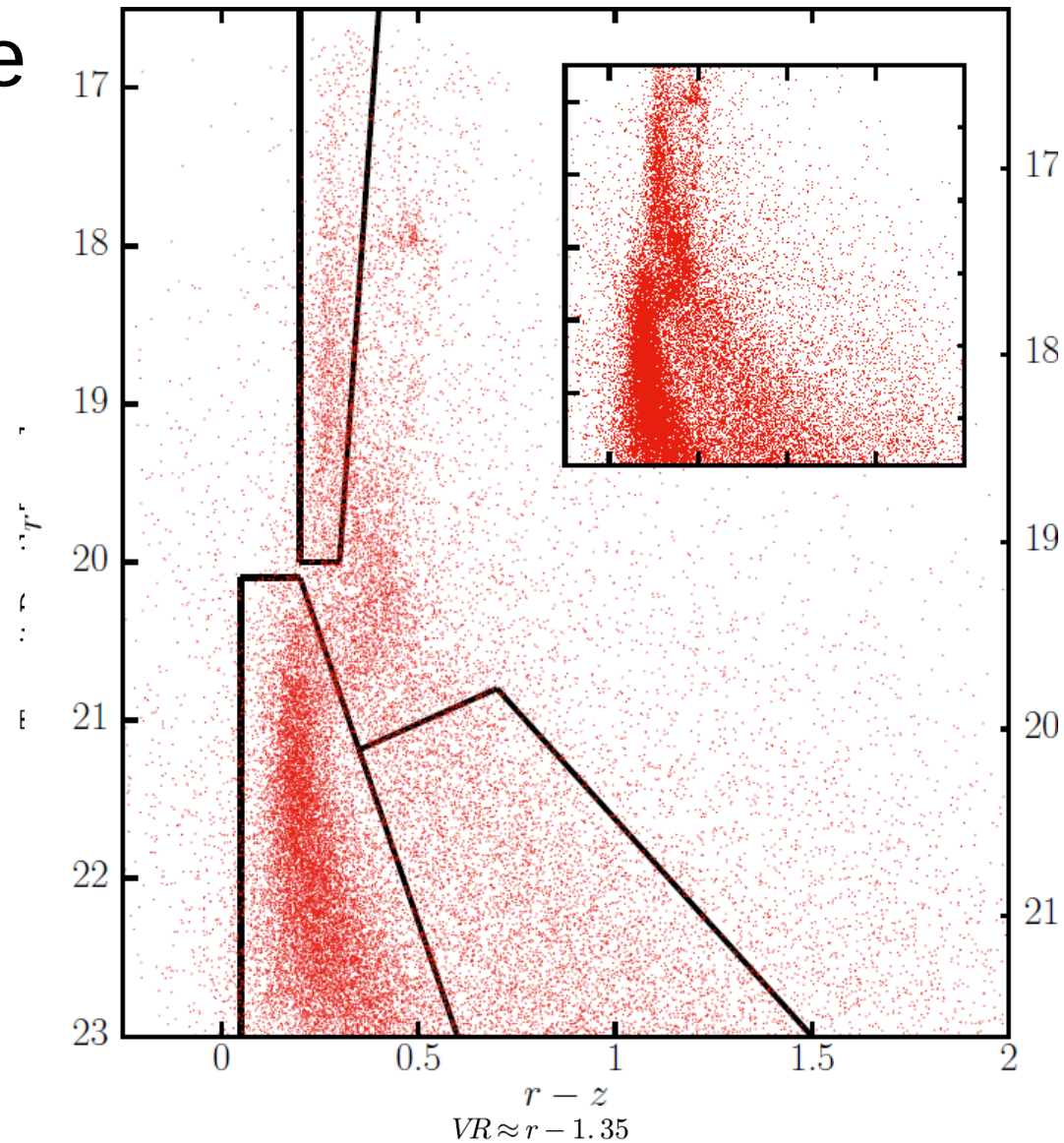
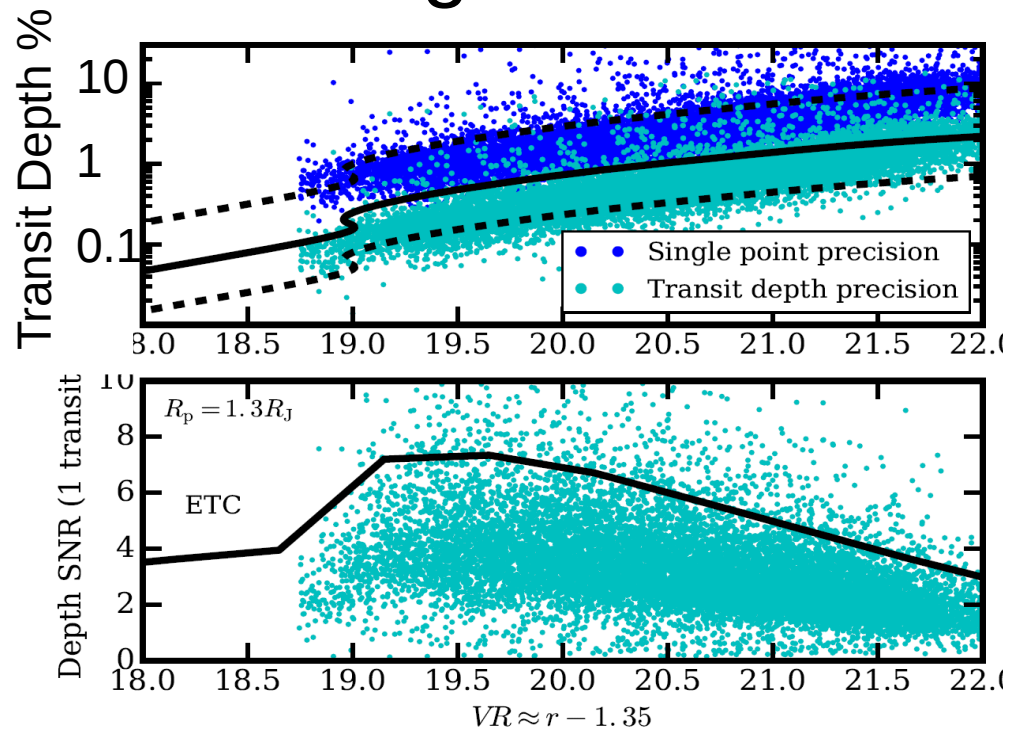
The 520 megapixel Dark
Energy Camera (DECam)



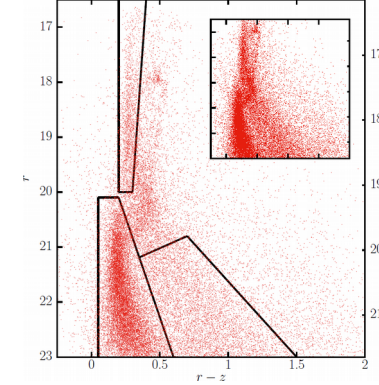
Transit Detectability

- Test run of survey-like observations & quick photometry

Sgr Dwarf



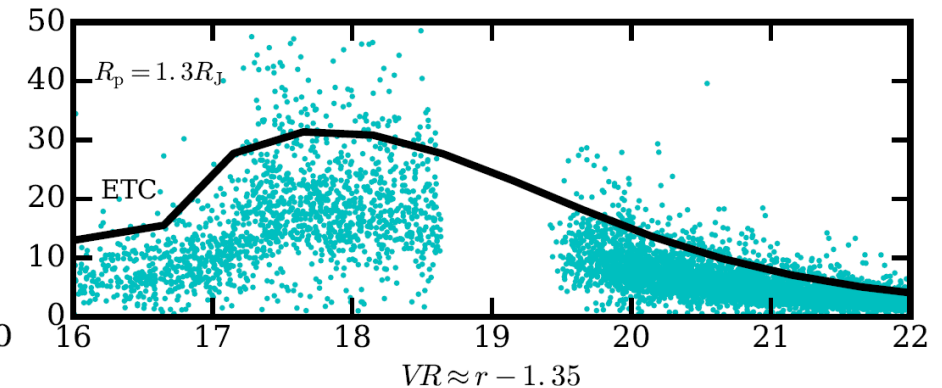
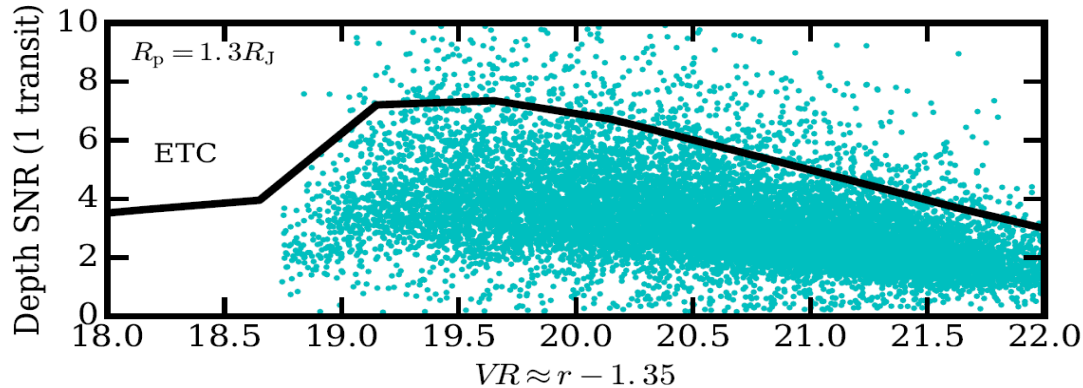
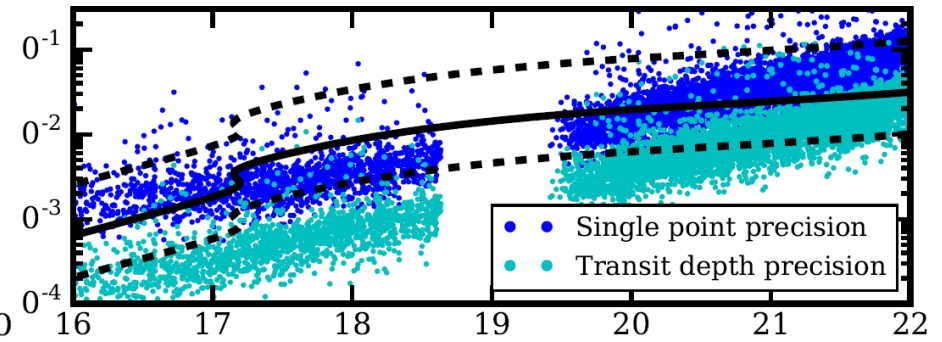
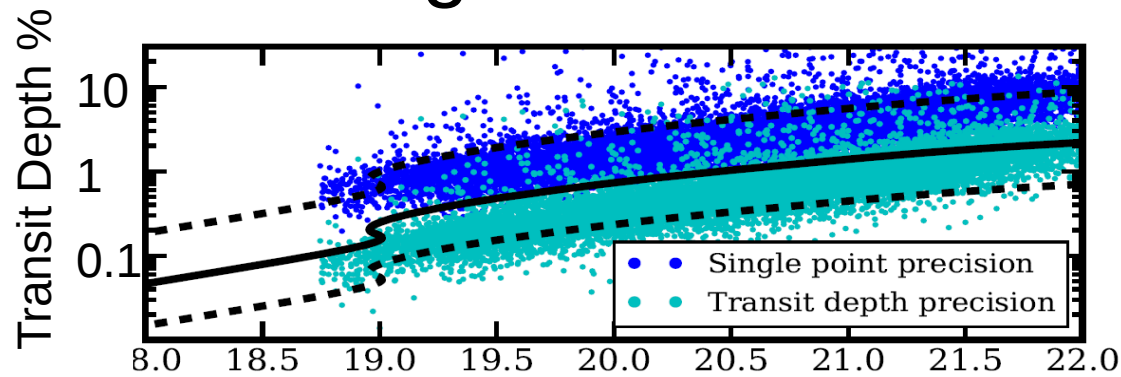
Transit Detectability



- Test run of survey-like observations & quick photometry
- Carter+2008 fisher matrix estimates (single transit)

Sgr Dwarf

Bulge



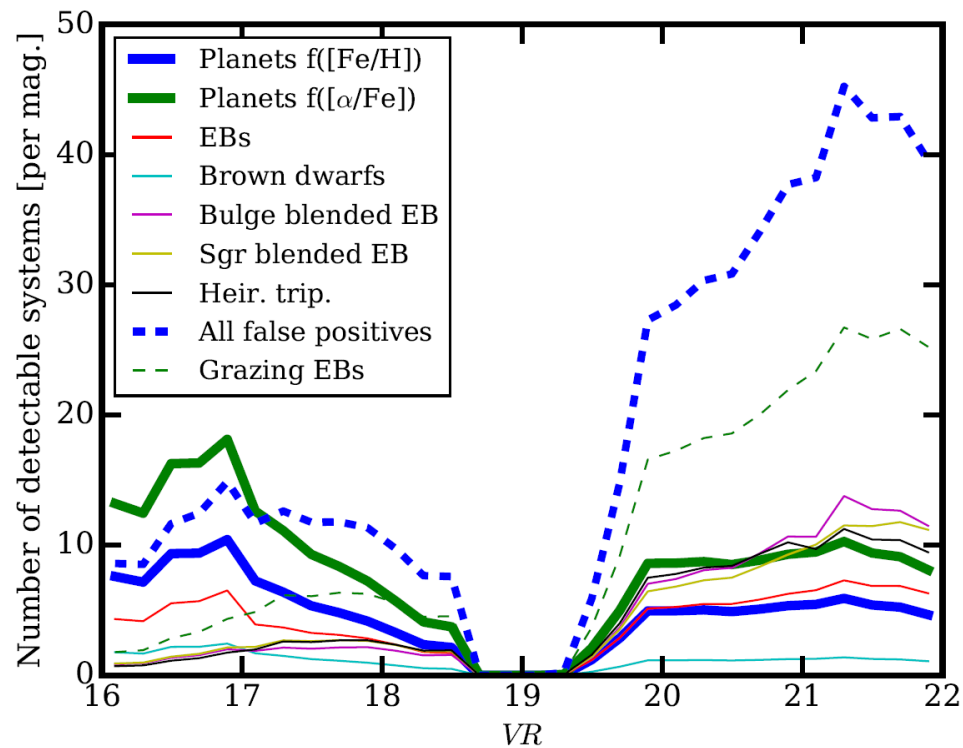
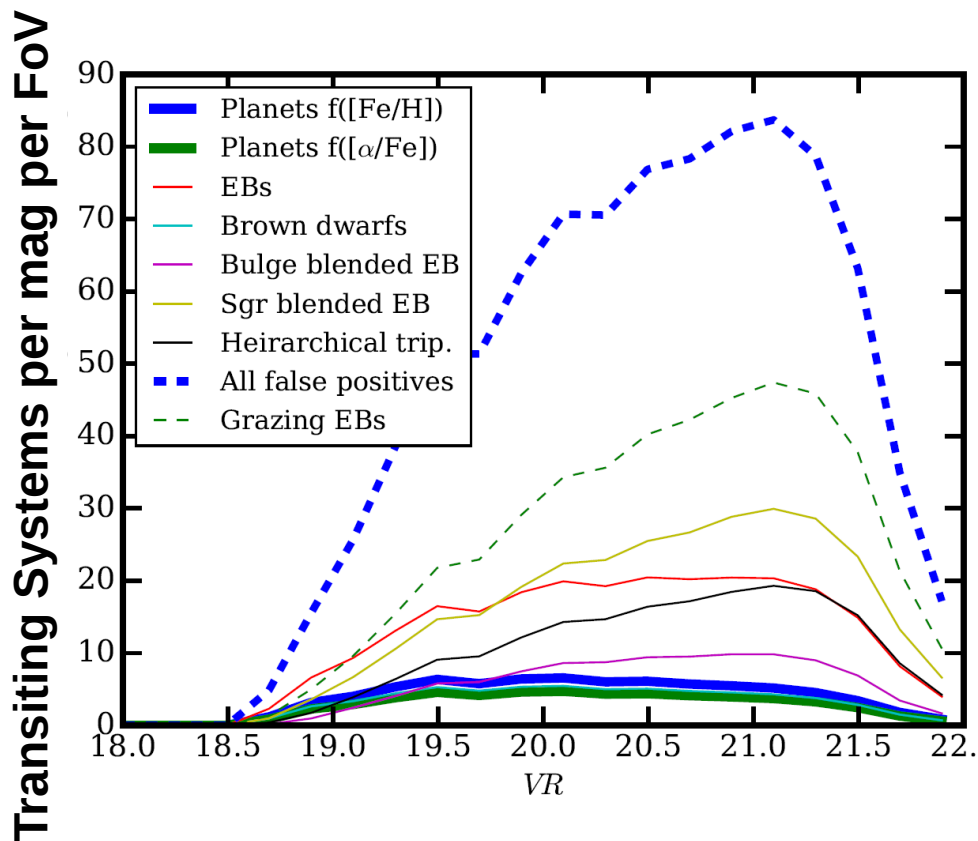
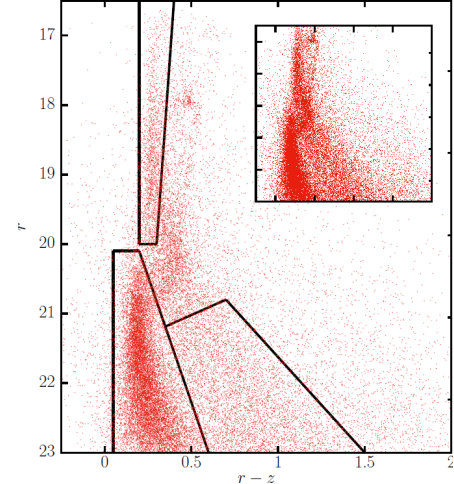
But what about False Positives?

- FGK-M eclipsing binaries
- Hierarchical Triples
- Blended Ebs (foreground/background or coincident)
- Grazing EBs
- Assume bulge and Sgr are thin screens with same age & metallicity
- Use isochrones + IMF + binary/planet frequencies to estimate the ratio of transiting planet impostors $(R_B/R_A)^2 \sim 0.01$

False Positives

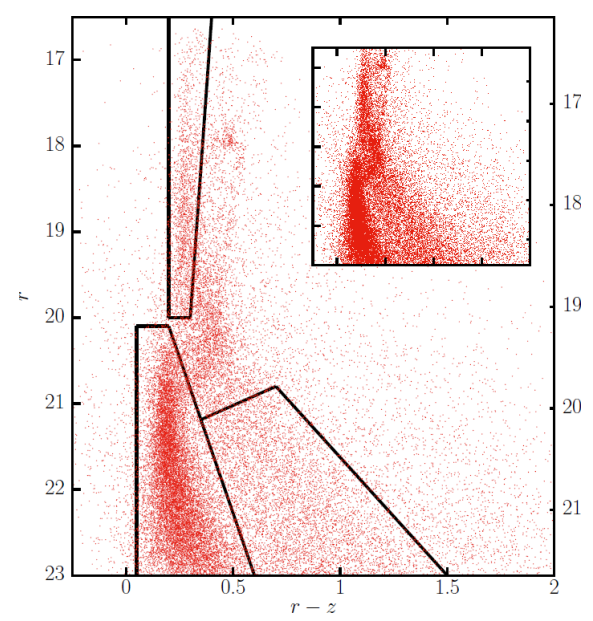
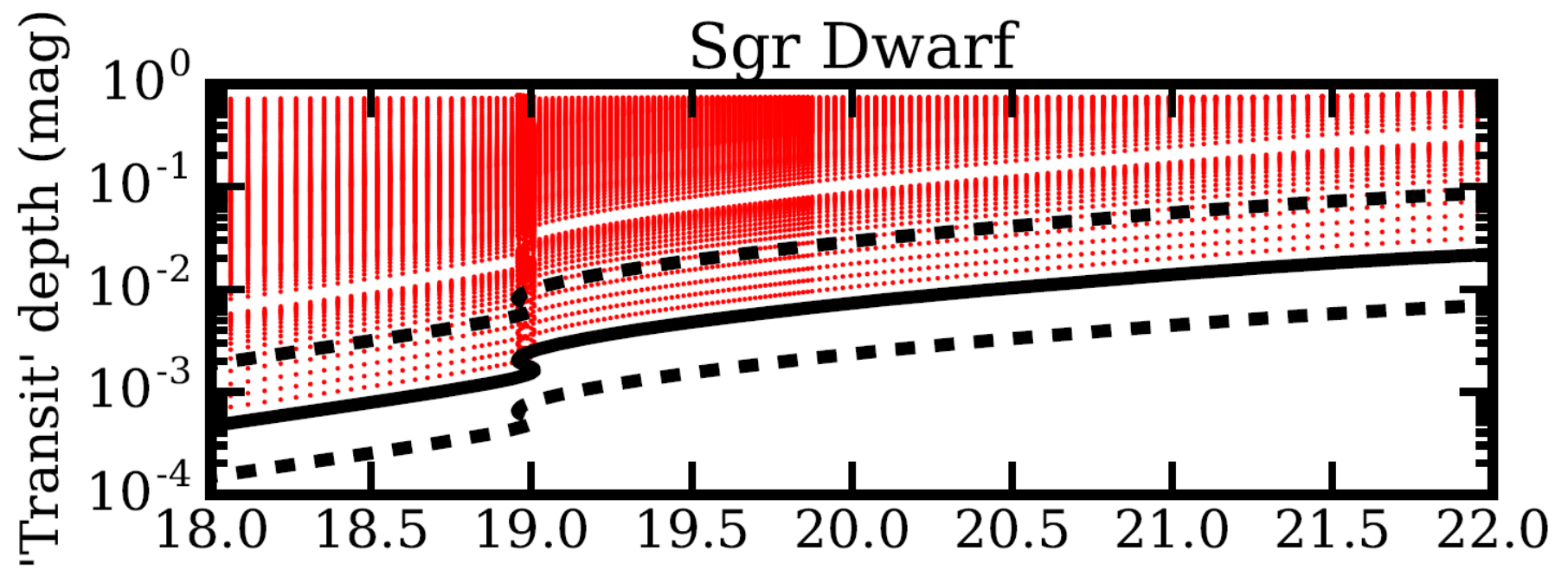
Sgr Dwarf

Bulge

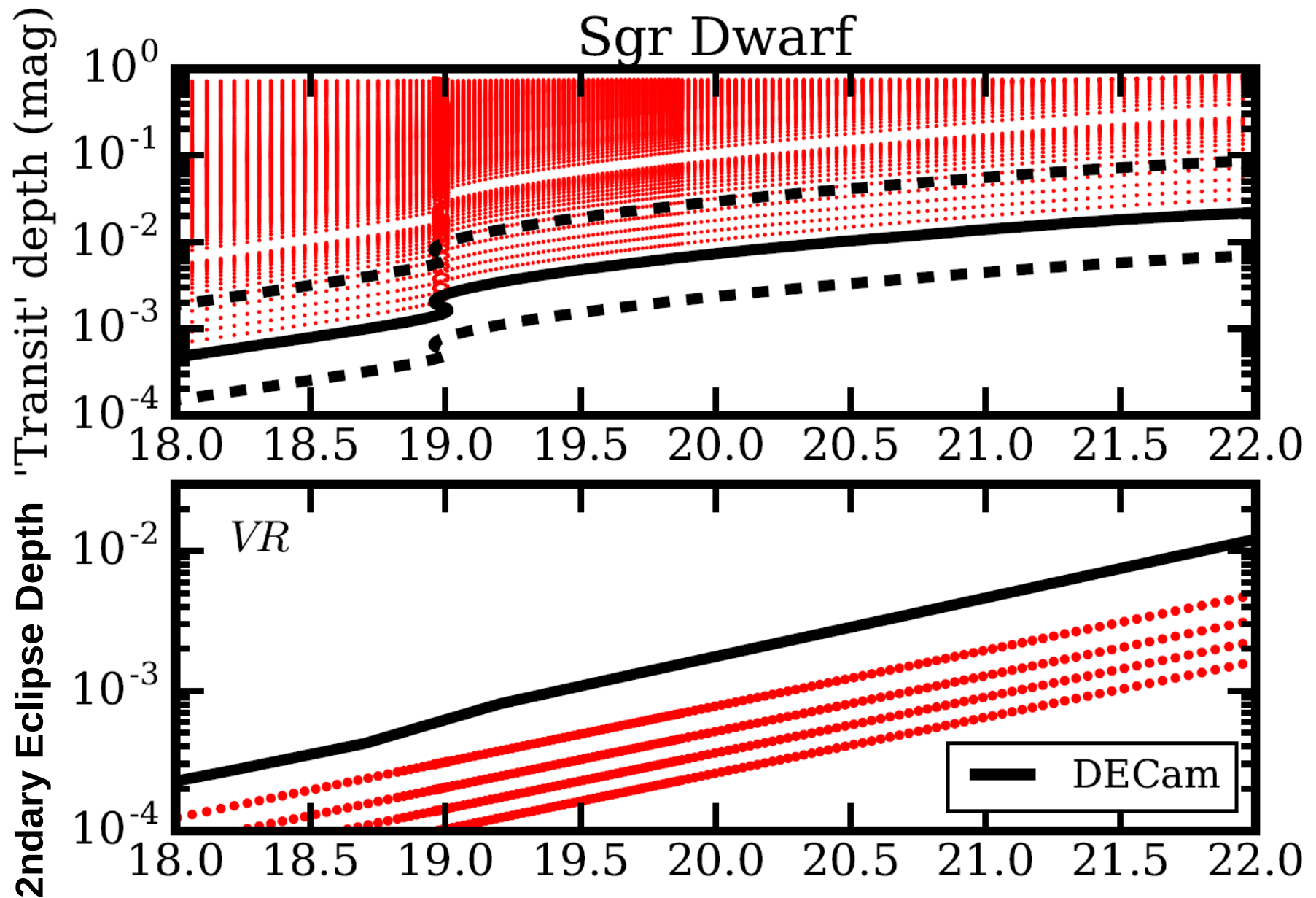


Based on stars detected in images (+ assumed incompleteness for blends)
Planet frequency assumes mean [Fe/H]

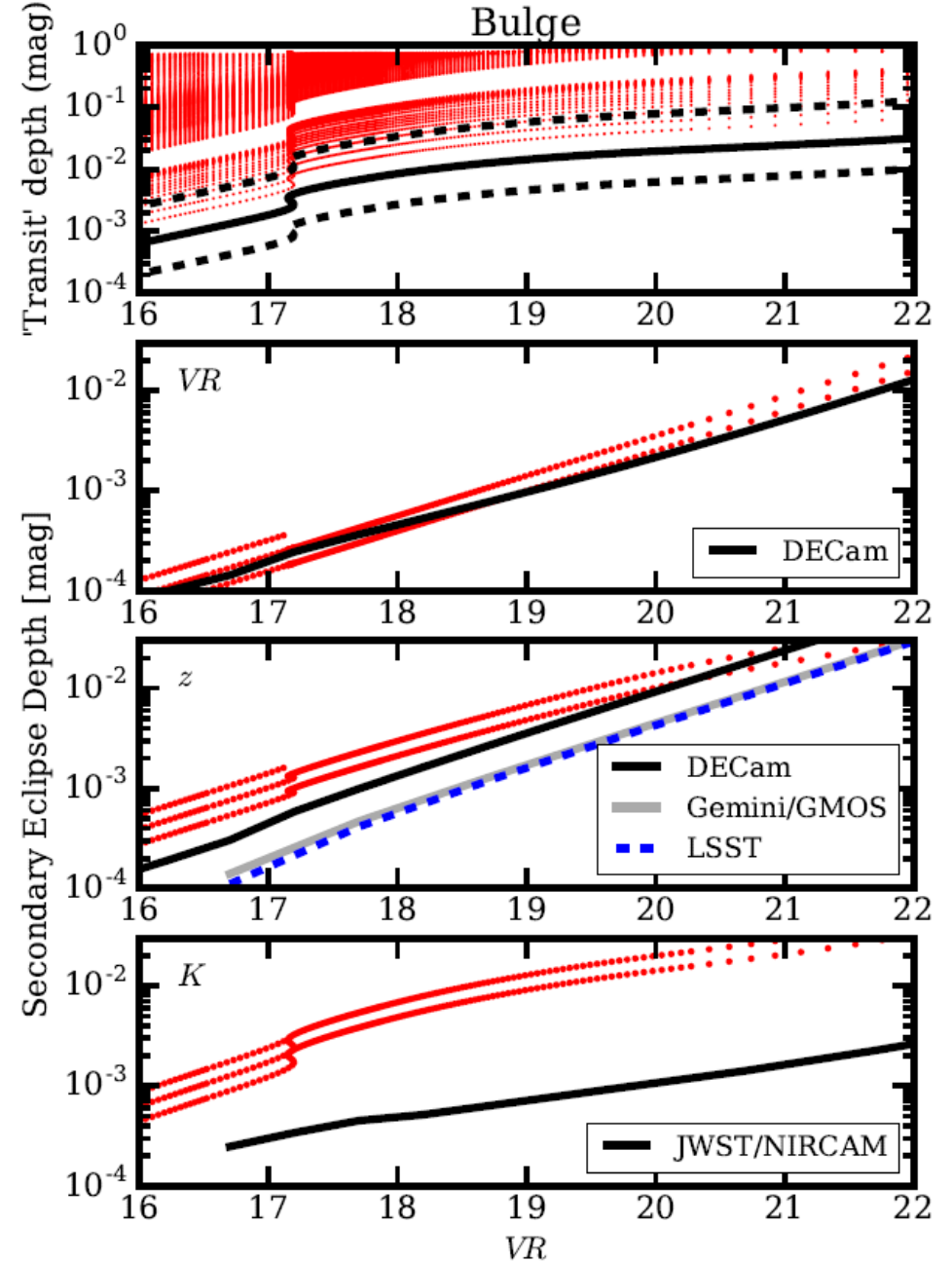
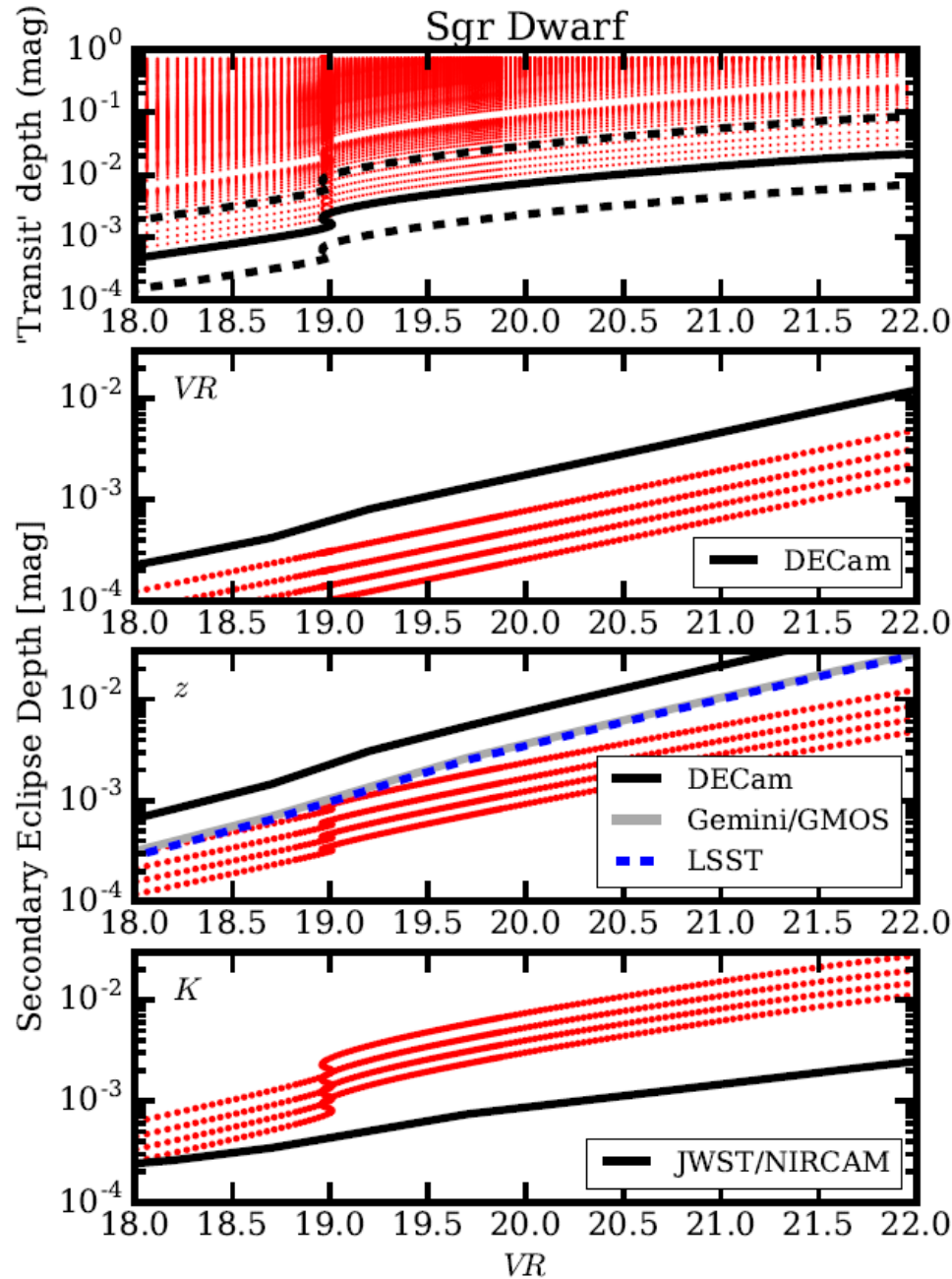
Transiting Planet Impostors: EBs



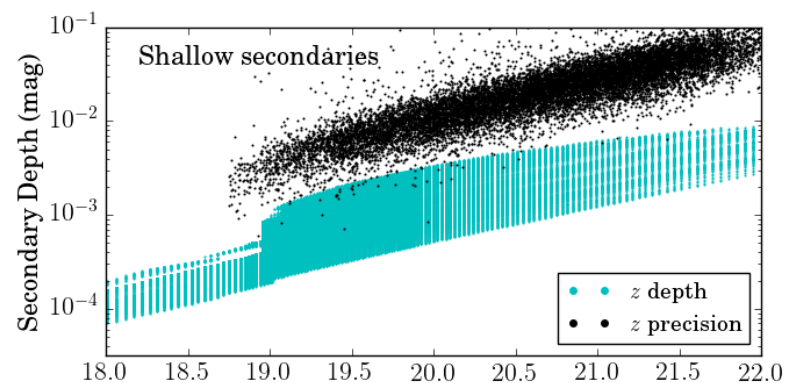
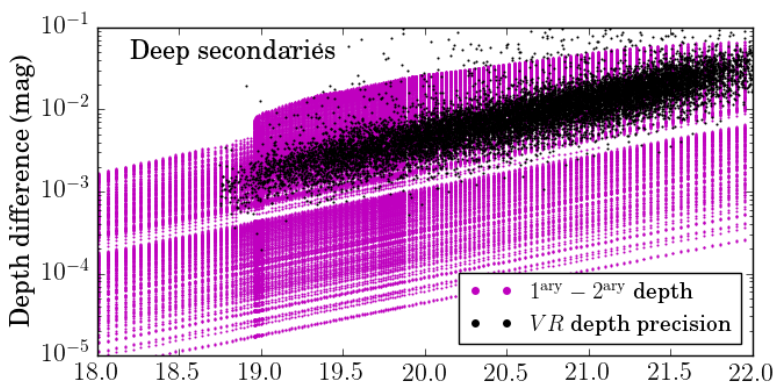
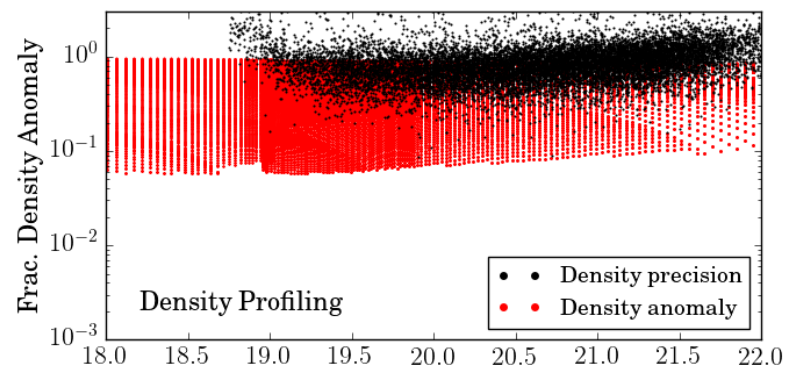
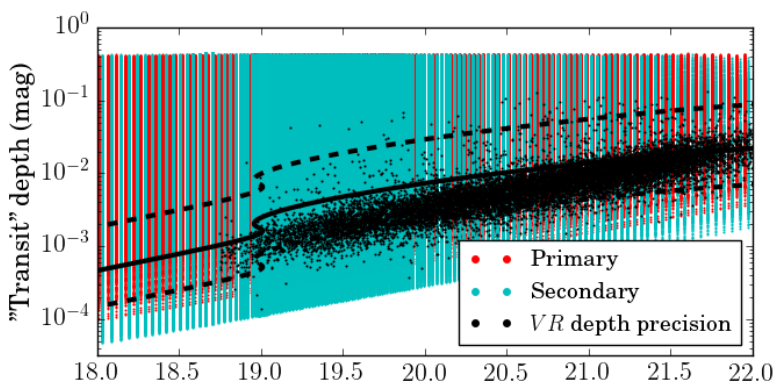
Transiting Planet Impostors: EBs



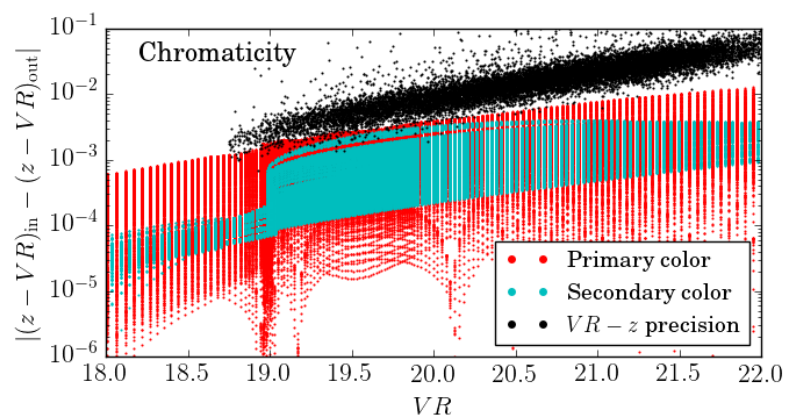
Ruling out Impostors: EBs



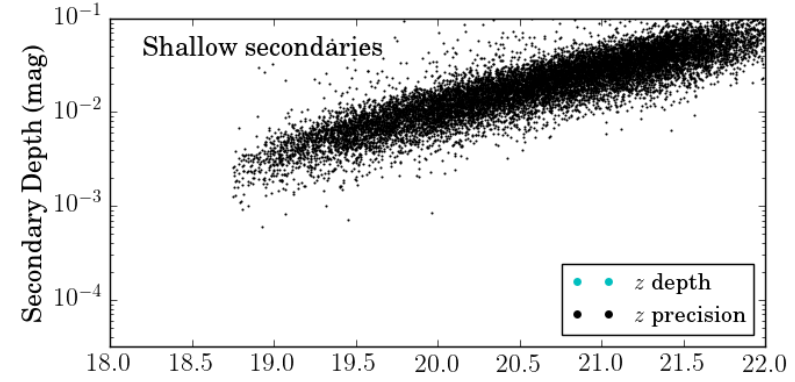
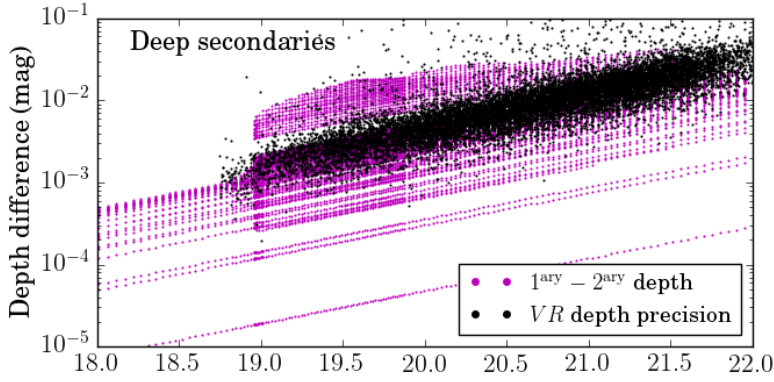
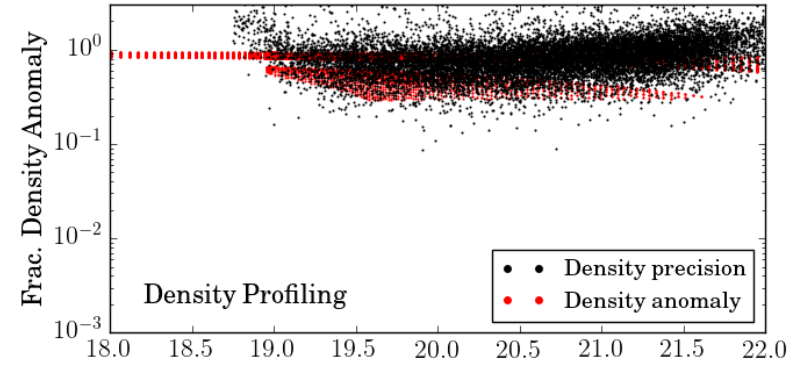
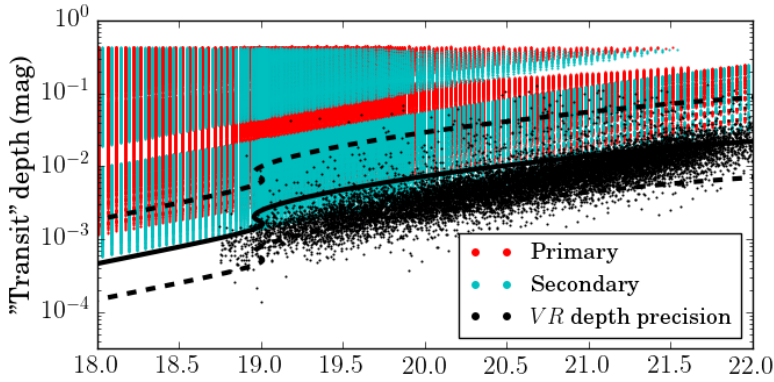
Blended EB = Heirarchical Triple $\pm \mu$



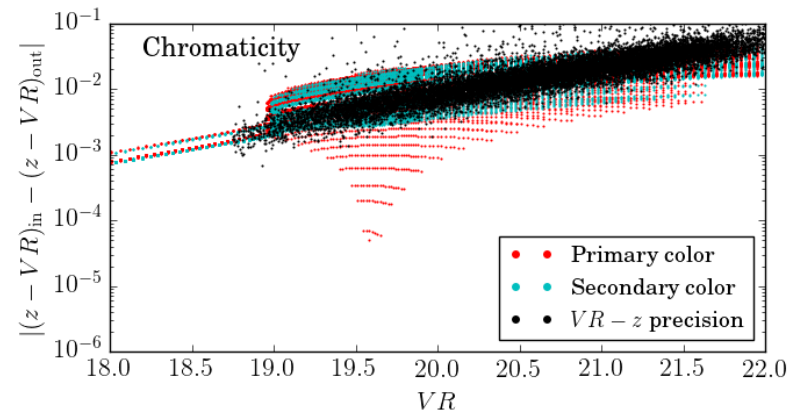
Sgr Targets
Sgr Blended EBs
Sgr Heirarchical Triples



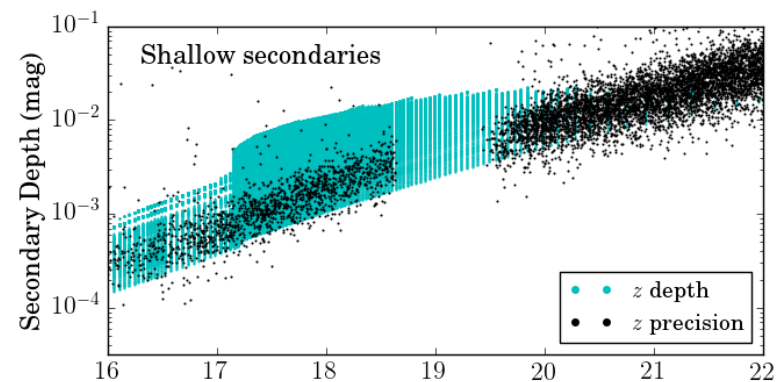
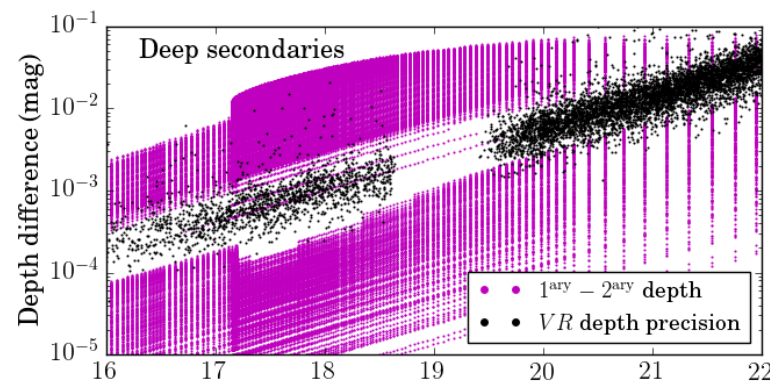
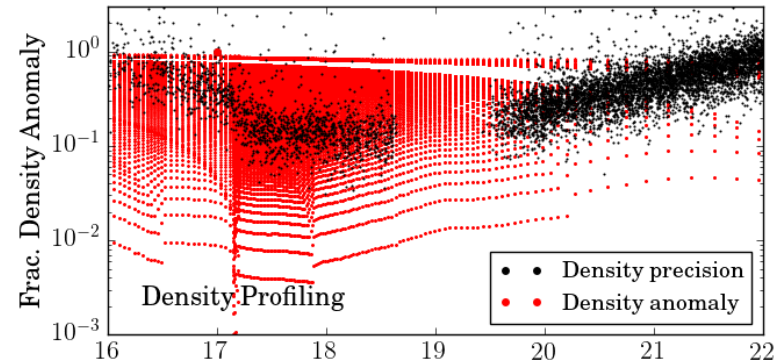
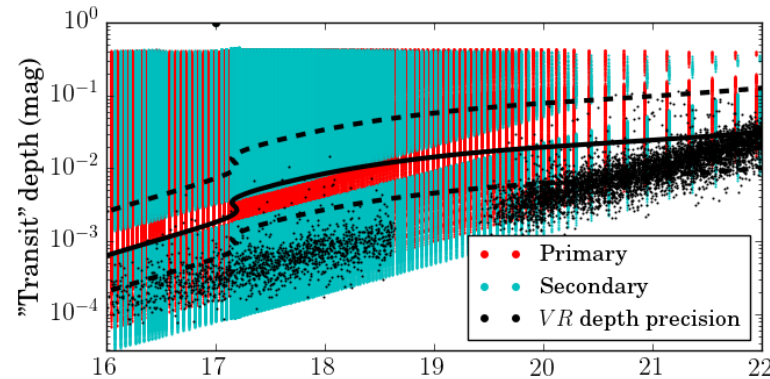
Blended EB = Heirarchical Triple $\pm \mu$



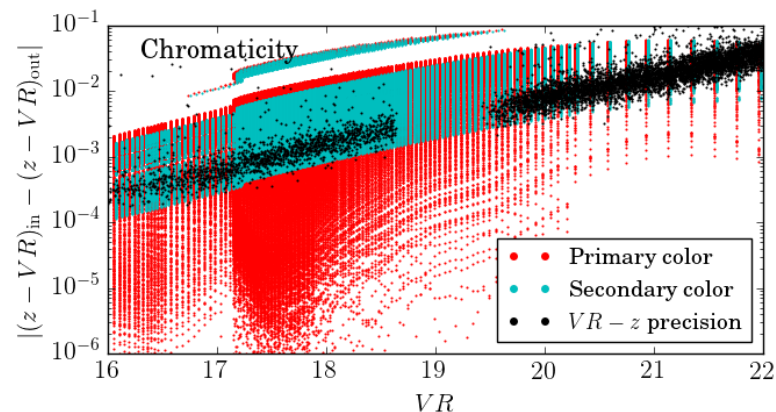
Sgr Targets
Bulge Blended EBs



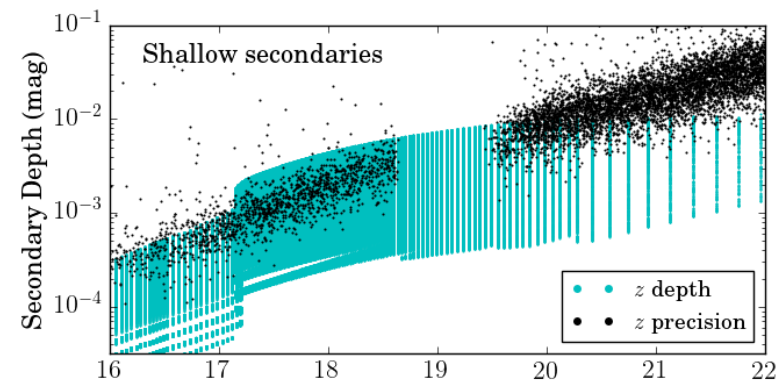
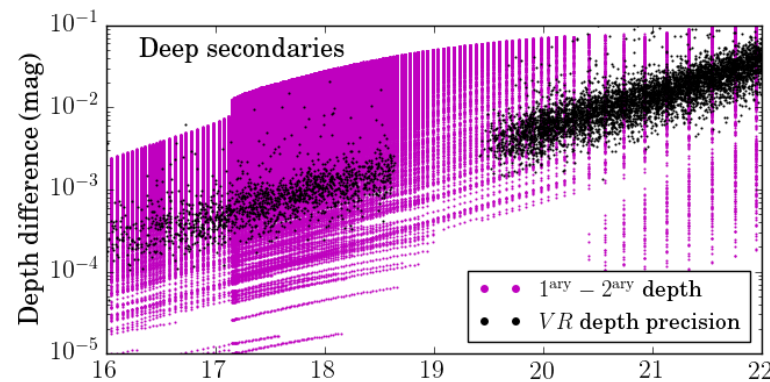
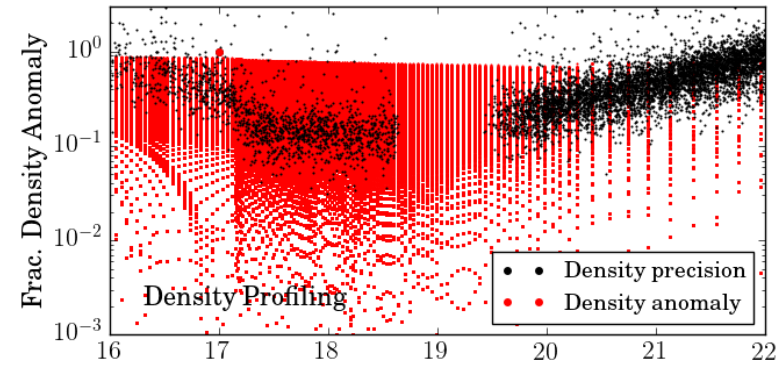
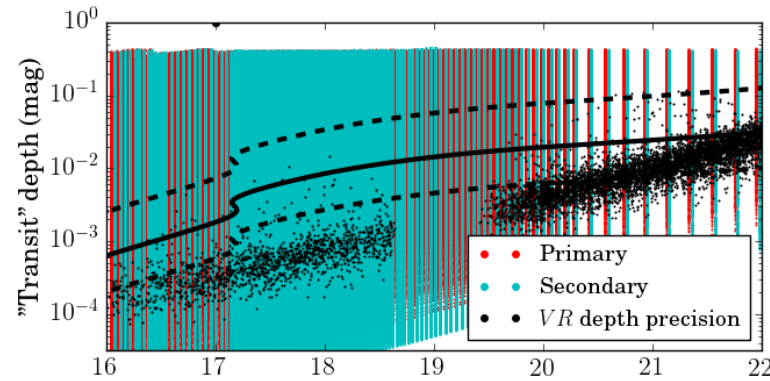
Blended EB = Heirarchical Triple $\pm \mu$



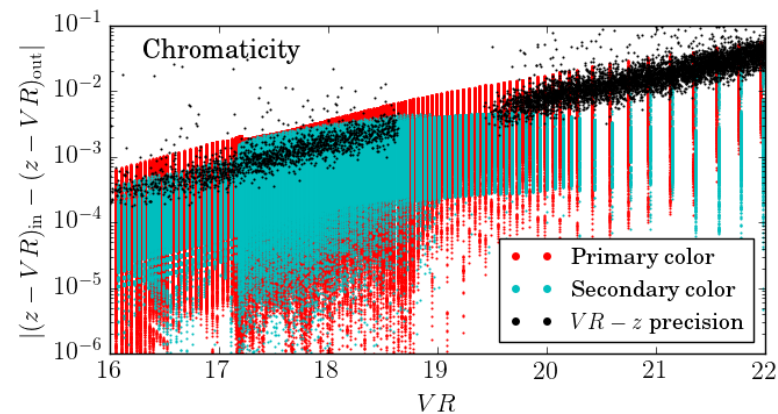
Bulge Targets
Bulge Blended EBs
Bulge Heirarchical Triples



Blended EB = Hierarchical Triple $\pm \mu$



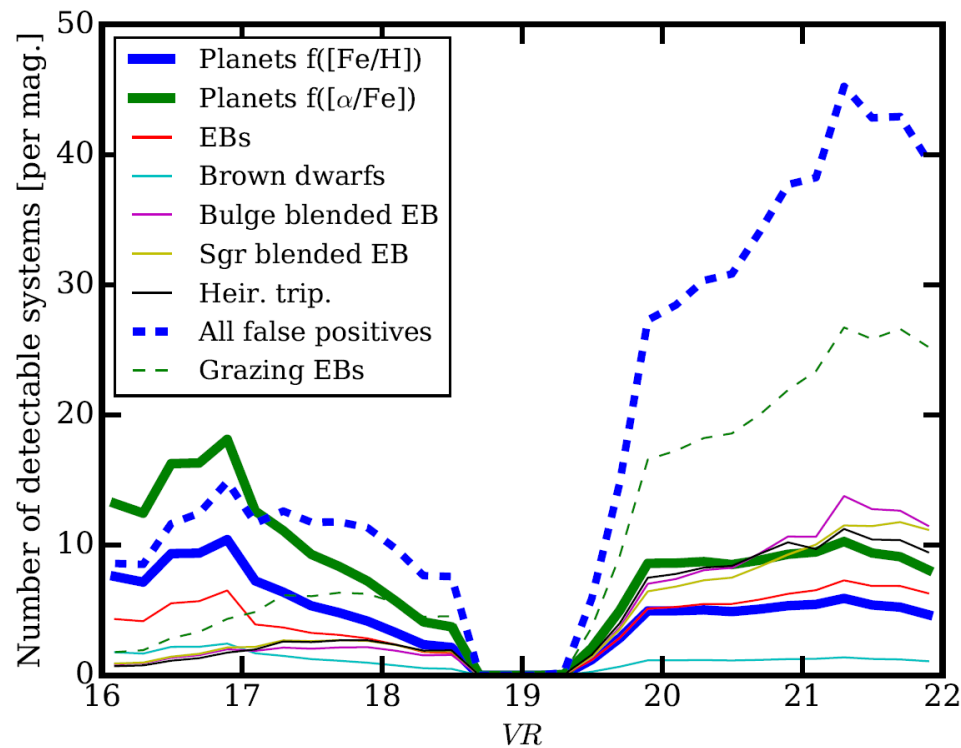
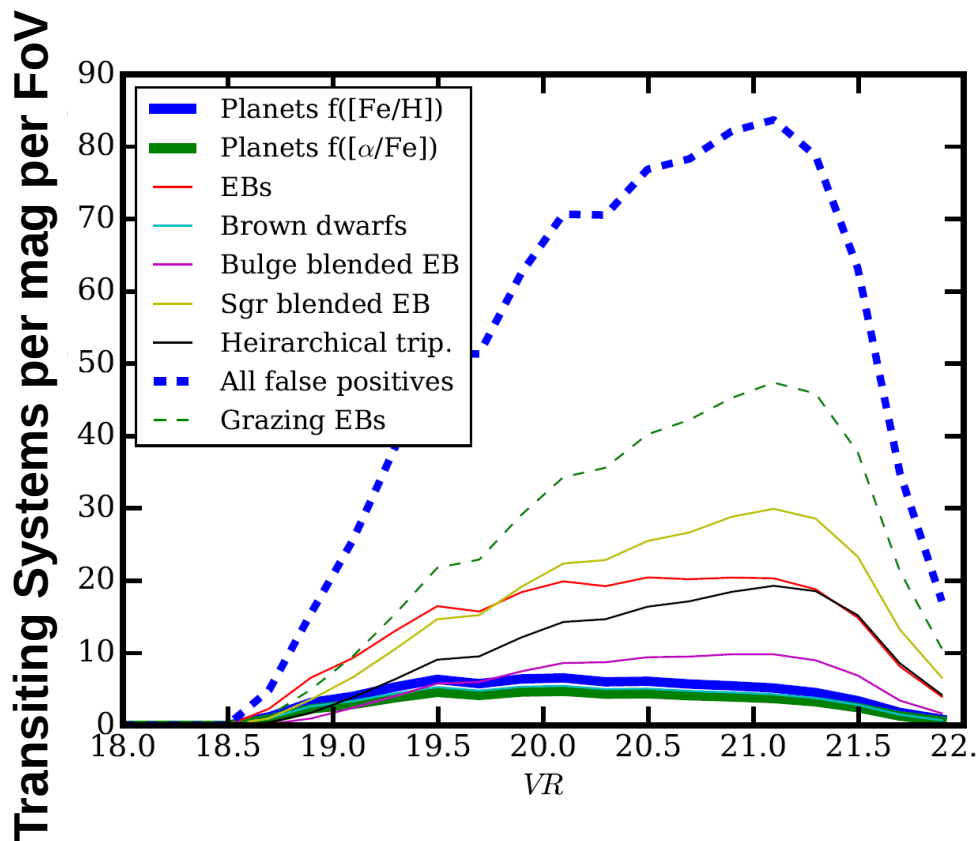
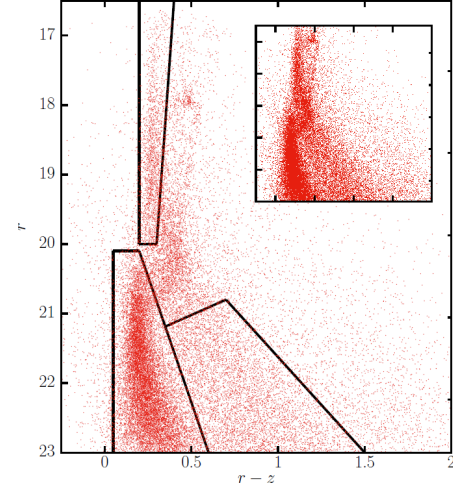
Bulge Targets
Sgr Blended EBs



False Positives

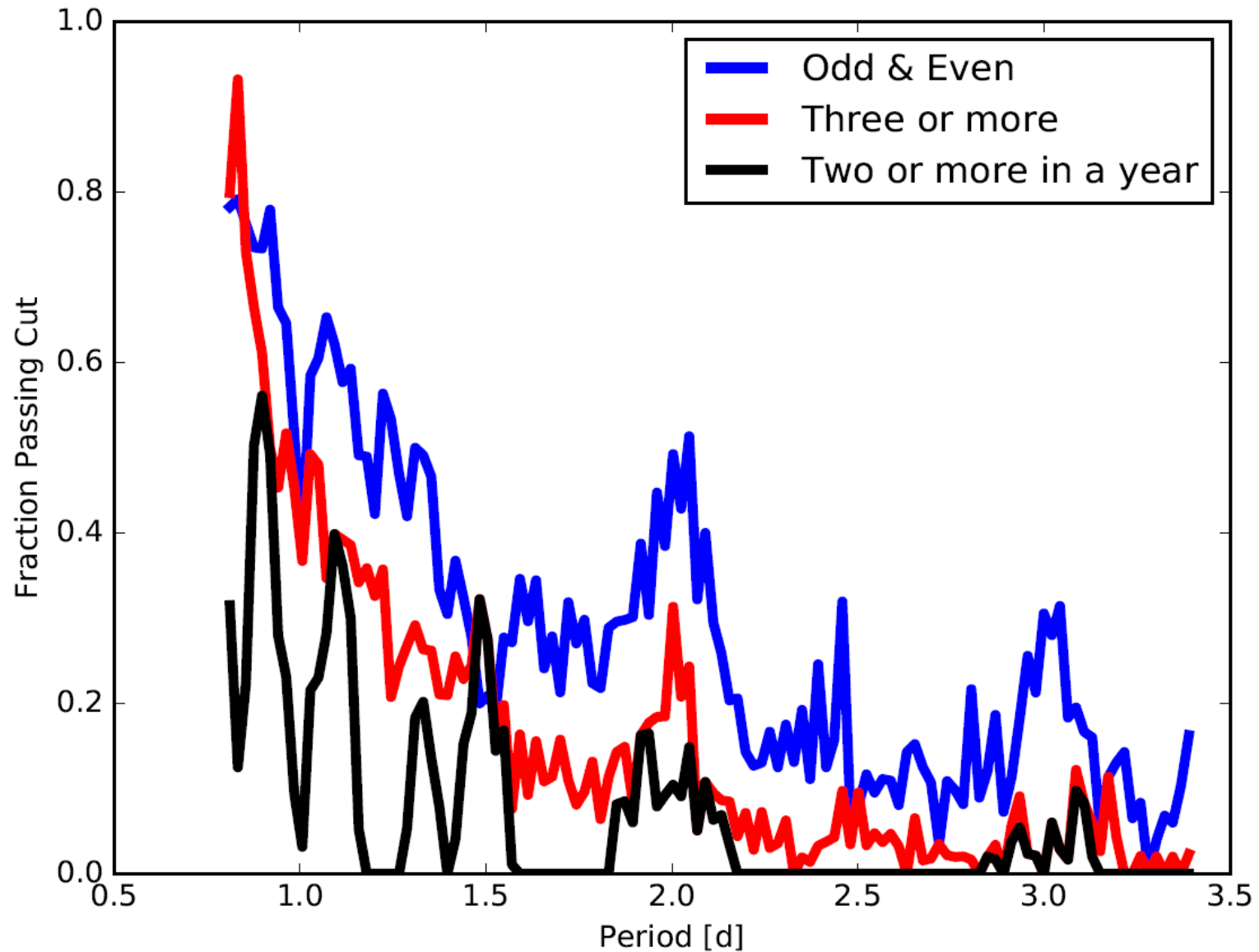
Sgr Dwarf

Bulge

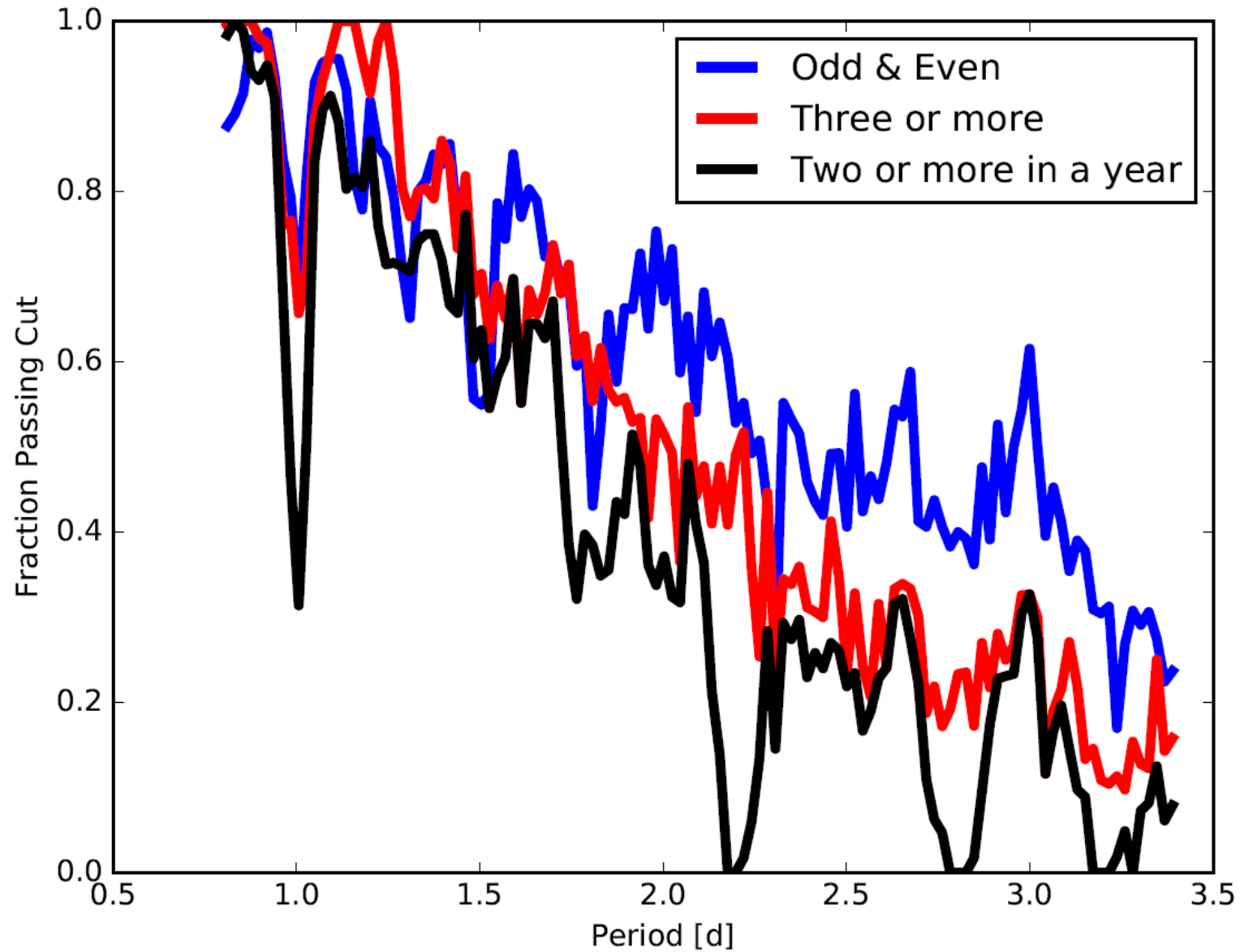


Based on stars detected in images (+ assumed incompleteness for blends)
Planet frequency assumes mean [Fe/H]

Closing the Window Function



Closing the Window Function



Conclusions

Sgr Dwarf

- Transits are detectable
- But, high false positive rate and lack of rejection in survey style observations makes project unfeasible
- RV virtually impossible

Bulge

- Transits easily detectable!
 - Lower false positive fractions
 - Many can be rejected with survey photometry only
 - RVs hard, but attainable (esp. recon)
- Bulge Transit surveys a promising route to probe high-Iron high-alpha planet formation
 - Can be done with modest amounts of telescope time