



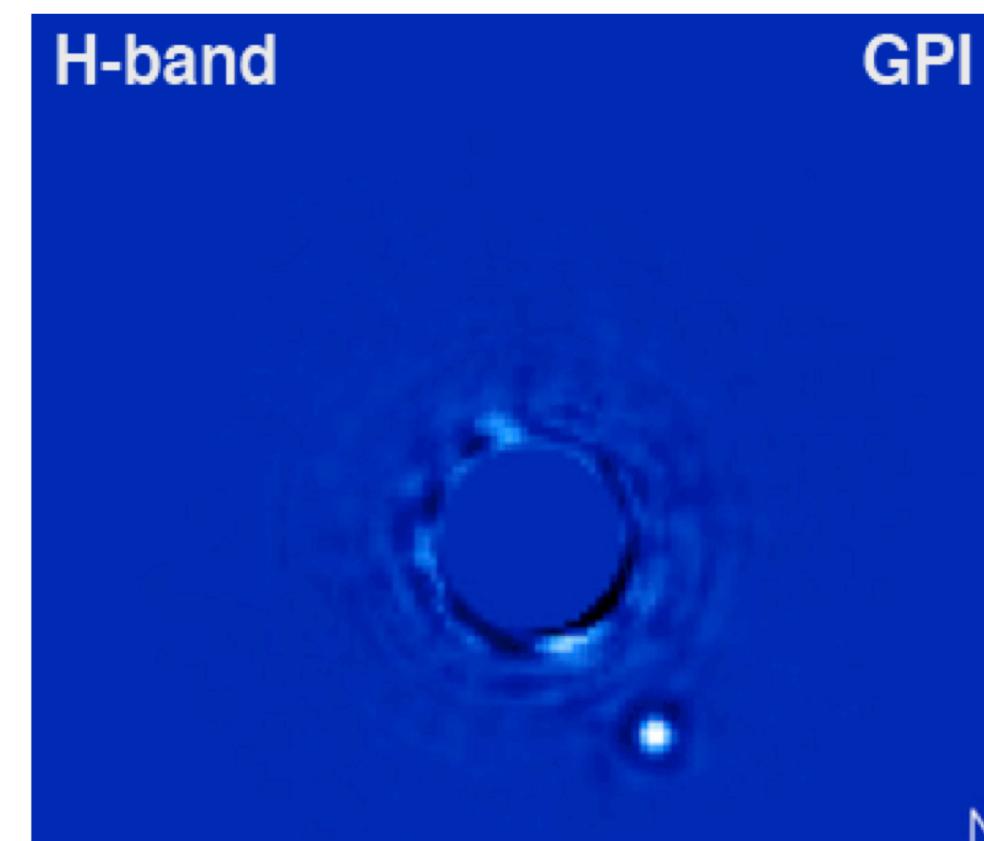
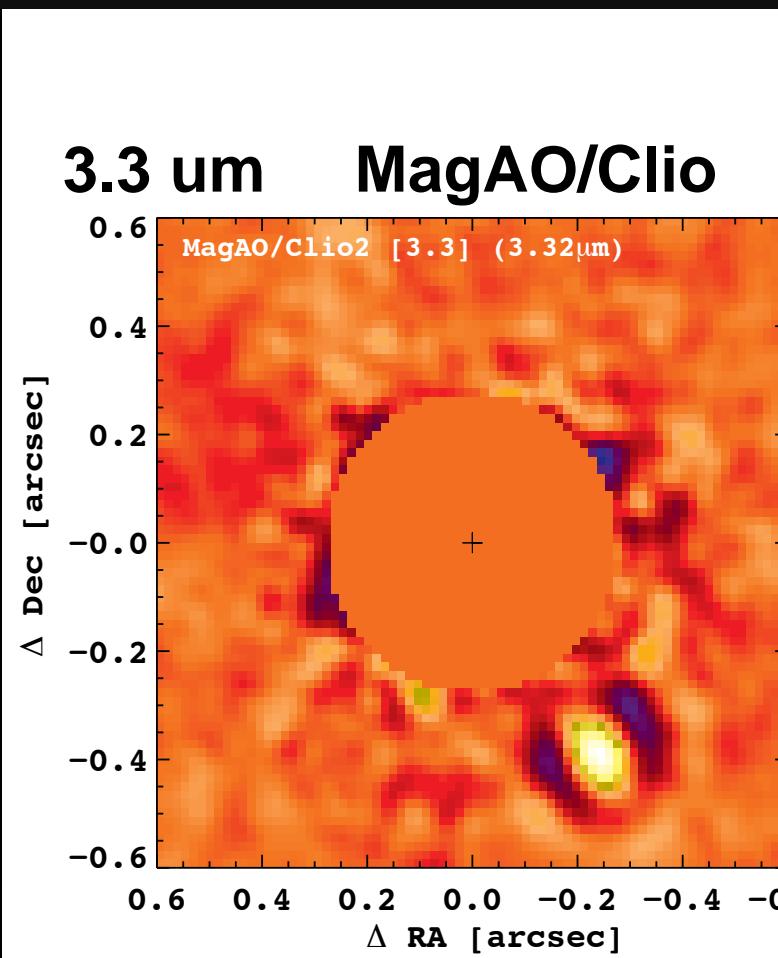
Using MagAO and GPI to obtain complete SEDs and empirical bolometric luminosities of young giant exoplanets

SEDs &
Lbols
of young
EGPs

Katie Morzinski

MagAO
Instrument
Scientist

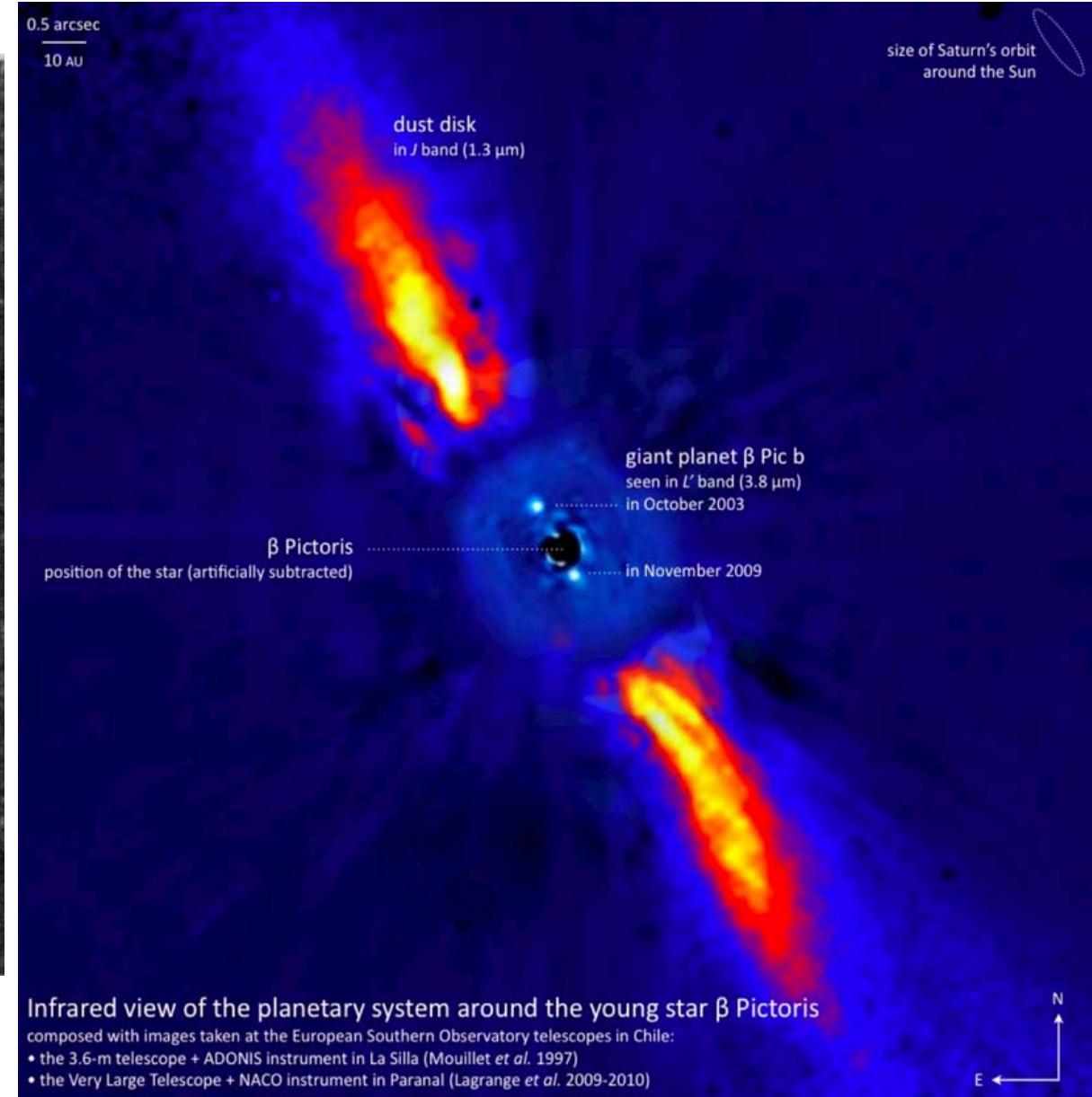
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Beta Pictoris – First star known to host bodies analogous to those seen in the Solar System

Using
MagAO



Magellan AO





Gemini Planet Imager

Using
MagAO



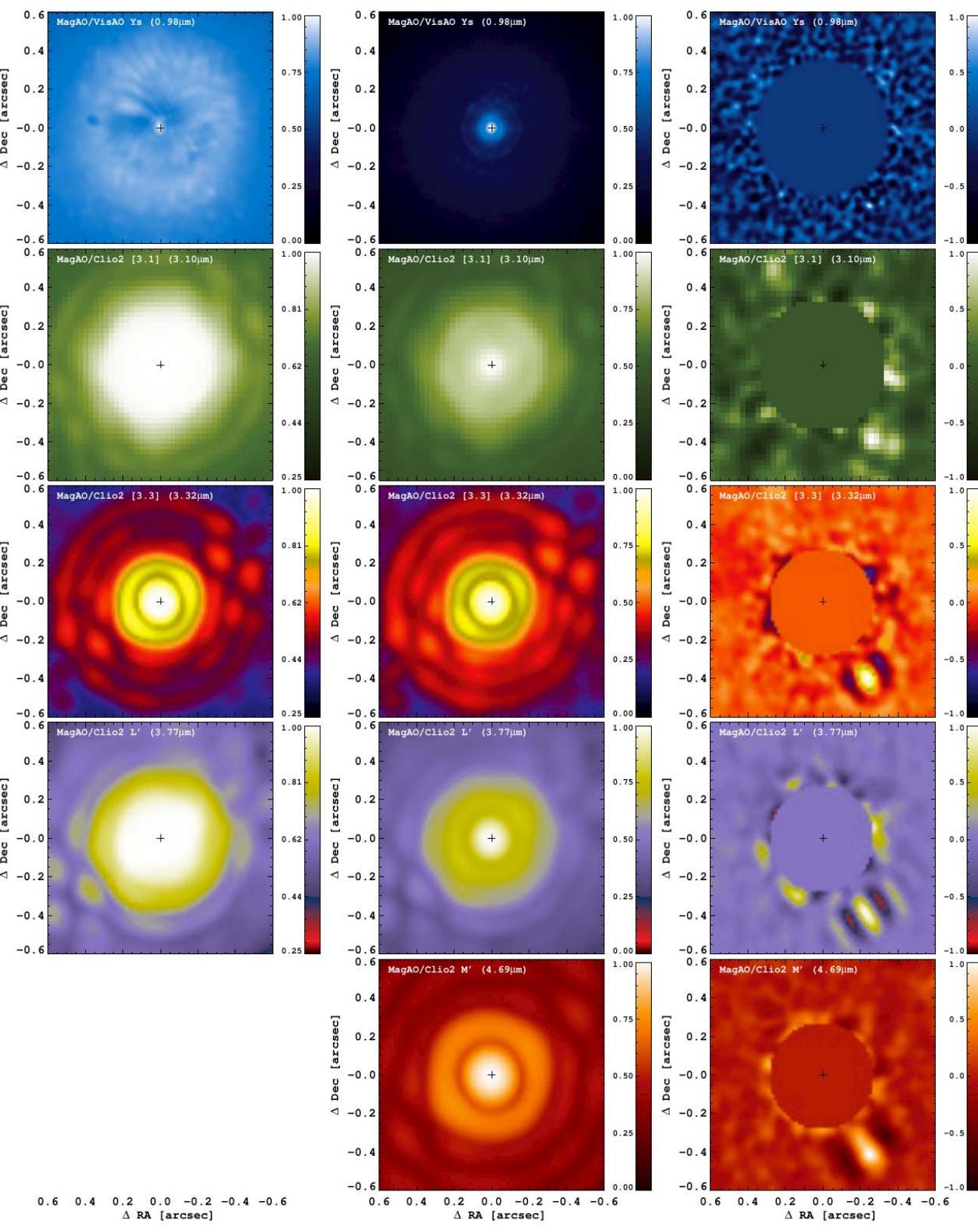


Using MagAO & GPI to obtain complete SEDs & Lbol of young EGPs

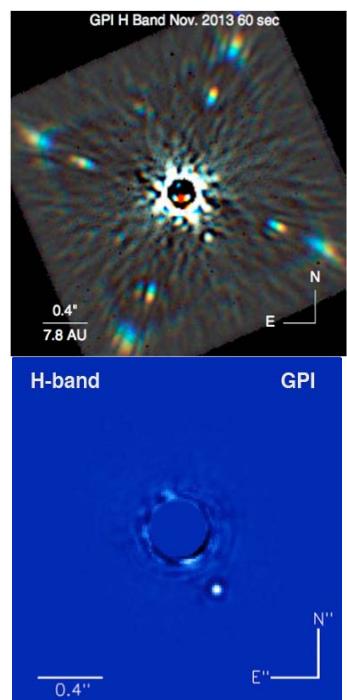
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Males et al 2014



Macintosh et al 2014

Morzinski et al 2015



Beta Pic b: L2.5

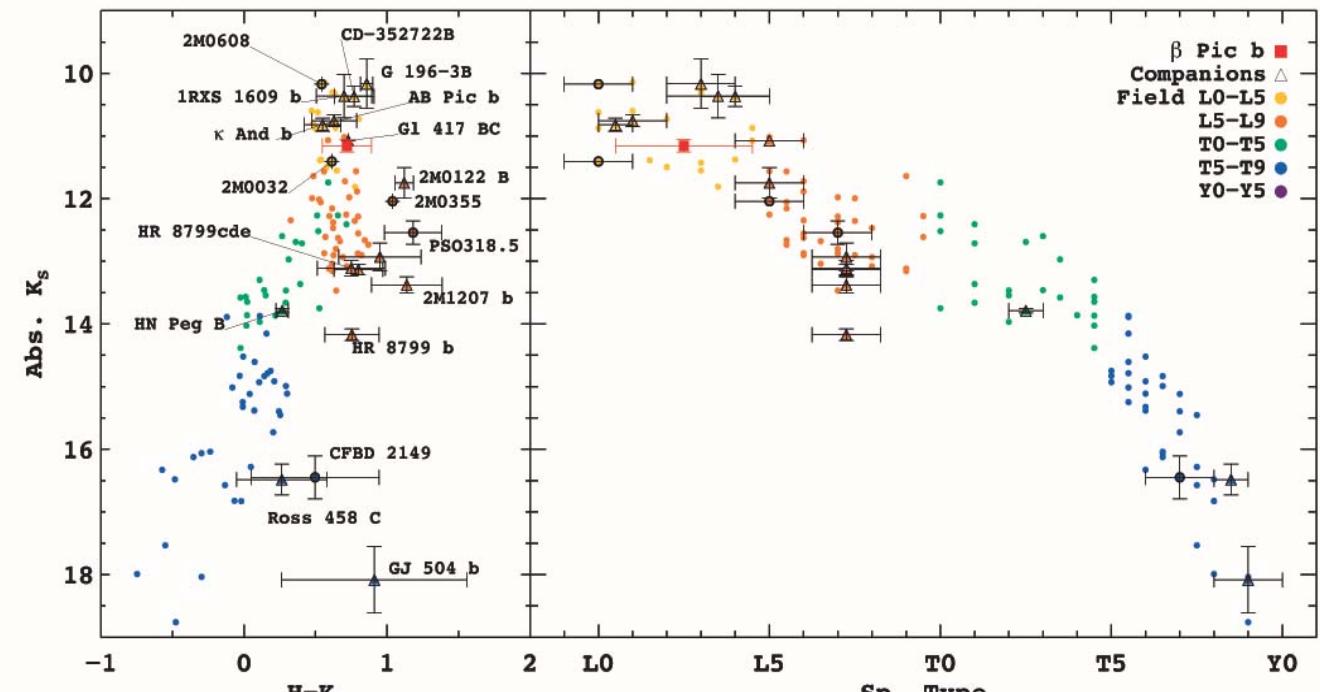
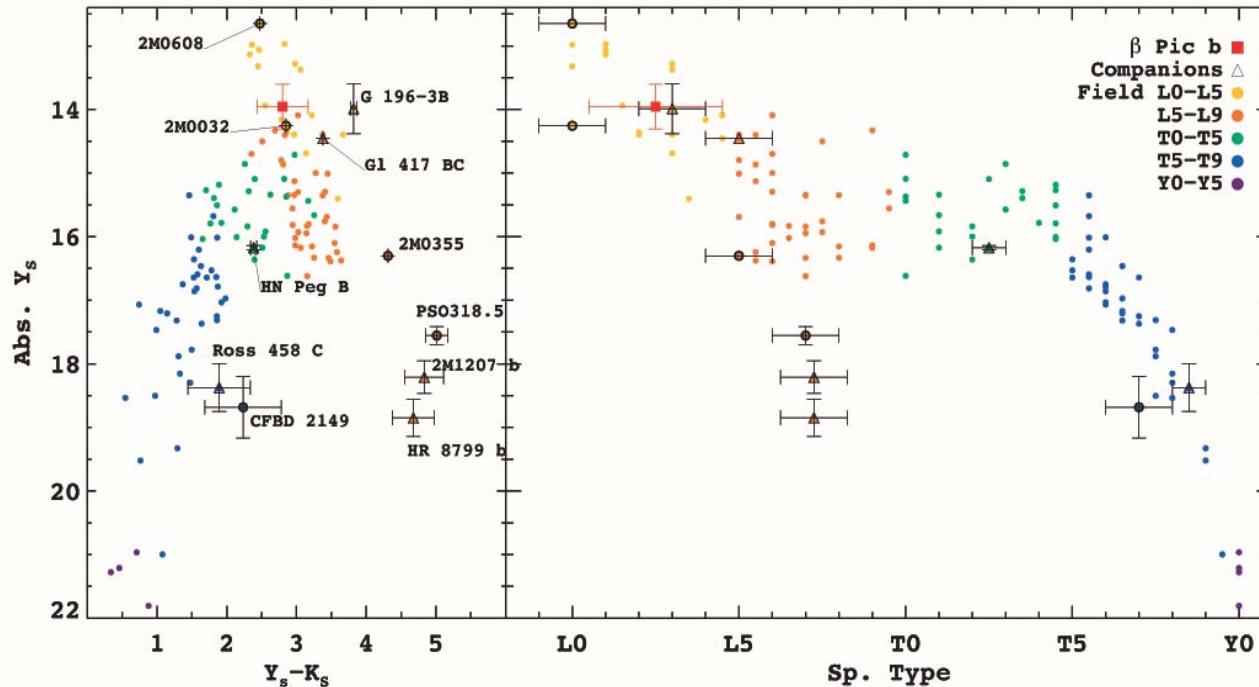
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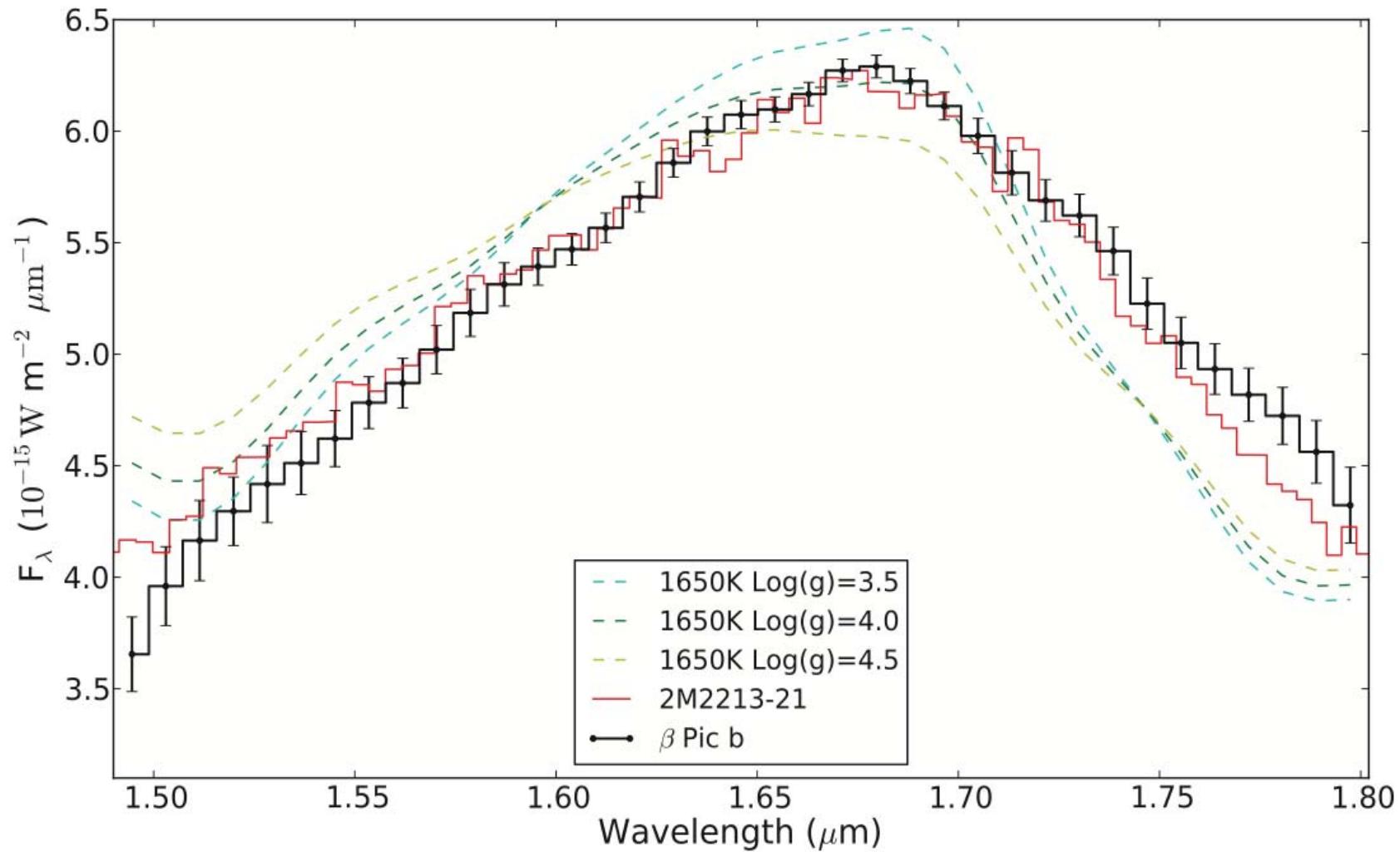
GPI H spectrum: Teff 1600-1700 K; logg 3-5

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Chilcote et al. 2015



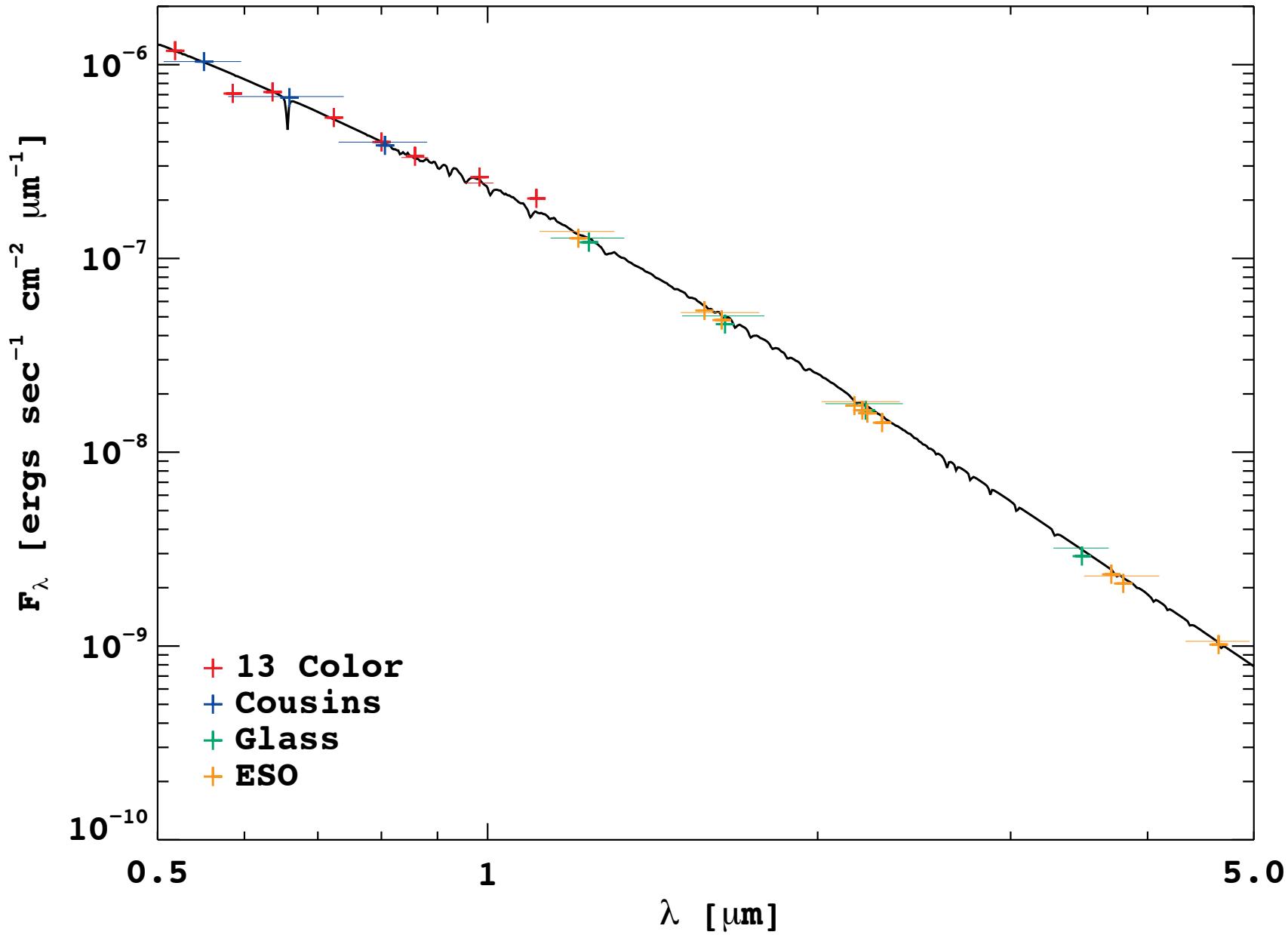
Spectrum of Beta Pic A

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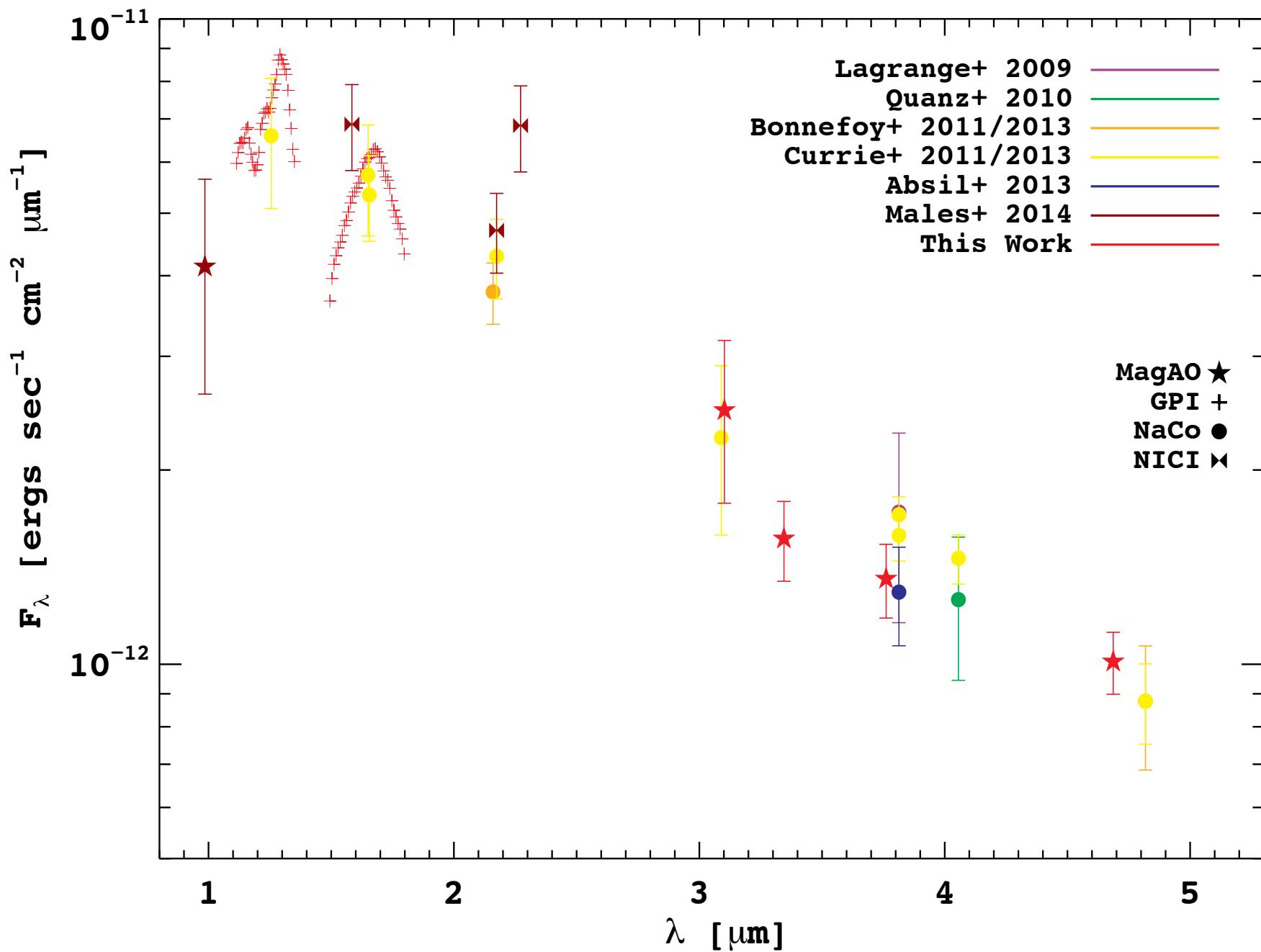
Beta Pic b Spectral Energy Distribution 0.9-5um

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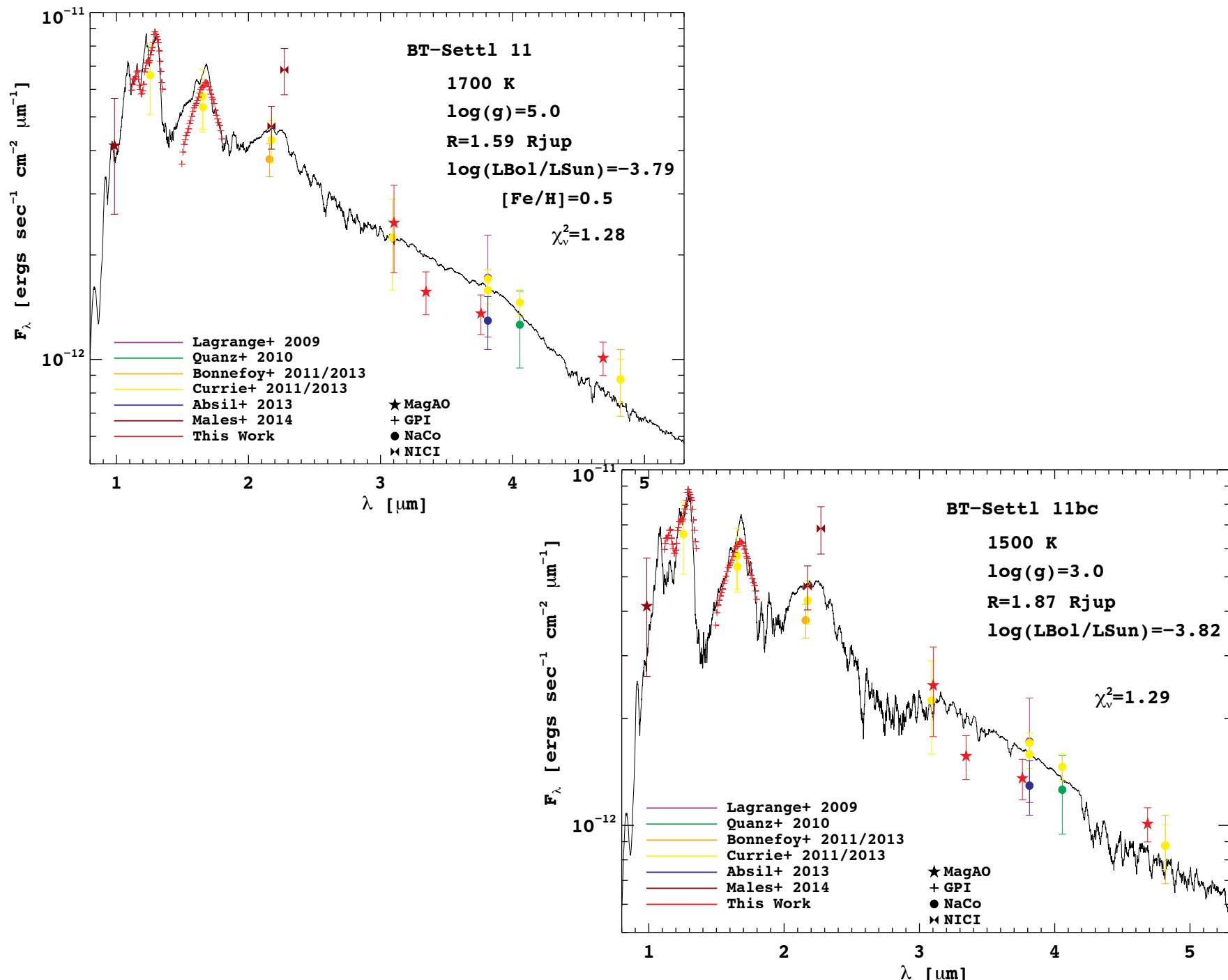
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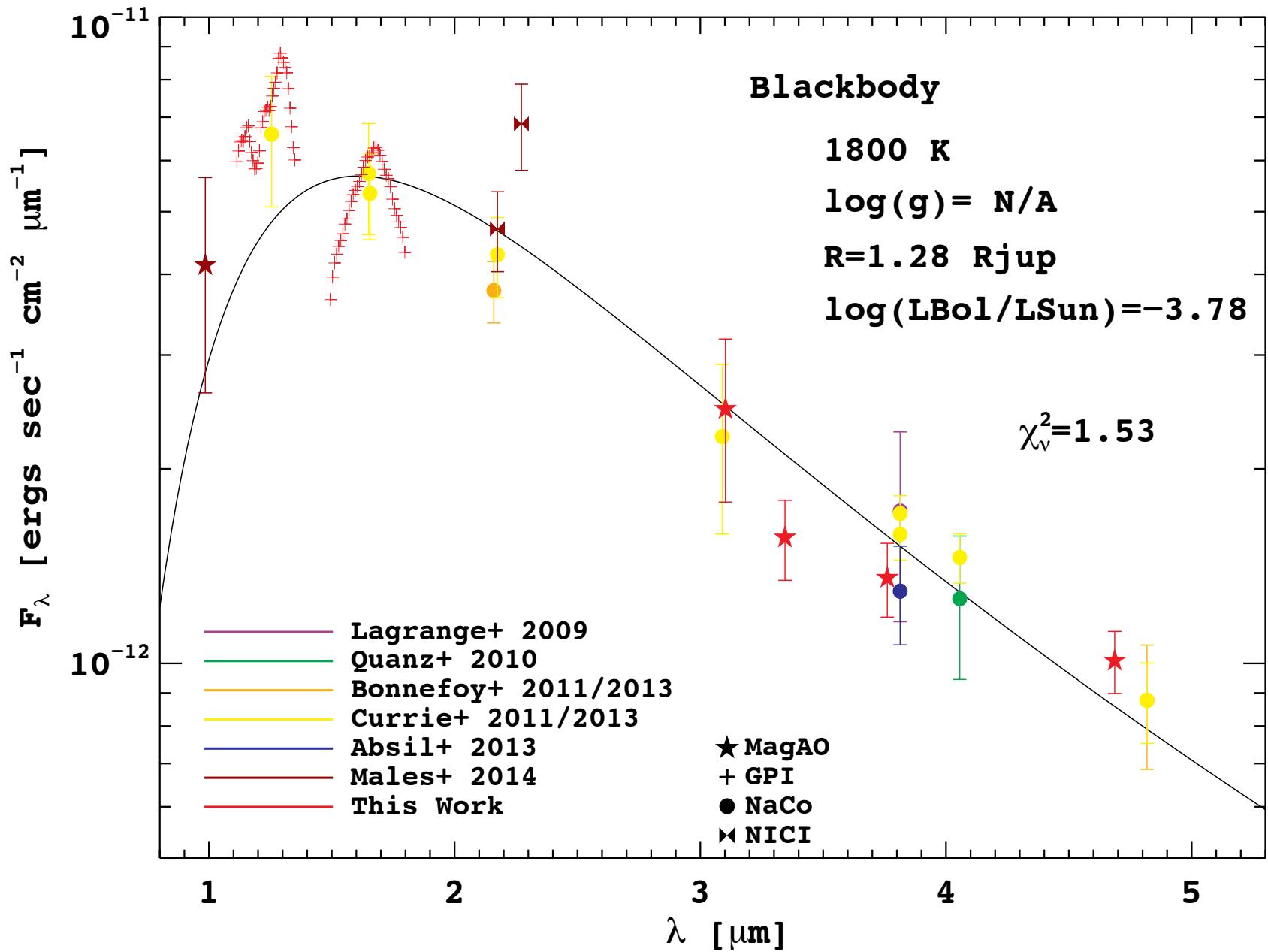
Blackbody fit: Teff 1800 K

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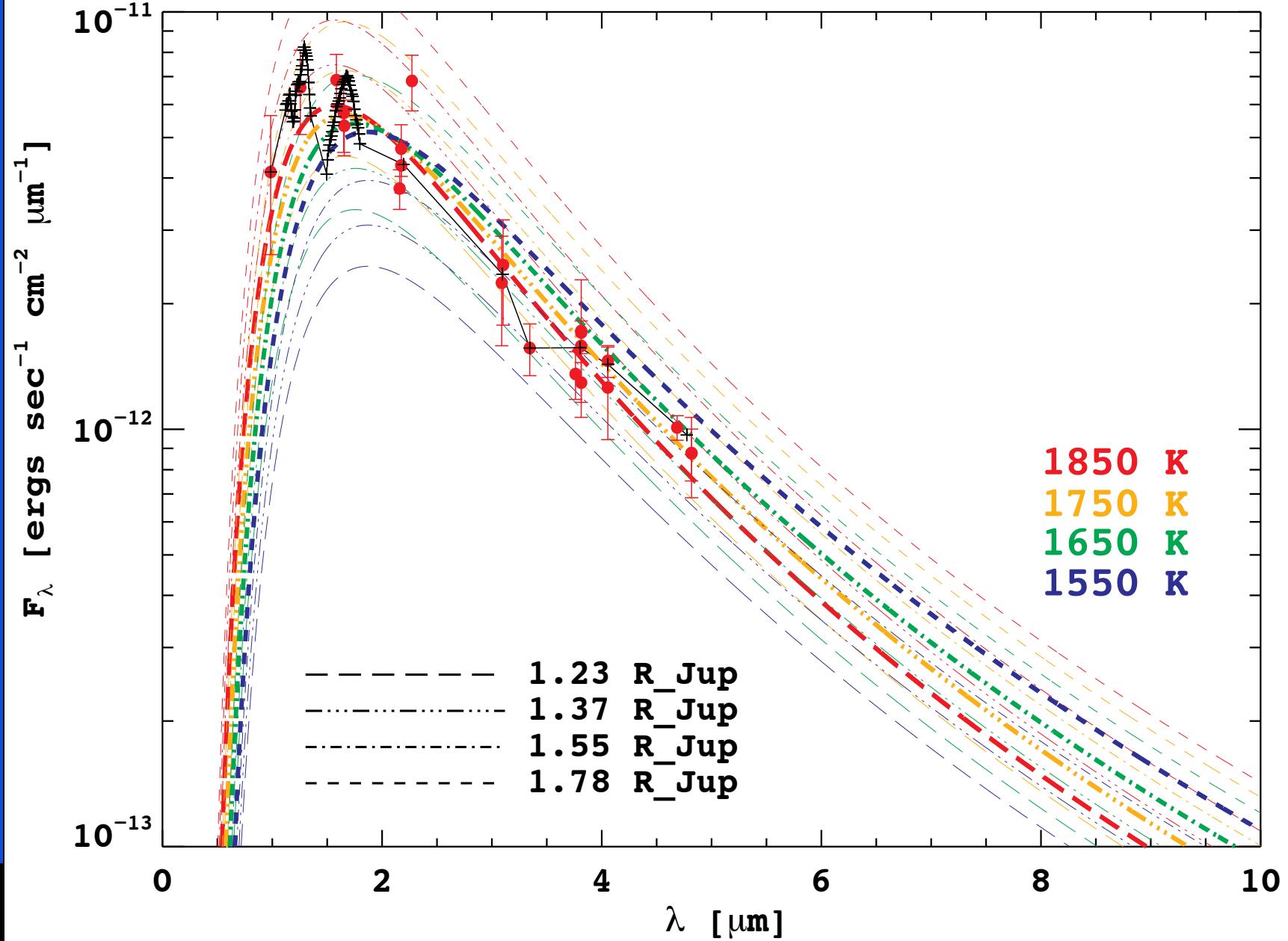
Blackbody families

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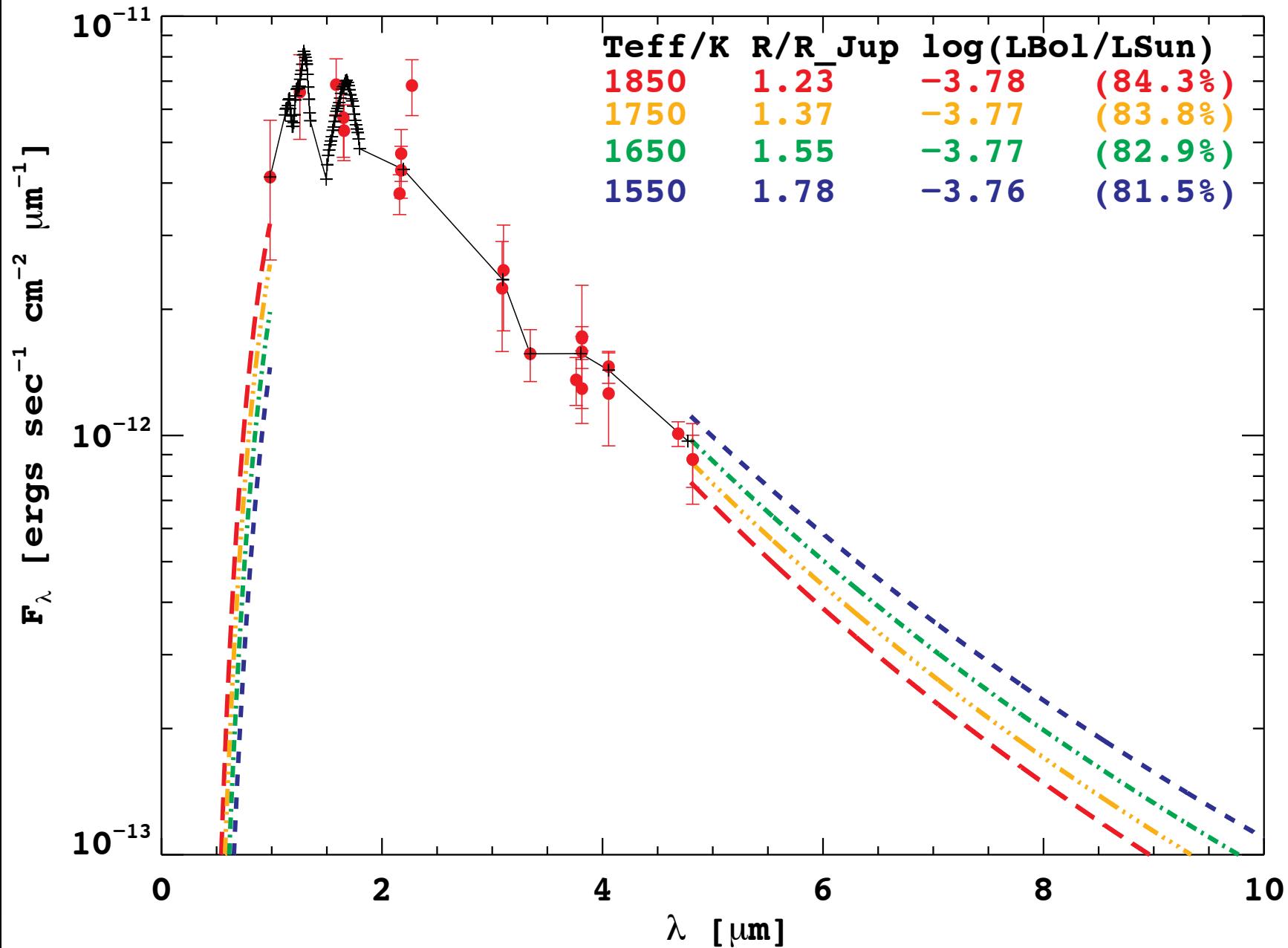
VisAO constrains Teff while Clio constrains R

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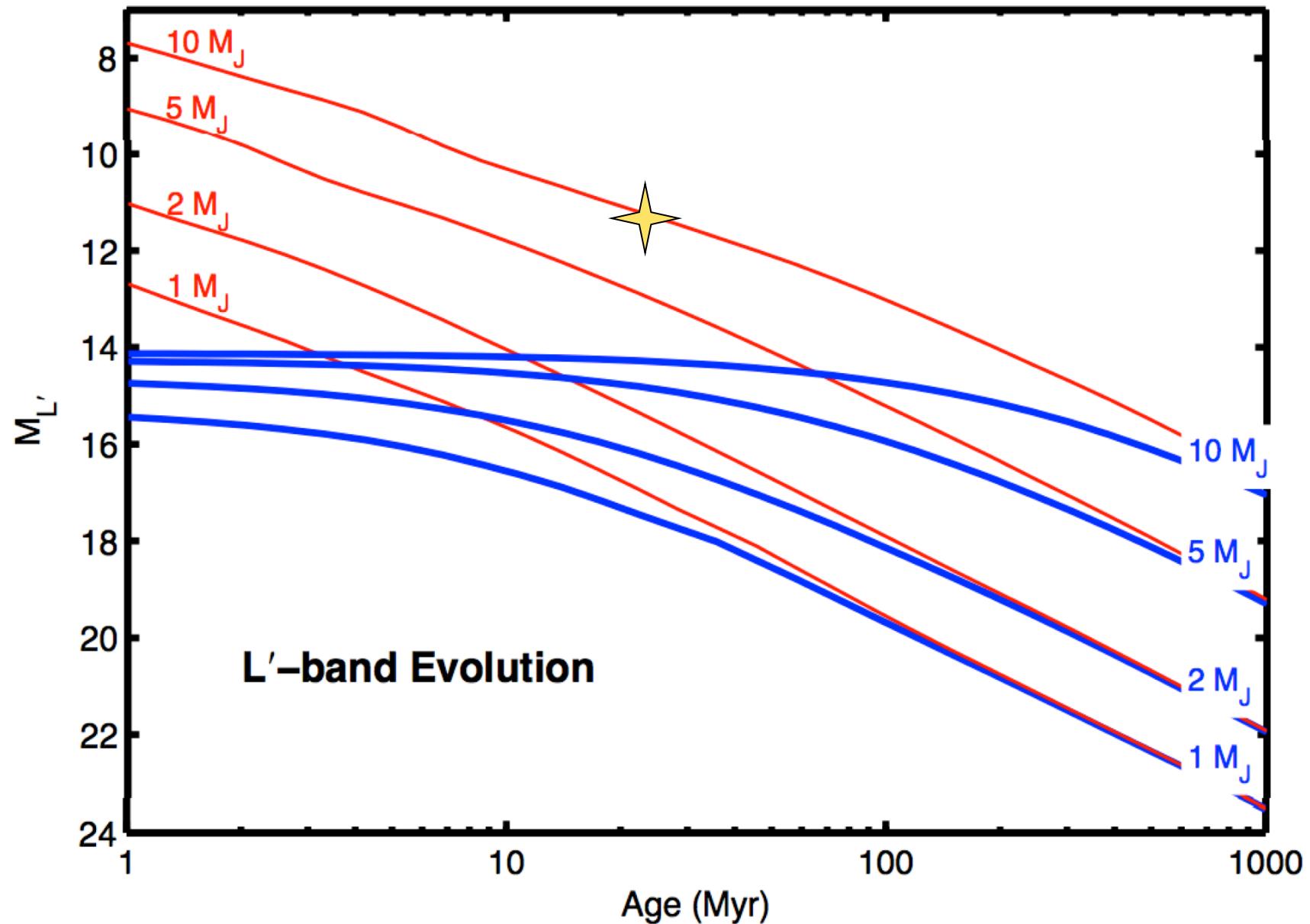
Beta Pic b fits the hot-start evolutionary models

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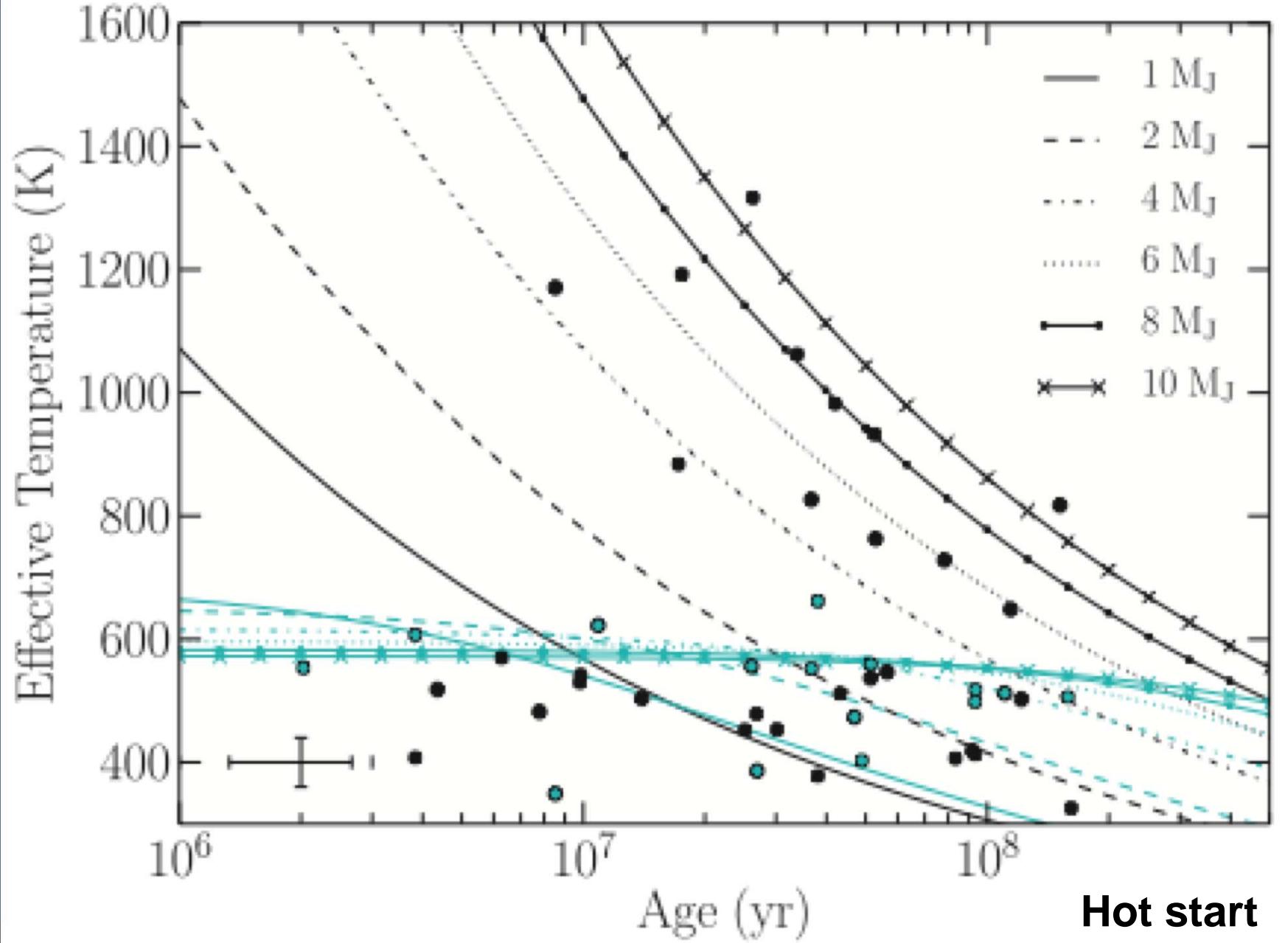
Measuring L_Bol for tens more planets with GPI & MagAO will fill in the formation picture

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McBride et al 2011

Hot start
Cold start



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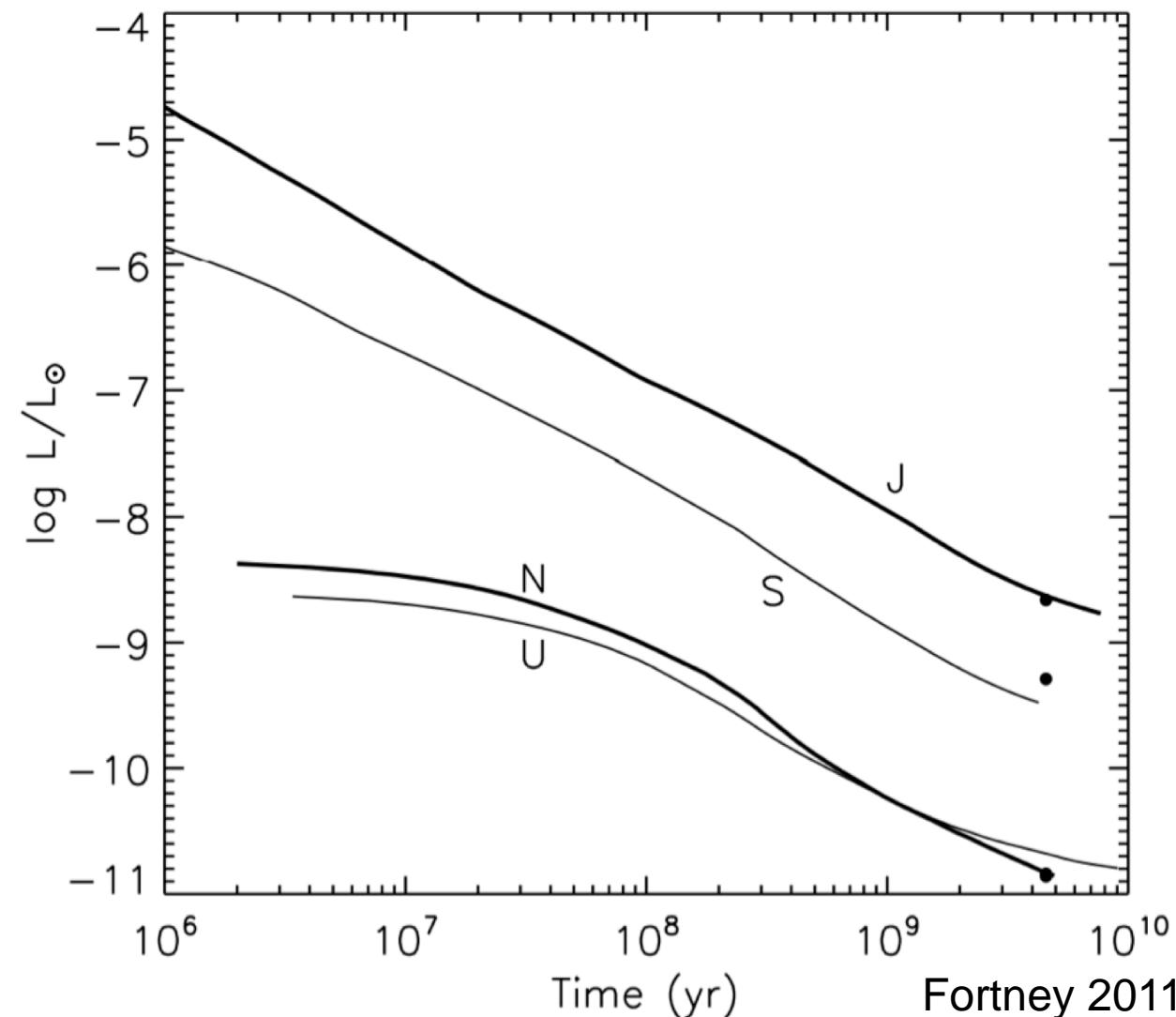
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Energy budget: to model cooling evolution

- Insolation/irradiation
- Gravitationally contracting
- Differentiation

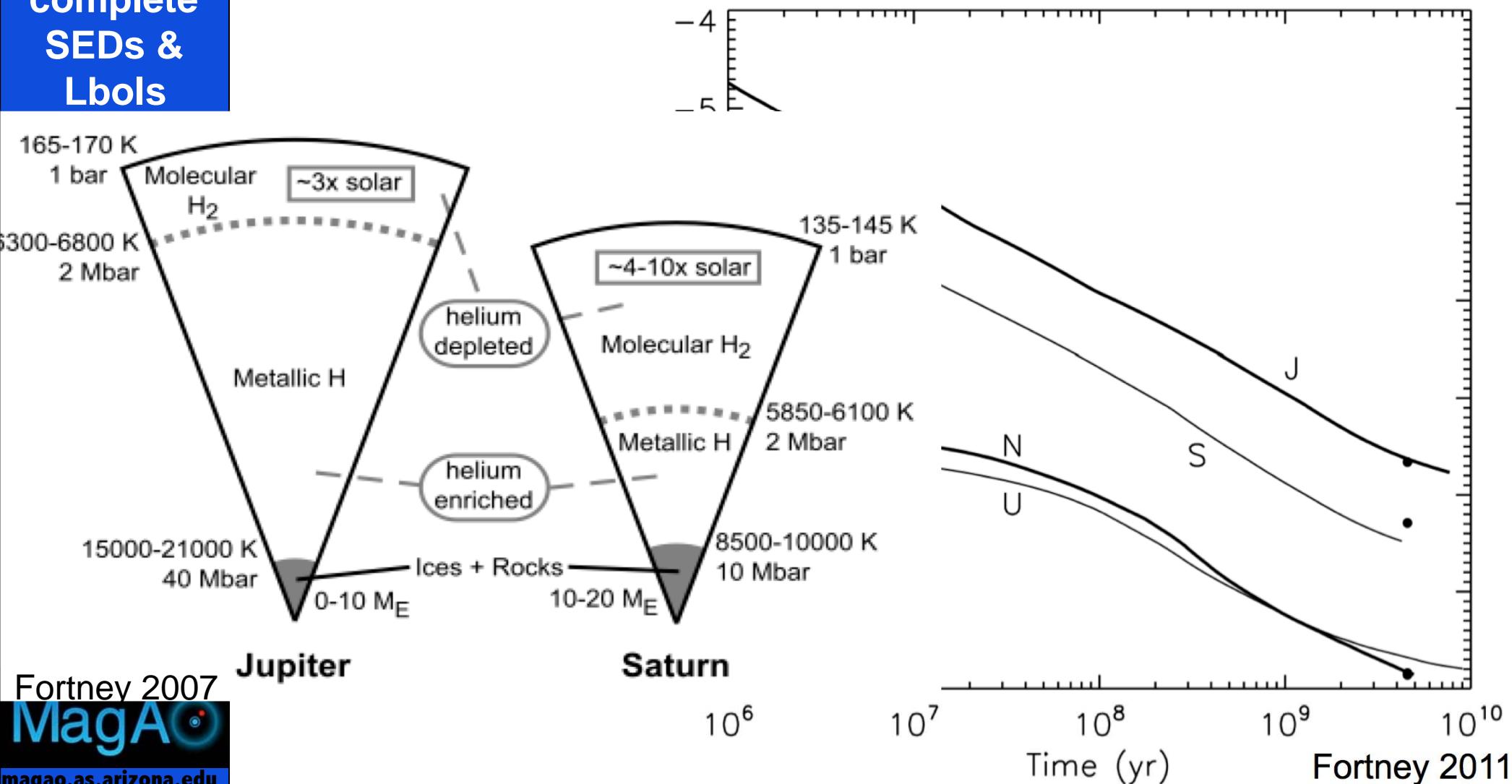




Cooling evolution: Interior models

Using
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- Insolation/irradiation
- Gravitationally contracting
- Differentiation





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- Fundamental parameters describing formation & evolution remain to be measured
- Broad spectral coverage of the full O/IR SED with MagAO, plus spectra with GPI, is a powerful tool to measure L_Bol and atmospheric properties
- Fully characterizing tens of Jupiter analogs with MagAO+GPI will give a better picture of the formation mechanisms
- This is still a data-starved field with a lot of exciting developments to come with cutting-edge instruments like MagAO, GPI, & JWST!

Thanks: Michelson Sagan NExSci Dawn
et al!

MagAO

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