

The Search for Exomoons

David M. Kipping^{1,2}

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- “For moons the effects are likely to be not just small, but minute—right on the hairy edge of what *Kepler* can do” Greg Laughlin, *The Economist*
- But, the potential for new understanding is extraordinary.

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- Extrinsic habitability
- Planet formation theory

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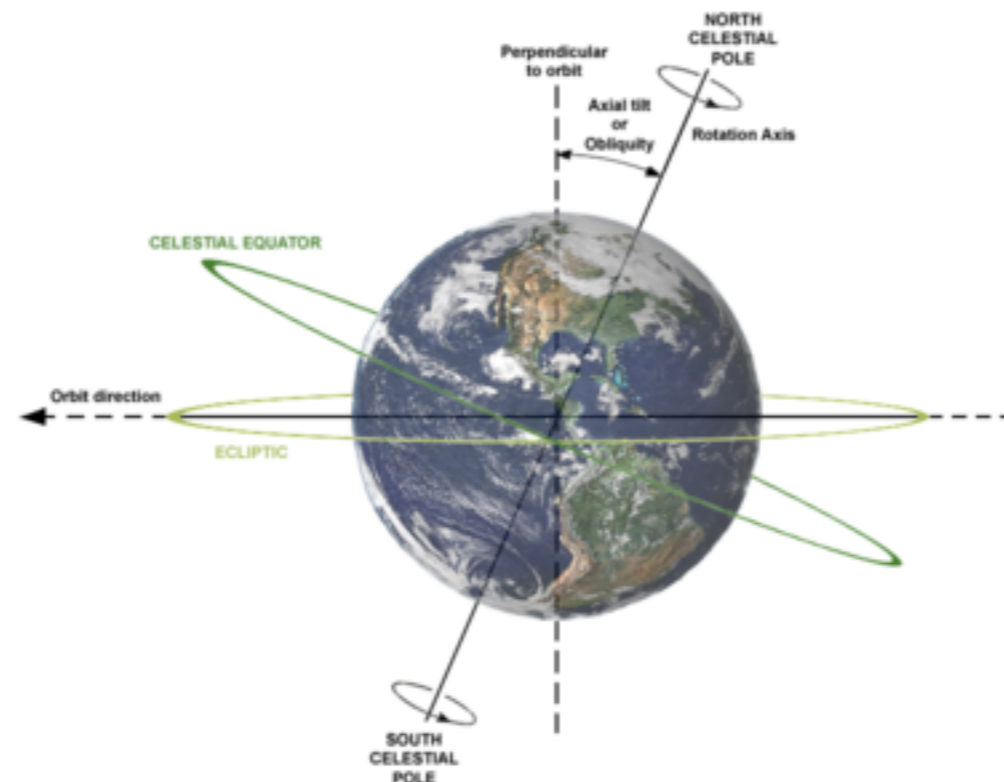


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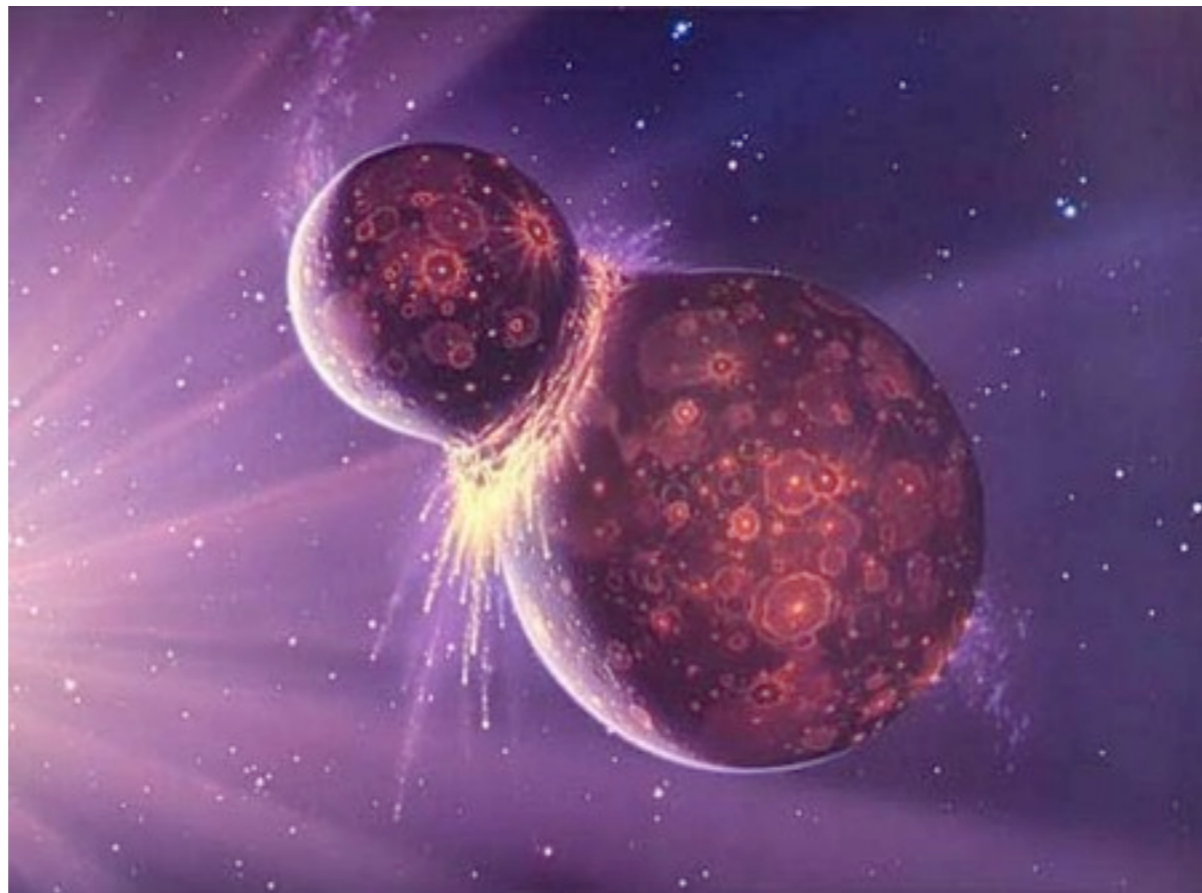


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Order-of-Magnitude Feasibility

- Roughly, Kepler is sensitive to $\sim 1 R_{\oplus}$ planets
- \Rightarrow Kepler is sensitive to $\sim 1 R_{\oplus}$ moons
- We may be able to detect Earth-sized/mass moons

Order-of-Magnitude Feasibility

No such moons in the Solar System, so we are searching for an alien class of object

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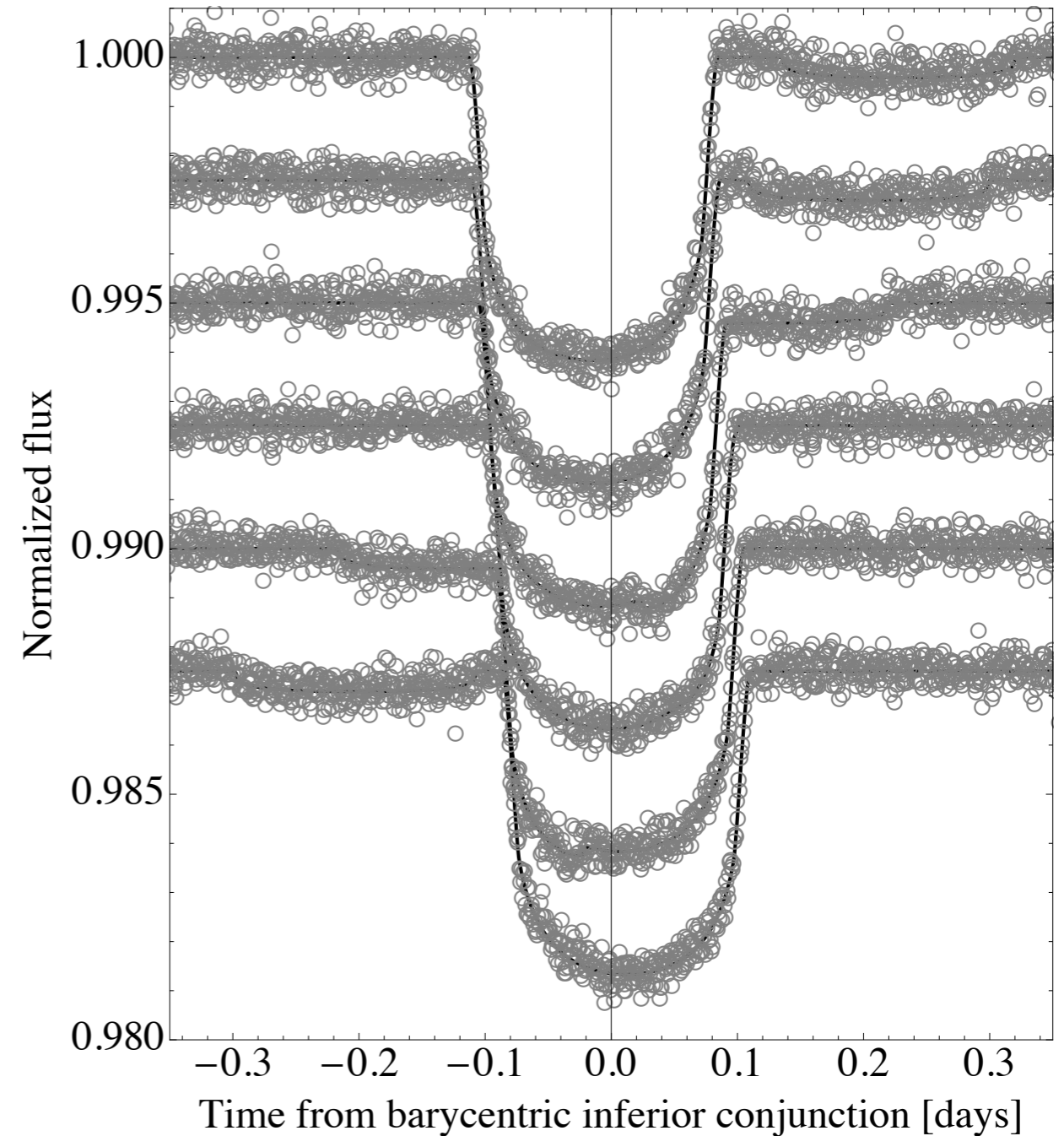
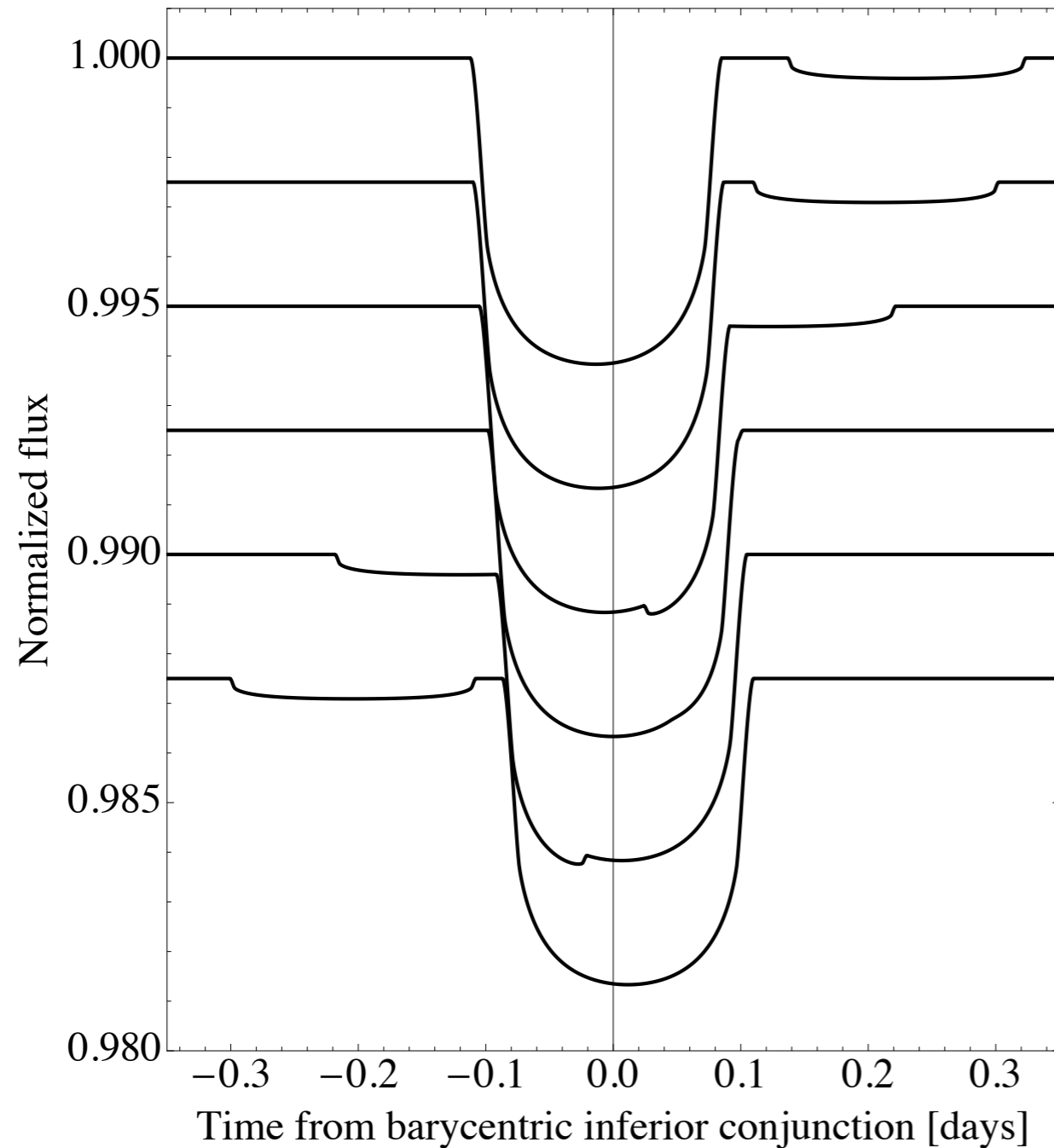
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LUNA

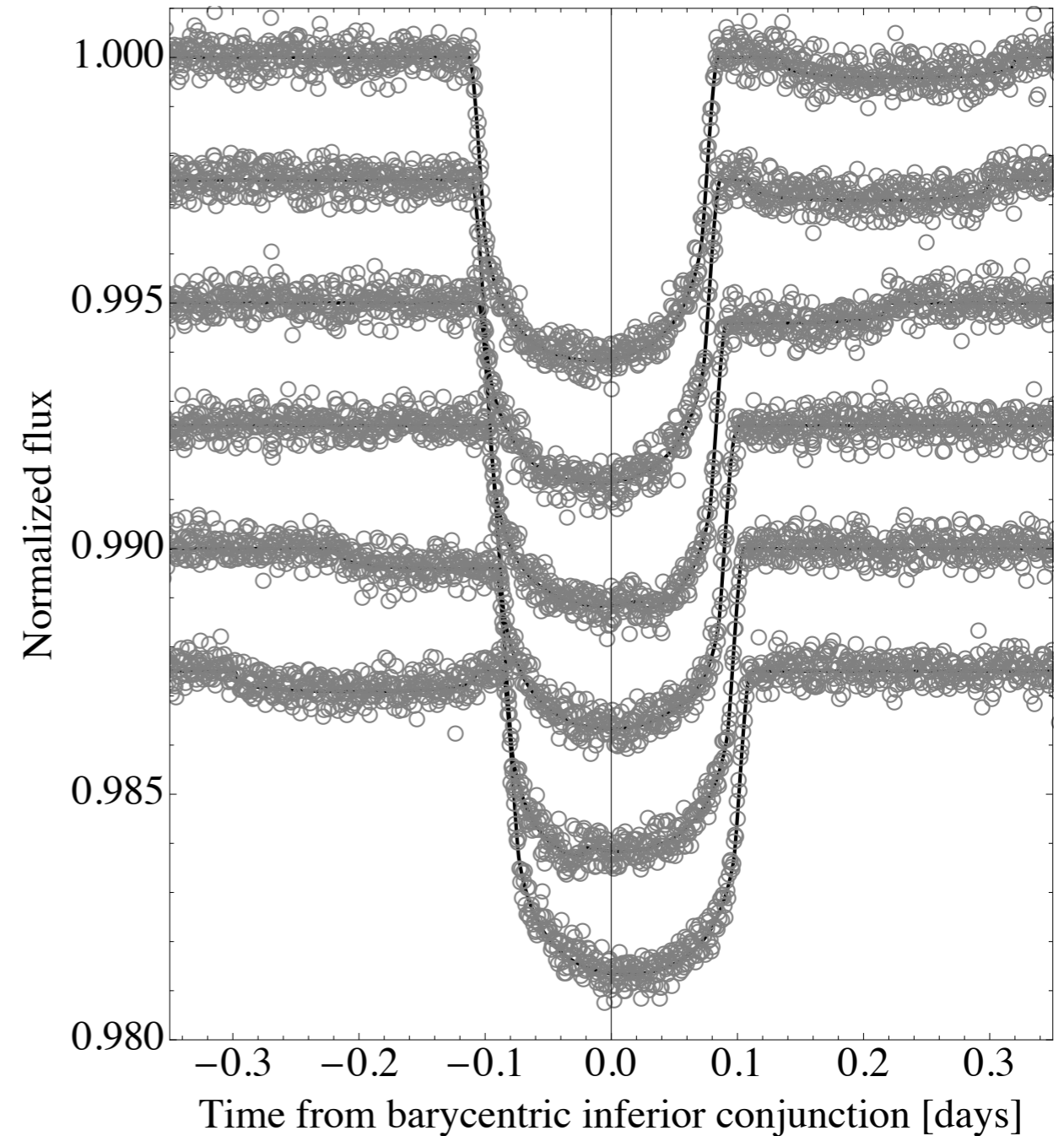
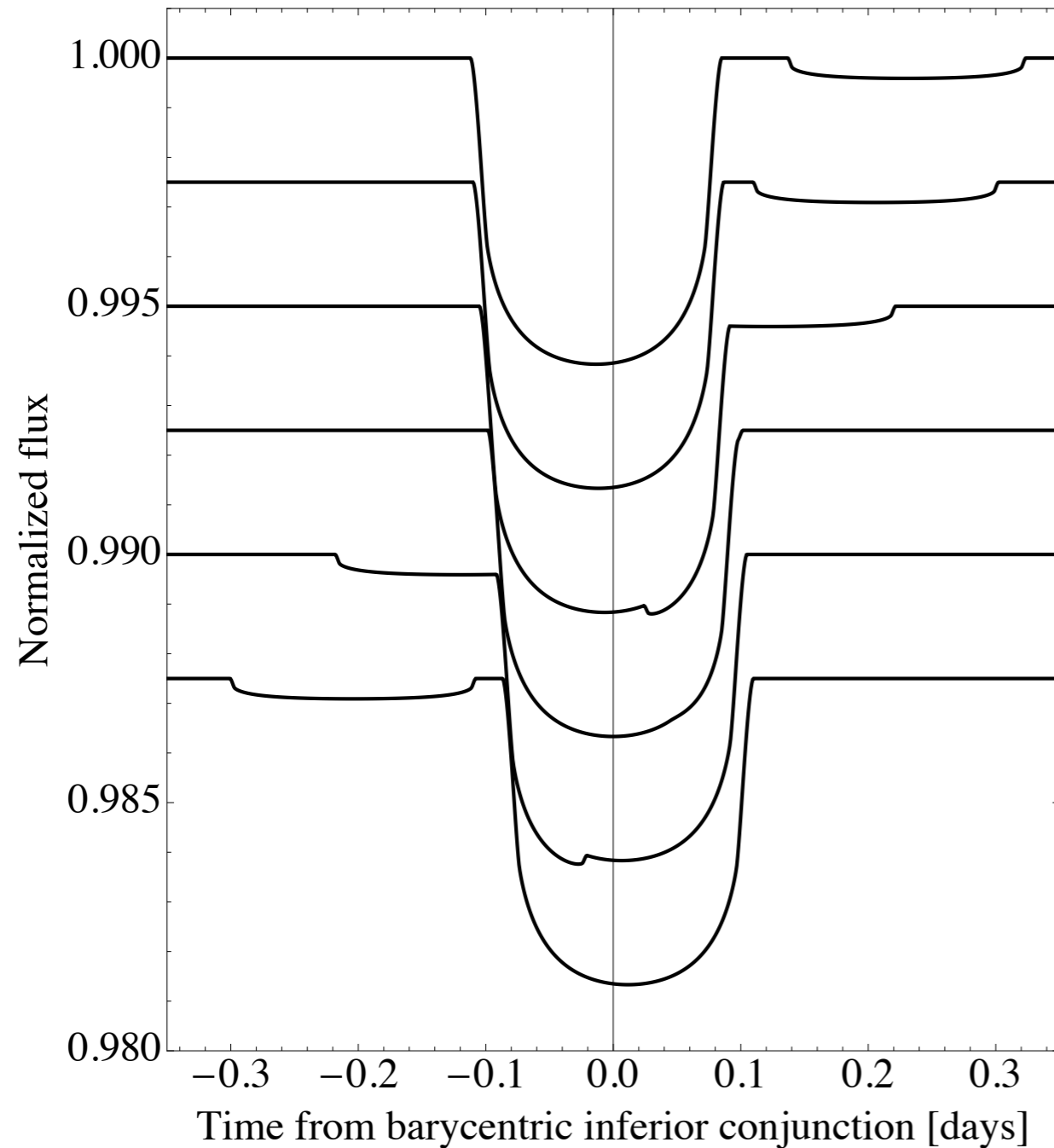
- *LUNA: An algorithm to generate dynamic planet-moon transits* (Kipping 2011)
- LUNA accounts for TTV/TDV-V/TDV-TIP as well as auxiliary transits/mutual events
- Fully dynamic and analytic => very fast
- Accounts for limb darkening too
- Even models previously unconsidered effects, such as ingress/egress asymmetry

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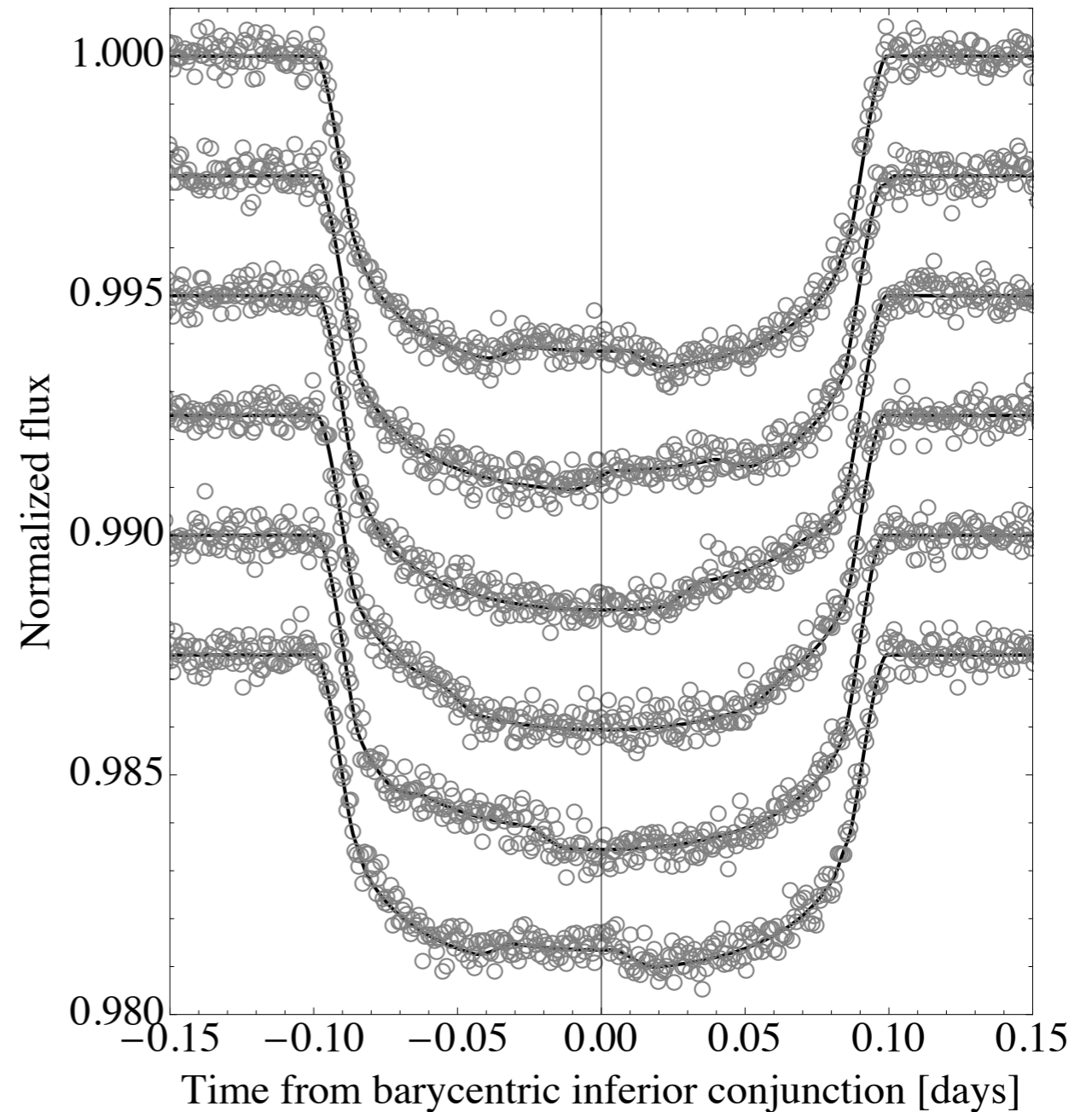
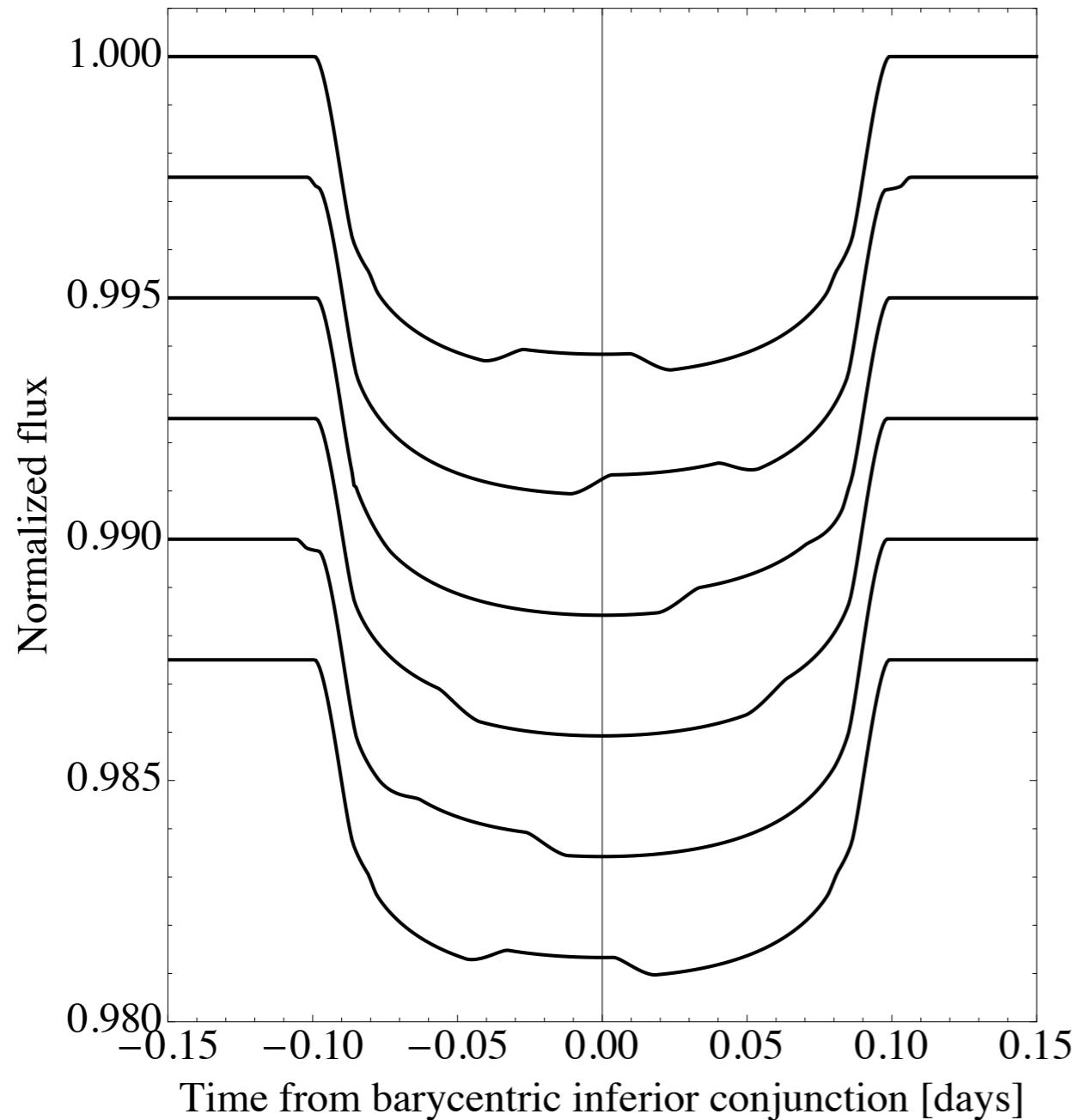
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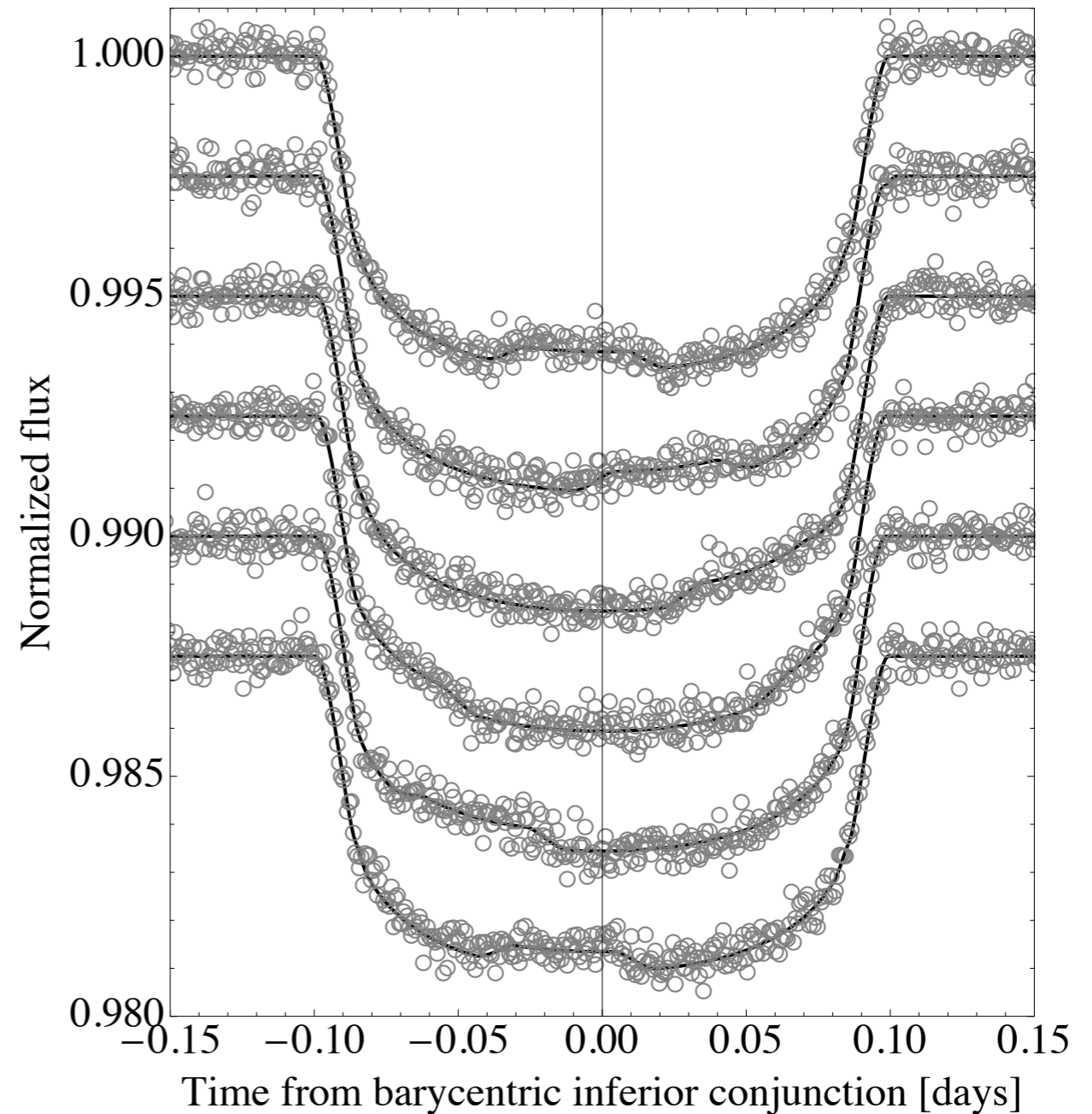
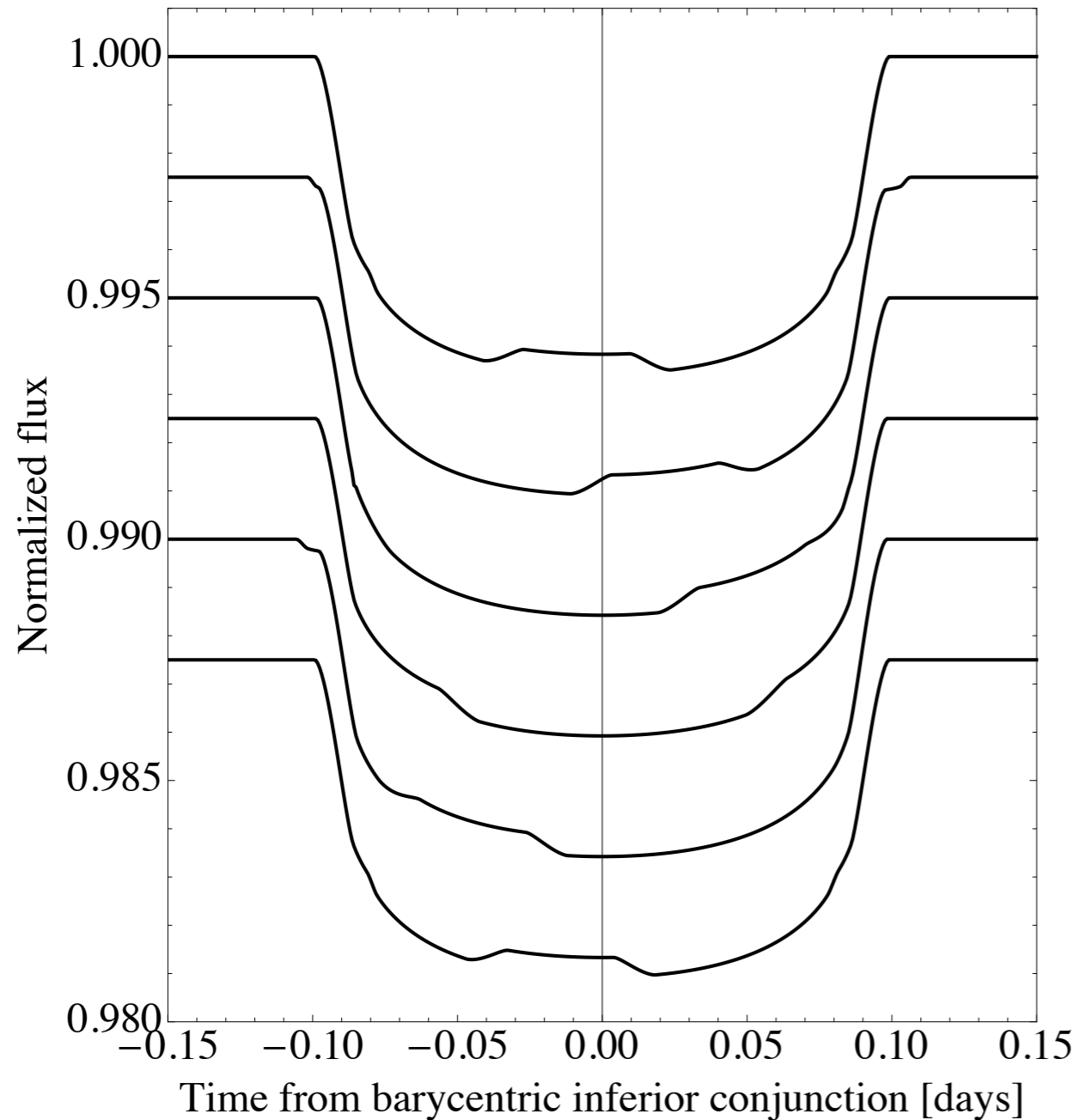
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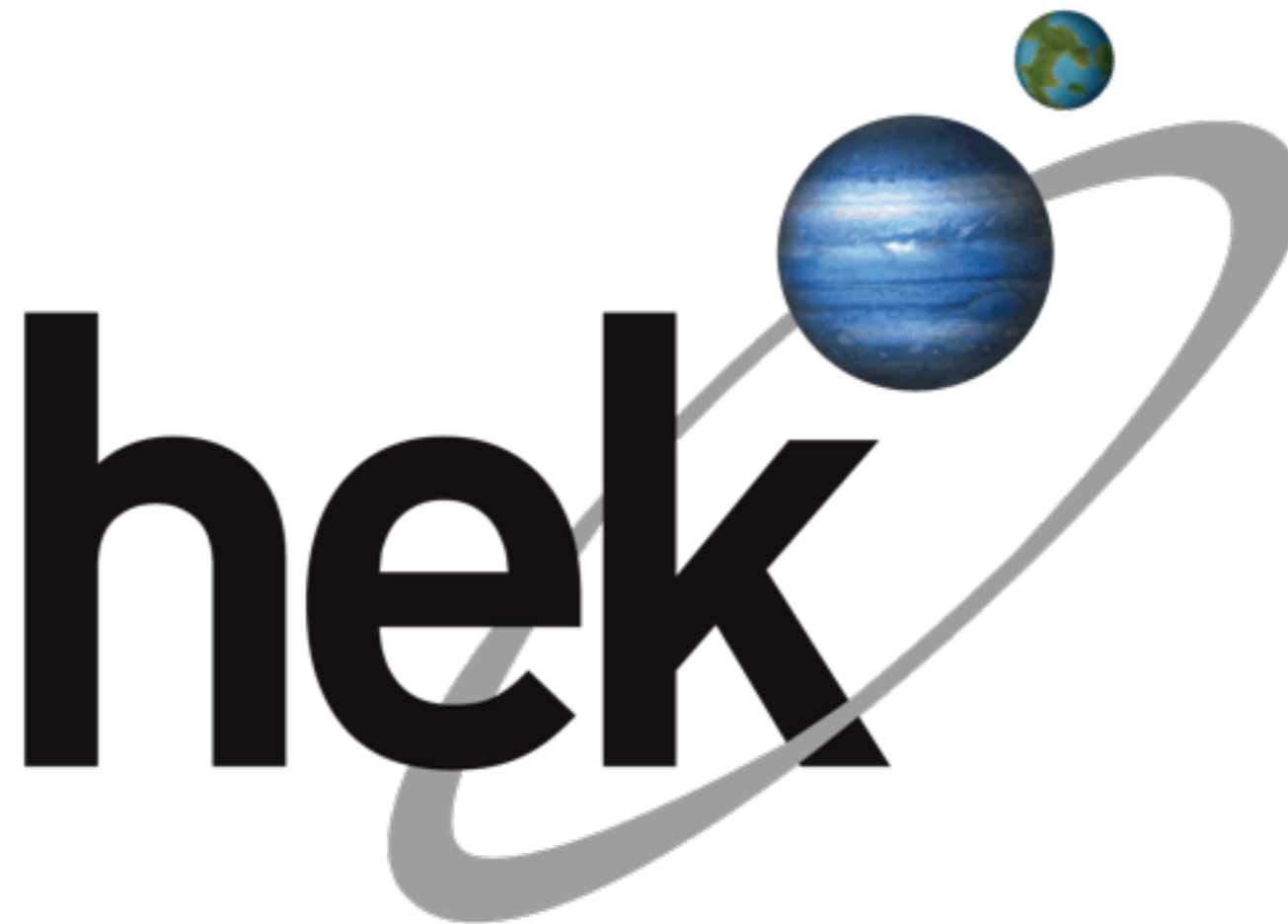


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The Hunt for Exomoons with Kepler

David M. Kipping^{1,2}, Gáspár Á. Bakos³, Lars Buchhave⁴,
David Nesvorný⁵, Joel Hartman³, Allan Schmitt⁶

¹ Harvard-Smithsonian Center for Astrophysics; ² Carl Sagan Fellow; ³ Princeton University;
⁴ Neils Bohr Institute; ⁵ Southwest Research Institute; ⁶ Citizen Scientist

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- Utilizing LUNA to identify exomoons
- Primary goal: detect a transiting exomoon(s)
- Secondary goal: obtain upper limits
- Tertiary goal: determine the frequency of large moons around viable planet hosts, η_{C}

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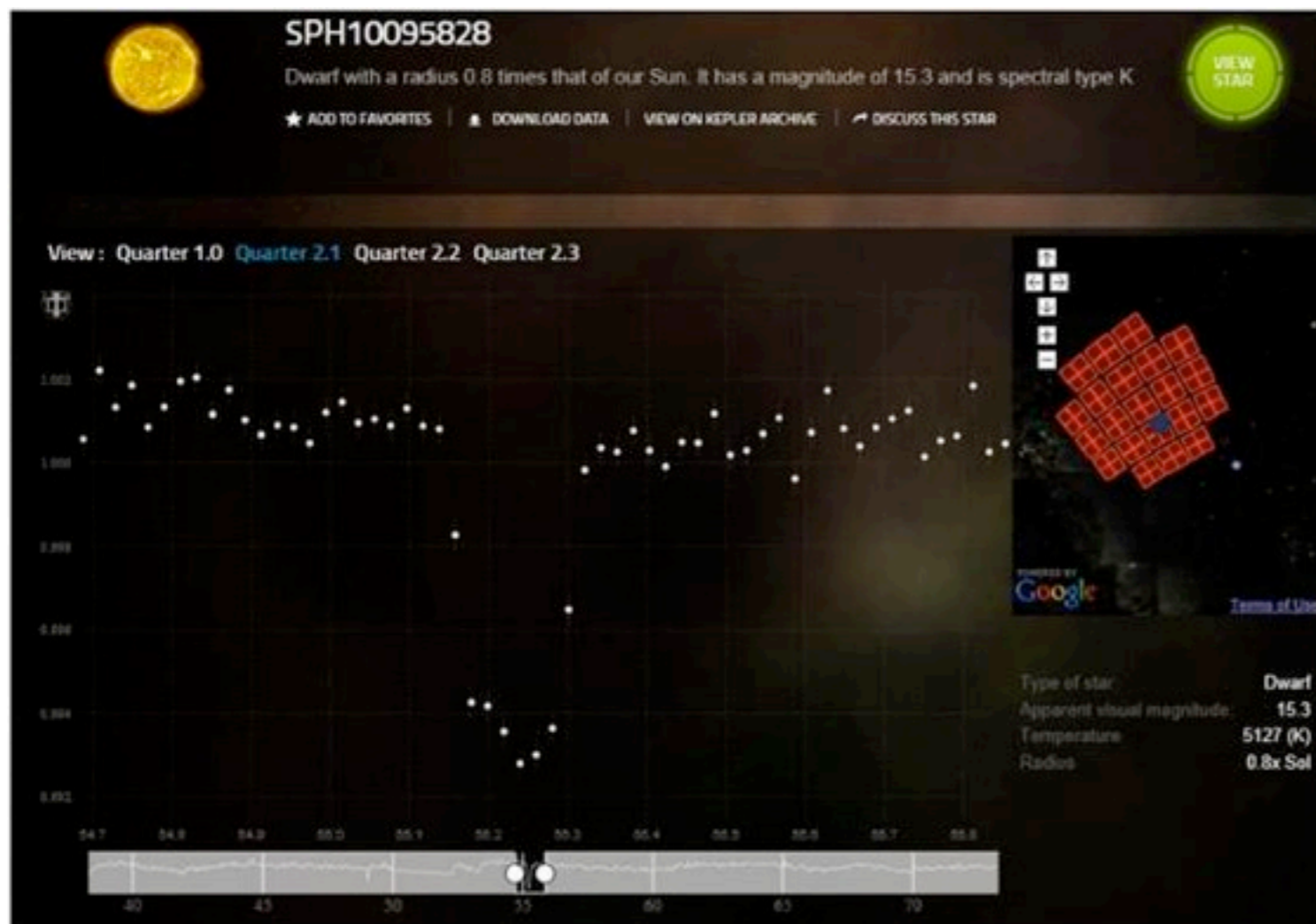
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- Parameter searches are highly multimodal
- Signals expected to be at the limit of Kepler's sensitivity
- Starspots, correlated noise and perturbing planets are challenging sources of false positive
- Bayesian model selection and inference is a must
- No-one has ever done this anything like this before
 - we are forced to invent everything as we go

KOI-872.01

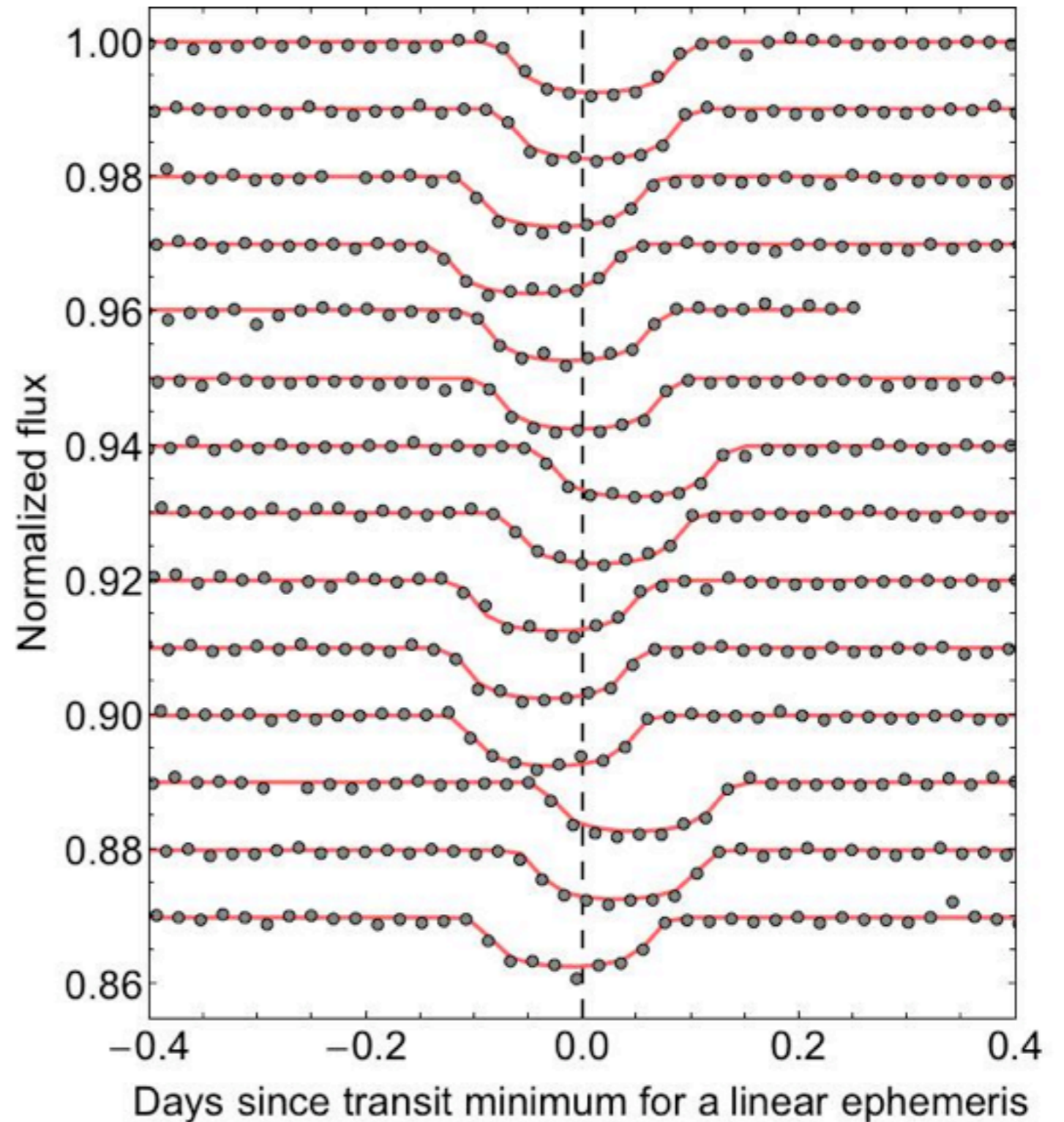
- One of our first TSVs (target selection visual), identified as showing in-transit anomalies:



Identified by Kepler (Borucki et al. 2010) as a candidate planet with $P=33.6$ d, Saturn-sized transit dip. Only one candidate in the system.

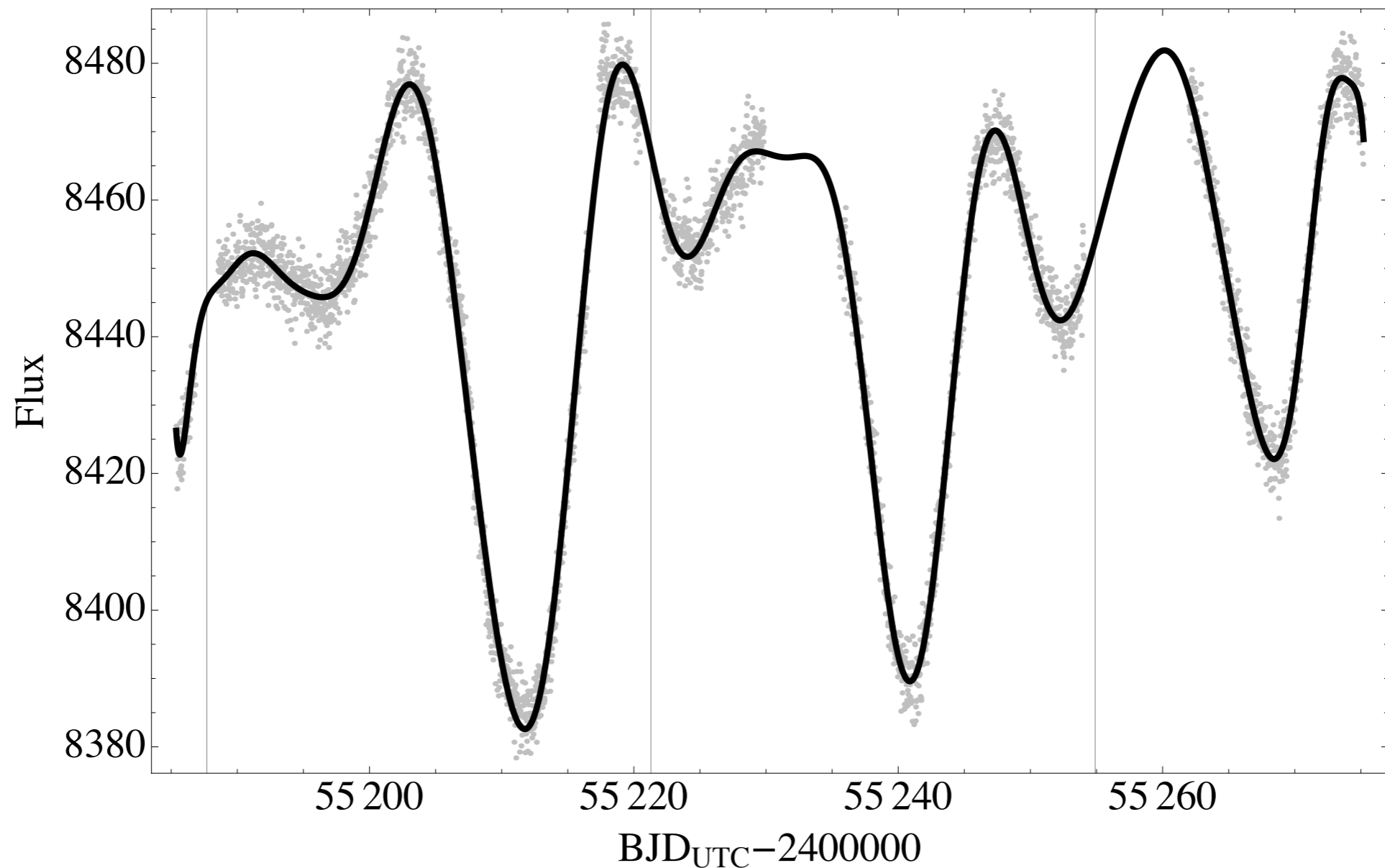
KOI-872.01

- More detailed study showed huge TTVs...



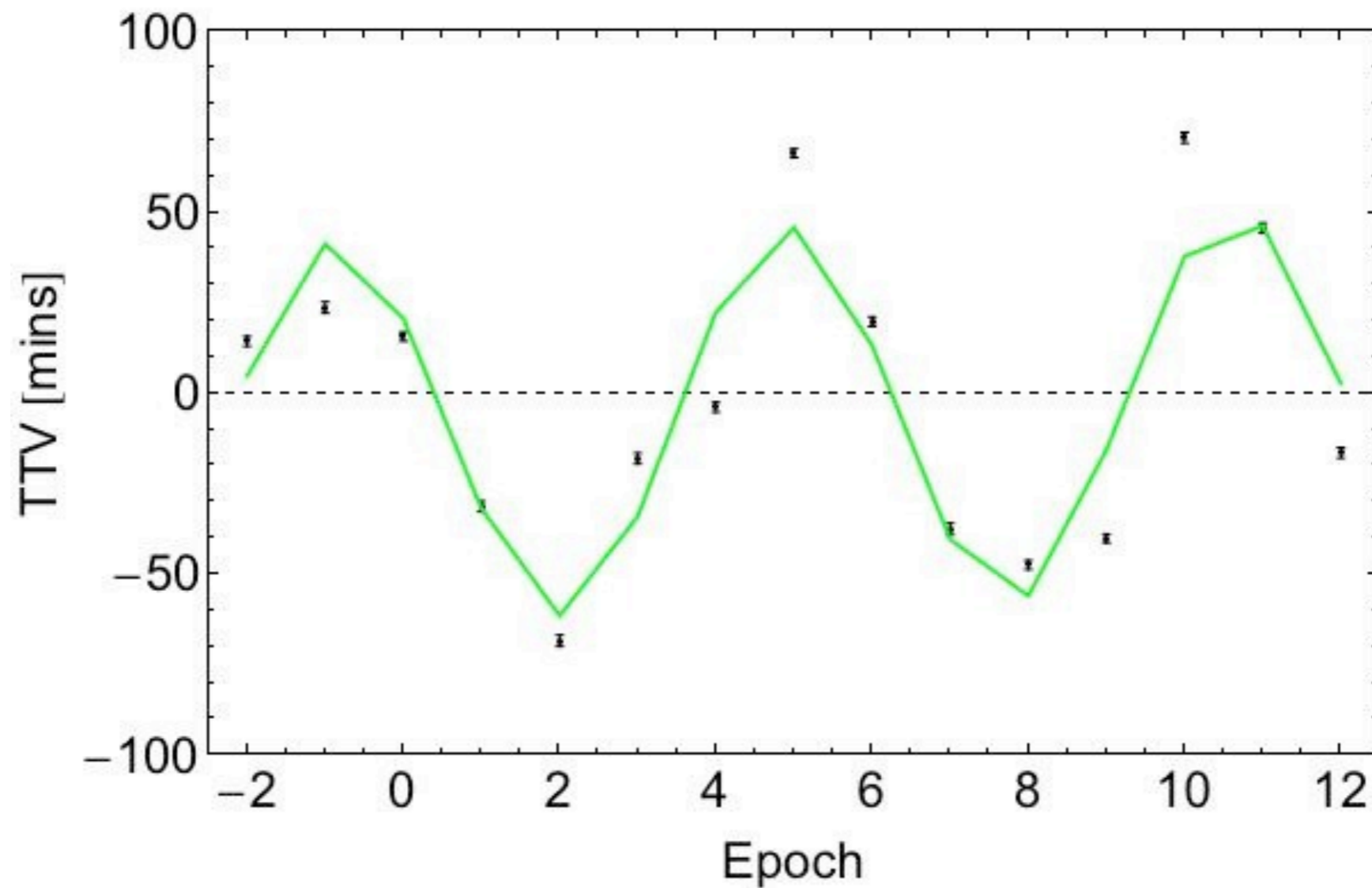
KOI-872.01

- And also stellar activity (like rotating spots)...



KOI-872.01

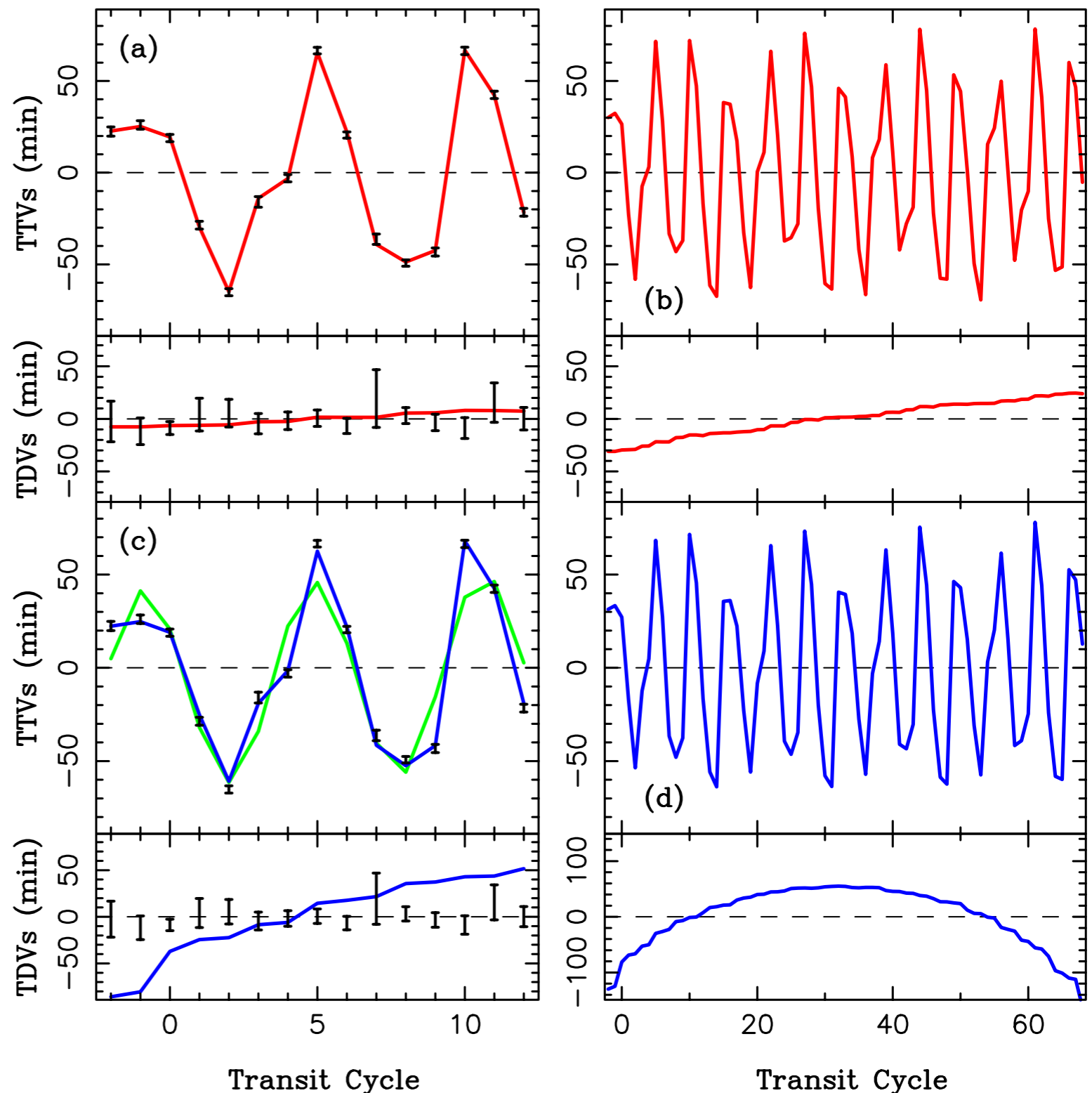
- But the TTVs are not poorly fit by a moon model...



- + lack of TDVs + spots = moon unlikely

KOI-872.01

- So we tried a second planet instead.
- 5:3 resonance (red) works very well.
- 5:2 does OK (blue) but requires an inclined planet => TDVs, which we do not see.



KOI-872.02 = KOI 872c

- The 2nd planet is very precisely pinned down.
- Dynamically measured mass ratio confirms the candidate is a real planet of Saturn-mass.

	KOI-872b	KOI-872c
τ_0 (BJD _{UTC})	2455053.2826 ^{+0.0013} _{-0.0014}	—
P_p (days)	33.60134 ^{+0.00021} _{-0.00020}	57.004 ^{+0.091} _{-0.100}
R_p/R_*	0.0887 ^{+0.0011} _{-0.0012}	—
b_p	0.759 ^{+0.022} _{-0.027}	3.1 ^{+1.1} _{-1.9}
a_p/R_*	44.9 ^{+2.1} _{-1.8}	63.9 ^{+2.9} _{-2.5}
i_p (°)	89.033 ^{+0.076} _{-0.069}	87.25 ^{+1.70} _{-0.95}
a_p (AU)	0.1967 ^{+0.0029} _{-0.0028}	0.2799 ^{+0.0041} _{-0.0040}
e_p	0.01 ^{+0.01} _{-0.01}	0.0145 ^{+0.0035} _{-0.0039}
Ω_p (°)	270	303 ⁺²⁰ ₋₃₄
ϖ_p (°)	—	329.4 ⁺¹¹ _{-9.2}
λ_p (°)	0	338.3 ^{+1.3} _{-1.4}
M_p/M_*	$<6.4 \times 10^{-3}$	$3.97^{+0.17}_{-0.14} \times 10^{-4}$
M_p (M_J)	<6	0.376 ^{+0.023} _{-0.020}
R_p (R_J)	0.812 ^{+0.043} _{-0.043}	—
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T_{eq} (K)	544 ⁺¹⁶ ₋₁₆	456 ⁺¹³ ₋₁₃

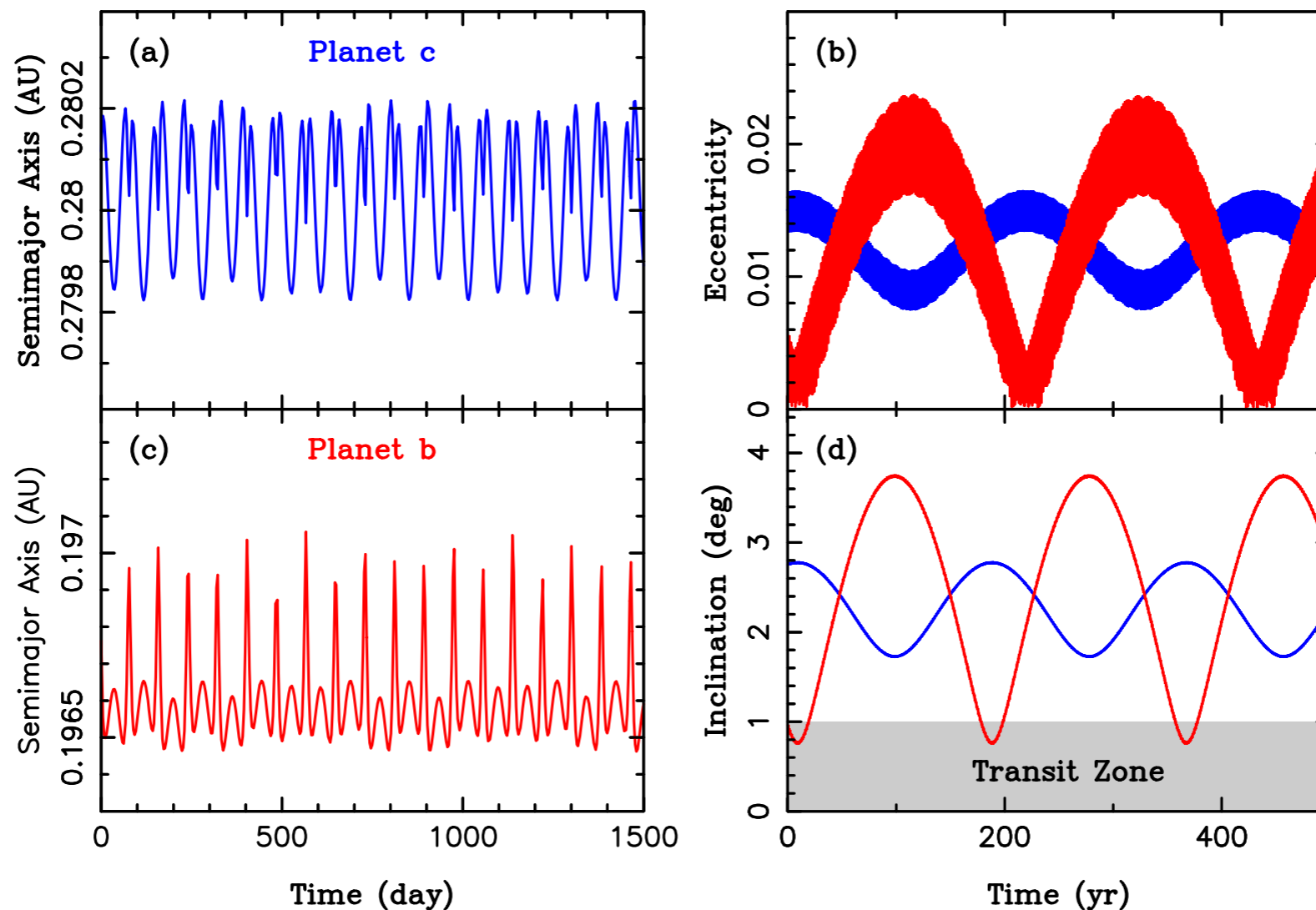
KOI-872.01 = KOI 872b

- TTVs do not allow us to measure mass of inner object (yet).
- However, dynamical stability requires mass is less than 6 Jupiter masses => a real planet

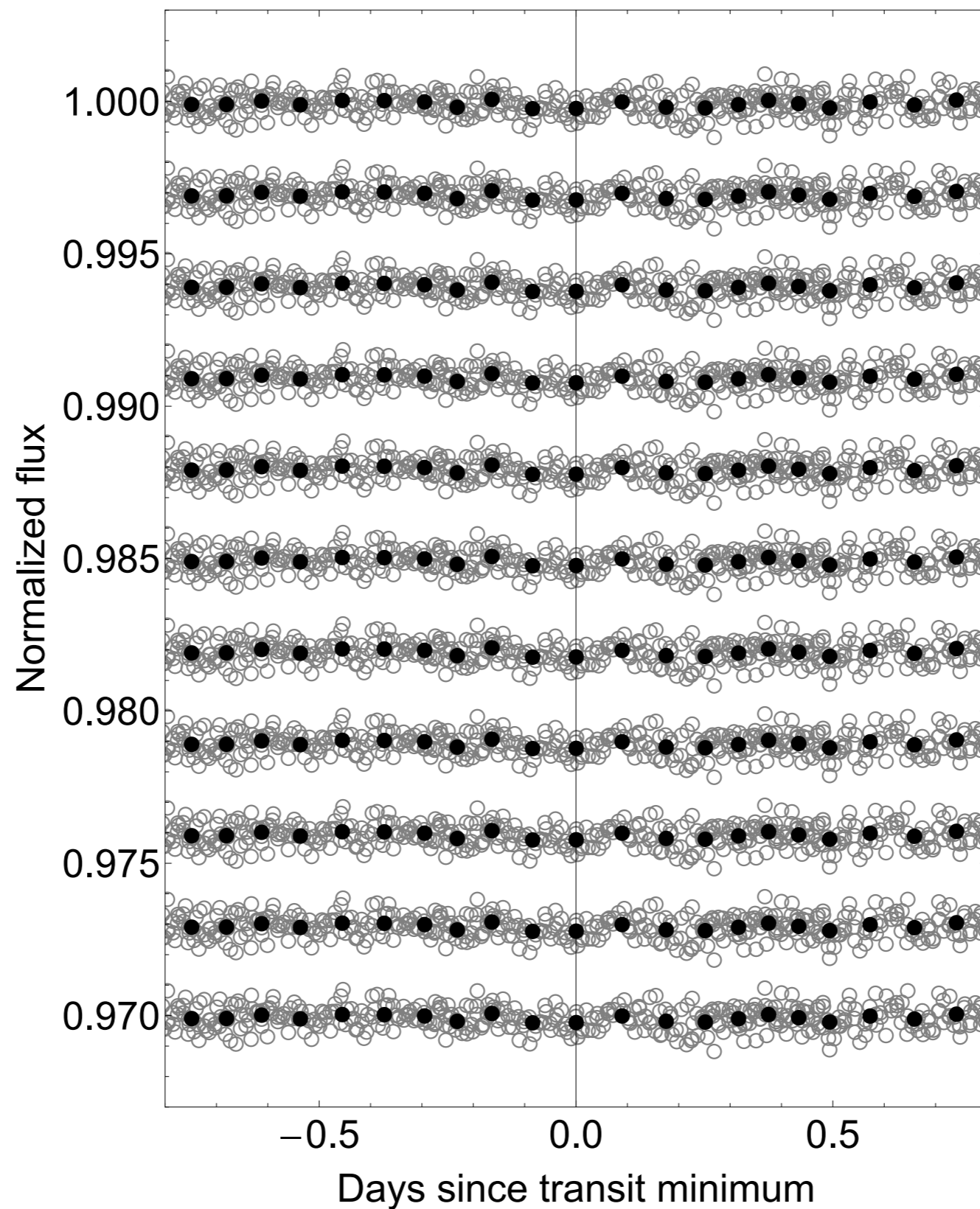
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Does KOI 872c transit?

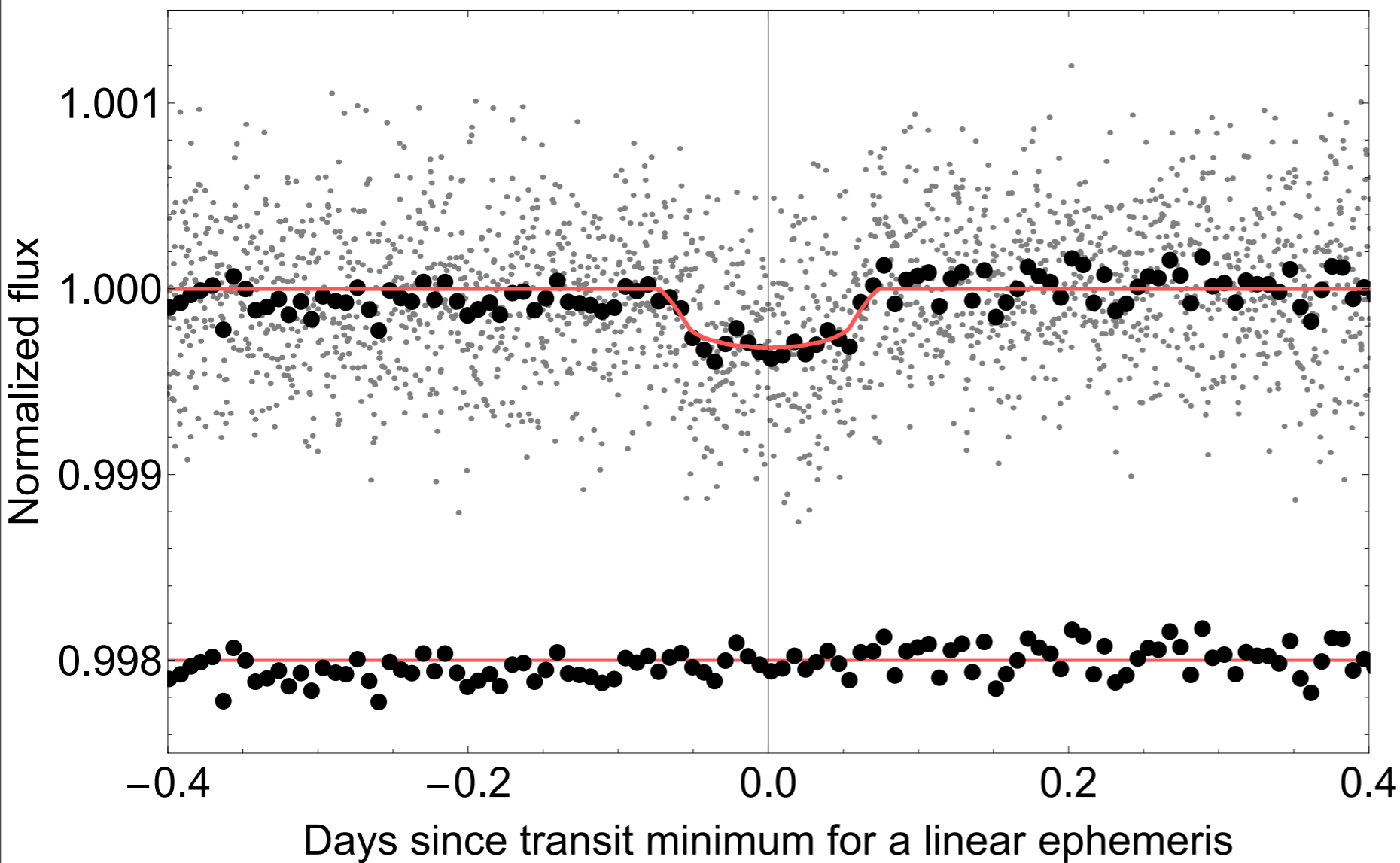
- TTVs imply c does not transit, now or ever.



KOI-872c does not transit

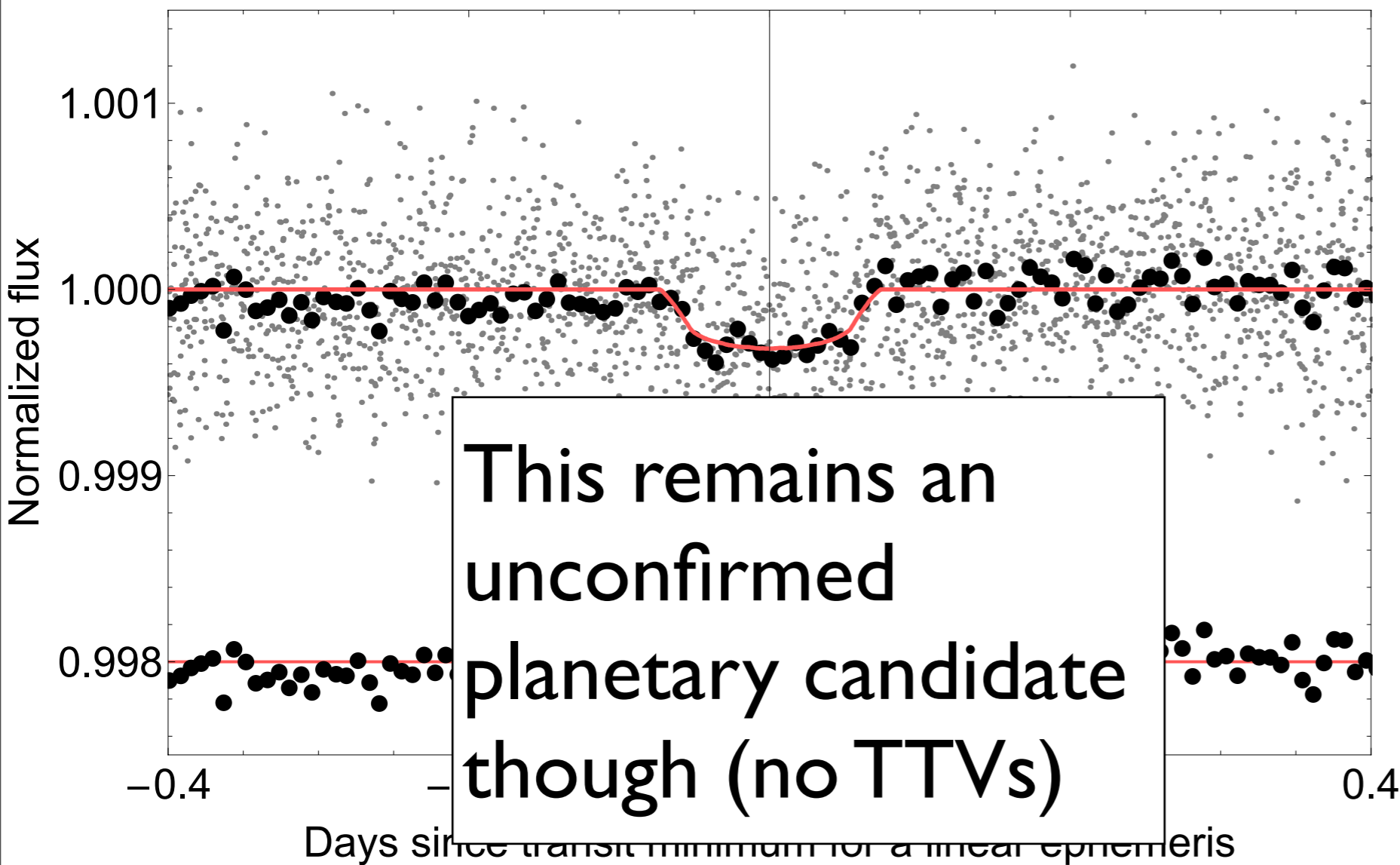


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- A 6.7d periodic signal.
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T_{eq} (K)	544 ⁺¹⁶ ₋₁₆	456 ⁺¹³ ₋₁₃	924 ⁺³⁷ ₋₃₆
M_{moon}/M_p	<0.021	—	—
KOI-872			
ρ_* (kg m^{-3})	1520 ⁺²²⁰ ₋₁₇₀		
M_* (M_\odot)	0.902 ^{+0.040} _{-0.037}		
R_* (R_\odot)	0.940 ^{+0.039} _{-0.040}		
$\log g_*$	4.445 ^{+0.041} _{-0.036}		
T_{eff} (K) (SPC)	5155 ± 105		
L_* (L_\odot)	0.559 ^{+0.0793} _{-0.0699}		
M_V	5.60 ^{+0.17} _{-0.17}		
Age (Gyr)	9.9 ^{+3.5} _{-3.1}		
Distance (pc)	857 ⁺⁶⁹ ₋₆₄		
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L_* (L_\odot)	0.559 ^{+0.0793} _{-0.0699}		
M_V	5.60 ^{+0.17} _{-0.17}		
Age (Gyr)	9.9 ^{+3.5} _{-3.1}		
Distance (pc)	857 ⁺⁶⁹ ₋₆₄		
(M/H) (SPC)	0.41 ± 0.10		

Final Parameters

	KOI-872b	KOI-872c	KOI-872.03
τ_0 (BJD _{UTC})	2455053.2826 ^{+0.0013} _{-0.0014}	—	2455255.2603 ^{+0.0032} _{-0.0031}
P_p (days)	33.60134 ^{+0.00021} _{-0.00020}	57.004 ^{+0.091} _{-0.100}	6.76671 ^{+0.00013} _{-0.00012}
R_p/R_*	0.0887 ^{+0.0011} _{-0.0012}	—	0.01656 ^{+0.00079} _{-0.00082}
b_p	0.759 ^{+0.022} _{-0.027}	3.1 ^{+1.1} _{-1.9}	0.39 ^{+0.19} _{-0.12}
a_p/R_*	44.9 ^{+2.1} _{-1.8}	63.9 ^{+2.9} _{-2.5}	15.58 ^{+0.51} _{-0.49}
i_p (°)	89.033 ^{+0.076} _{-0.069}	87.25 ^{+1.70} _{-0.95}	88.56 ^{+0.48} _{-0.69}
a_p (AU)	0.1967 ^{+0.0029} _{-0.0028}	0.2799 ^{+0.0041} _{-0.0040}	0.0680 ^{+0.0030} _{-0.0030}
e_p	0.01 ^{+0.01} _{-0.01}	0.0145 ^{+0.0035} _{-0.0039}	0 (assumed)
Ω_p (°)	270	303 ⁺²⁰ ₋₃₄	—
ϖ_p (°)	—	329.4 ⁺¹¹ _{-9.2}	—
λ_p (°)	0	338.3 ^{+1.3} _{-1.4}	—
M_p/M_*	$<6.4 \times 10^{-3}$	$3.97^{+0.17}_{-0.14} \times 10^{-4}$	—
M_p (M_J)	<6	0.376 ^{+0.023} _{-0.020}	—
R_p (R_J)	0.812 ^{+0.043} _{-0.043}	—	0.1513 ^{+0.0085} _{-0.0089}
ρ_p (kg m^{-3})	<14000	—	—
T_{eq} (K)	544 ⁺¹⁶ ₋₁₆	456 ⁺¹³ ₋₁₃	924 ⁺³⁷ ₋₃₆
M_{moon}/M_p	<0.021	—	—
KOI-872			
ρ_* (kg m^{-3})	1520 ⁺²²⁰ ₋₁₇₀		
M_* (M_\odot)	0.902 ^{+0.040} _{-0.037}		
R_* (R_\odot)	0.940 ^{+0.039} _{-0.040}		
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No moon, but when
life gives you lemons,
make lemonade!



hek

The Hunt for Exomoons
with Kepler

Questions?