

PTFO 8-8695b: A Transiting T-Tauri Planet Candidate

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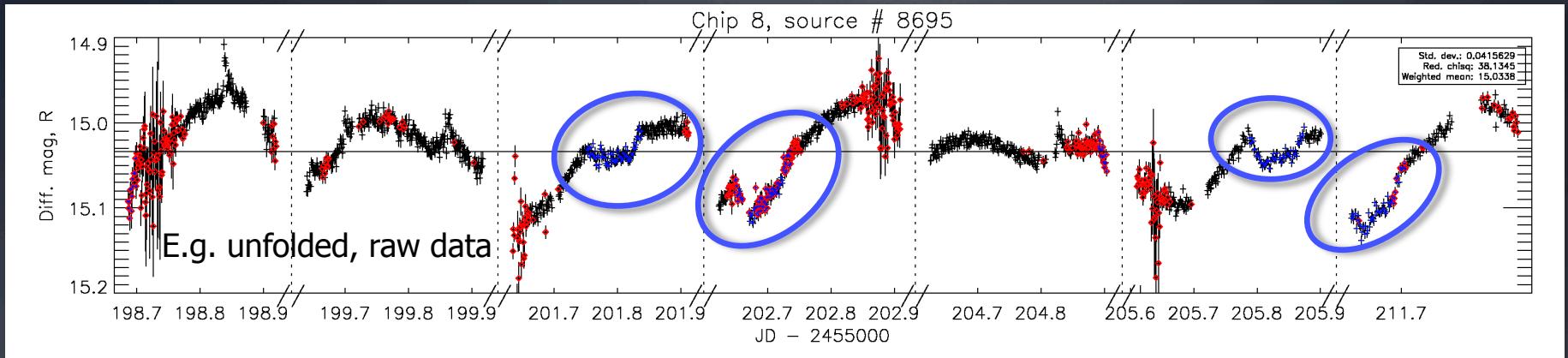
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PTF Orion

- Young transit search using Palomar Transient Factory
- Centered on 7-10Myr old 25-Ori region
- R-band, ~8os cadence, diff. photometry
- Observed ~17 clear nights Dec '09, 7 clear nights Dec '10
- ~7,000 exposures, ~110,000 light curves total

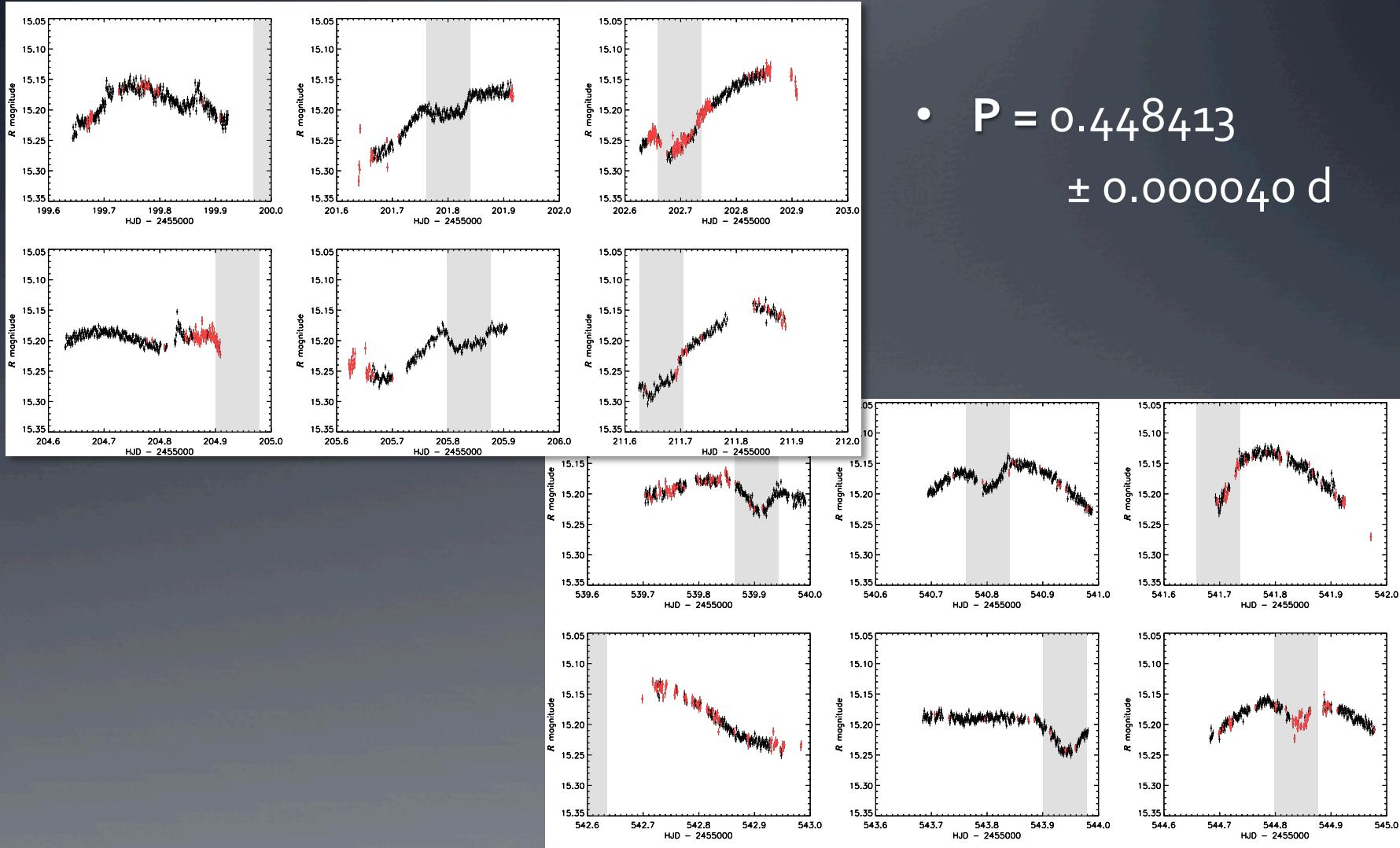


PTFO 8-8695



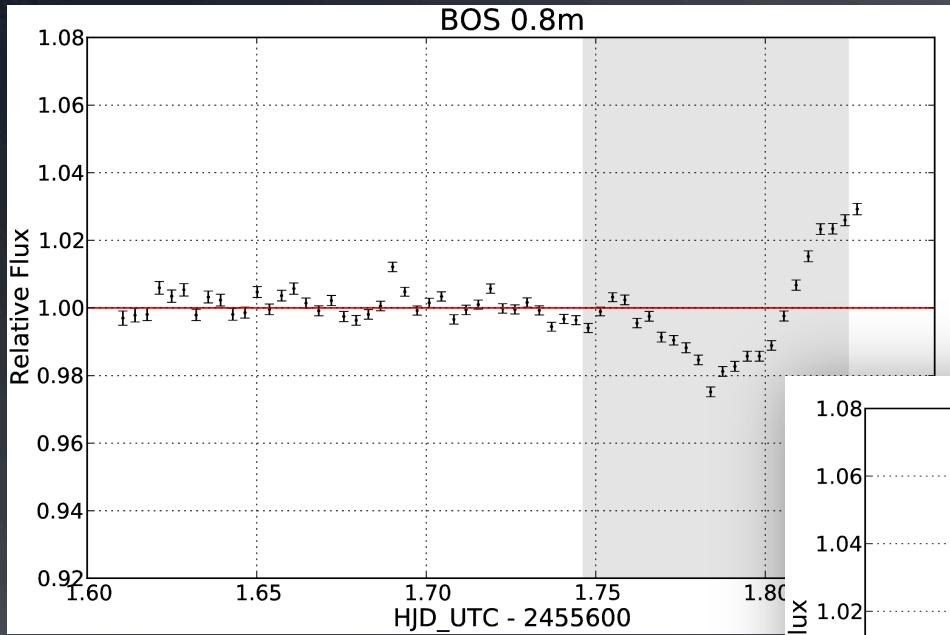
- Known WTTS (Briceno et al., AJ, 2005)
- $V = 16.3$ mag, median $R = 15.2$ mag
- M_3 , ~ 2.7 Myr old
- $\sim 0.4 M_\odot$, $1.39 R_\odot$

Regular transit window

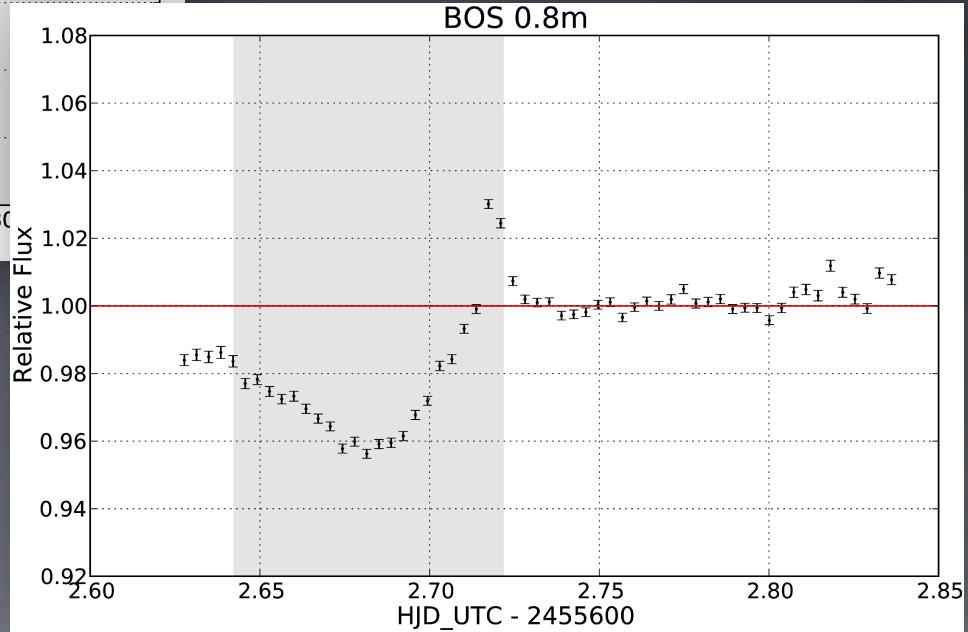


- $P = 0.448413 \pm 0.000040 \text{ d}$

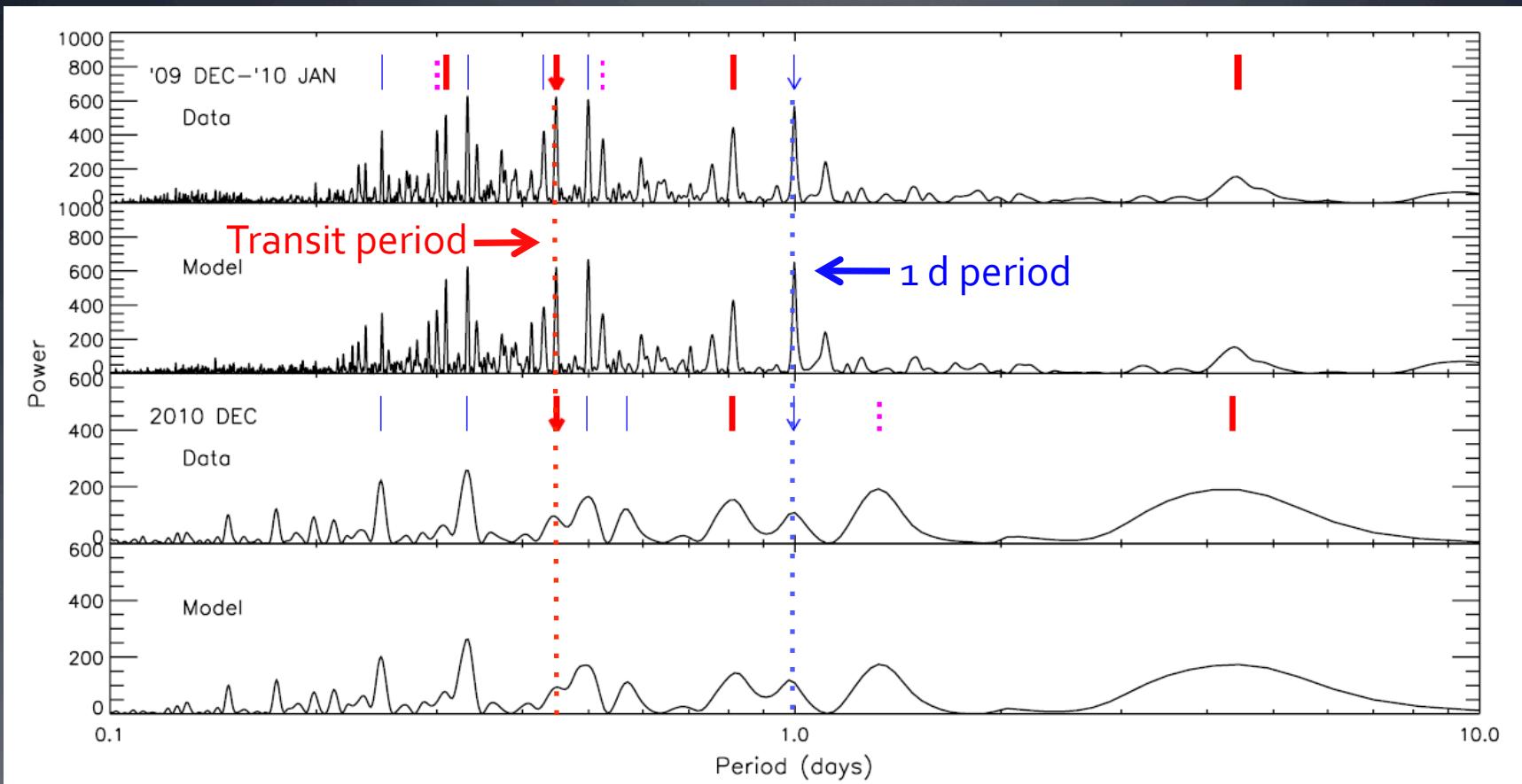
LCOGT follow-up



Clear filter, Byrne
Observatory 0.8m

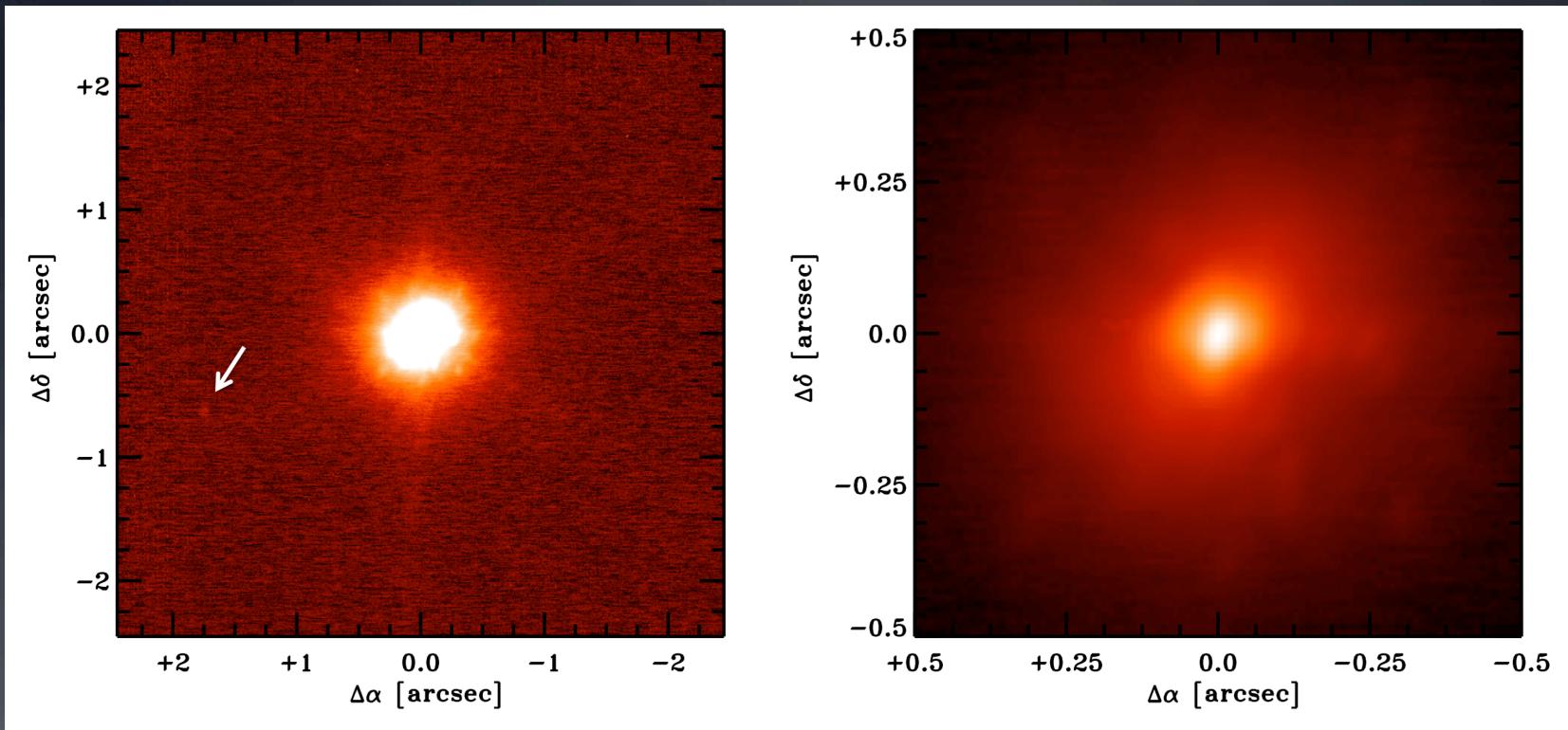


Periodogram analysis



- All other peaks are aliases
- Effect of transit in models is small + no other periodicities evident
- => **Star/orbit co-rotation – prob. not background binary**

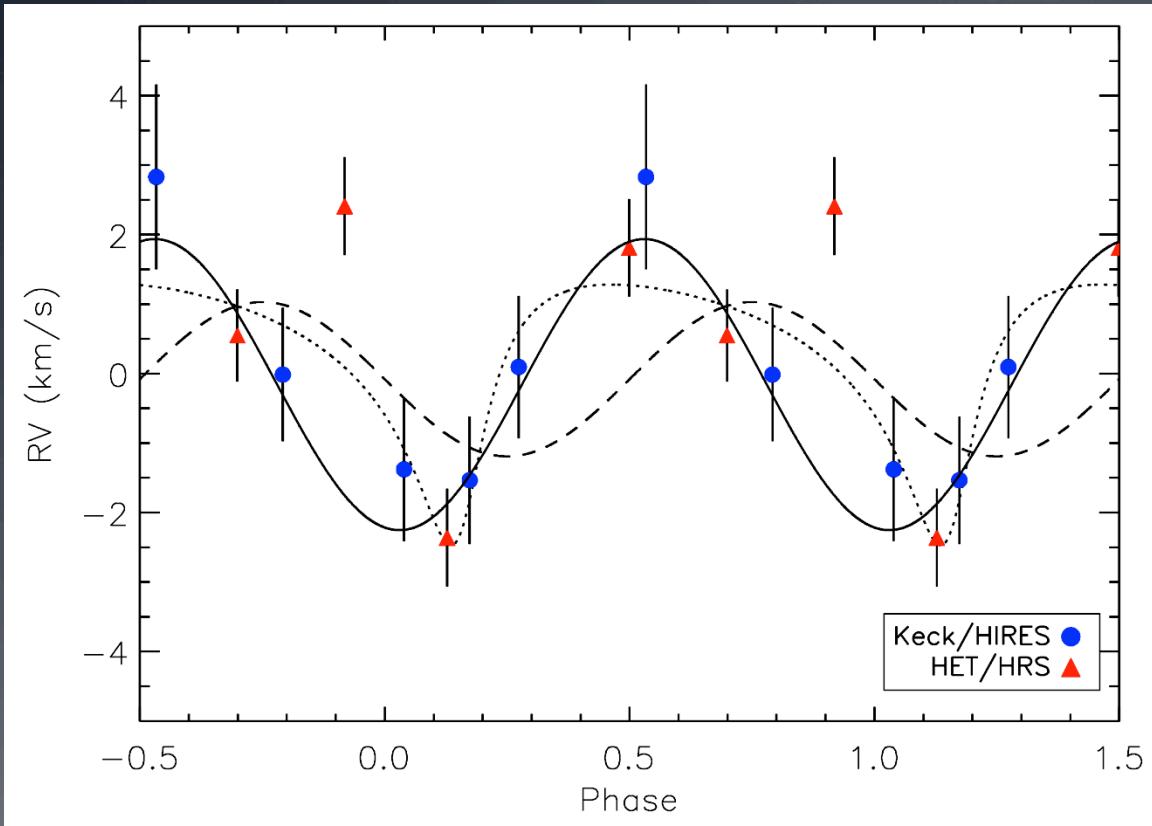
Keck AO follow-up



- Single faint additional source (left)
- $1.8''$ separation, 6.96mag fainter
- *Source cannot account for transits*

Keck+HET RV follow-up

RV Folded on Transit Period



- Signal probably dominated by star spots since out of phase
- Upper limit, $M_p \sin i \leq 4.8 \pm 1.2 M_{\text{Jup}}$

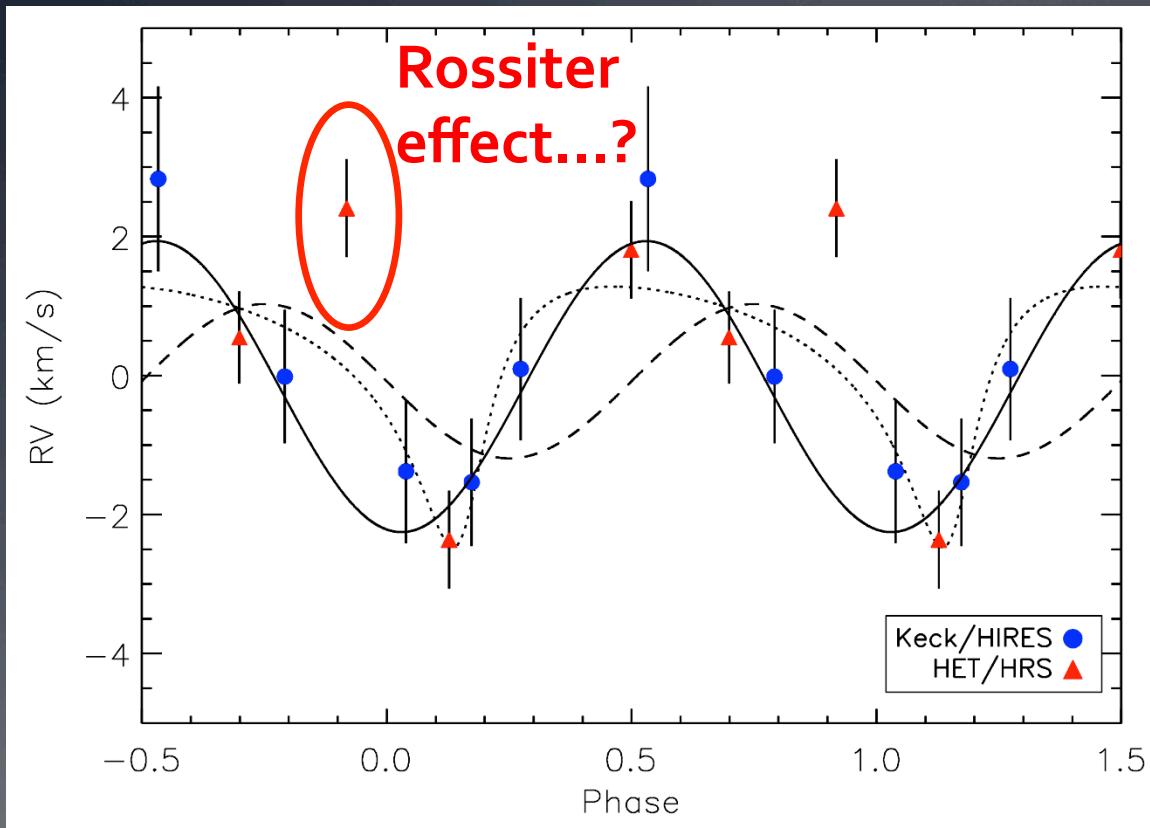
Circular fit, fixed to transit phase

Eccentric fit, fixed to transit phase – *maybe unlikely?*

Sinusoidal (circular) fit, floating phase

Keck+HET RV follow-up

RV Folded on Transit Period



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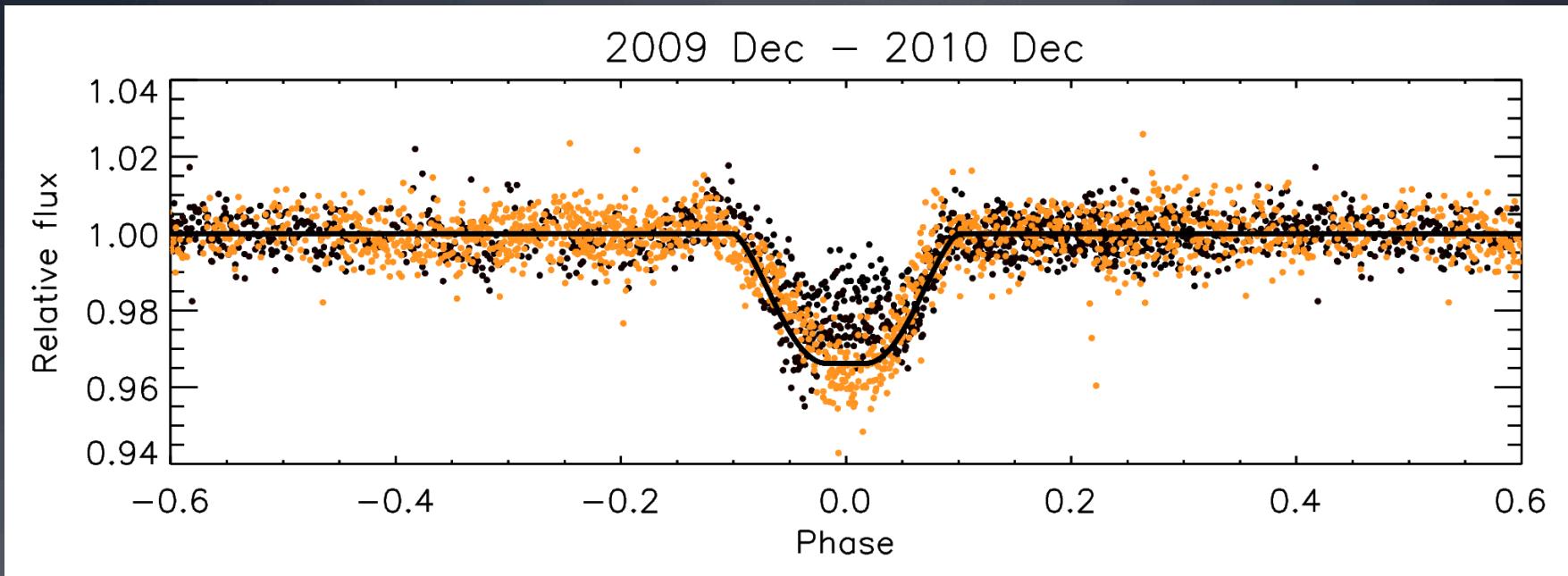
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Sinusoidal (circular) fit, floating phase

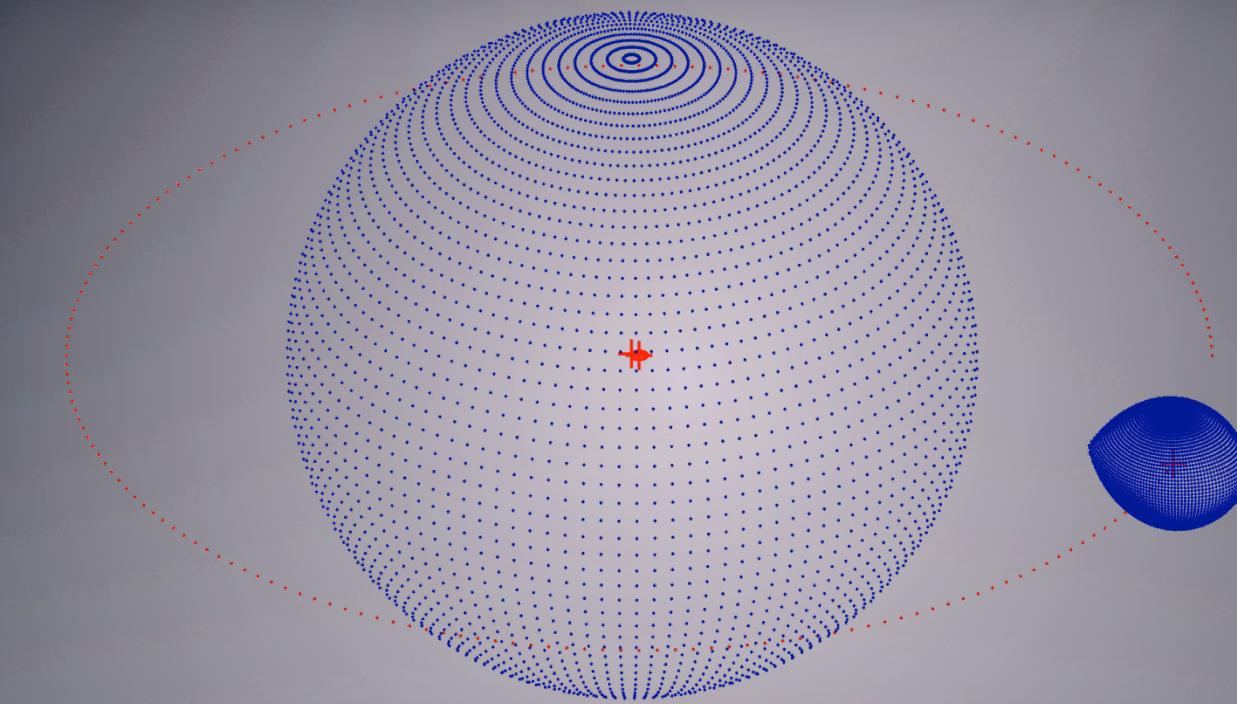
Transit fitting

Detrended and folded light curve



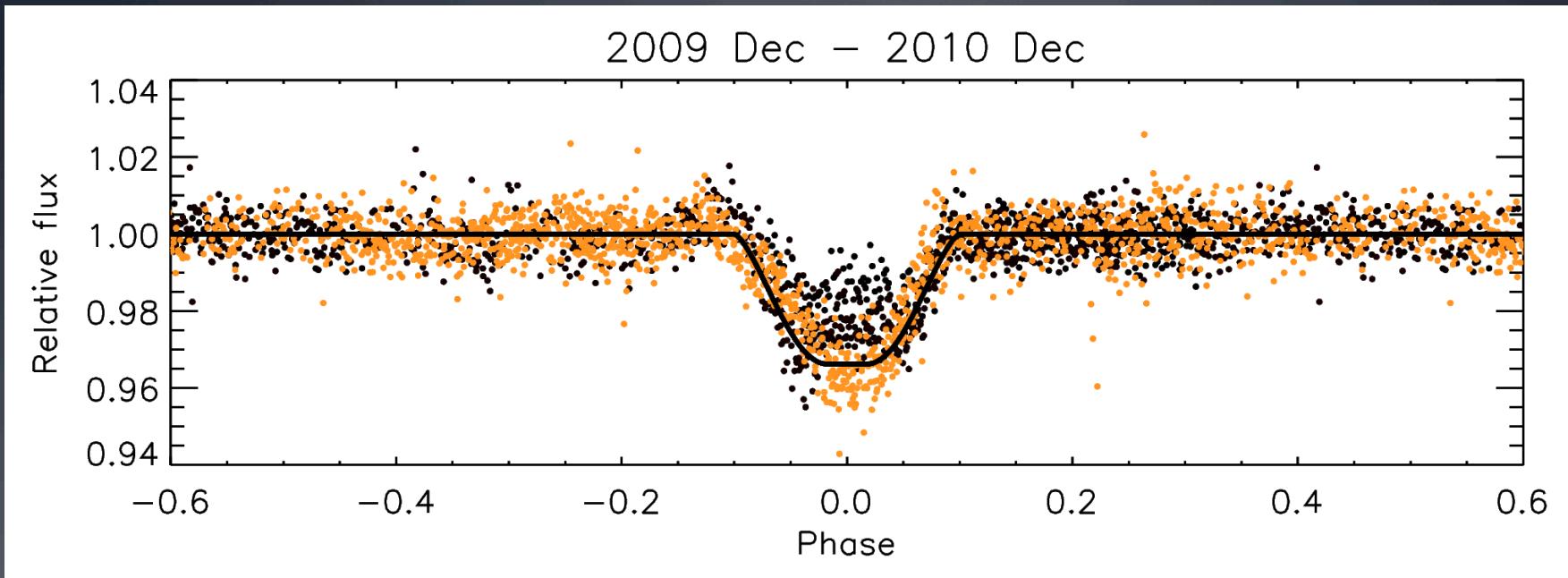
- **Transit shape is v. difficult to model with star spots**

Vital Stats



- 2.7Myr M3 WTTS
- $P \approx 0.4484\text{d}$
- $R_p \approx 1.9R_{\text{Jup}}$, $M_p \leq 5.5M_{\text{Jup}}$
- $i \approx 62^\circ$
- Co-rotating star
- V. close to Roche limit – *may be actively losing mass.*

A puzzle – transit shape change?

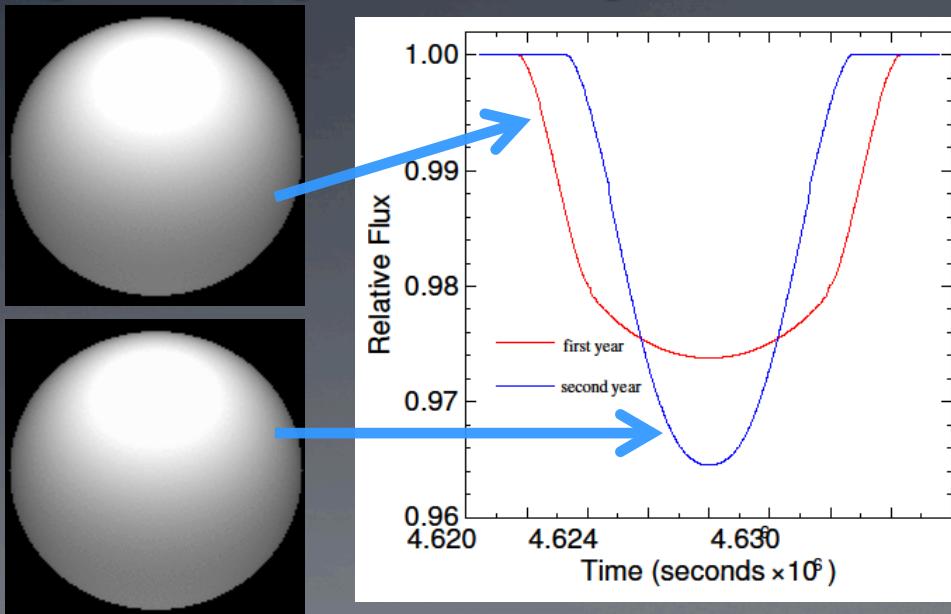


- Black = 1st yr, orange=2nd
- Change in orbital geometry...?

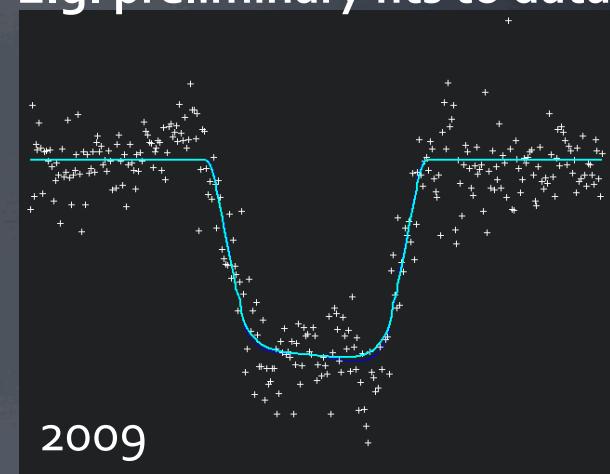
Precession + Gravitational Darkening...?

- Star is fast rotator => *oblate*
- Significant grav. darkening
- If orbit oblique, precession expected on timescale ~100 d – about right...

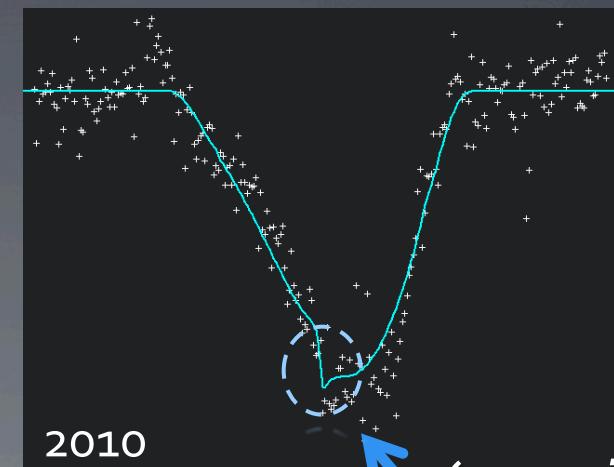
E.g. effect of grav. darkening w/ fast rotator



E.g. preliminary fits to data



2009



2010

(numerical artifact)

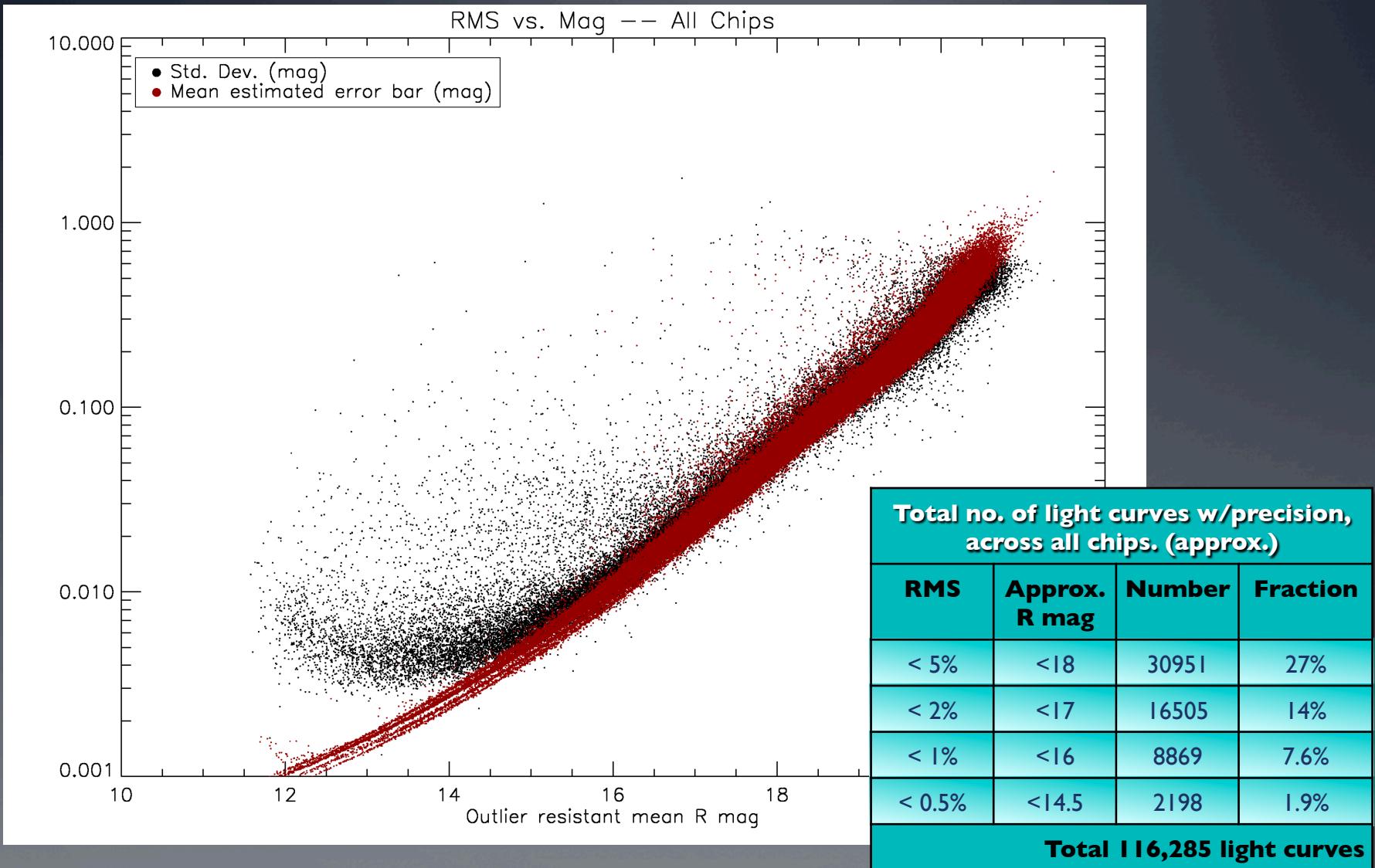
Courtesy Jason Barnes, U. Idaho

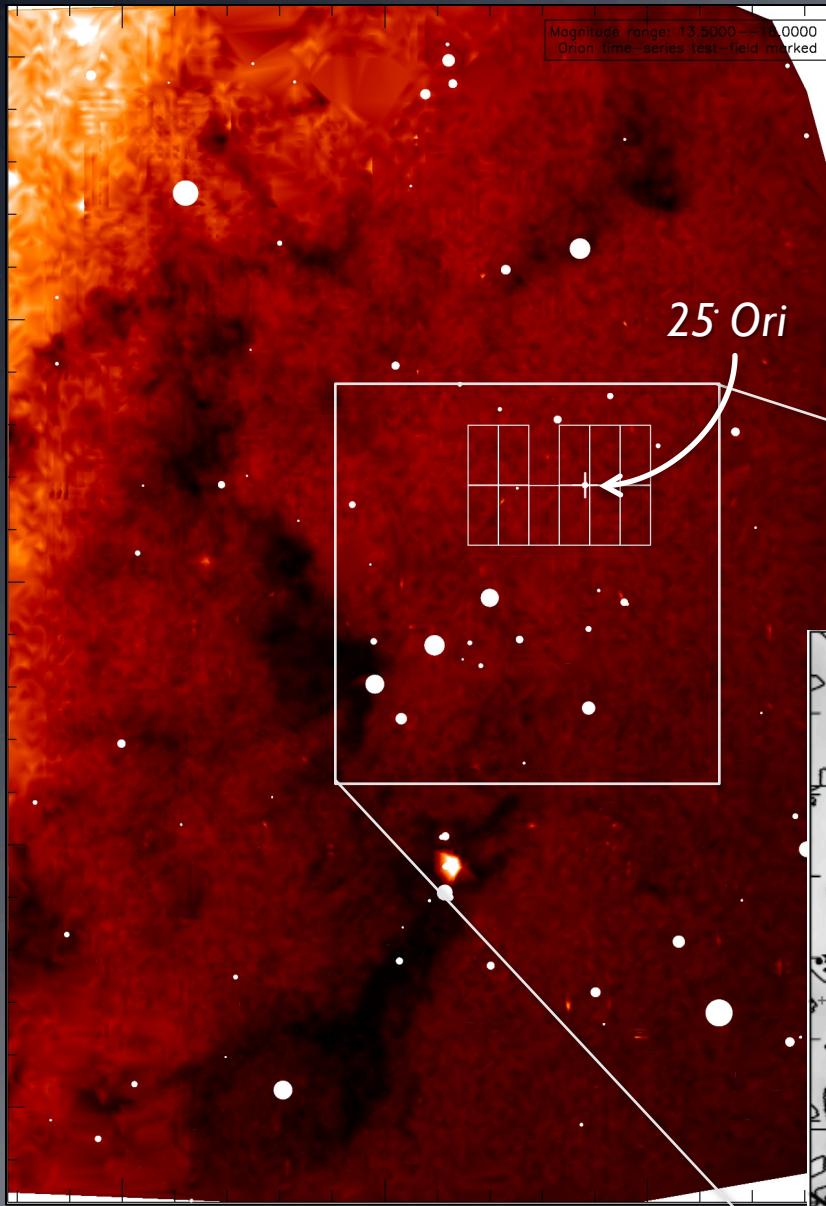
Summary

- PTFO 8-8695b – a candidate T-Tauri planet
 - ~2.7Myr-old
 - $P \approx 0.4484\text{d}$, $R_p \approx 1.9R_{\text{Jup}}$, $M_p \leq 5.5M_{\text{Jup}}$
 - AO -> no significant blends
 - Co-rotating => background binary unlikely
 - VERY close to Roche limit – **poss. losing mass/evaporating!**
 - Change in transit shape in ~1yr? Why...?
 - More follow-up multi-band photometry may help mitigate star-spots
 - Follow-up RV → catch Rossiter McLaughlin effect?
 - Investigating transit models including precession/gravitational darkening.
 - **Potentially a unique object for constraining formation theories**

Backup slides

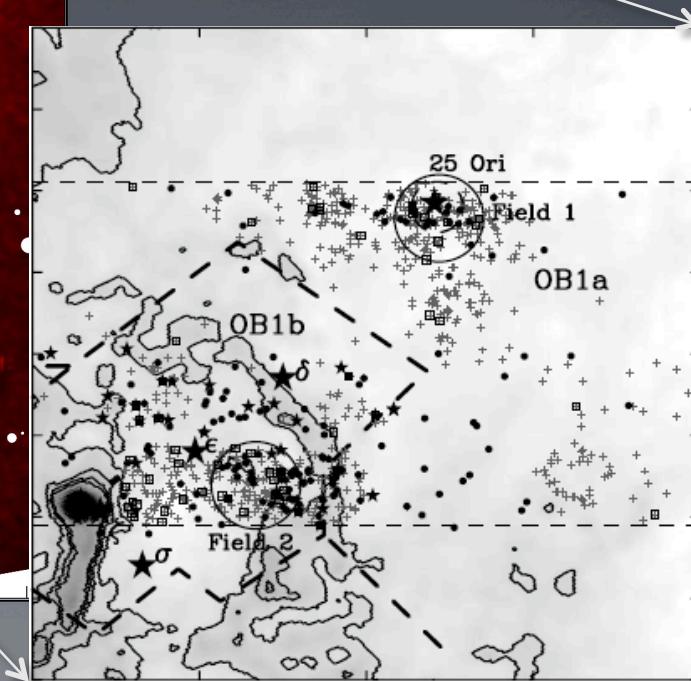
Results - rms precision





PTF Orion

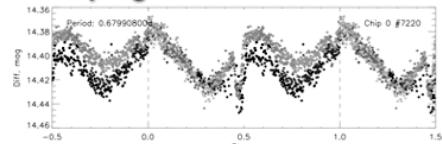
- Primarily search for young planets
- Field centred on 25 Ori
- ~7-10Myr old region
- Known grouping of TTS



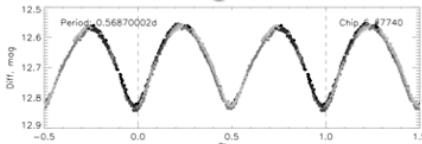
Briceño et al., ApJ 2007

E.g. periodic variables (rms > 3x error bar)

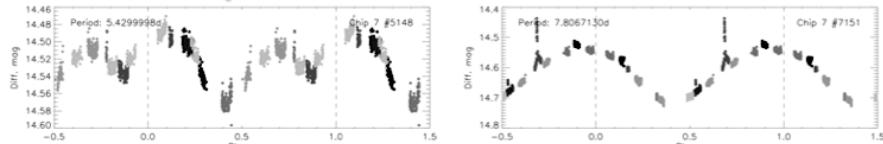
Eclipsing binaries



Sinusoidal/ambiguous



Spotted/TTS/unidentified, etc.



Total variable light curve types from visual inspection

Type	No. found
Binary	82
Pulsating	18
Clear flaring	44
Sinusoidal, ambiguous	68
Periodic, non-sinusoidal, (spots/unidentified/etc.)	95
Irregular	31
Currently unclassified	~200
Total	~530

Pulsating

