

Modeling Spectro-astrometric Observations of CO in Circumstellar Disks

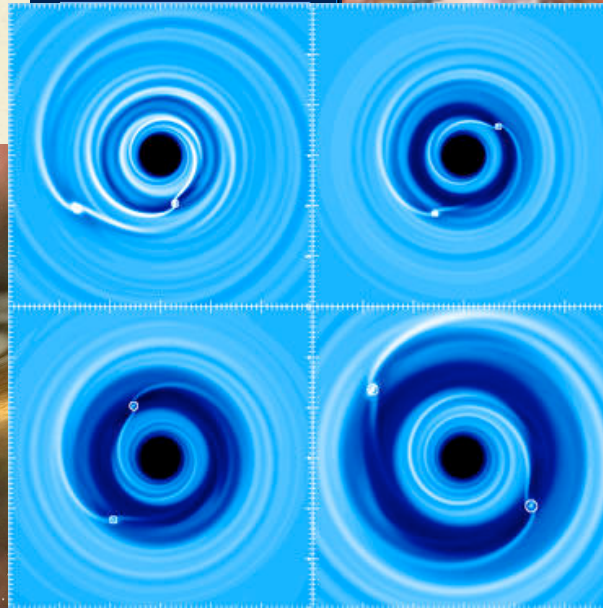
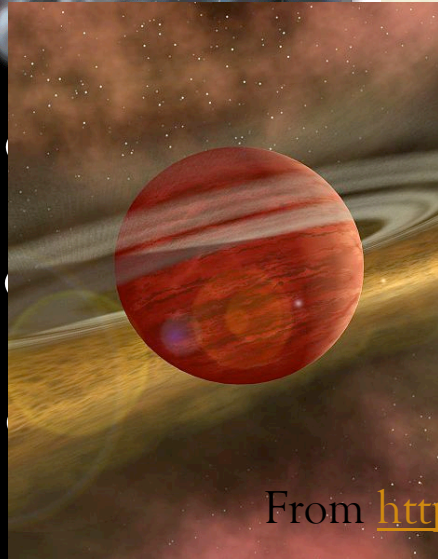
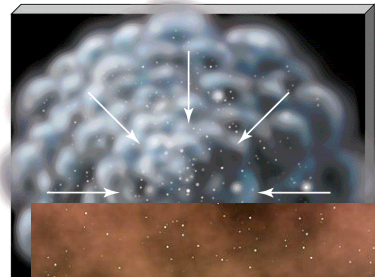
Matt Troutman - Clemson University
Sagan/Michelson Fellows Symposium
November 12, 2009



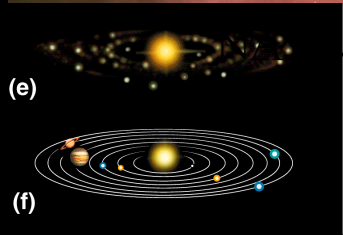
Planet Formation

TWO PLANET FORMATION SCENARIOS

Accretion model



From <http://www.star.qmw.ac.uk/~masset/intertmr.html>



JPL-C

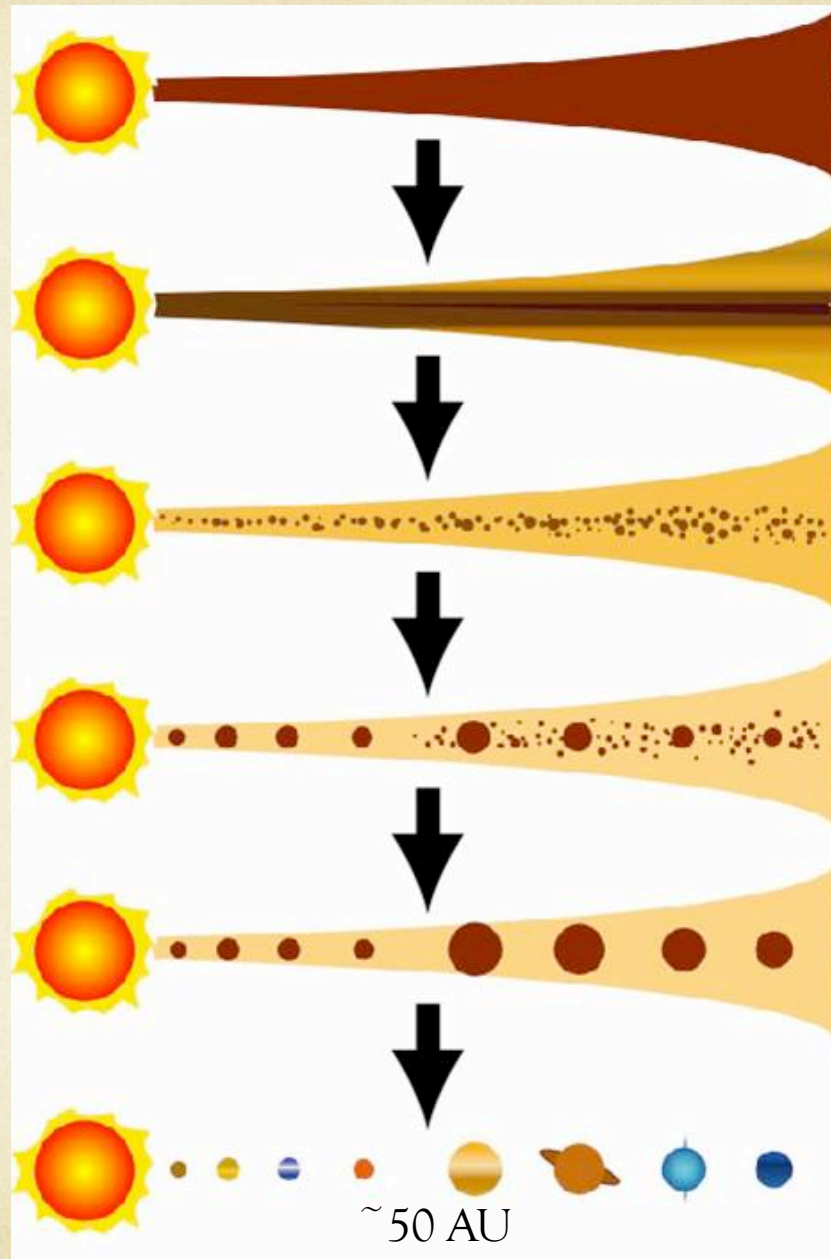
Gas-giant planets scatter or accrete the remaining planetesimals and embryos.

NASA/ESA and A.



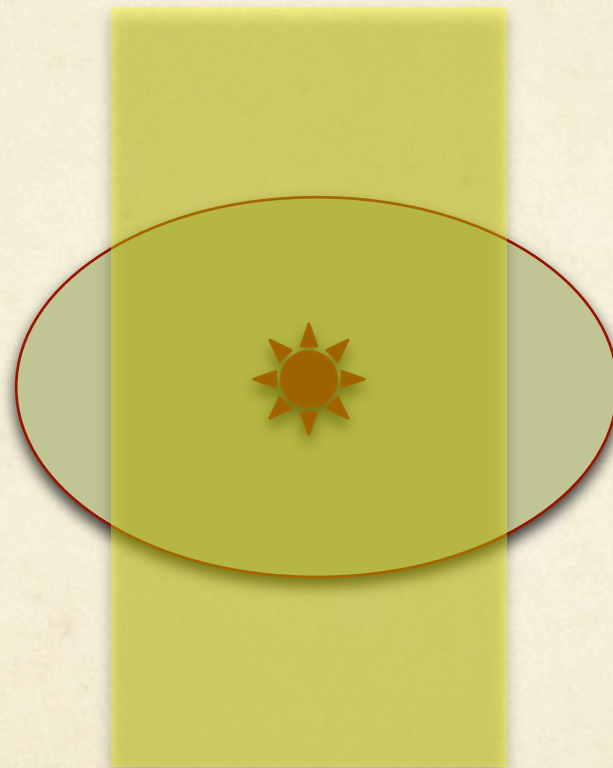
From lpl.arizona.edu

From astro.psu.edu



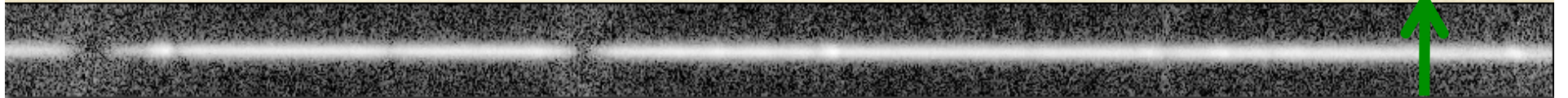
From planet.sci.kobe-u.ac.jp

Observing

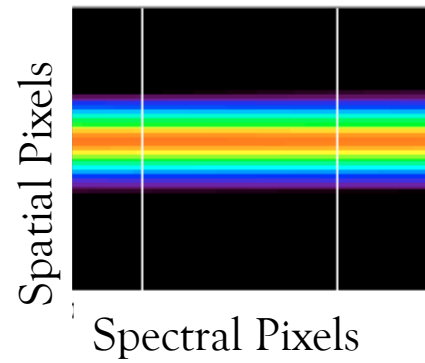
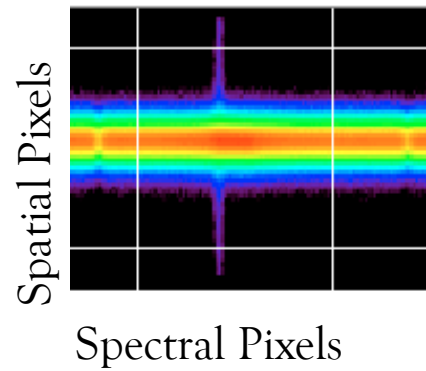


Wavelength →

Spatial ↑

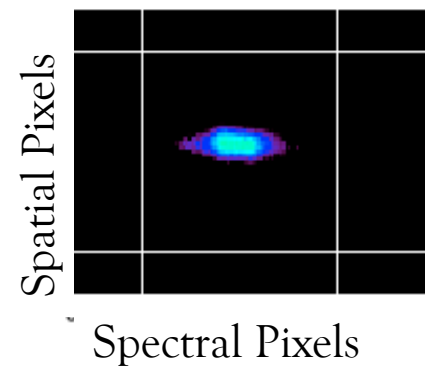
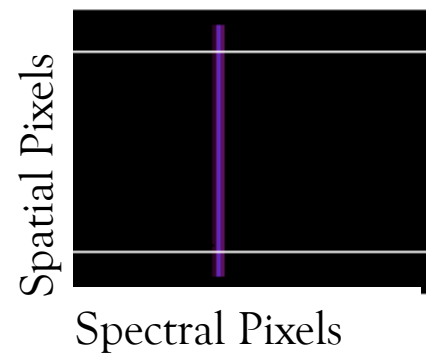


Raw Data →



← Continuum

Sky Line →

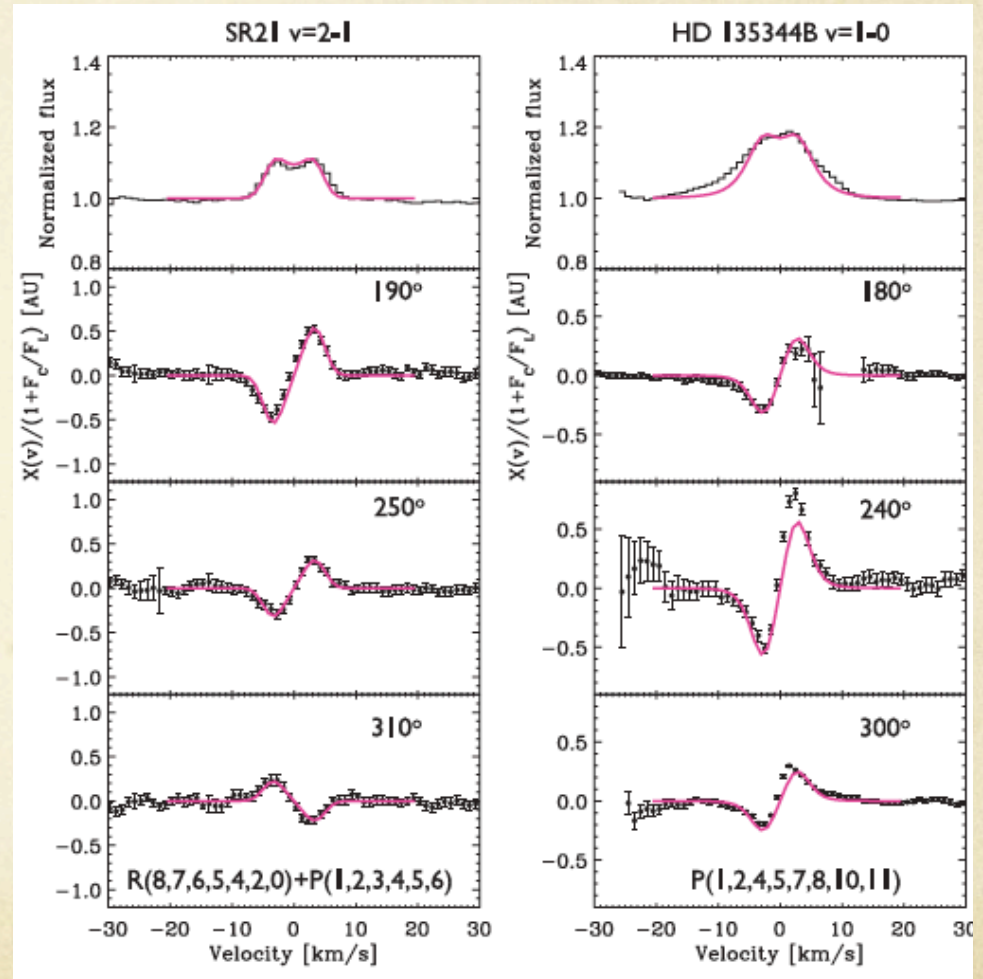


← Signal

Acke & van den Ancker (2006)

Spectro-astrometry

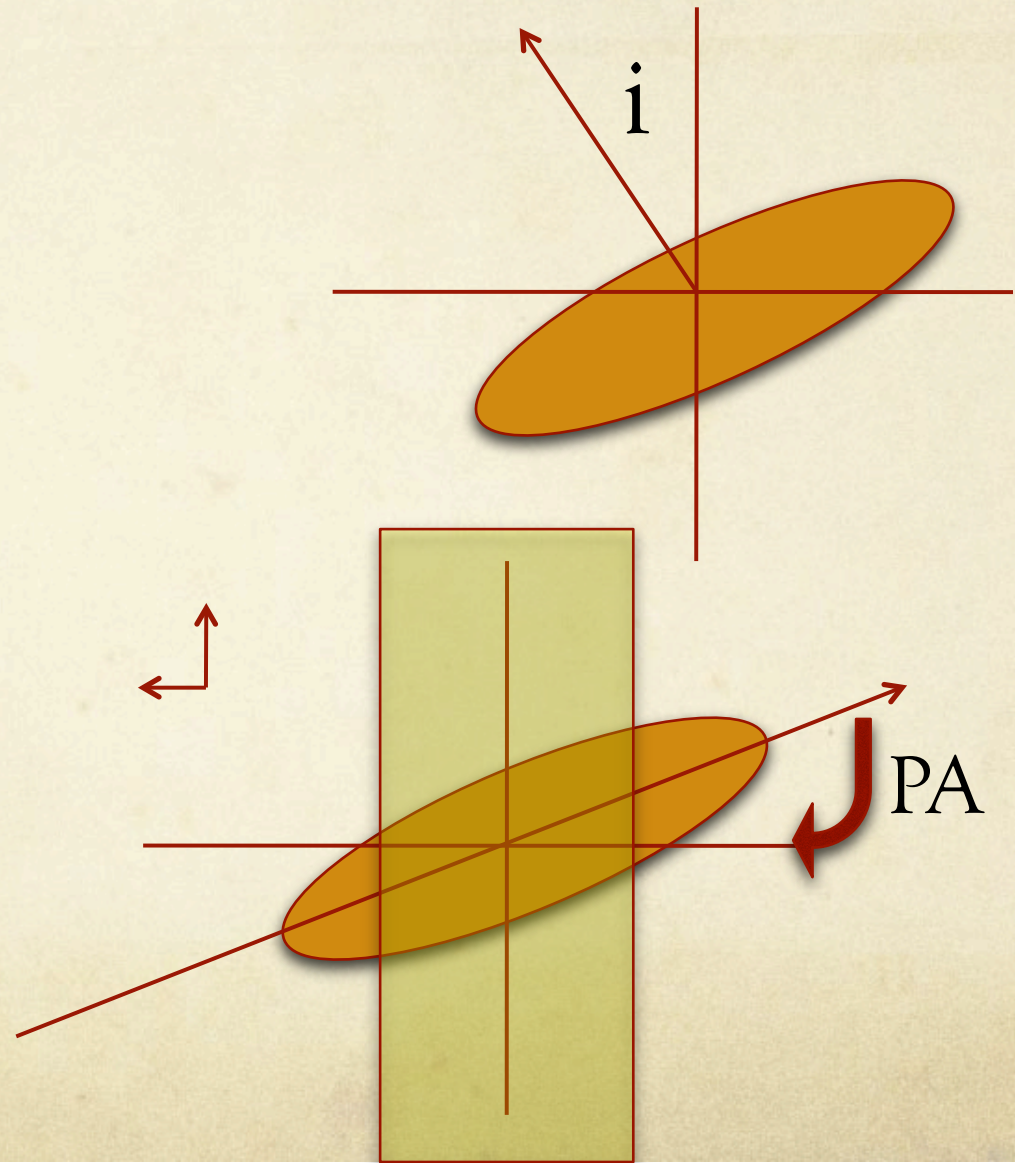
- If spatially extended, can find spatial and velocity information about extended object
- Used to observe BD companions, jets, binaries, etc.
- Can also be used for circumstellar disks



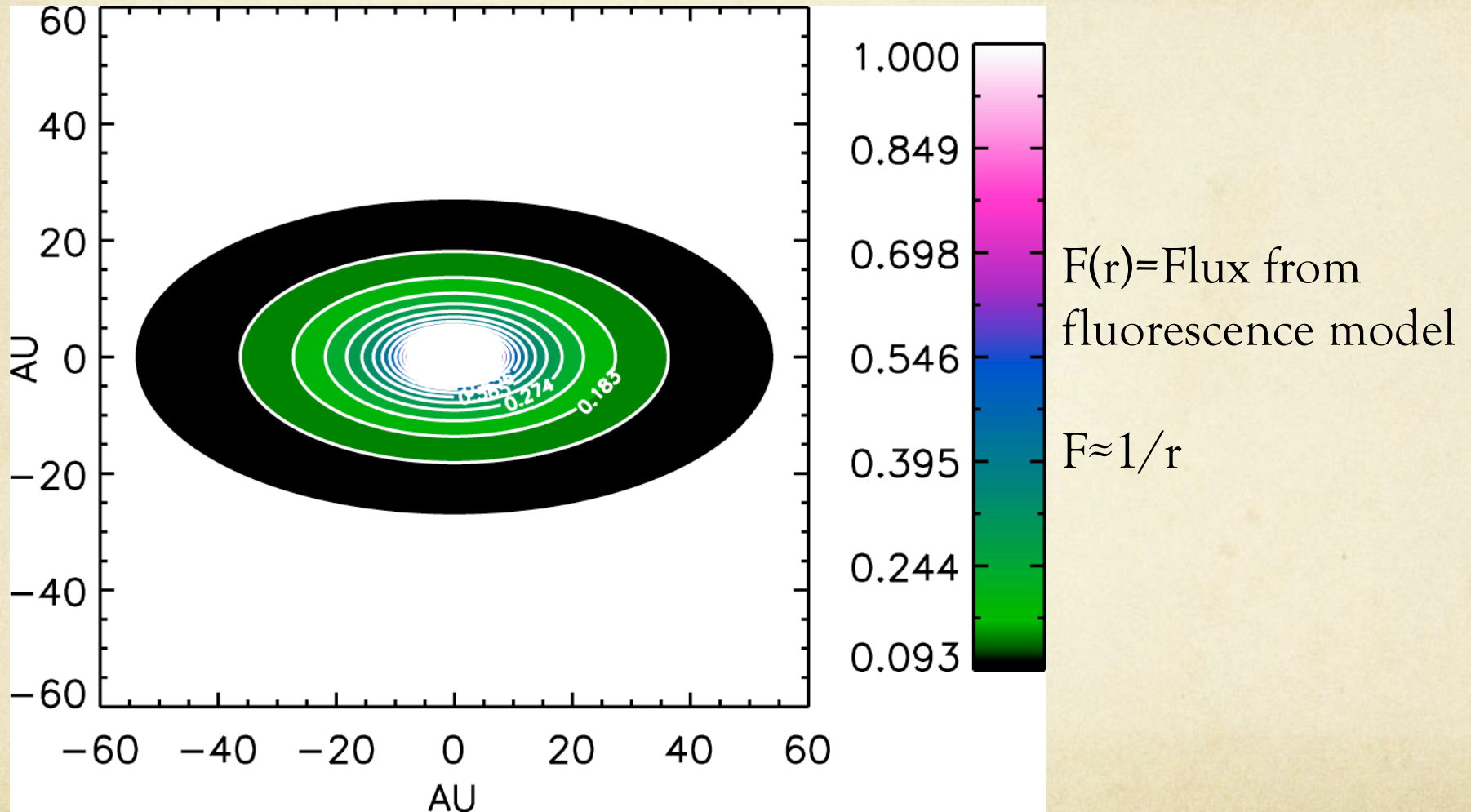
Pontoppidan et al. (2008)

Parameters

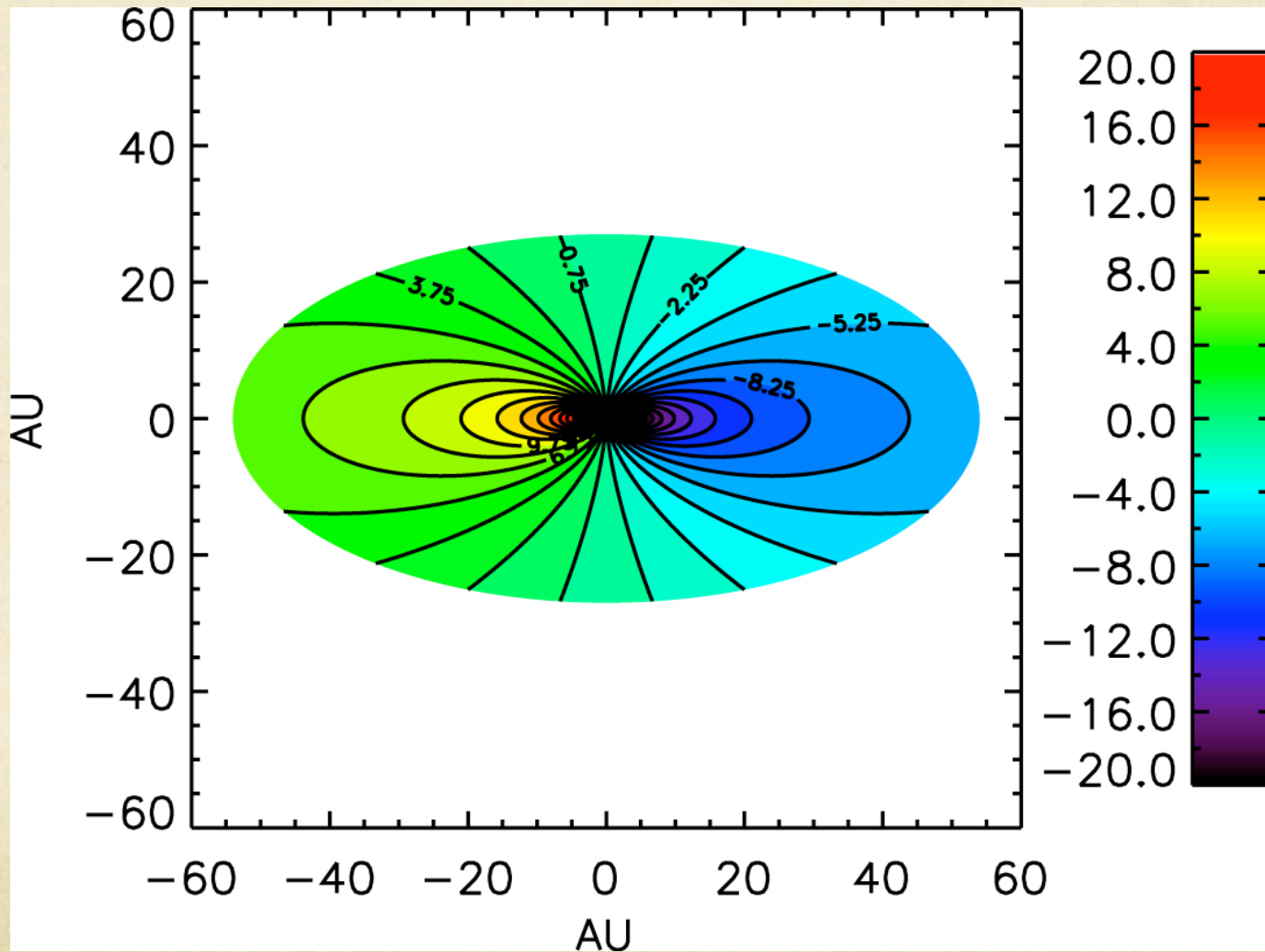
- Disk size
- Inclination - i
- Position angle - PA
- Distance
- Mass of star
- Disk flux
- Slit size



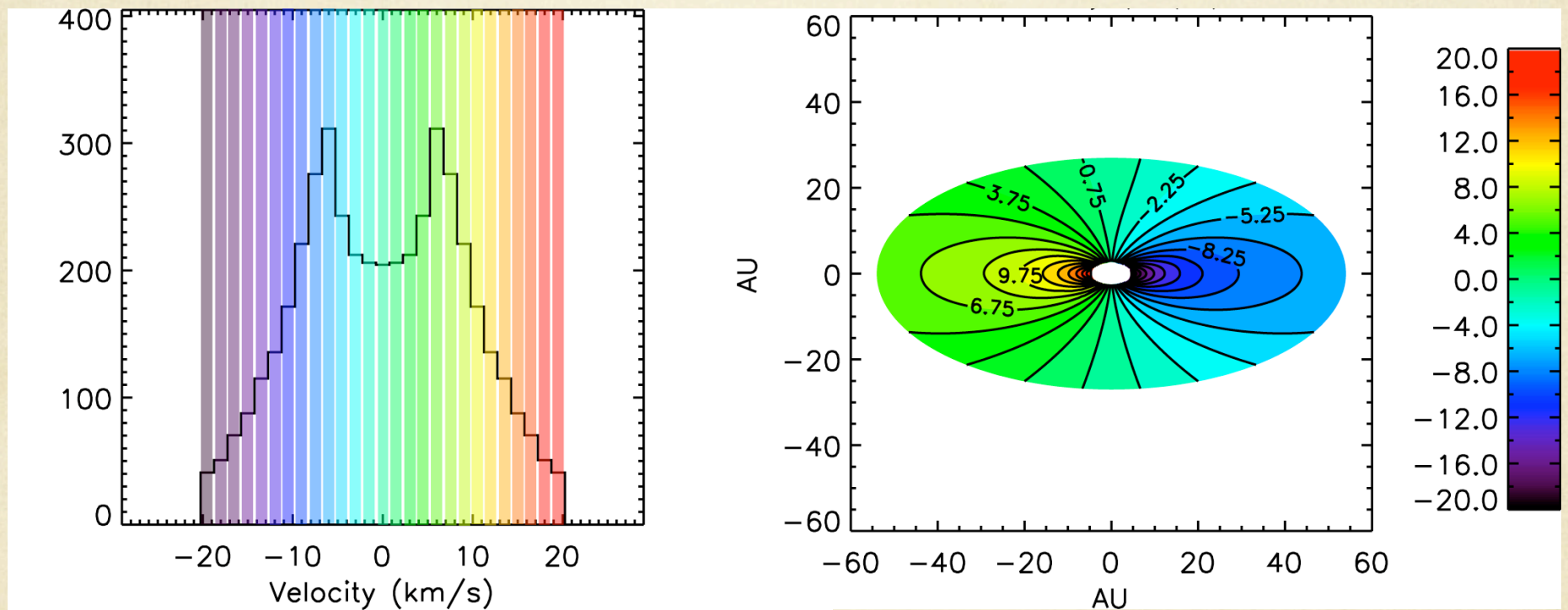
Disk Flux



Disk Velocity Field

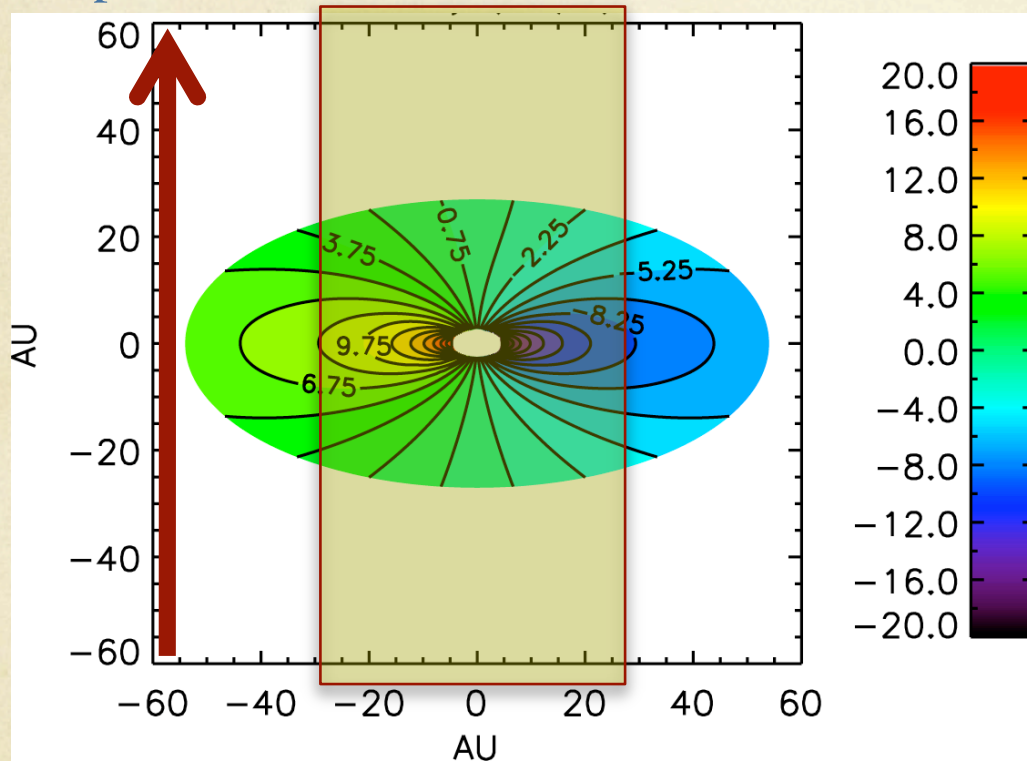


Connecting to Flux



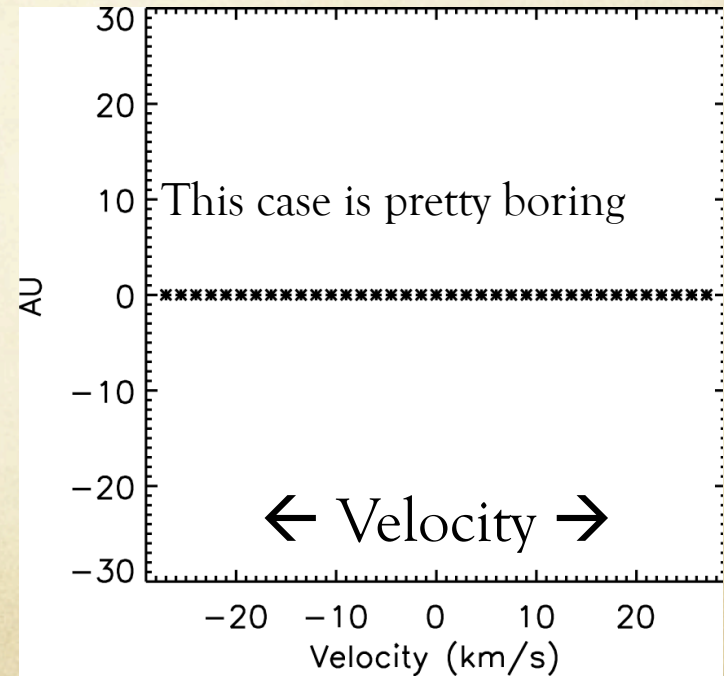
Connecting to SA Signal

Spatial

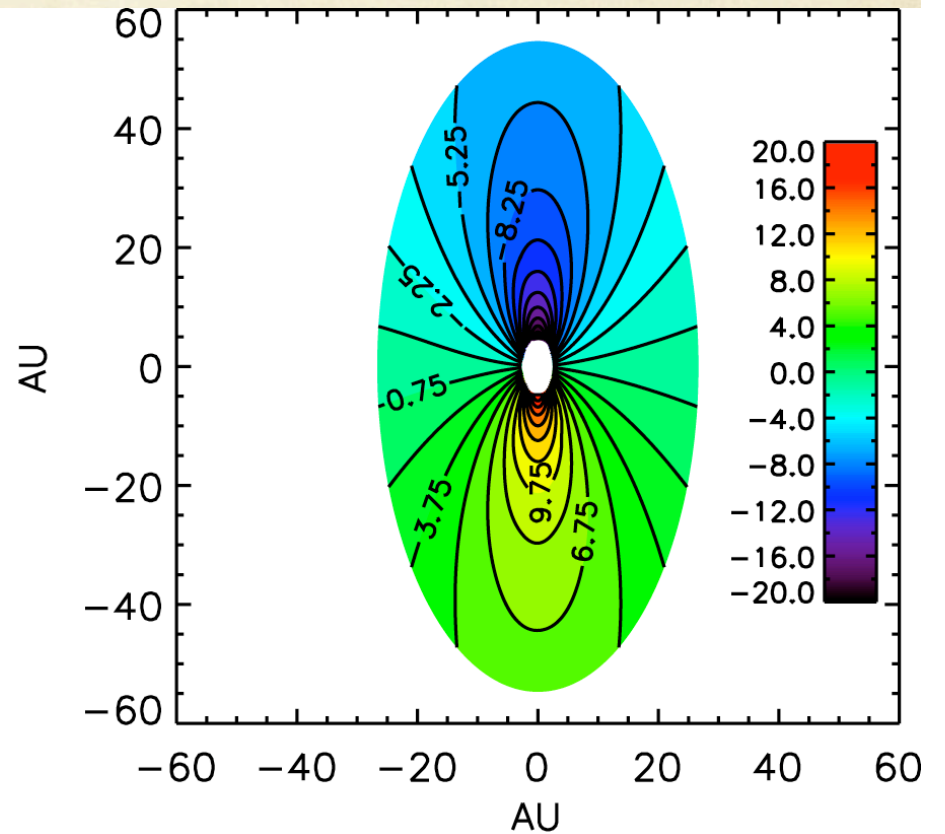
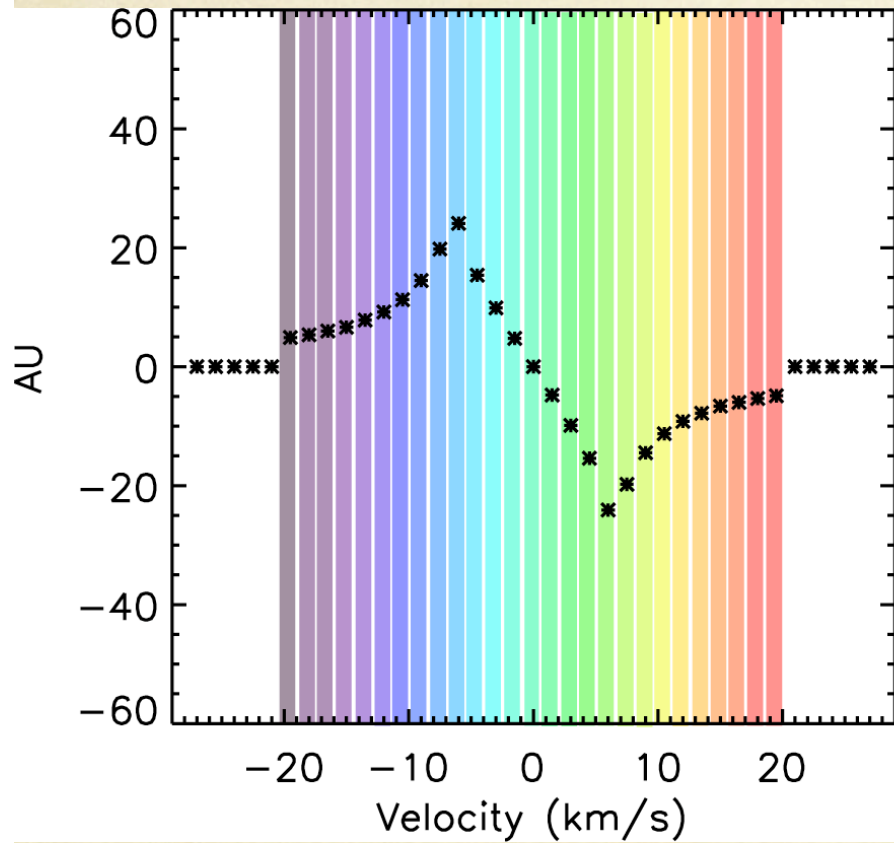


Signal is taking “COM” of the flux from the disk for a specific velocity:

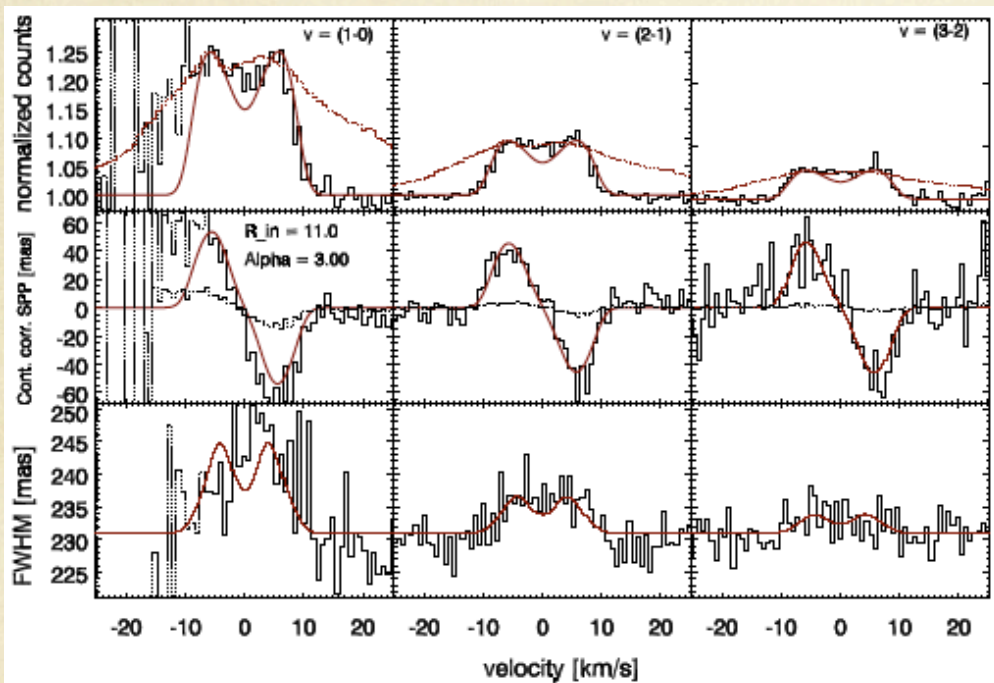
$$\frac{\sum F_i \times y}{\sum F_i}$$



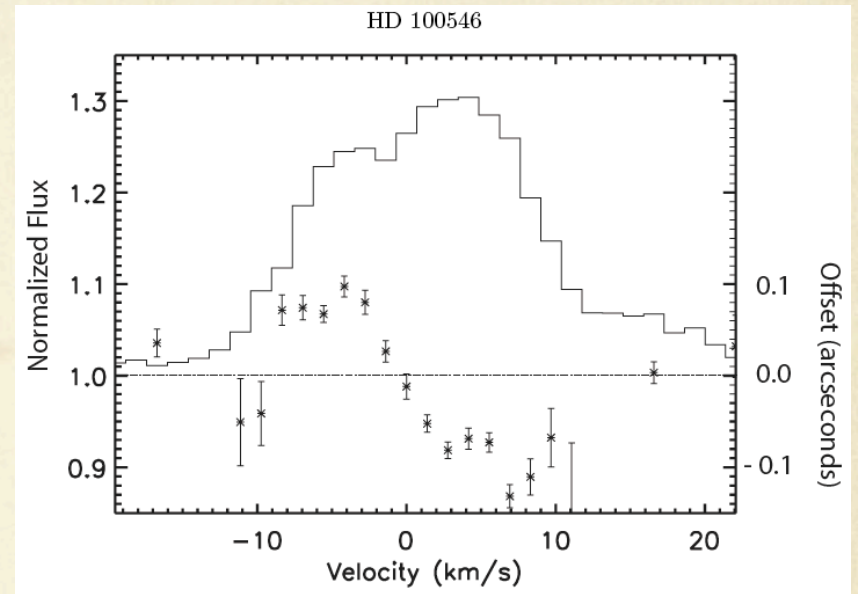
Velocity and Spectro-astrometry



Example Data



van der Plas et al. (2009)



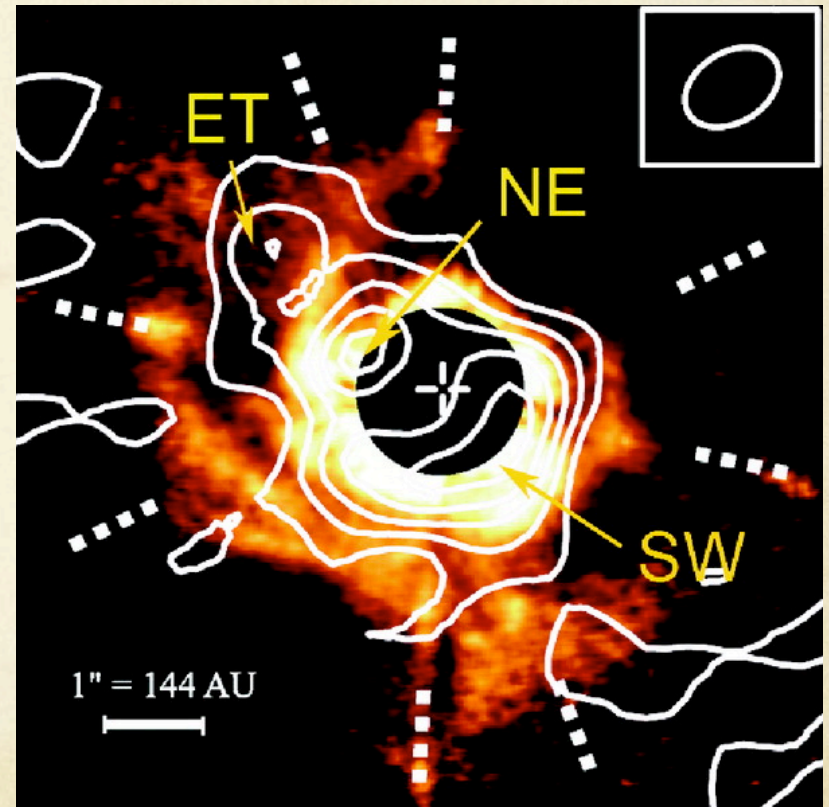
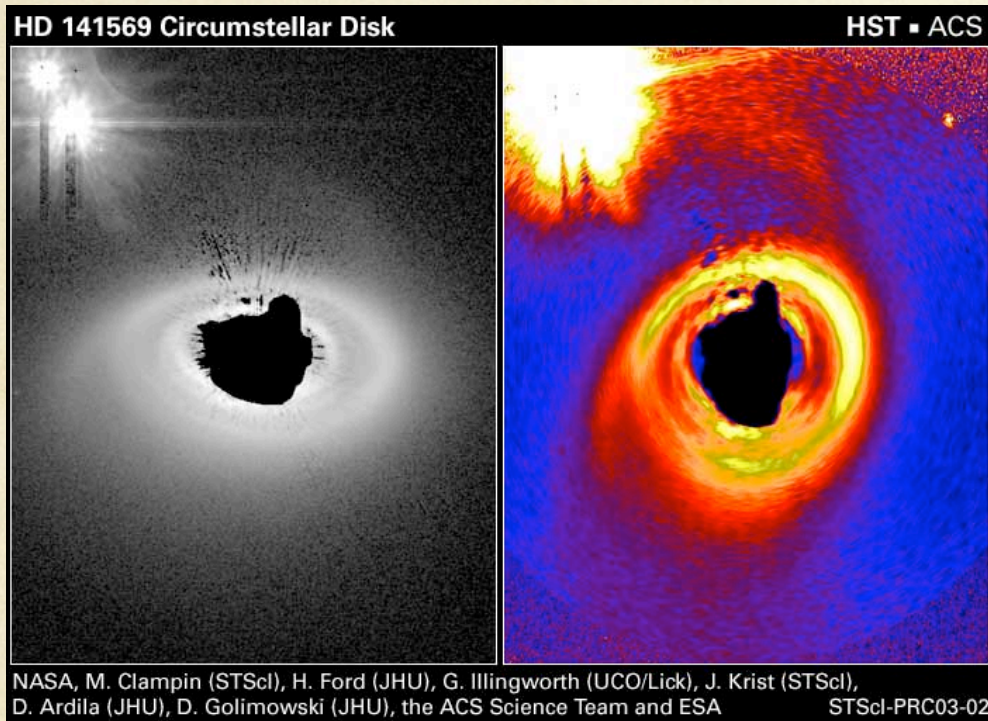
Brittain et al. (2009)

So Why is this important?

- Measure of velocity field independent of Keplerian assumption
- Can independently measure:
 - inclination
 - position angle
 - stellar mass
- Also...

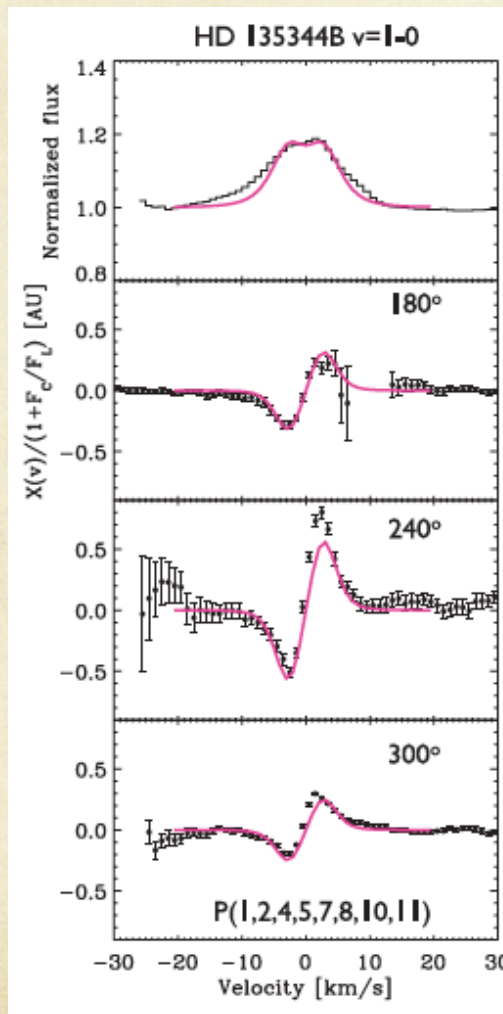
Other Effects

AB Aur

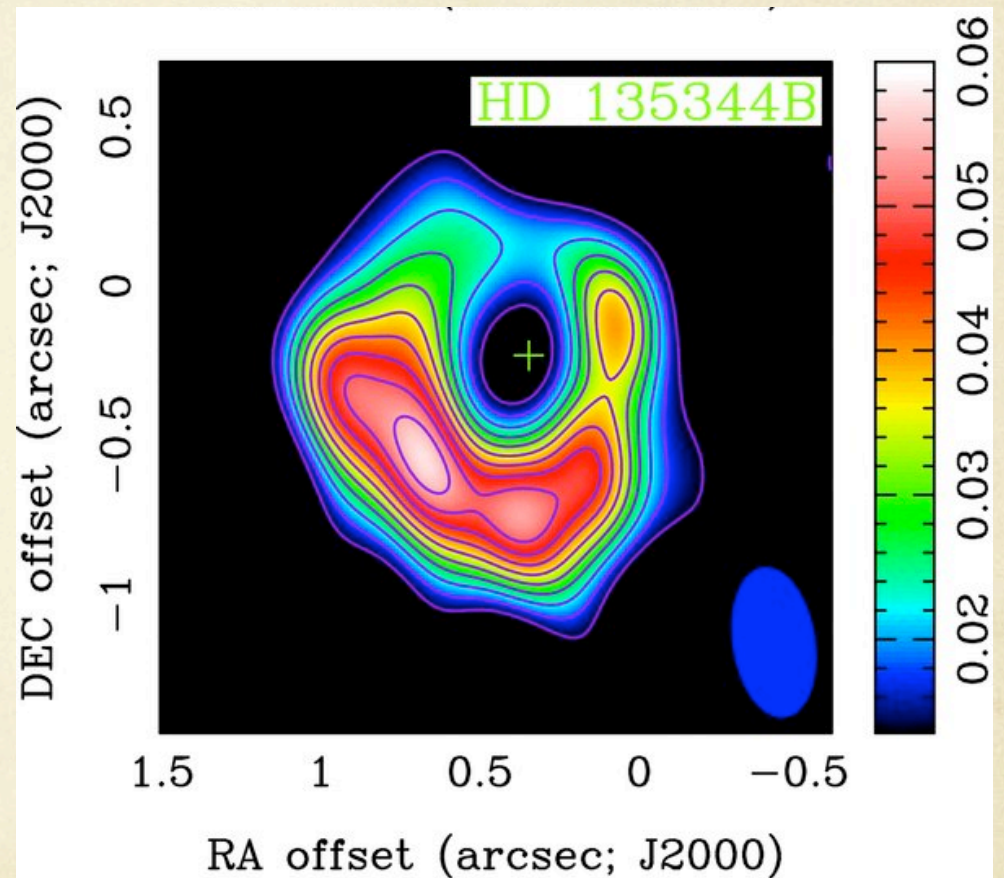


Lin et al. (2006)

Evidence for Asymmetry?



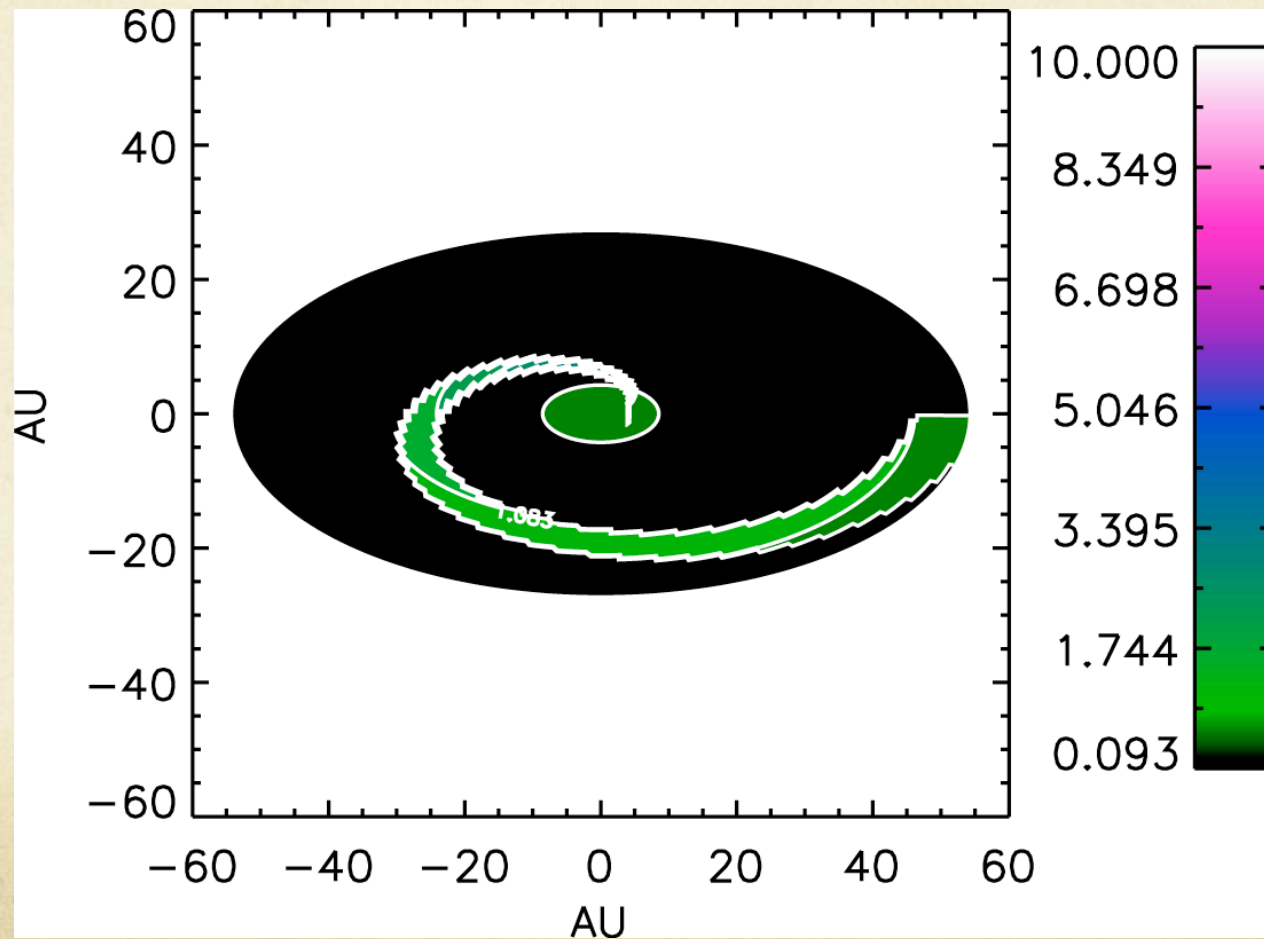
Pontoppidan et al. (2008)



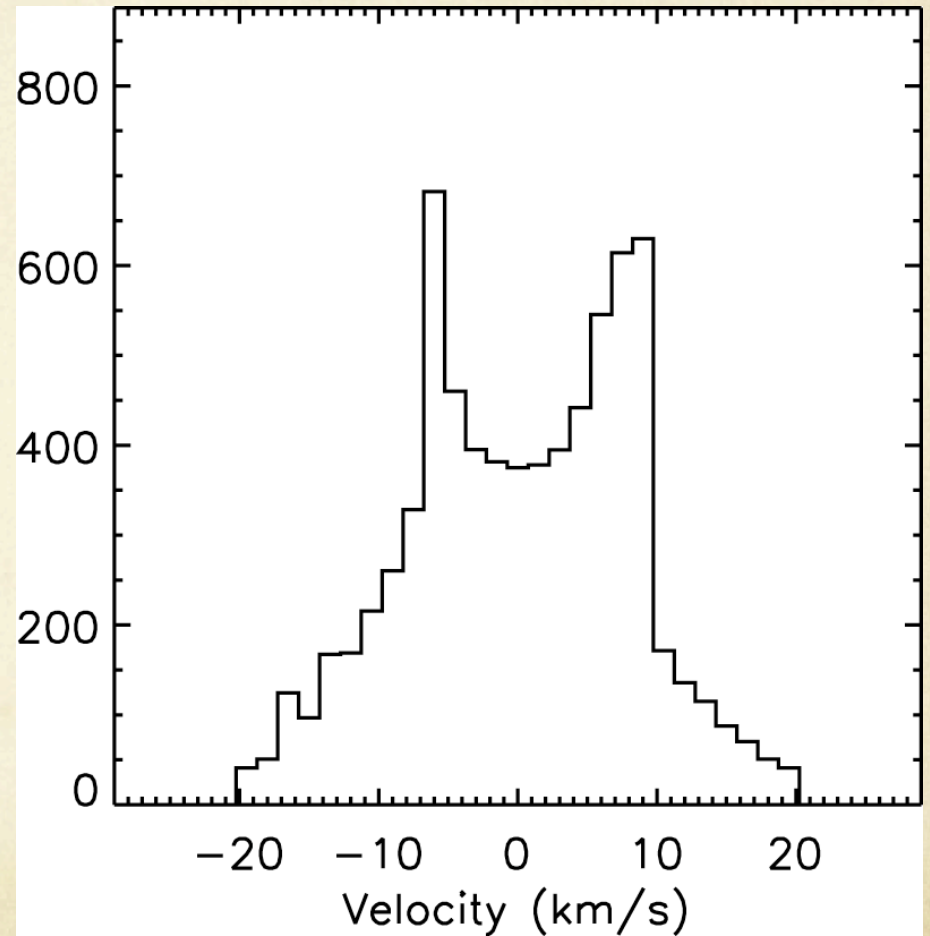
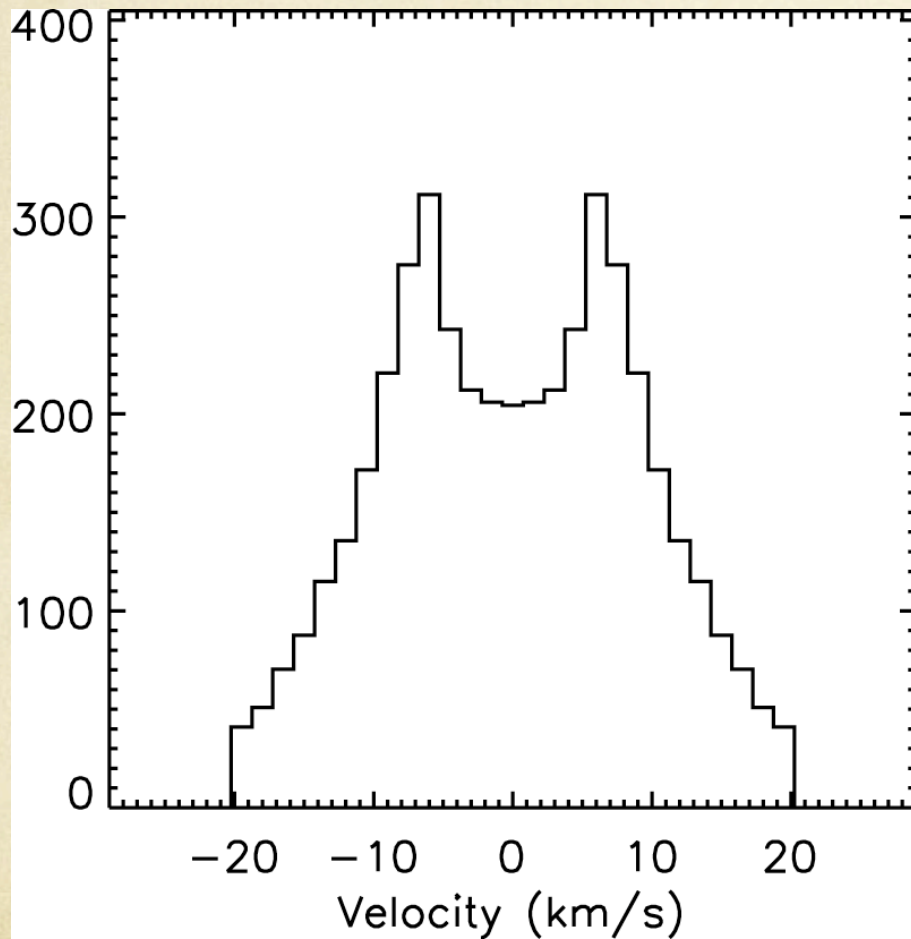
Brown et al. (2009)

Adding a Spiral Arm

- $F(s) = 10 F(ns)$
- $r = 7 \theta$

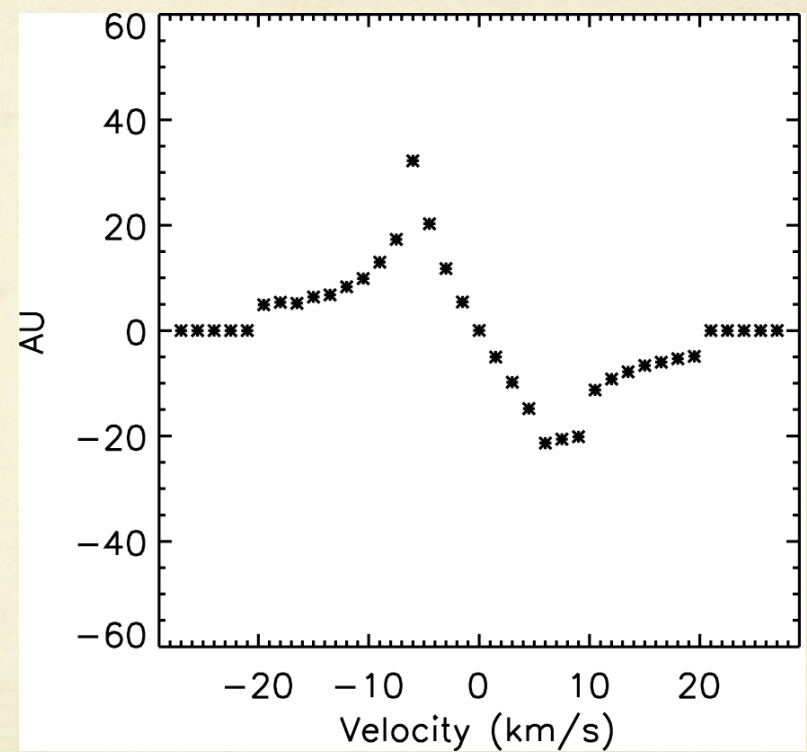
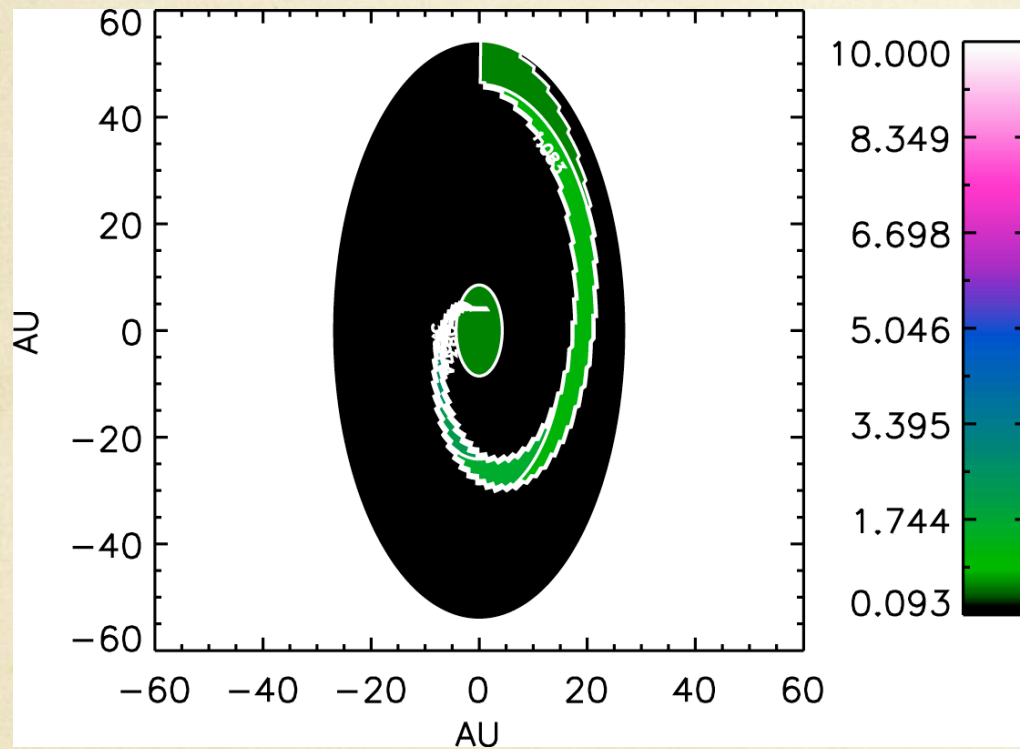


Flux Comparison

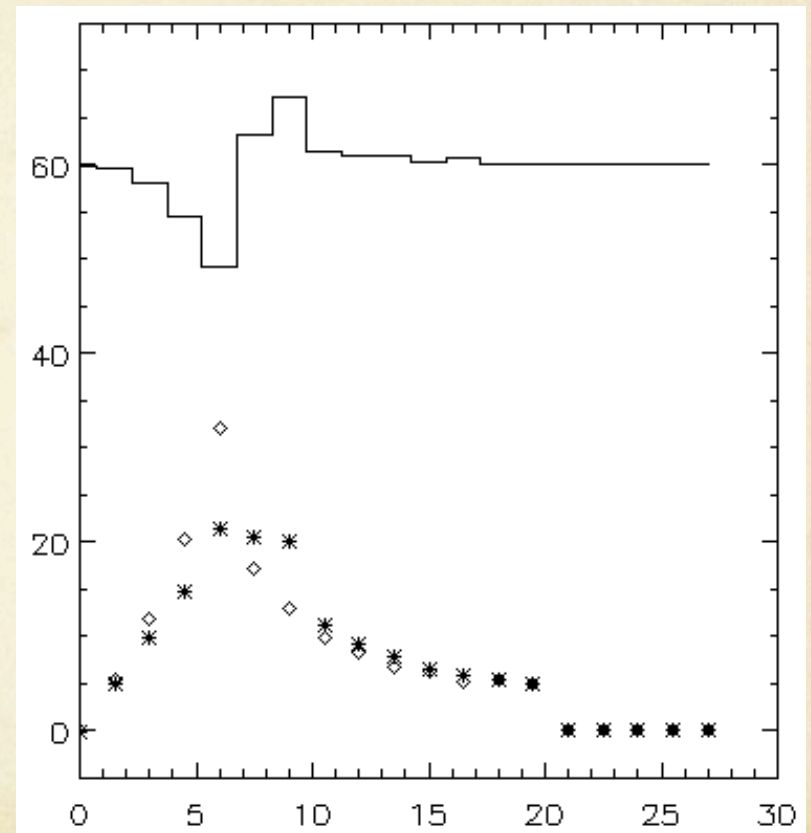
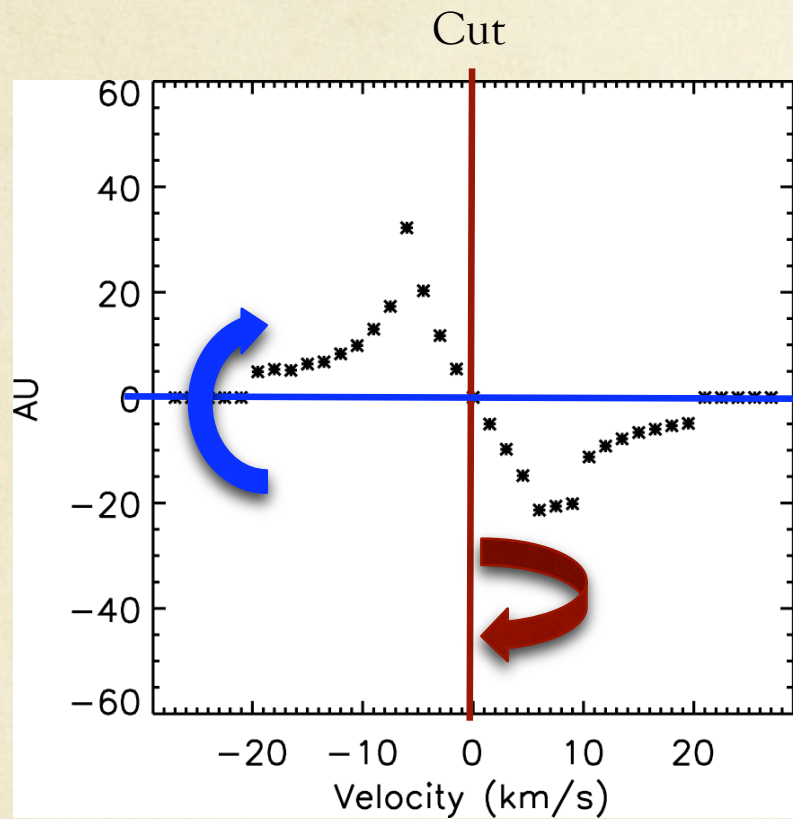


May not see this in flux signal!!

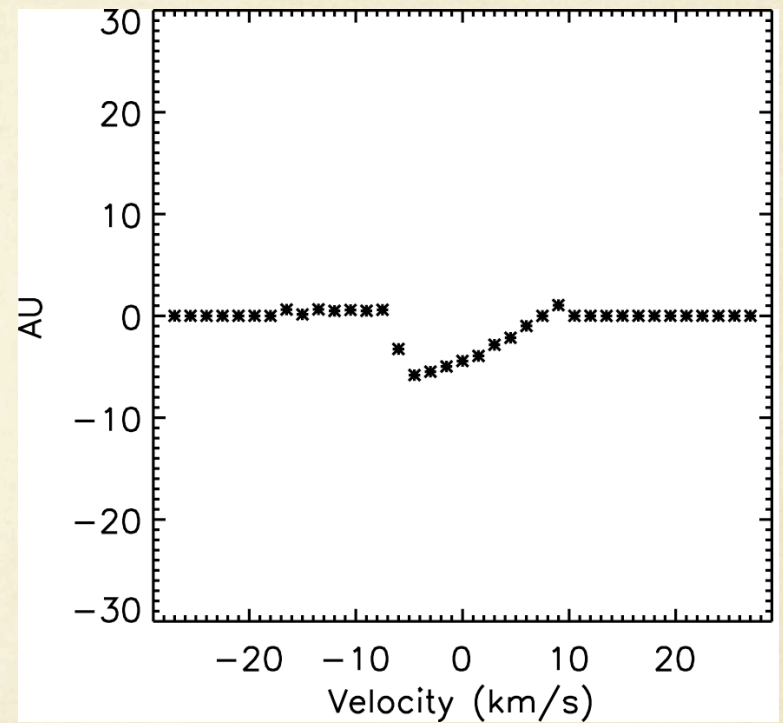
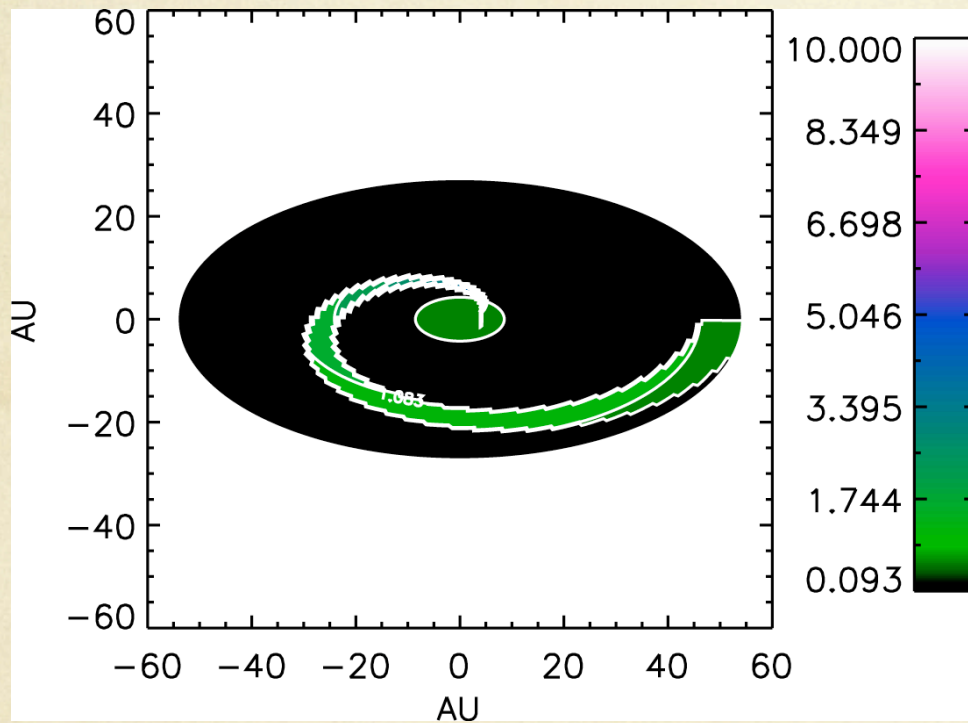
Spectro-astrometric Signal



Technique for Comparison



Spectro-astrometric Signal



Still observable with “non-optimal” PA

Summary

- Spectro-astrometry is an extremely unique way to gain high spatial resolution from the disk, and find structures that cannot be probed with other techniques.