# **Precision Multi-Object Radial Velocity Surveys** with a Dispersed Fixed Delay Interferometer

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# **ET: The Exoplanet Tracker**



Erskine & Ge 2000; Ge, Erskine, Rushford 2002

# **ET: The Exoplanet Tracker**

**Slit Direction** 

**Dispersion Direction** 

Instead of looking at line shifts like a conventional spectrograph, ET will look at the drifts in the interference fringes and convert this drift to velocity

Thorium-Argon Reference

Aldebaran

## **The Basic Principle: Information Heterodyned**



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### Prototype Multi-Object RV instrument on Sloan 2.5m March, April 2005



25 Fibers 20 Spectra, V=8-11 in 30 min. exposure

## Prototype Multi-Object RV instrument on Sloan 2.5m March, April 2005



Prototype low efficiency (1-3%) compared to KPNO ET. Worse stability. Yet,  $\sigma = 10-20$ m/s! **Proof of Concept!** 

### **Multi-Object Radial Velocity Instrument**



Wavelength Coverage =1000 Angstroms.  $\sigma = 3m/s$  for V=8,  $\sigma = 20m/s$  for V=12

**Surveying Open Clusters with the Multi-object Instrument** 

Keck Hyades Cluster Survey –*Cochran et al. 2002* 

~100 F, G, K, M stars surveyed using many nights of Keck time

~ 15 minute per star,  $\sigma = 3-6$  m/s



### Searching the *KEPLER* mission FOV.



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Can also discover non-transiting hot Jupiters by Reflected Light

## Searching the *KEPLER* mission FOV. April 2005



#### Searching the *KEPLER* mission FOV.



SDSS Spectrographs used simultaneously with the multi-object instrument-640 spectra in the Kepler field.





**Goal** 2000 stars searched 10-20 hot Jupiters 30-60 other planets Spectra for 10-20k stars.