

Probing the inner circumstellar structures of T Tauri and Herbig stars

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Vink et al., 2005b, MNRAS accepted, astro-ph/0502535

Outline

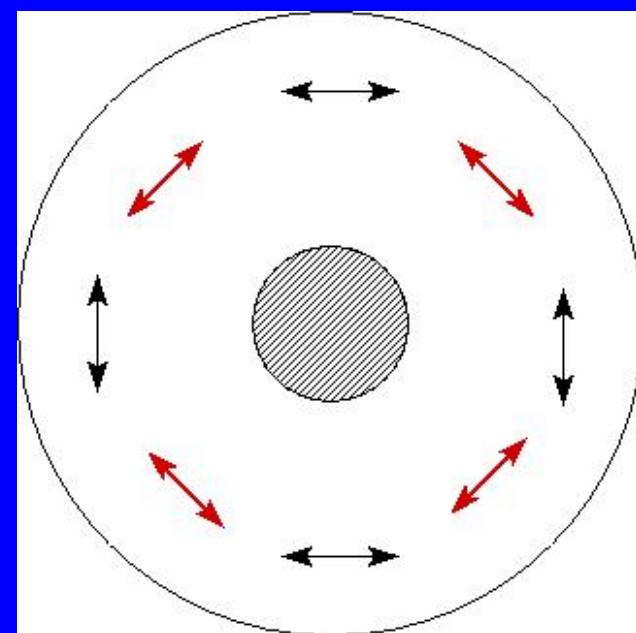
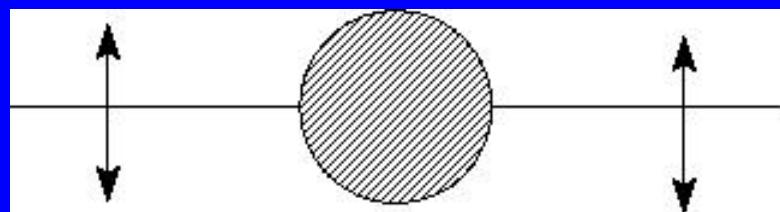
- Introduction on Spectropolarimetry
- Data
 - T Tauri 1 Msun
 - Herbig Ae 3 Msun
 - Herbig Be 10 Msun
- Disc scattering models
 - inner hole
 - undisrupted
- Conclusions

Polarimetry – from disks

$$\begin{aligned} I &= \text{Total Intensity} \\ U &= \text{Upward Polarization} - \text{Downward Polarization} \\ Q &= \text{Rightward Polarization} - \text{Leftward Polarization} \end{aligned}$$

$$P = \sqrt{(U^2 + Q^2)}$$

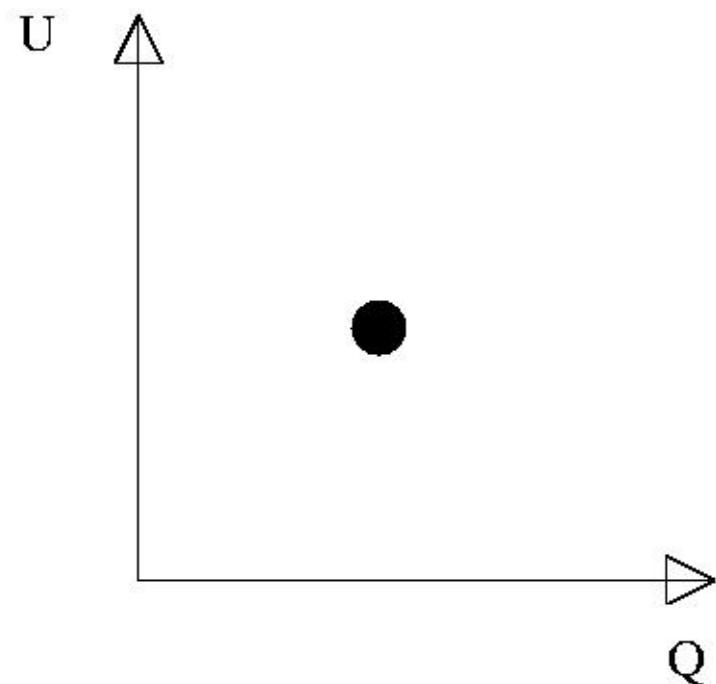
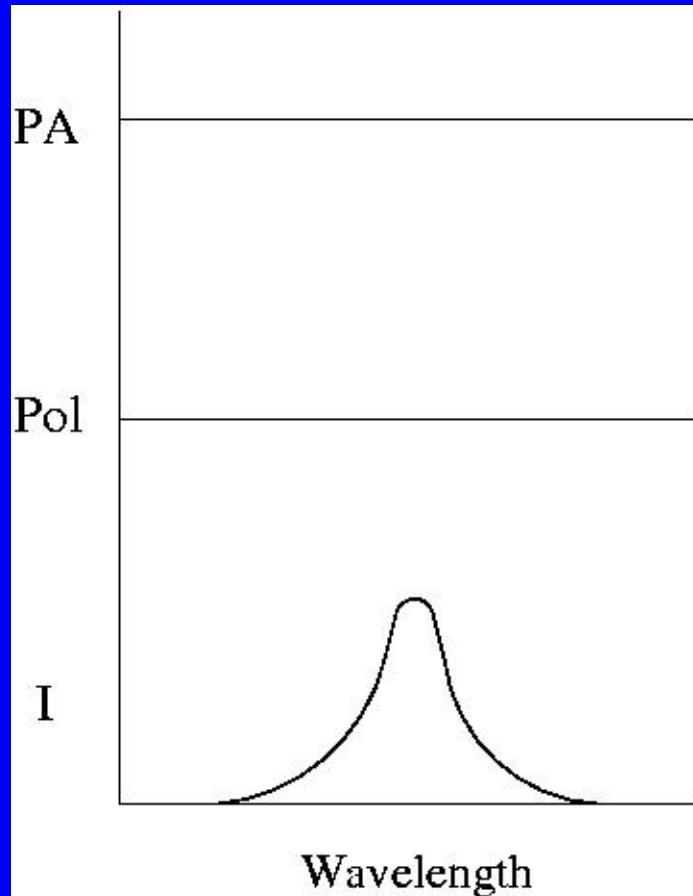
$$\theta = \frac{1}{2} \arctan\left(\frac{U}{Q}\right)$$



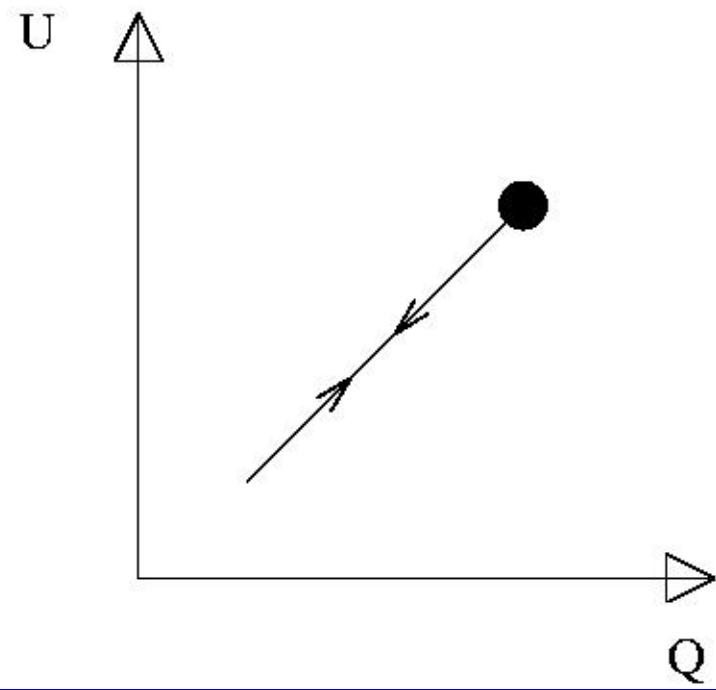
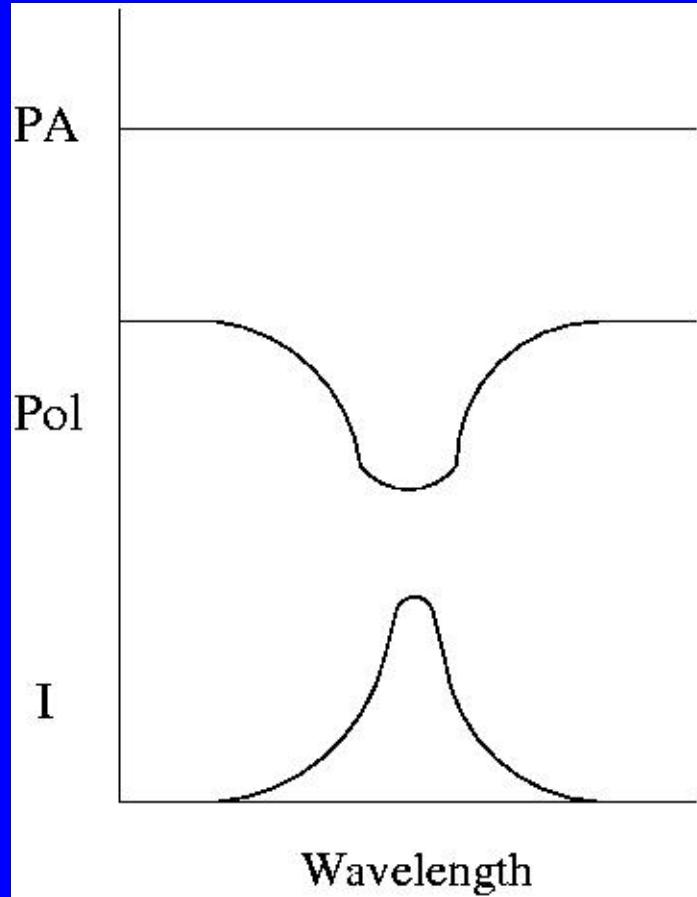
Polarisation across line?

1. No change
2. Depolarisation
3. LINE Polarisation

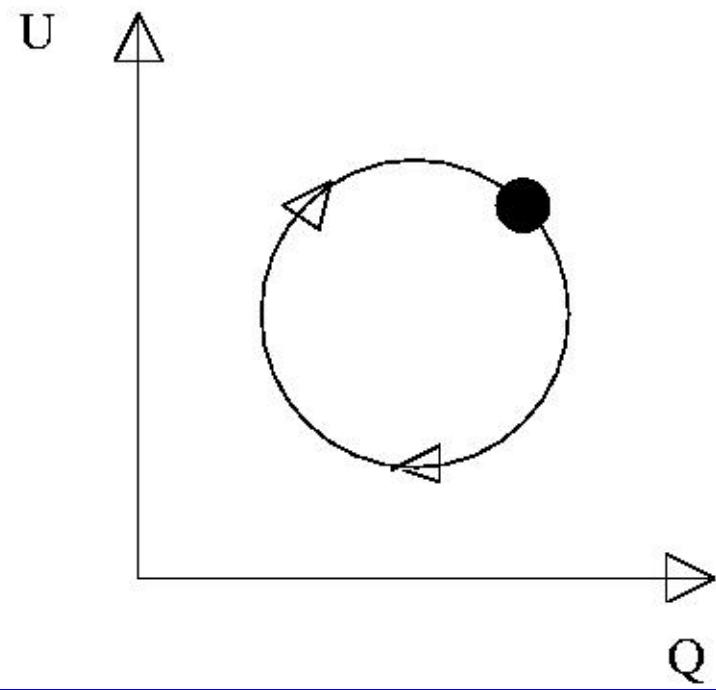
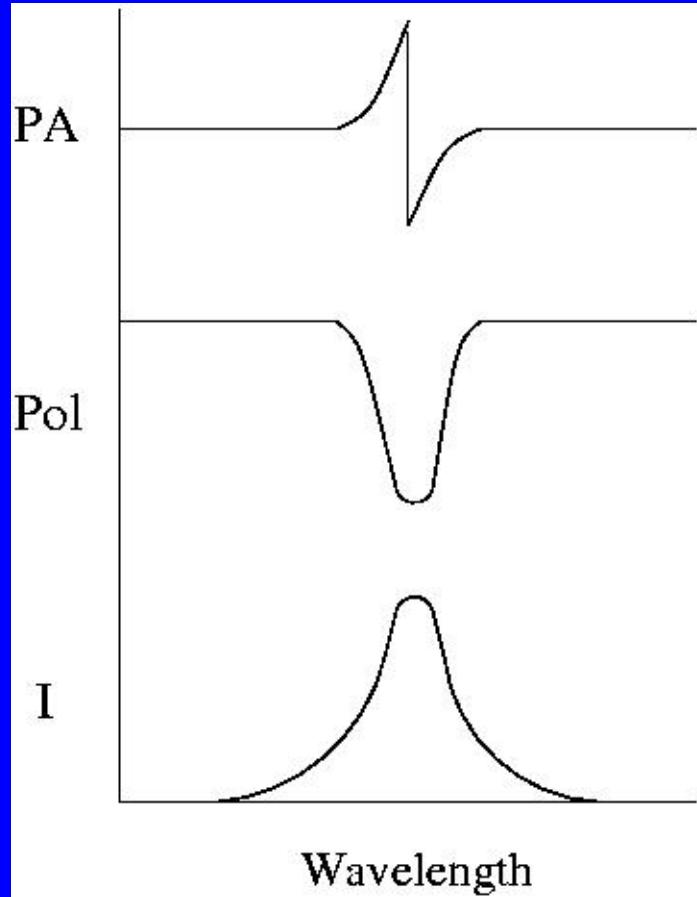
No Polarisation



Depolarisation



Line Polarisation – PA Flip



Survey Herbigs and T Tauris

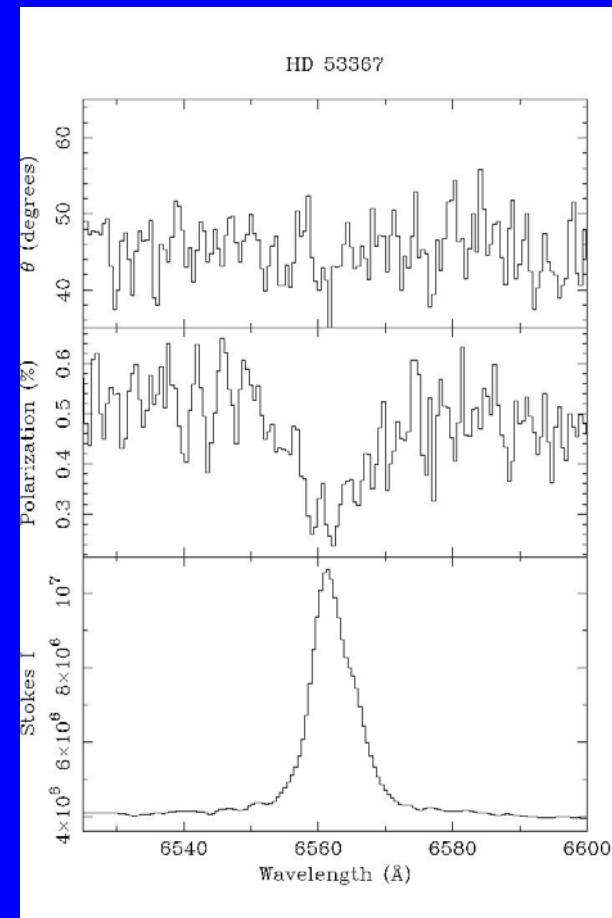
- Herbig Be stars: 12
- Herbig Ae stars: 11
- T Tauri stars: 10

Data: Herbigs and T Tauris

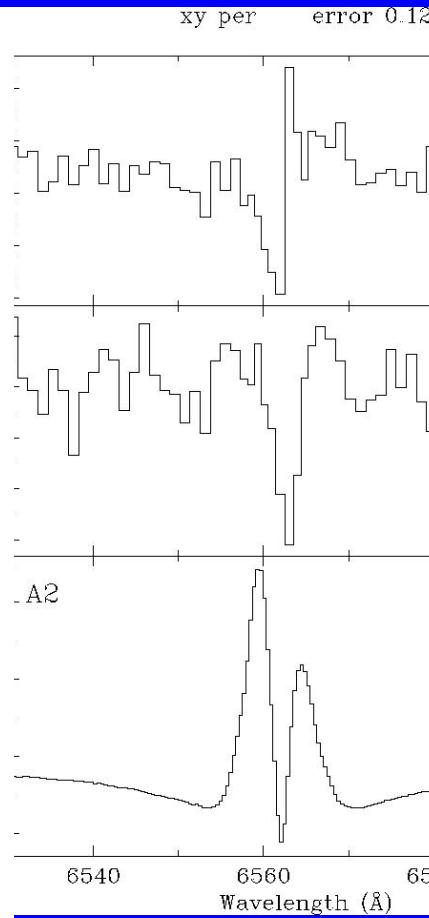
PA

Pol

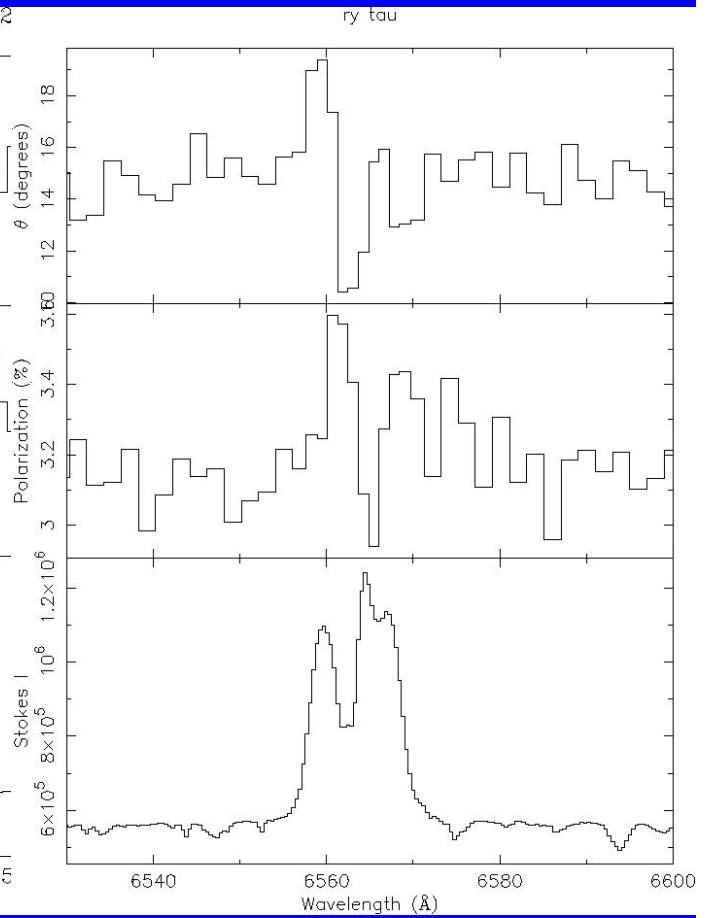
I



Herbig Be



Herbig Ae



T Tauri

Polarisation across line?

1. No change
2. Depolarisation
3. LINE Polarisation

→ Herbig Be: 7/12



Herbig Ae: 9/11

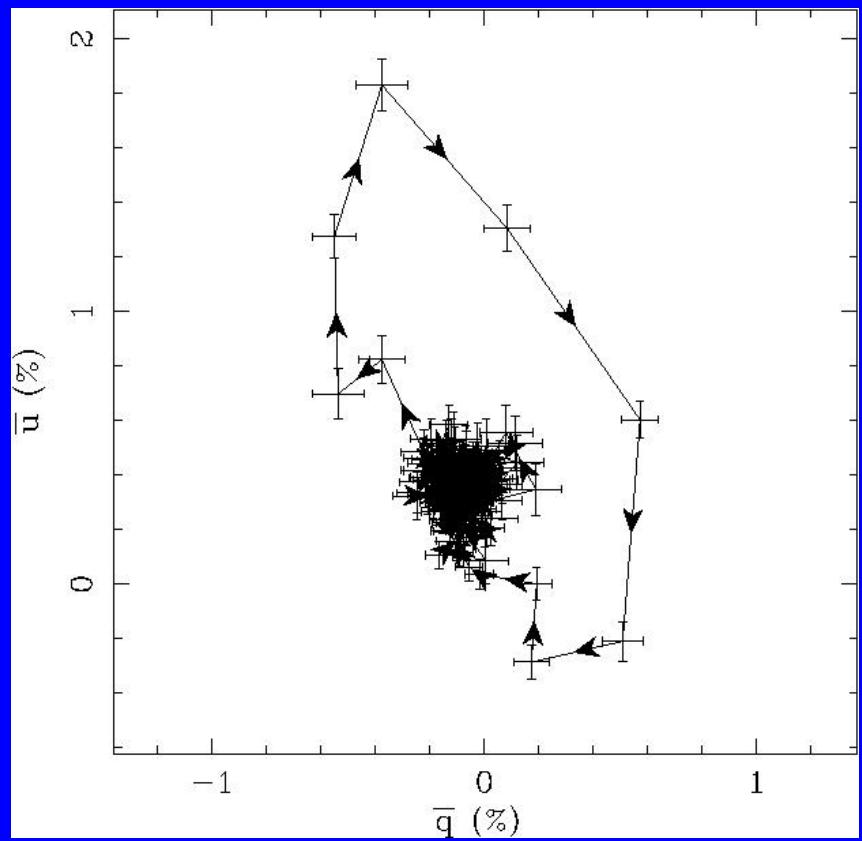
(Vink et al. 2002, MNRAS)

T Tauri: 9/10

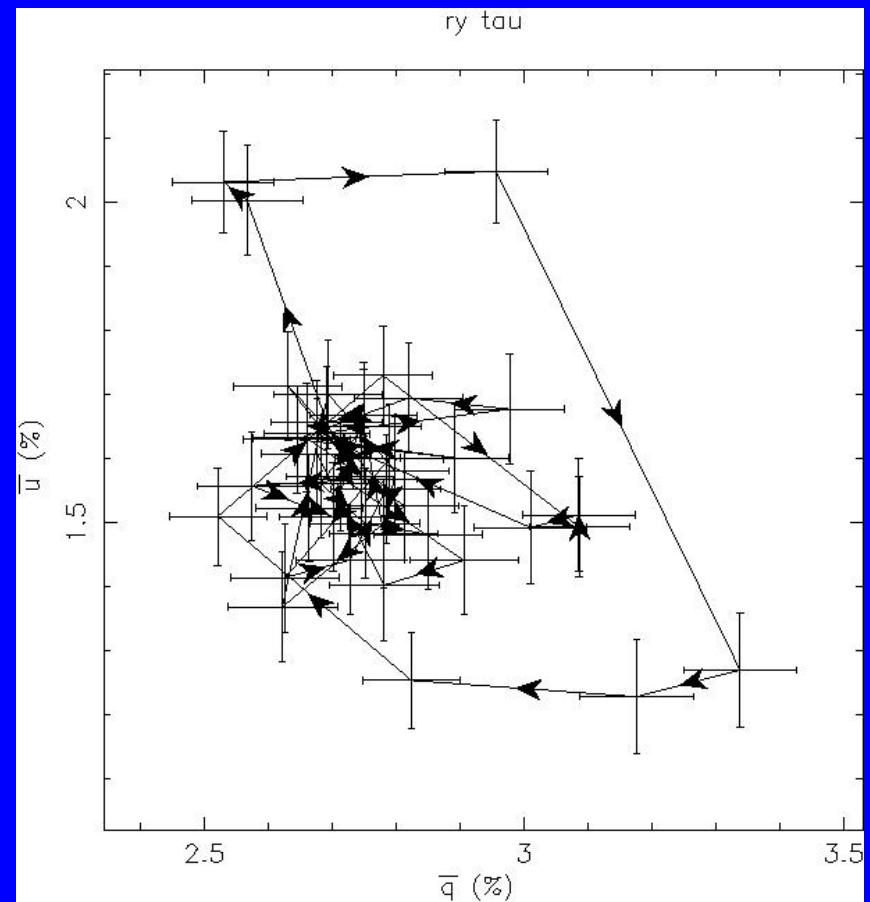
(Vink et al. 2003, A&A)

(Vink et al. 2005b)

QU: Herbig Ae and T Tauri star



MWC 480

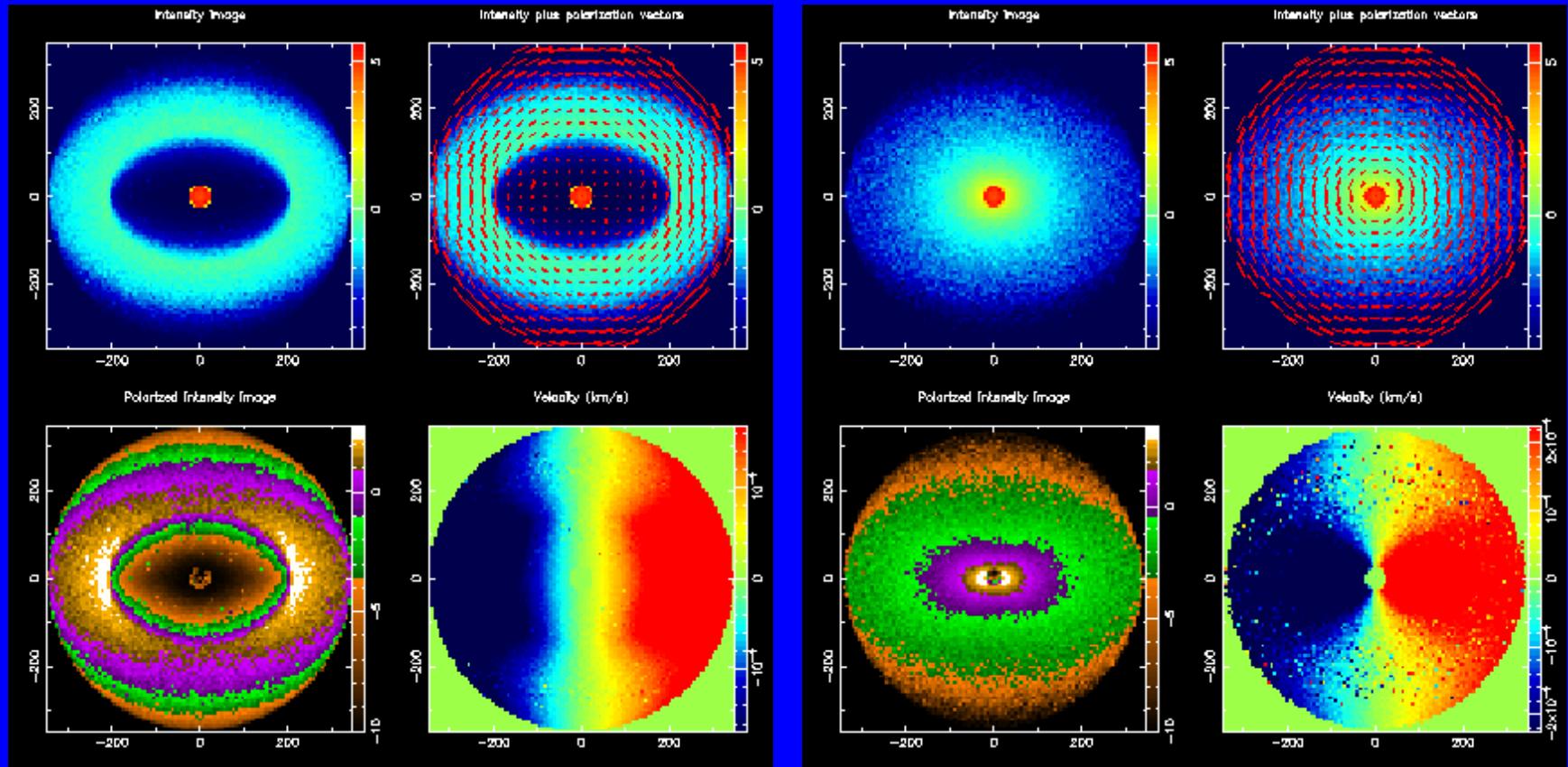


RY Tau

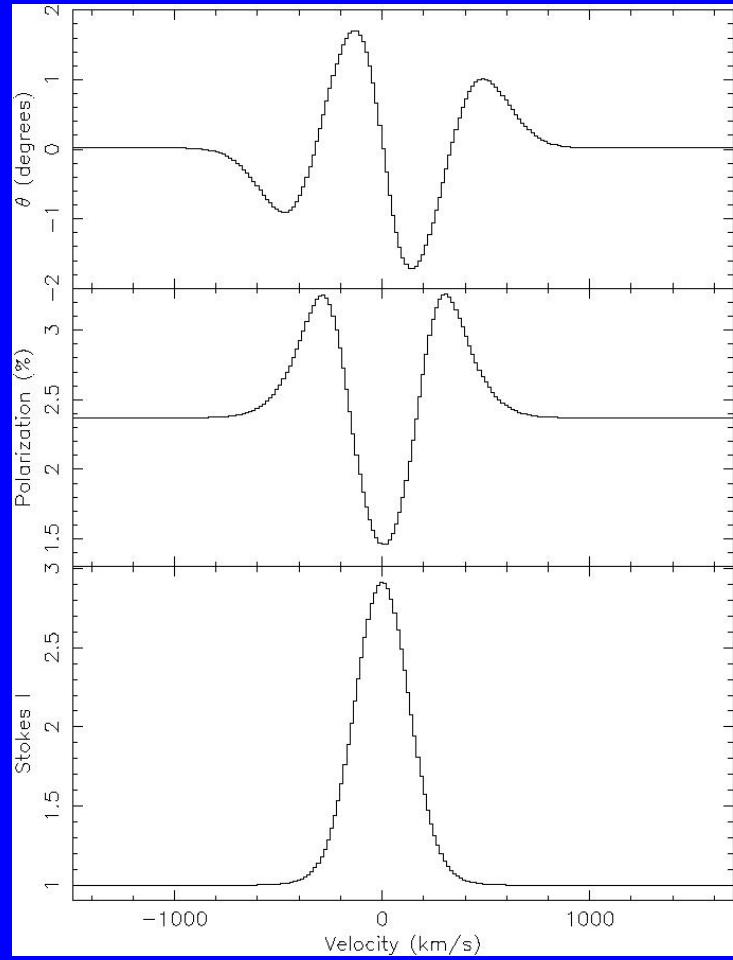
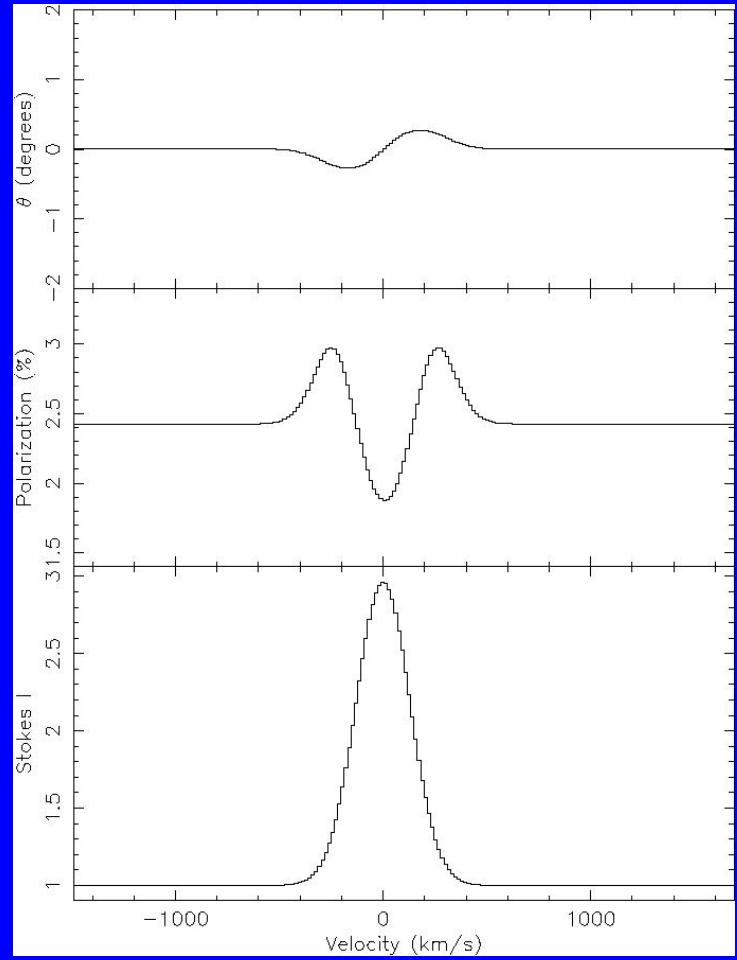
Models of COMPACT line emission

- 3D Monte Carlo
- Keplerian rotating disk
- Flat or constant opening angle
- Scattering only – no line transfer
- With and without an inner hole

With/without an inner hole

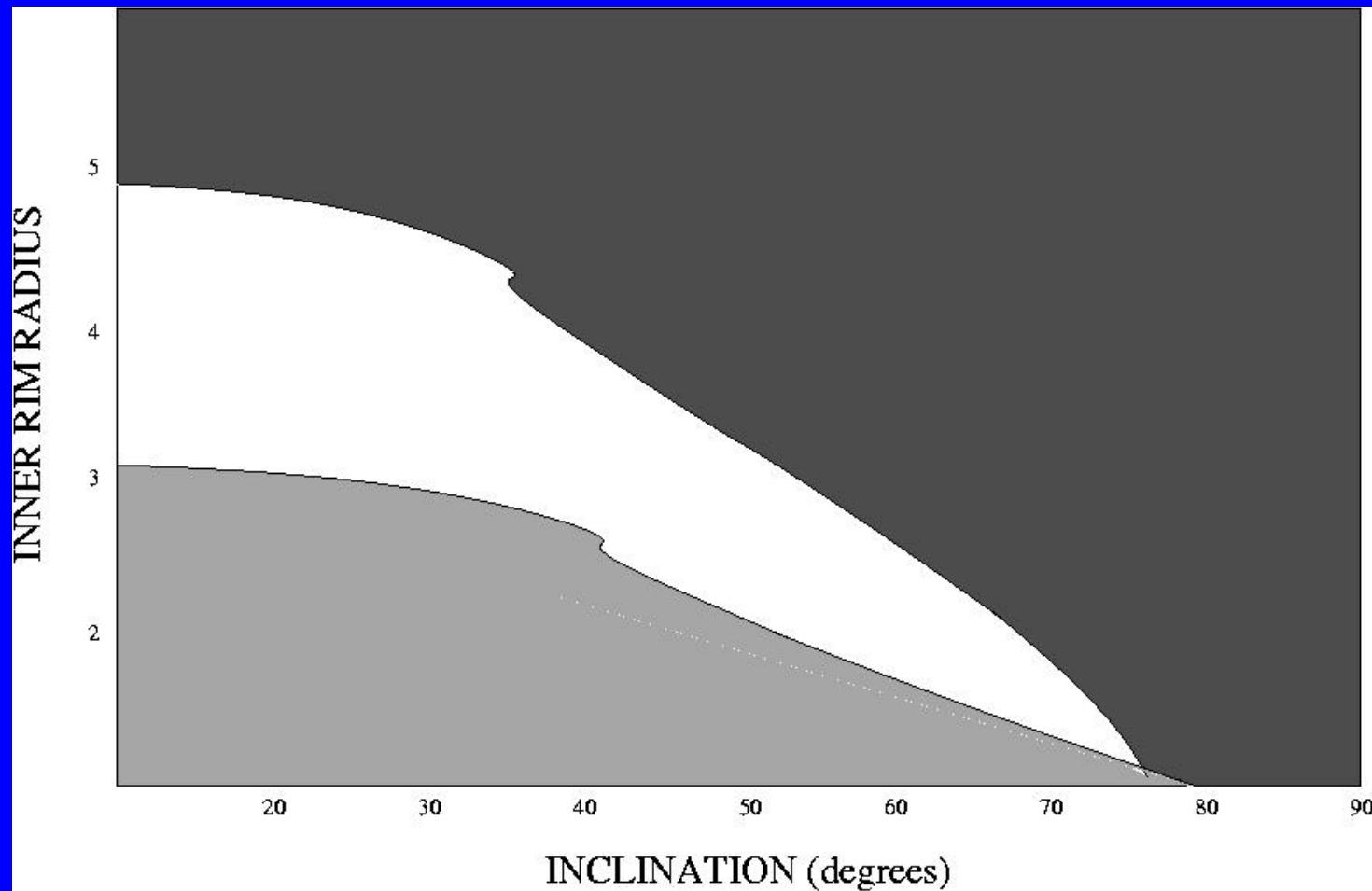


With/without a hole



Vink et al., 2005a, A&A accepted

Constraining the inner disk radius



Constraining the inner hole size:

Single PA flip; known inclinations

- AB Aur Inner rim > 5 Rstar
- CQ Tau Inner rim > 4 Rstar
- SU Aur Inner rim > 3 Rstar

Gradual PA change

- GW Ori Inner rim 3 or 4 Rstar

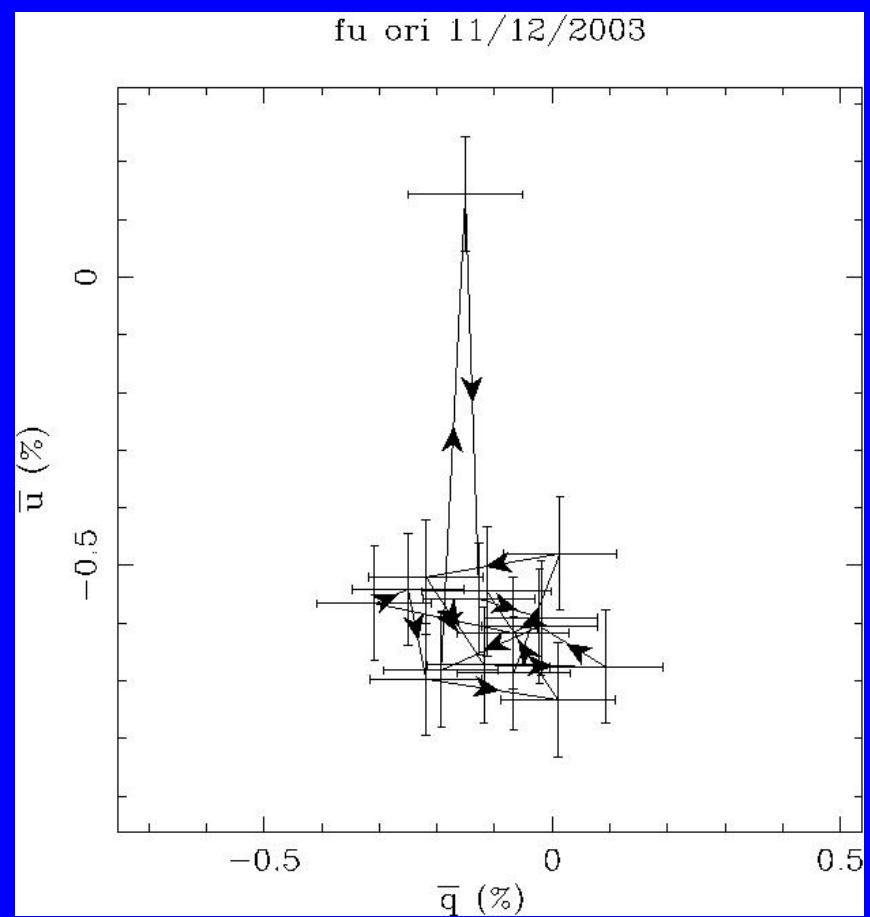
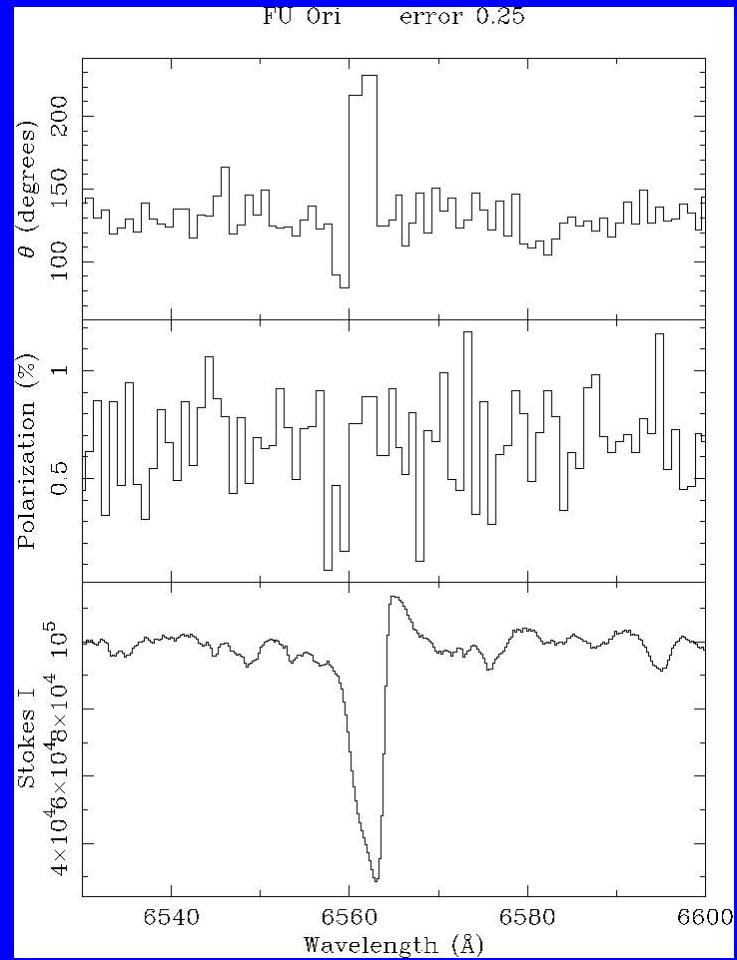
(Vink et al. 2005a, 2005b)

McLean effect in FU Ori

PA

Pol

I



Accurate measurement of intrinsic polarisation PA

Imaged disks: position angles

	PA line pol	PA direct	Delta PA
AB Aur	160	60/80	80/100
MWC 480	55	150	95
CQ Tau	20	105	95
RY Tau	163	62	101
FU Ori	45	47	2
SU Aur	130	127	3
DR Tau	120	128	8

(Vink et al. 2005b)

Conclusions

- Herbig Be: disks on small scales
- Herbig Ae/T Tau: rotating accretion disks
- Inner rim sizes 3 – 5 stellar radii
- Herbig Ae: polarisation PA perpendicular to disk PA → optically inner thin
- T Tauris: polarisation PA parallel to disk PA → optically thick