The Radius Valley as a By-Product of Planet Formation: The Core-Powered Mass-Loss Mechanism





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Recent Observations

- Most common planets known are 1-4 Earth radii in size
- Lack of small, close-in exoplanets around 1.5-2.0 Earth radii, i.e., a radius `valley' separating a population of super-Earths (smaller planets) and sub-Neptunes (larger planets)
- Compositions?
 - Super-Earths have higher densities \rightarrow consistent with \mathbb{R} planets having rocky 'Earth-like' composition
 - Sub-Neptunes have lower densities \rightarrow consistent with planets engulfed in H/He envelopes





Background: The Core-Powered Mass-Loss Mechanism



• Valley's slope: $t_{loss}^B = t_{cool} \rightarrow t_{loss}^B \propto \exp(R_p^3 T_{eq}^{-1}) \propto \exp(R_p^3 M_s^{-13/12}) \rightarrow d\log(R_p)/d\log(M_s) = (3\alpha - 1)$ 2)/36. For CKS-II dataset, $\alpha \approx 5 \rightarrow d \log(R_p) / d \log(M_s) = 0.36 \rightarrow \text{consistent with observations [3,7]}$.

- For observations, as shown in the figure, $d \log(R_p) / d \log(M_s) \sim 0.35$
- Physical understanding: more massive/luminous stars \rightarrow higher planetary equilibrium temperature \rightarrow more massive cores can get stripped of their atmospheres \rightarrow valley moves up in planet size
- Photoevaporation predicts a roughly flat radius valley as a function of stellar mass (provided that the orbital period distribution is independent of stellar mass [3])
- Therefore, a linear correlation between planet and stellar mass has been invoked to understand these observations under photoevaporation [8]
- In contrast to photoevaporation models: Core-powered mass-loss predicts no significant correlation between planet and stellar mass

Planet size vs. Insolation flux, as a function of Stellar mass



- **Sub-Neptune**: if a planet does not has enough internal energy to overcome the gravitational binding energy of its atmosphere OR if a planet's cooling timescale (t_{cool}) is shorter than its mass-loss timescale (t_{loss}) , i.e., if $t_{cool} < t_{loss}$.
- **Super-Earths**: : if a planet has enough internal energy to overcome the gravitational binding energy of its atmosphere AND if $t_{cool} > t_{loss}$
- Slope of the Valley $\equiv t_{cool} = t_{loss}^B$
- Mass-loss timescale: $t_{loss}^B = \frac{M_{atm}}{4\pi R_s^2 c_s \rho_s} \propto \exp(GM_p / (c_s^2 R_{rcb})) \propto \exp(R_p^3 T_{eq}^{-1} \rho_{c*}^{4/3})$
- Exponential dependence \rightarrow valley's slope strongly depends on the terms in the exponent

Results from Gupta & Schlichting, 2019a, 2019b

Comparison with observations and trends in planet size distribution with stellar prop.

Planet size vs. Orbital period



- Good agreement with observations in reproducing the location, slope and shape of the radius valley and relative planetary occurrences Results corroborate gravitational compression of

30001000-300100 $10 \ 3000$ $1000 \quad 300$ 100 - 30 $10 \ 3000 \ 1000$ 300100 Insolation flux relative to Earth



Results show that super-Earth and sub-Neptune populations move to higher insolation flux with increasing stellar mass, since the observed orbital period distribution is roughly independent of stellar mass [3]

Observational Predictions

- No significant, i.e., weaker than linear, correlation between planet and stellar mass
- Slope of the valley is a function of stellar mass, i.e., it depends on the stellar-mass luminosity relation
- Relative abundance of super-Earths to sub-Neptunes changes with age even after the first 500 Myrs
- Existence of giga year old planets in the radius valley which are losing mass



[1] Gupta & Schlichting, 2019a, MNRAS, 487, 24. [2] Gupta & Schlichting, 2019b, arXiv: 1907.03732. [3] Fulton & Petigura, 2018, AJ, 156, 264. [4] Fulton et al. 2017, AJ, 154, 109. [5] Van Eylen et al. 2018, MNRAS, 479, 4786. [6] Martinez et al. 2019, ApJ, 875, 29. [7] Johnson et al. 2017, AJ, 154, 108. [8] Wu 2019, ApJ, 874, 91.