

- Photometric variability has many flavors
 - Intrinsic, e.g., pulsating variables
 - Extrinsic, e.g., binary stars (eclipsing, ellipsoidal)
 - Have a look at Debosscher et al. (2009, A&A, 506, 519)
- Focus on a particular class of variable stars: those showing activity-induced variations
 - Tell us about the existence, appearance and evolution of stellar surface features
- Space time-series are very powerful (CoRoT):
 - Silva-Valio et al. (2010, A&A, 510, 25)
 - Lanza et al. (2009, A&A, 506, 255; 2009, A&A, 493, 193)
 - Mosser et al. (2009, A&A, 506, 245)





Gyrochonology / stellar activity

- Besides understanding activity, modulations can be used to extract additional info
- Slow-down of stars with convective envelope as they age (magnetized mass loss)
- For solar-type stars the relationship is well defined and quite tight
- Noticed by Skumanich (1972) ⇒ Prot ∝ t^{1/2}



OUR SUN THROUGHOUT THE ÁGES Age, Rotation, Spot coverage and Coronal X-ray Emission



Lx ~ 5-10 E+29 erg/s ~ 5-10 E+28 ~ 1E+28 ~ 1E+27

P(rot) 2 -4 d; 10% spots ~8 d; 2- 5% spots ~14d; ~1% spots ~25d ; 0.2% spots

 Calibration refined by Barnes (2007) for different spectral types

 $P_{rot} = 0.7225 (B_0 - V_0 - 0.4)^{0.601} t^{0.519}$

- Error claimed to be around 15%
- BUT, important discrepancies found by Mamajek & Hillenbrand (2008)

$$P_{rot} = 0.407 (B_0 - V_0 - 0.495)^{0.325} t^{0.566}$$





- Calibrations for older ages are still TBD
- Some progress made



From Guinan & Engle (2009)

The project

- Use Prot-age relationships and period-search tools in photometric time series to determine the ages of stars with planet candidates
- Tools:
 - Kepler archive
 - NStED toolbox (light curve retrieval, periodogram calculation)
 - Your own codes & favorite visualization software

Data source: Borucki et al. (2010, arXiv:1006.2799) Catalog of 306 planetary candidates



Questions:

- 1. What are the ages of stars with planet candidates?
- 2. Are they representative of the solar neighborhood?
- 3. Is there any correlation with planet size?
- 4. Do you see differential rotation?
- 5. What is the fraction of doublysynchronized systems?
- 6. Can you put constraints on the mass of the planet candidate from tidal evolution?