Biomarkers evolution in the nonresolved visible spectrum of the Earth

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Motivation

- Before the arrival of instruments that will secure the <u>first spectra of terrestrial exoplanets</u>, it is crucial to get a good understanding of what can be learned from those spectra.
- A good starting point is to study the spectrum of the only inhabited planet we know, the Earth.

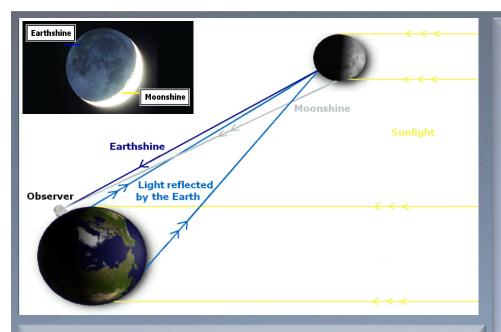
The research project

- We obtain <u>spectra of the Earth</u> by <u>observing the Earthshine</u> with the <u>1.6 m telescope</u> at the Observatoire du mont Mégantic (Québec, Canada).
- Our goal is to follow over a year the <u>spectral evolution of the signature of various molecules</u> (water, oxygen, ozone) that, taken together, could be <u>indicators of habitability and/or biological</u> <u>activity.</u> Of great interest also is the <u>Vegetation Red Edge</u>, a spectral signature of vegetation that might provide a direct evidence of life.









Earthshine: Weak glow of the portion of the Moon's disk that is lit by sunlight reflected from the Earth. **Earthshine spectrum**

- Provides unresolved information on the Earth.
- <u>Varies</u> according to the portion of the Earth that reflects the light toward the Moon (depends on the <u>lunar phase</u> and the <u>longitude of the observer</u>).

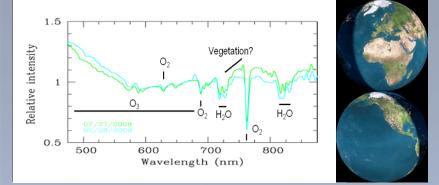
Observatoire du mont Mégantic

- Located in Quebec (45°3N, 71°9W), it provides a <u>unique longitude location</u> for observations.
- A <u>lot of telescope time available</u>: the evolution of the spectrum is followed over different timescales (hour, month, year).
- <u>Large spectral coverage</u>: from 400 nm to 2.5 µm.



Results

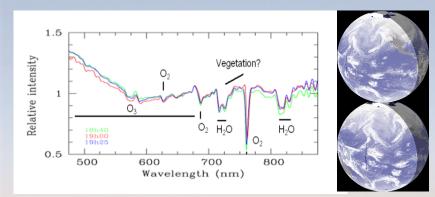
Here are some sample visible spectra obtained:



Left: Earth's spectrum on July 28th, 2008, around 3:50AM, and on Feb. 28th, 2009, around 7:00PM.

Upper Right: The portion of the Earth contributing to the spectrum taken in July, after New Moon. Africa and part of the Atlantic ocean are in sight.

Bottom right: Portion contributing to the February 28th spectrum, before New Moon. the portion contributing is America and part of the Pacific ocean.



Left: Earth's spectrum on February 28th, 2009 at 6:40PM (local time), 7:00PM and 7:25PM. Upper Right: The global cloud cover around 6:30PM. Bottom Right: The global cloud cover around 8:30PM.

References

- Arnold, L., et al. 2002, A test for the search for life on extrasolar planets: looking for the terrestrial vegetation signature in the Earthshine spectrum, A&A, 392, 231.
- European Space Agency. Home Page. Venus Express searching for life on Earth. October 10th, 2008. http://www.esa.int/esaCP/SEMUOW4N0MF_index_0.html.
- Hamdani, S., et al. 2006, *Biomarkers in disk-averaged near-UV to near-IR Earth spectra using Earthshine observations*, A&A, 460, 617.
- Montañés-Rodriguez, P., et al. 2005, *Globally integrated measurements of the Earth's visible spectral albedo*, ApJ, 629, 1175.
- Montañés-Rodriguez, P., et al. 2006, Vegetation signature in the observed globally integrated spectrum of Earth considering simultaneous cloud data: applications for extrasolar planets, ApJ, 651, 544.
- Sagan, C., et al. 1993, A search for life on Earth from the Galileo spacecraft, Nature, 365, 715.
- Seager, S., et al. 2005, *Vegetation's red edge: a possible spectroscopic biosignature of Extraterrestrial plants*, Astrobiology, 5, 372.
- Woolf, N. J., et al. 2002, *The spectrum of a Earthshine: A pale blue dot observed from the ground,* ApJ, 574, 430.
- Turnbull, M. C., et al. 2007, *Spectrum of a habitable World: Earthshine in the near-infrared*, ApJ, 644, 551.
- Earth simulation images are from http://www.fourmilab.ch/earthview/ vplanet.html