Astrometric Surveys: Modern Astrometric Catalogues

Carlos E. Lopez Universidad de San Juan, Argentina and Yale Southern Observatory

What is Astrometry?

Astrometry is that part of astronomy dealing with the positions, motions and trigonometric distances of celestial objects.

What is an Astrometric Survey?

- An astrometric survey can be defined as a search oriented to finding those celestial objects meeting a given accuracy in their absolute or relative positions, motions or trigonometric distances.
- Some of these surveys have been conducted with groundbased meridan circles, astrolabes or astrographic telescopes. Others have used or will use space-based instrumets.

Detectors Used in the Construction of Astrometric Catalogues



The Very Early Beginning

The compilation of catalogues including positions and proper motions is among the best known activities in astrometry. In the early beginnings of astronomy, providing positions of celestial bodies was a real necessity for many of the ancient cultures. With roots dating back to pre-Greek civilizations, astrometry provided the information for computing solar and lunar eclipses and, more importantly, for the determination of time.

The First Astrometric Catalogues

- One of the first catalogues ever compiled was prepared by Hipparchus. It was completed around 129 BC.
- Previous compilations by Greek and Chinese astronomers date back to roughly 300 BC and even to 360 BC.
- During the Middle Ages the most important compilation worth mentioning is called the *Alfonsine Tables*, compiled under the reign of Alfonso the Wise and first published –probably- after 1277. The real purpose of the *Tables* was to offer a means to calculate the positions of the planets; one of the main users of the *Alfonsine Tables* was Copernicus.

The Start of Systematic Surveys

Tycho Brahe is among the first astronomers who started to make accurate systematic observations (mainly of Mars) during the pre-telescope years.

Tycho was the last astronomer living under the geocentric paradigm. This prevented him from further pursuing the determination of stellar parallaxes.

Time for the Telescope

- With the introduction of the telescope, many zone catalogues started to be observed by means of meridian circles. Most of the early to mid 1800s were limited to some specific areas and there was no agreement on the "system" that should be used (the convenience of having a system -in the modern sense of the term to which positions have to be refered- was first proposed in the late 1800s).
- One of the first attempts to carry out a comprehensive survey was started by Argelander with the observation of the zone from +80 to -2 degrees. The zones from -2 to the south pole were added later. This first survey -known as the Durchmusterungs- was made in three parts: Bonner, Córdoba, and Cape Photographic Durchmusterung.

Modern Times

The most important work of the early years of the 20th century was the General Catalogue -GC- (published by Boss in 1937), a preliminary version of which was published in 1910. This version was called Preliminary General Catalogue (PGC). What is important to highlight about the PCG / GC project is Lewis Boss' ultimate aim: the construction of a fundamental system. This goal was finally accomplished by his son, Benjamin Boss.

The Photographic Plate Era

When the photographic plate was finally accepted by the astronomical community as a detector (around 1880), the photographic astrometric surveys started to develop. Curiously, the first observing program with the newly introduced astrographic telescope (the accompanying instrument to the photographic plate) was the observation of the biggest international project of the last years of the 19th century: the Astrographic Catalogue (AC).

The reign of the photographic plate lasted from 1880 to around 1990.

Astrometry's Major Landmarks

- 129 BC: Hipparchus publishes the first catalogue
- 1609: Galileo starts to use the telescope in astronomical research
- 1718: Halley* introduces the concept of proper motions
- 1802: Herschel notes the relative orbital motion of Castor (beginning of double stars)
- 1837: Bessel determines the trigonometric parallax of 61 Cyg
- 1887: The photographic plate is adopted as a detector
- 1989: ESA launches the Hipparcos satellite

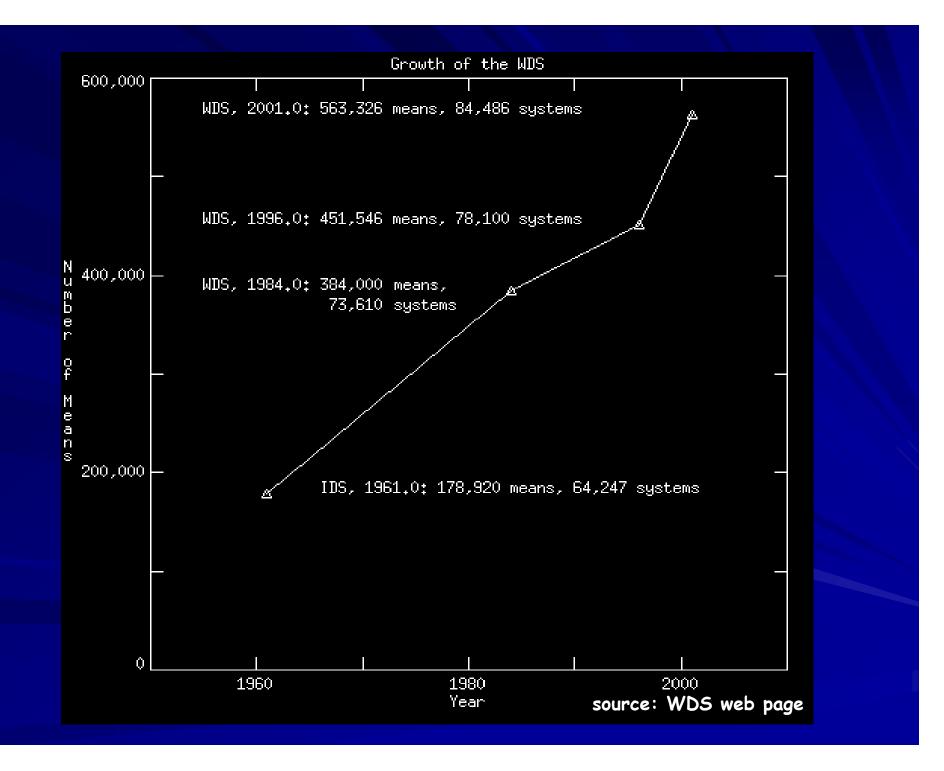
* There is some evidence indicating that it was actually Nicholas of Cusa (circa 1400) who first mentioned the proper motion of stars

Double Stars

- There is a first reference to v1 and v2 Sgr in the Almagest
- 1650 The duplicity of Mizar is annouced by Riccioli
- By the end of the XVII century the duplicity of α Cen and α Cru is announced
- 1779 Mayer publishes the first catalogue of double stars (80 entries)
- 1802 W. Herschel admits that the changes in Castor may be the consequence of an orbital motion
- 1827 Struve introduces the θ and ρ parameters. Catalogue with 3134 pairs

Double Stars (cont.)

- 1900 A new catalogue with 15,000 stars is published
- 1919 First interferometric observations by K. Schwarzschild
- 1921 Hertzsprung starts observations using photographic plates
- 1970 Speckle interferometry observations
- 1980 CHARA (Center for High Angular Resolution Astronomy) is formed
- 2002 The Tycho Double Stars Catalogue is announced by Fabricius et al.



The Washington Double Stars (WDS) Catalogue in numbers

WDS Catalog (~12.5Mb)
00-06 hour section (~2.6Mb)
06-12 hour section (~3.6Mb)
12-18 hour section (~2.3Mb)
18-24 hour section (~4.0Mb)
Format of the current WDS
Notes file for the WDS (~1.0Mb)
References and discoverer codes (~0.7Mb)

Total of 99798 systems

The "Neglected stars"

List Set I: <u>Northern List (Dec > +20, num = 3072)</u> <u>Equatorial List (-20 < Dec < +20, num = 2331)</u> <u>Southern List (Dec < -20, num = 1039)</u>

List Set II: <u>Northern List (Dec > +20, num = 1860)</u> <u>Equatorial List (-20 < Dec < +20, num = 2710)</u> <u>Southern List (Dec < -20, num = 2062)</u>

List Set III: <u>Northern List (Dec > +20, num = 21,061)</u> <u>Equatorial List (-20 < Dec < +20, num = 11,491)</u> <u>Southern List (Dec < -20, num = 13,216)</u>

Total of 58842 systems

Identification of LDS systems in 2MASS and SSS(R) (sample table)

| LDS# | cat | | RA | . (| (J20) | 0.00 |) I | Dec | | Epoch | | | | Magnitude | | т |
|--------|------|----|-----|-----|-------|------|-----|-----|------|----------|-------------|-------|-----------------|------------------|------------------|------------------------|
| 5312 A | 2000 | 01 | 0.4 | 17 | 920 | _02 | 25 | 22 | 90 | 1998.892 | (ma: | s∕yr) | (J∕B) 12.409 | (H∕R1) 11.804 | (K⁄R2) 11.573 | I |
| 5312 A | | | | | | | | | | 1988.764 | 111. | -44. | 16.308 | 99.999 | 14.251 | 12.996 |
| 5312 B | | | | | | | | | | 1998.892 | | | 11,987 | 11.366 | 11.136 | 12.770 |
| 5312 B | | | | | | | | | | 1988.764 | 117. | -24. | 16.960 | 99.999 | 14.768 | 12.882 |
| 5313 A | 2mas | 01 | 06 | 46 | 270 | -03 | 45 | 50 | . 30 | 1998.798 | | | 13.828 | 13.251 | 13.123 | |
| 5313 A | | | | | | | | | | 1995.795 | 56. | -120. | 19.137 | 99.999 | 17.130 | 15.236 |
| 5313 B | | | | | | | | | | 1998.798 | | | 15.677 | 15.172 | 15.261 | |
| 5313 B | sssr | 01 | 06 | 56. | 379 | -03 | 47 | 12 | . 40 | 1995.795 | 48. | -81. | 20.280 | 99.999 | 18.098 | 17.115 |
| 3233 A | 2mas | 01 | 06 | 48. | 800 | +01 | 48 | 45 | . 20 | 2000.815 | | | 16.696 | 16.324 | 15.767 | |
| 3233 A | | | | | | | | | | 1986.577 | 219. | -28. | 17.489 | 99.999 | 17.242 | 17.144 |
| 3233 B | 2mas | 01 | 06 | 42. | 490 | +01 | 46 | 17 | . 60 | 2000.815 | | | 13.678 | 13.115 | 12.848 | 10147-002101-00265-002 |
| 3233 B | sssr | 01 | 06 | 42. | 257 | +01 | 46 | 17 | . 60 | 1986.577 | 254. | -13. | 19.514 | 99.999 | 17.431 | 15.345 |
| 5318 A | 2mas | 01 | 07 | 42. | 850 | -11 | 42 | 45 | .00 | 2000.892 | | | 12.357 | 11.710 | 11.499 | |
| 5318 A | sssr | 01 | 07 | 42. | 755 | -11 | 42 | 45 | . 28 | 1988.764 | 231. | -18. | 16.396 | 99.999 | 14.337 | 12.668 |
| 5318 B | | | | | | | | | | 2000.892 | | | 15.470 | 14.846 | 14.644 | \$2783 \$278227 |
| 5318 B | sssr | 01 | 07 | 39. | .339 | -11 | 44 | 55 | .36 | 1988.764 | 254. | -58. | 20.278 | 99.999 | 18.166 | 16.462 |
| 5319 A | 2mas | 01 | 07 | 54. | 920 | +01 | 15 | 34 | .50 | 2000.755 | | | 11.324 | 10.814 | 10.728 | 1854-534118-0000-001 |
| 5319 A | | | | | | | | | | 1986.577 | 71. | -67. | 13.880 | 12.658 | 12.415 | 11.921 |
| 5319 B | | | | | | | | | | 2000.755 | | | 12.561 | 11.980 | 11.748 | |
| 5319 B | sssr | 01 | 07 | 43. | 881 | +01 | 15 | 33 | . 62 | 1986.577 | 73. | -139. | 17.982 | 99.999 | 16.155 | 14.009 |
| 5321 A | 2mas | 01 | 11 | 06. | 440 | -11 | 08 | 06 | . 90 | 2000.825 | | | 11.669 | 11.034 | 10.820 | 8797 - 470000 |
| 5321 A | | | | | | | | | | 1988.764 | -137. | 28. | 15.421 | 99.999 | 13.111 | 11.439 |
| 5321 B | 2mas | 01 | 11 | 04. | 170 | -11 | 08 | 26 | . 40 | 2000.825 | | | 13.206 | 12.619 | 12.346 | |
| 5321 B | sssr | 01 | 11 | 04. | 299 | -11 | 08 | 26 | . 93 | 1988.764 | -150. | 29. | 18.079 | 99.999 | 15.800 | 14.042 |
| 3244 A | 2mas | 01 | 12 | 26 | 430 | -13 | 08 | 33 | 50 | 1998.602 | | | 11.363 | 10.760 | 10.555 | |
| 3244 A | | | | | | | | | | 1991.672 | 241. | -70. | 17.628 | 99,999 | 13.873 | 12.177 |
| 3244 B | | | | | | | | | | 1998.602 | 10000000000 | | 12.116 | 11.555 | 11.310 | |
| 3244 B | | | | | | | | | | 1991.672 | 215. | -94. | 17.087 | 99.999 | 14.955 | 13.100 |

Parallaxes

1580: First attempt by Tycho Brahe

- 1781: "...the displacement due to the parallax must be less than 1 arcsec" said James Bradley when he attempted to determine the γ Dra parallax
- 1837: Parallax of 61 Cyg determined by Bessel
- 1838: Parallax of Vega determined by Struve
- 1903: Systematic observations with photographic plates started by Schlesinger
- 1924: First edition of the YPC (~1680 star)
- 1983: Monet & Dahn determined the first parallaxes using CCD
- 1995: Fourth edition of the YPC (~8100) star
- 1997: Hipparcos results
- 1998: Nearby Star Databse established
- 2003: Teegarden et al. dicovery of a star between 2.7 and 3.6 pc (using SkyMorph)
- 2005: Jao et al. present new results from the CTIOPI program

Proper Motions

- 1718: Halley announces the proper motion of the stars
- 1775: Mayer publishes the first proper motions catalogue (998 stars)
- 1783: Herschel suggests the idea of solar motion
- 1887: the Astrographic Catalogue is started
- 1916: Barnard discovers his famous star
- 1926: Schlesinger starts the Yale Zones
- 1947: Wright starts the NPM
- 1950: Luyten starts to survey both hemispheres for high proper motion stars
- 1960: Giclas starts his surveys
- 1965: Brower, Schill and Cesco start the SPM (YSO)
- 1989: ESA launches the Hipparcos satallite
- 2000: Monet announces the USNO B

| Year | Catalogue | # of Objects |
|--------|-------------------------|--------------|
| 360 BC | Chinese | ? |
| 260 BC | Aristillus & Timocharis | 850? |
| 129 BC | Hipparchus | 850 (1,080?) |
| 150 DC | Ptolemy's Almagest | 1,080 |
| 1277 | Alfonsine Tables | ? |
| 1437 | Ulugh-Beg | 1,018 |
| 1594 | Rothman & Wilhelm | 1,004 |

| Year | Catalogue | # of Objects |
|------|----------------|--------------|
| 1601 | Tycho Brahe | 1,005 |
| 1690 | Hevelius | 1,564 |
| 1725 | Flamsteed | 3,310 |
| 1751 | Lacaile | 9,766 |
| 1760 | Lalande | 50,000 |
| 1792 | Piazzi | 7,646 |
| 1847 | British Assoc. | 47,390 |

| Year | Catalogue | # of Objects |
|------|------------------------|--------------|
| 1850 | Durchmusterung | 300,000 |
| 1887 | Astrographic Catalogue | ~4,000,000 |
| 1910 | PGC | 30,000 |
| 1926 | Yale Zones | ~150,000 |
| 1937 | General Catalogue | 33,000 |
| 1950 | N30 | 5,000 |
| 1966 | SAOC | 257,997 |

| Year | Catalogue | # of Objects |
|------|------------|--------------|
| 1984 | FK5 | 1,535 |
| 1990 | GSC 1.0 | 20,000,000 |
| 1992 | 4 Millions | 4,000,000 |
| 1993 | PPM | 350,000 |
| 1996 | USNO A1.0 | 488,000,000 |
| 1997 | Hipparcos | 118,218 |
| 1997 | Tycho – 1 | 1,058,332 |

| Year | Catalogue | # of Objects |
|------|------------------------|---------------|
| 1998 | USNO A2.0 | 500,000,000 |
| 1998 | AC2000 | 4,000,000 |
| 2000 | GSC II | 1,000,000,000 |
| 2000 | Tycho – 2 | 2,500,000 |
| 2001 | SPM | 30,000,000 |
| 2002 | UCAC | 40,000,000 |
| 2002 | SuperCosmos Sky Survey | 1,000,000,000 |

| Year | Catalogue | # of Objects |
|-------|------------------------------------|---------------|
| 2003 | USNO B1.0 | 1,000,000,000 |
| 2017? | Gaia | 1,000,000,000 |
| ? | Origins Billion Star Survey | 1,000,000,000 |

The Compilation of the Smithsonian Astrophysical Observatory Catalogue (SAO)

| Declination Zone | Catalogue / Observatory |
|------------------|-------------------------|
| +90 a +85 | Yale |
| +85 a +80 | AGK2 – Greenwich AC |
| +80 a +60 | AGK2 – AGK1 |
| +60 a +50 | Yale |
| +50 a +30 | AGK2 – AGK1 |
| +30 a -30 | Yale |
| -30 a -40 | CPC |
| -40 a -52 | Cape Astrographic |
| -52 a -64 | CPC |
| -64 a -90 | Me3 – Me4 |

The MEGA Catalogues

During the past 15 years four astrometric catalogues -with well over a million entries- have been published. They are called MEGA Catalogues:

Astrometric

✓ Guide Star Catalogue (GSC)
 ✓ USNO (A and B series)
 ✓ SuperCosmos Sky Survey (SSS)
 ✓ UCAC

Non-Astrometric

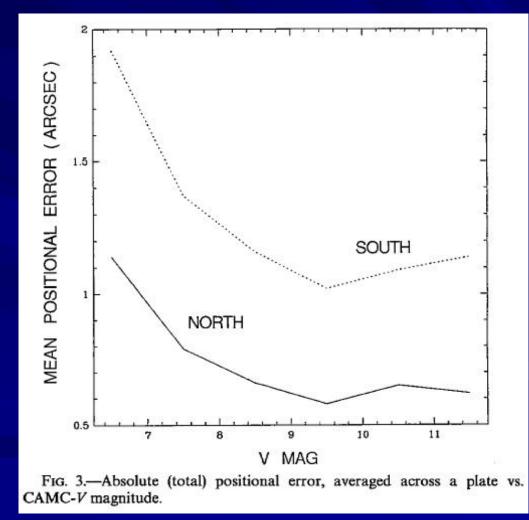
✓ 2MASS✓ DENIS✓ SDSS

Guide Star Catalogue: the pioneer

- The GSC was constructed as support for the HST.
- Over 19 million objects in the 6th to 15th mag. range.
- Astrometry is available at the epochs of the individual plates used in the GSC (no proper motions).
- Reference stars were slected from AGK3, SAOC, and CPC.
- Extensive analysis against the Carlsberg Automatic Meridian Circle data, showed that GSC absolute positional errors from plate center to edge vary from 0.5" to 1.1" in the north and from 1.0" to 1.6" in the south.
- Different improvments has yield the following realeses:

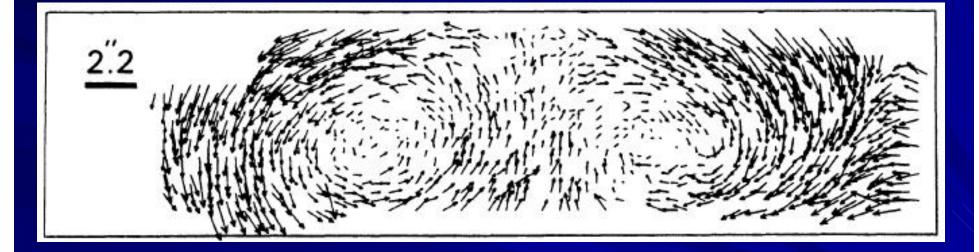
GSC I (1.0, 1.1, and 1.2) GSC II (2.0, 2.1, 2.2, and 2.3)

Comparison GSC 1.0 vs. CAMC (1, 2, and 3)



Taff, L. et al. 1990. ApJ 353, L45 – L48

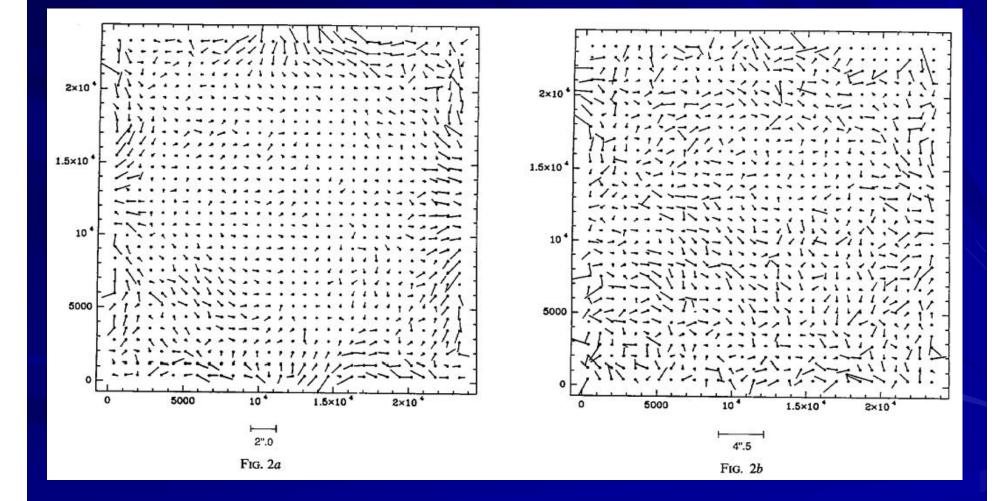
Properties of the GSC I



Typical astrometric residuals in the overlaping areas of a plate pair (about 6° x 1.5°)

Russel, J. et al., 1990, AJ 99, 2059.

The GSC I: Mean residuals map



Comparison GSC 1.0 vs. GSC 1.2

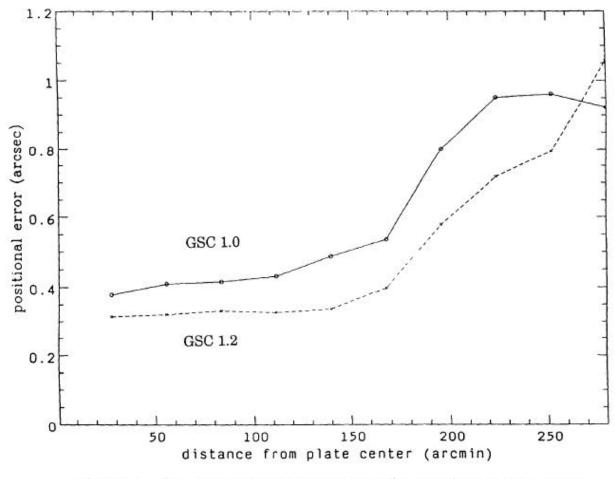


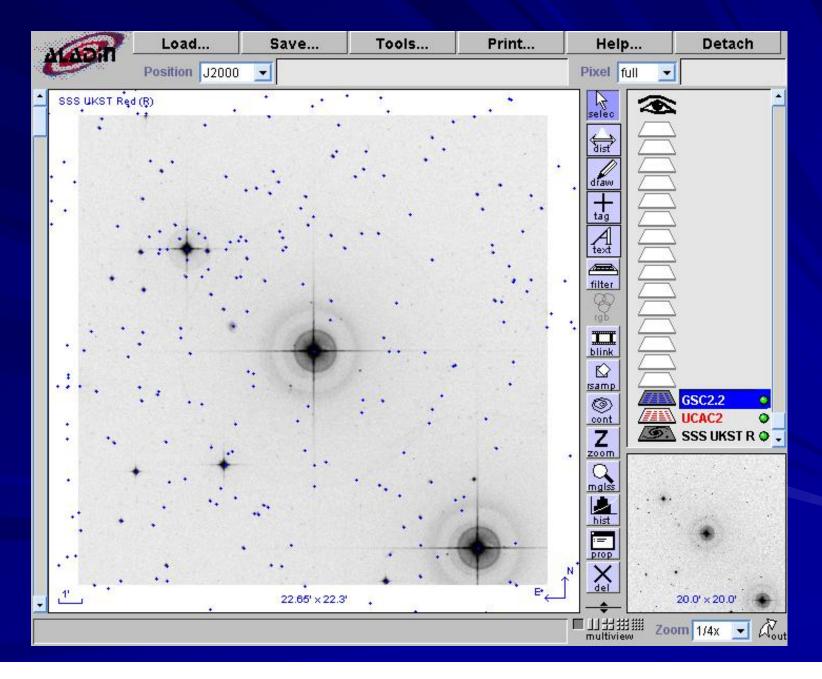
Figure 2. Average positional errors vs. distance from plate center.

Bucciarelli, B., et al. 1994. IAUS 161, 277.

Quick Fact Sheet on GSC 1.1 vs. GSC 2.2

| Item | GSC 1.1 | GSC 2.2 | |
|------------------------------------|--|--|--|
| Total Objects | 18,819,291 | 435,457,355 | |
| Mean epoch of catalog positions | 1981.8 | 1992.2 | |
| Magnitude limit | 15.5 | 18.5 F, 19.5 J | |
| Complete to magnitude limit | No | Yes | |
| Bandpass: North | V | F and J | |
| Bandpass: South |] | F and J | |
| Digitization resolution | 1.7 arcsec | 1 arcsec | |
| Source for bright stars | Hipparchos Input Catalog | Tycho 2 | |
| Astrometry: | | | |
| Reference frame | FK4 | ICRF | |
| Reference catalogs | AGK3, SAO: 10^2 calibrators per plate | ACT, Tycho2: 10^3 calibrators per plate | |
| Errors | 0.5" - 1.2" errors increase near plate edges | 0.3" - 0.75" errors increase near plate edges for F > 16 | |

Zone of Hip 5164 (UKST-R frame) as shown by Aladin and GSC 2.2 detections



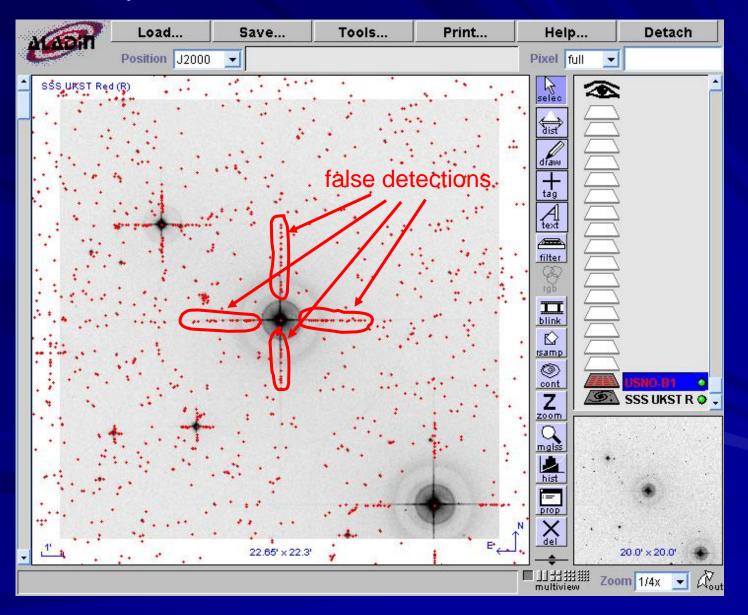
The USNO Series

| Release | Number of Objects |
|------------|-------------------|
| USNO A1.0 | 491,848,883 |
| USNO A2.0 | 526,280,881 |
| USNO SA2.0 | 54,787,624 |
| USNO B1.0 | 1,045,913,669 |

Summary of the USNO-B1

- Monet, D.; Levine, S.; Canzian, B. et al. 2003 AJ 125, 948
- Number of sources: 1,045,913,669
- Number of observations: 3,643,201,733
- Number of plates: 7,435
- Completeness down to V= 21
- 0.2 arcsec accuracy at J2000
- 0.3 mag accuracy in up to 5 colors

Zone of Hip 5164 (UKST-R frame) as shown by Aladin plus USNO B1.0 detections



The SuperCosmos Sky Survey (SSS) and The SuperCosmos Science Archive (SSA)

- The database contains over 1 billion multi-colour, multi-epoch sources and covers the southern celestial hemisphere (δ < +3.0) in three passbands (BRI), with one colour (R) represented at two epochs.
- All SSA global astrometry is tied to the Hipparcos-Tycho reference frame via the Tycho-2 and ACT catalogues.
- Astrometry is globally good between 0.2 and 0.3 arcsec.
- New proper motions (with respect to the ones in SSS) have been computed using all available positions. Up to four different epochs have been used.
- The SSS and SSA are based on the same underlying data with the main differences arising in the construction of the SSA merged source table.



Plate Material

| Survey | Dec centres | Plate limit | Dates of observation | Reference |
|----------------------------|--------------------------------------|--------------------|-------------------------|-------------------------------------|
| Southern hemisphere | survey: | | | |
| SERC-J/EJ ¹ | $\delta \leq 0^{\circ}$ | $B_{\rm J}\sim 23$ | 1974 to 1994 | Cannon (1984) |
| SERC-ER/AAO-R ² | $\delta \stackrel{-}{\leq} 0^\circ$ | ${ m R}\sim22$ | 1984 to 2000 | Cannon (1984); Morgan et al. (1992) |
| SERC-I | $\delta \leq 0^\circ$ | $\rm I\sim19$ | $1978 \longrightarrow$ | Hartley & Dawe (1981) |
| $ESO-R^3$ | $\delta \leq -20^{\circ}$ | $R\sim 22$ | 1978 to 1990 | West (1984) |
| $POSS-I E^3$ | $-18^\circ \leq \delta \leq 0^\circ$ | $R\sim 20$ | 1949 to 1958 | Minkowski & Abell (1963) |
| | | | | |
| Putative northern hen | nisphere survey: | | | |

| $POSS-II B^3$ | $\delta \geq 0^\circ$ | $B_J \sim 22.5$ | 1987 to 1999 | Reid et al. (1991) |
|---------------|-----------------------|------------------|------------------------|--------------------------|
| $POSS-II R^3$ | $\delta \ge 0^\circ$ | $R\sim 20.8$ | 1987 to 1999 | Reid et al. (1991) |
| $POSS-II I^4$ | $\delta \geq 0^\circ$ | $\rm I\sim 19.5$ | $1989 \longrightarrow$ | Reid et al. (1991) |
| $POSS-I E^3$ | $\delta \geq 0^\circ$ | $R\sim 20$ | 1949 to 1958 | Minkowski & Abell (1963) |

Notes:

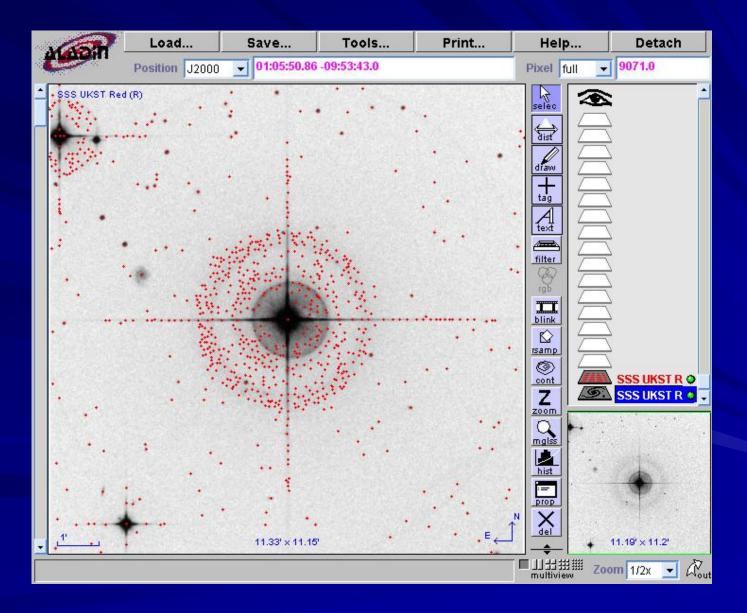
¹Original glass survey plates scanned with the exception of the following 6 fields: 31, 102, 167, 330, 555, 575 (replacement glass originals).

²Original glass survey plates scanned with the exception of the following 32 field numbers: 52, 54, 111, 114, 119, 244, 296, 298, 355, 412, 413, 472, 473, 476, 483, 540, 541, 549, 550, 611, 619, 632, 679, 686, 691, 758, 760, 765, 828, 831, 837, 838 (film originals).

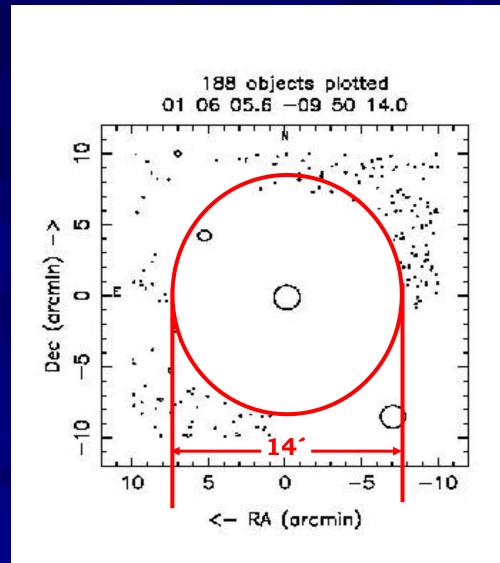
³Glass atlas copies of survey glass originals.

⁴Film atlas copies of survey glass originals.

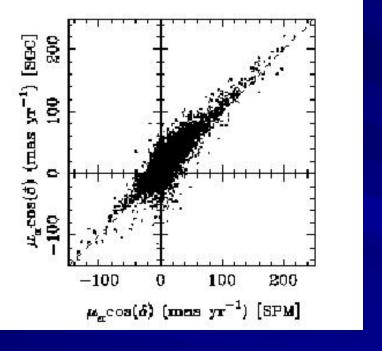
SSS Detections (as shown by Aladin) around Hip 5164



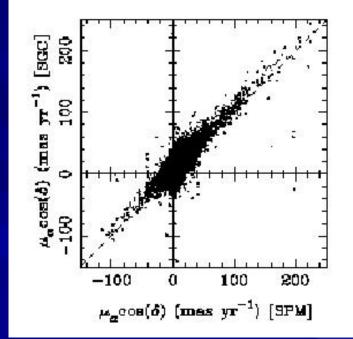
SSS Chart around Hip 5164



The SuperCosmos Sky Survey: Proper Motion Comparison with the SPM



10.0<V<14.0



14.0<V<16.0

Hambly, N. C. et al. 2001. MNRAS 326, 1315.

The UCAC Project

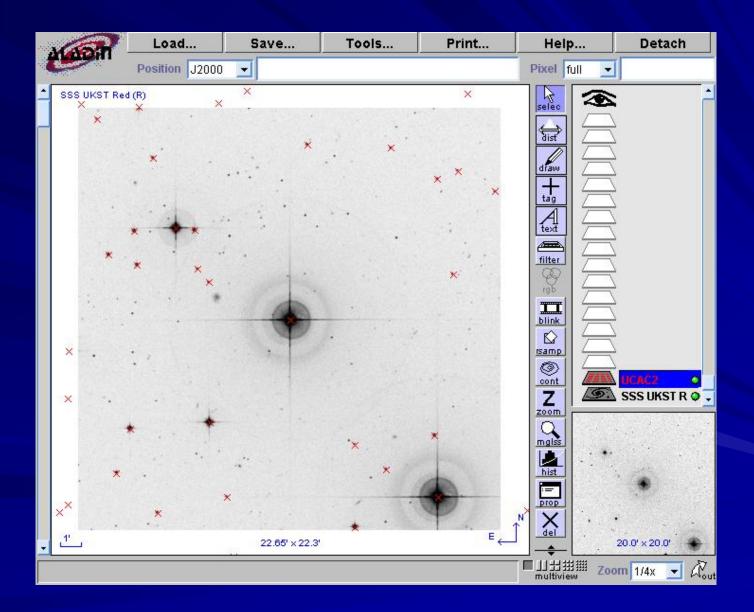
UCAC project goals

- ✓ densification of the reference frame beyond Hipparcos/Tycho
- ✓ improve accuracy of positions of faint end Tycho-2 stars
- improve link between Hipparcos and the International Celestial Reference Frame

The UCAC2

- Positions and proper motions for over 48 million sources (mostly stars).
- Precision on the positions is 15-70 mas (depending on magnitude).
- Proper motions are derived by using over 140 ground- and space-based catalogues. With errors about 1-3 mas yr⁻¹ for stars to 12th mag., and about 4-7 mas yr⁻¹ for fainter stars to 16th mag.
- Current epoch positions are obtained from observations with the USNO 8-inch Twin Astrograph equipped with a 4k CCD.
- The catalogue covers from -90° up to +48° (to +52° in some areas) and supersedes UCAC1 released in 2001.

UCAC2 Detections (as shown by Aladin) around Hip 5164



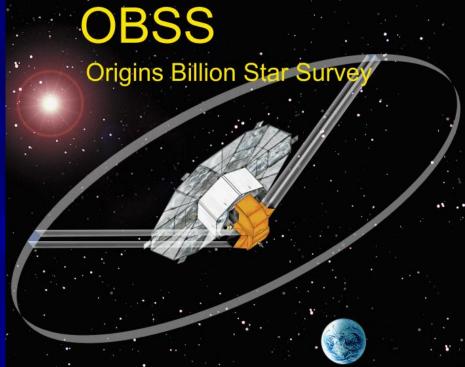
NOMAD

(Naval Observatory Merged Astrometric Database)

- Astrometric and photometric data for over 1 billion stars.
- Source catalogues (for astrometry and optical photometry): Hipparcos, Tycho-2, UCAC2, and USNO-B.
- Photometry supplemented by 2MASS.
- NOMAD is not a compiled catalogue; that is, if a given star is presented in more than one of the above mentioned catalogues, only one catalogue entry is chosen.
- All source catalogues astrometric data are on the ICRF.
- 100 GB of data.

The Next Two (Last?) Steps





"If the Almighty had consulted me before he embarked on creation, I should have recommended something simpler"

Alfonso the Wise

