



The Impact of Escaping Hydrogen Atmospheres on super-Earth Interiors James Rogers

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Solid core accretion

Gas accretion

Rogers, Owen & Schlichting (2023)





Boil-off

Photoevaporation + Core-powered mass-loss



Petigura, Rogers et al. (2022)



Sub-Neptunes

Super-Earths



Core Density







"Observe" Planets with Kepler







"Observe" Planets with Kepler





Hierarchical Inference Model





Mean Core





Mean Core

Mean Core



suggests that small planet interiors are slightly under-dense when compared to Earth

(A good example: TRAPPIST-1)

Why?

Can hydrogen *itself* explain this?



Rogers, Schlichting and Young (in prep.)

See Schlichting and Young (2021), Young et al. (2023)





Rogers, Schlichting and Young (in prep.)



Rogers, Schlichting and Young (in prep.)

Step 1: Use atmospheric evolution models to determine the H₂ atmospheric mass at the 'time of final global equilibrium (solid magma ocean surface)

Step 2: Find global chemical equilibrium state that produces this amount of H₂ atmospheric mass





Rogers, Schlichting and Young (in prep.)





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What is left behind?

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Rogers, Schlichting and Young (in prep.)





What is left behind?





Some speculation...



Some speculation...



Conclusions

- Super-Earth interiors can be slightly under-dense when compared to Earth.
- As H₂ escapes, it is also sequestered into the interior, reducing overall bulk density.
- This produces abundant H₂O, and steam-dominated atmospheres.



Rogers, Schlichting and Young (in prep.)